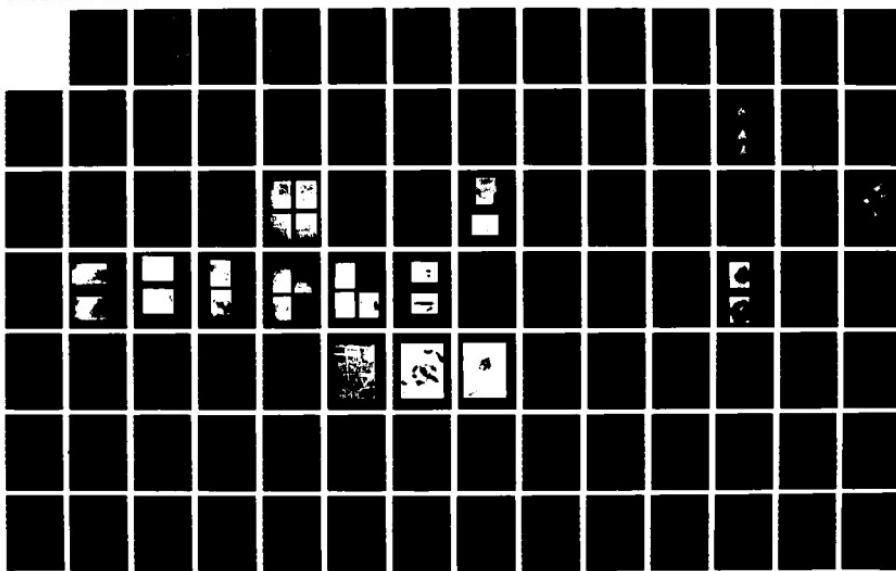
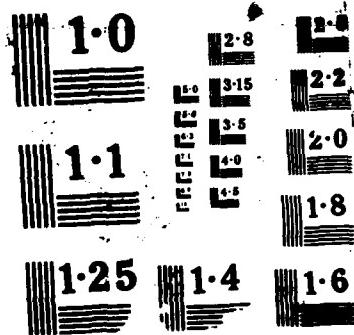


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# **TECHNICAL PAPERS PRESENTED AT THE DEFENSE NUCLEAR AGENCY GLOBAL EFFECTS REVIEW**

**Volume II**

**Kaman Tempo  
Alexandria Office  
Huntington Building  
2560 Huntington Avenue  
Alexandria, VA 22303-1410**

**15 May 1986**

**Technical Report**

**CONTRACT No. DNA 001-82-C-0274**

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**Prepared for  
Director  
DEFENSE NUCLEAR AGENCY  
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## PREFACE

The Defense Nuclear Agency has collected and printed the attached papers from the February 25-27 1986 Global Effects review as a service to the community. The Defense Nuclear Agency takes this opportunity to express its gratitude to the numerous participants in the Global Effects review.

The technical papers enclosed include all those which were received by DNA prior to the closing date of 28 April 1986. Where papers are missing their place is occupied by the abstract received prior to the meeting.

The inclusion of a paper in this proceeding does not necessarily imply endorsement of the results of the research reported or conclusions which might be drawn from that research. It is the opinion of the Defense Nuclear Agency that, while good progress is being made in improving our understanding of Global Effects, the results to date are tentative and preliminary and should not be used for planning beyond the planning of future research.



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**SECTION 1**  
**CHEMISTRY AND MICROPHYSICS**

LABORATORY STUDY OF WET COAGULATION

GLOBAL EFFECTS PROGRAM TECHNICAL MEETING

FEBRUARY 26, 1986

GEORGE CARRIER\*  
FRANK FENDELL  
DAN KWOK  
BRUCE LAKE

\*Harvard University, Cambridge, Massachusetts 02138

TRW SPACE AND TECHNOLOGY GROUP  
ONE SPACE PARK  
REDONDO BEACH, CALIFORNIA 90278

## OBJECTIVE

Do laboratory experiment:

- (1) To see if light extinction by soot particles changes after confinement in a cloud
- (2) To measure size distribution and number density of soot particles before and after confinement in a cloud

## APPROACH

- (1) Introduce soot particles and moist air into cloud chamber
- (2) Measure extinction, number density and size distribution of soot particles
- (3) Expand cloud chamber adiabatically to form cloud
- (4) Sit for sometime for scavenging, wet coagulation to take place
- (5) Compress cloud chamber to evaporate cloud
- (6) Measure extinction, number density and size distribution of soot particles again

## PROBLEMS TO BE SOLVED

### (1) Cloud Chamber

Wall temperature must closely match gas temperature  
to prevent condensation, convection

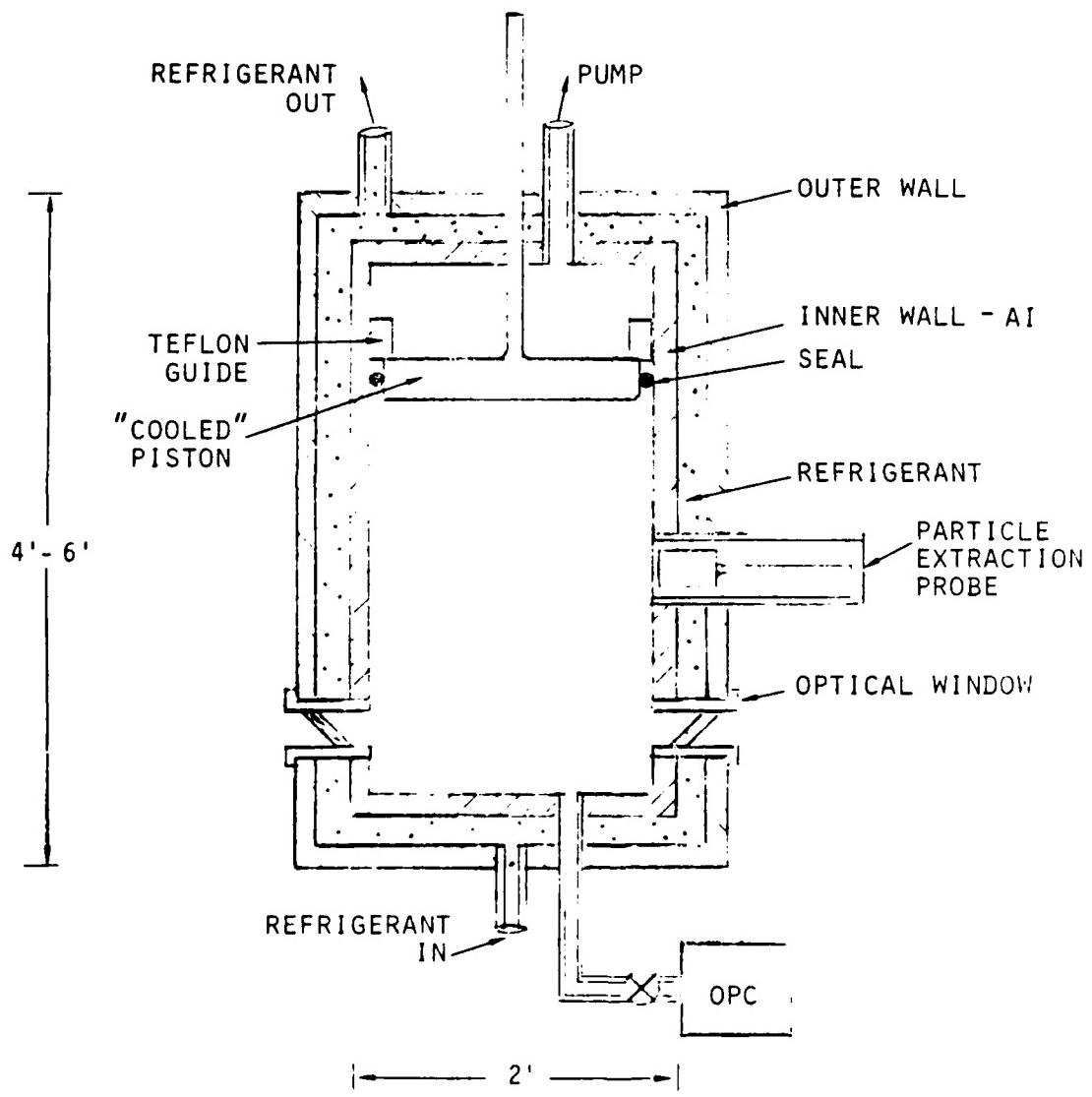
### (2) Extinction Measurement

Low number density of soot particles  $\rightarrow$  low  
extinction which is hard to measure

### (3) Soot Particles' Number and Size Measurements

Submicron particle size and number measurements  
are non-trivial

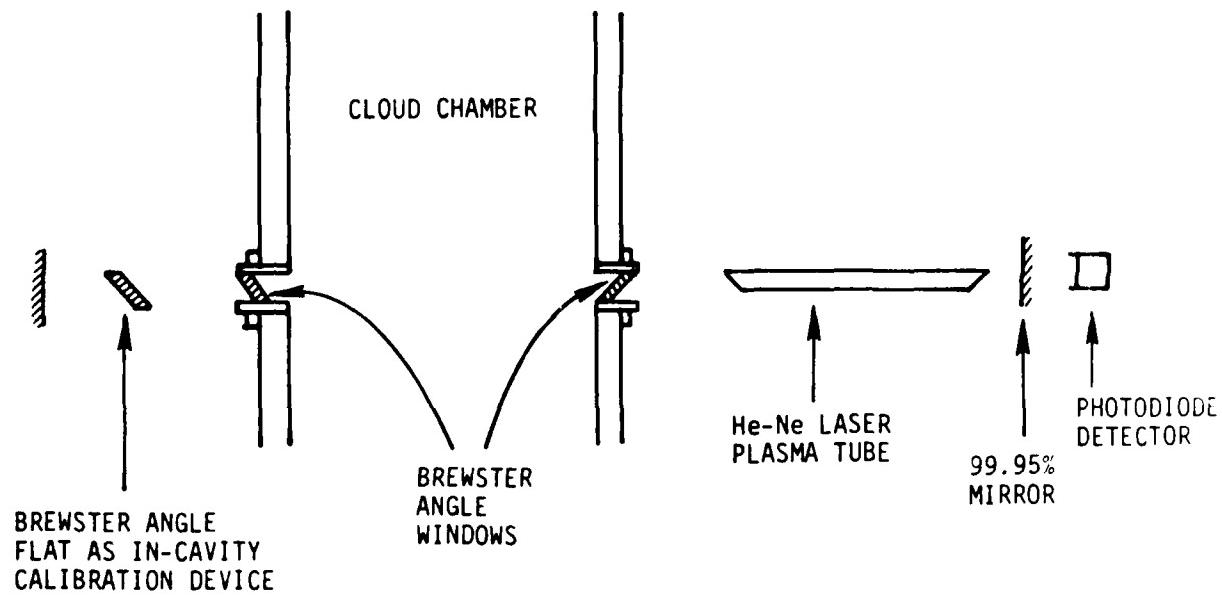
## CLOUD CHAMBER: SCHEMATICS



## CLOUD CHAMBER: APPROACH

- Wall cooling by simply passing refrigerant through at controlled cold temperature and flow rate. Cool from R.T. to  $\sim -60^{\circ}\text{C}$  in 1/2 hour.
- Gas cooling by adiabatic expansion using piston. Expansion ratio  $\sim 1:3$
- Minimize leakage by piston by evacuating and equalizing pressure on back side of piston
- Monitor wall and gas temperature by thermocouple or transistor thermometer
- Make gas temperature to follow wall's
- Aim for wall  $0.1\text{-}0.3^{\circ}\text{C}$  warmer than gas to prevent condensation
- Make top of chamber warmer than bottom to suppress convection

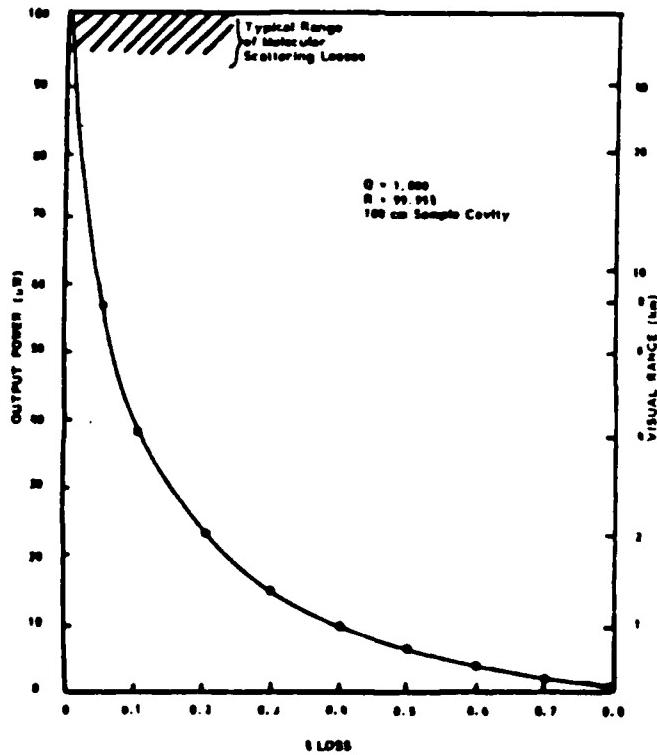
EXTINCTION MEASUREMENT: SCHEMATICS



## EXTINCTION MEASUREMENT: APPROACH

- $10^4\text{-}10^5$   $0.1 \mu$  particles/c.c.  $\rightarrow 10^{-4}\text{-}10^{-3}/\text{m}$  extinction coefficient
- Use "laser-cavity extinction" method:
  - make soot-laser air part of laser cavity
  - laser output highly non-linear function of cavity loss
  - equivalent measurement pathlength  $\sim$  kilometers
  - use variable Brewster angle flat as in-cavity calibration device
  - evacuate chamber either before or after dust measurement for calibration

## LASER OUTPUT VS. CAVITY LOSS



**FIGURE 2.** Visual range and output power versus percent cavity loss. The prototype uses a 100-cm sample cavity length. With current state-of-the-art high reflectivity mirrors, the sample cavity length can be reduced to less than 30 cm with a  $Q$  of 2,000 to 3,000.

BREWSTER ANGLE FLAT AS IN-CAVITY CALIBRATION STANDARD

<u>Incident Angle</u>	<u>Loss</u>
56.0	$4 \times 10^{-7}$
56.5	$6.4 \times 10^{-5}$
57.0	$2.4 \times 10^{-4}$
57.5	$5.3 \times 10^{-4}$
58.0	$9.7 \times 10^{-4}$
58.5	$1.53 \times 10^{-3}$
59.0	$2.25 \times 10^{-3}$
60.0	$4.18 \times 10^{-3}$

## SOOT PARTICLE NUMBER AND SIZE MEASUREMENTS

Primary Measurement — Extraction + electron  
microscopy

Secondary Measurement — Optical scattering methods

## EXTRACTION AND ELECTRON MICROSCOPY

### Methods of Extraction:

- (1) Impaction — many lose small particles X
- (2) Suctioning and filtration — agglomeration may occur in suctioning probe or filter? X
- (3) "Grabbing" and gravity settling — disturb particles the least.  $0.1 \mu$  particles settle in 3 hours ✓

### Methods of Electron Microscopy:

- (1) Scanning electron microscopy (SEM) —  $0.1 \mu$  resolution may be difficult, need to coat particles with  $100-300 \text{ \AA}$  gold layer?
- (2) Transmission electron microscopy (TEM) — sample preparation more involved but well developed. Mount Collidian film on screen or slide.

## OPTICAL METHODS AVAILABLE FOR MEASURING NUMBER DENSITY AND SIZE

Ensemble Measurement  
(beam measurement)

- [-] Scattering extinction ratio
- [-] 2 angle intensity ratio
- [-] diffusion broadening spectroscopy
- [-] 2 wavelength extinction ratio
- [-] polarization ratio

Particle Measurement  
(focused measurement)

- [-] 2 angle intensity ratio
- [-] 2 beam interferometry
- [-] white light integrated forward scattering

## COMMON SOURCES OF ERROR FOR ALL OPTICAL TECHNIQUES

1. Non-sphericity

maximum error ~ 40%

2. Index of Refraction Uncertainty

for highly absorptive particles like soot,  
a few percent

3. Depth of Field (for focused measurements)

broadens size distribution, < 30%

TWO-ANGLE INTENSITY RATIO - PARTICLE MEASUREMENT  
COMPARISON WITH SEM

---

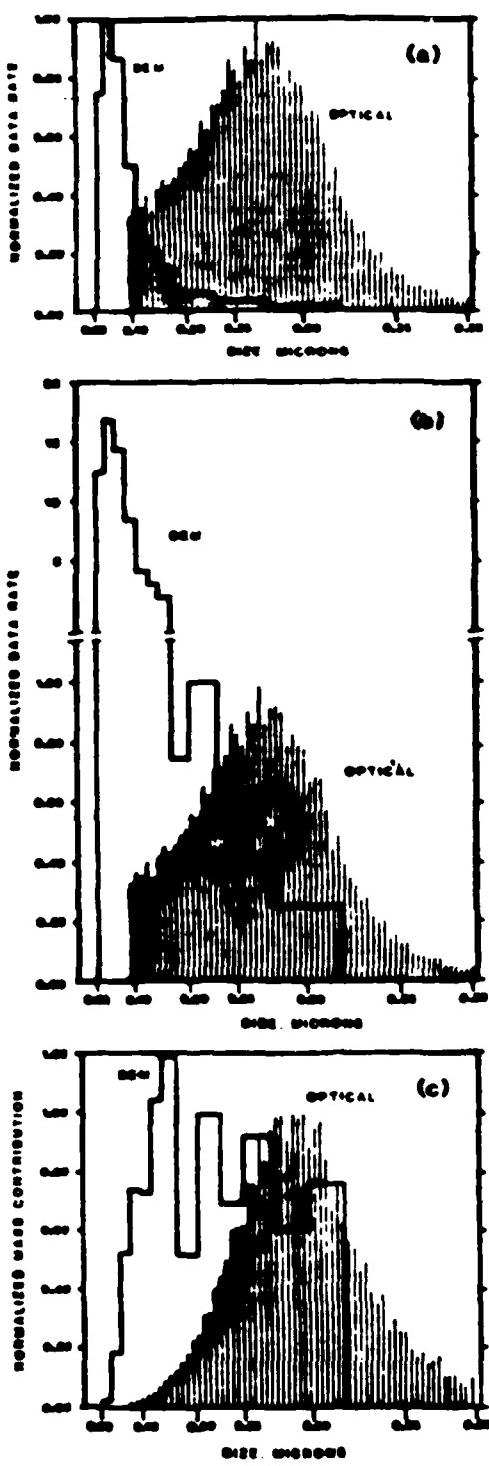


Fig. 8 Variation [1] Concentration,  $\phi = 0.4$ , proportionality (a)  
Mass density, (b) number density normalized; (c) mass density

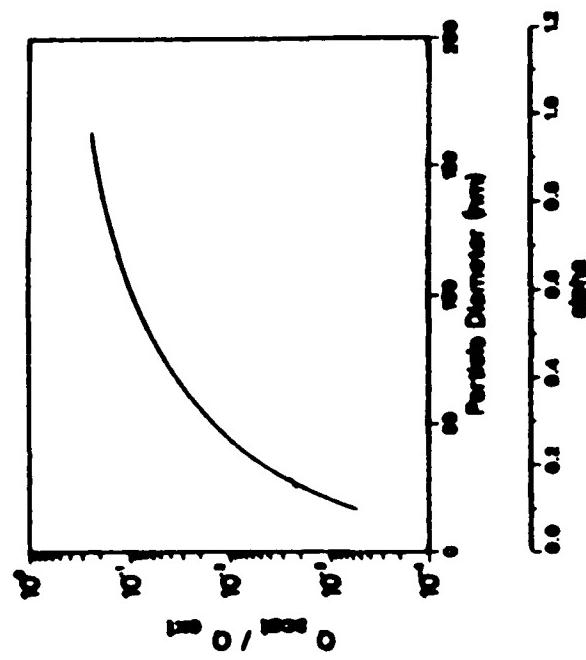
## OPTICAL METHODS SELECTED

For small particles ( $< 0.1 \mu$ ) — extinction/scattering ratio or 2 angle intensity ratio (large angle pair)

For large particles ( $>$  several  $\mu$ ) — 2 angle intensity ratio (small angle pair)

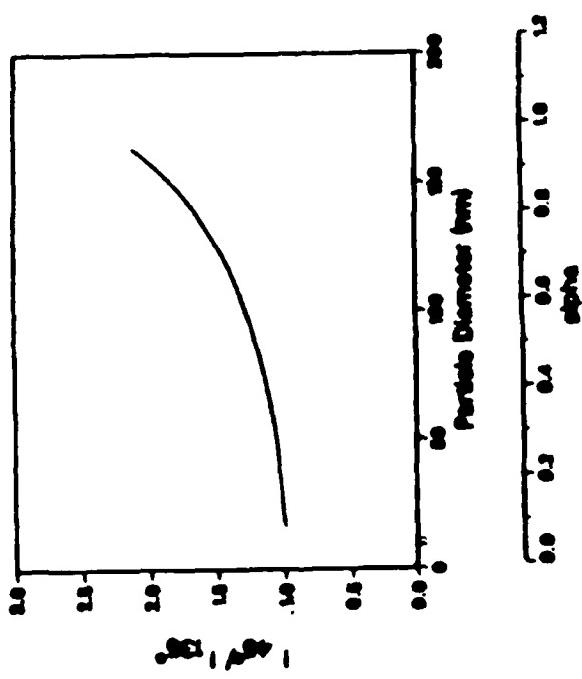
## ME SCATTERING

Scattering/Extinction Ratio



## ME SCATTERING

Two-Angle Intensity Ratio



## SUMMARY OF WHAT WE PROPOSE TO DO

- (1) Measure extinction accurately
- (2) Measure particle size distribution and number density accurately by extraction and TEM
- (3) Use other optical techniques for real time monitoring and as secondary measurements

MICROPHYSICAL PROPERTIES AND REMOVAL PROCESSES OF SOOT

John Hallett, Barry Gardiner, Dennis Lamb,  
Fred Rogers, Richard Pitter, Jim Hudson

Desert Research Institute  
Reno, Nevada 89506

Viewgraphs - February 1986

NASA AMES DNA Meeting

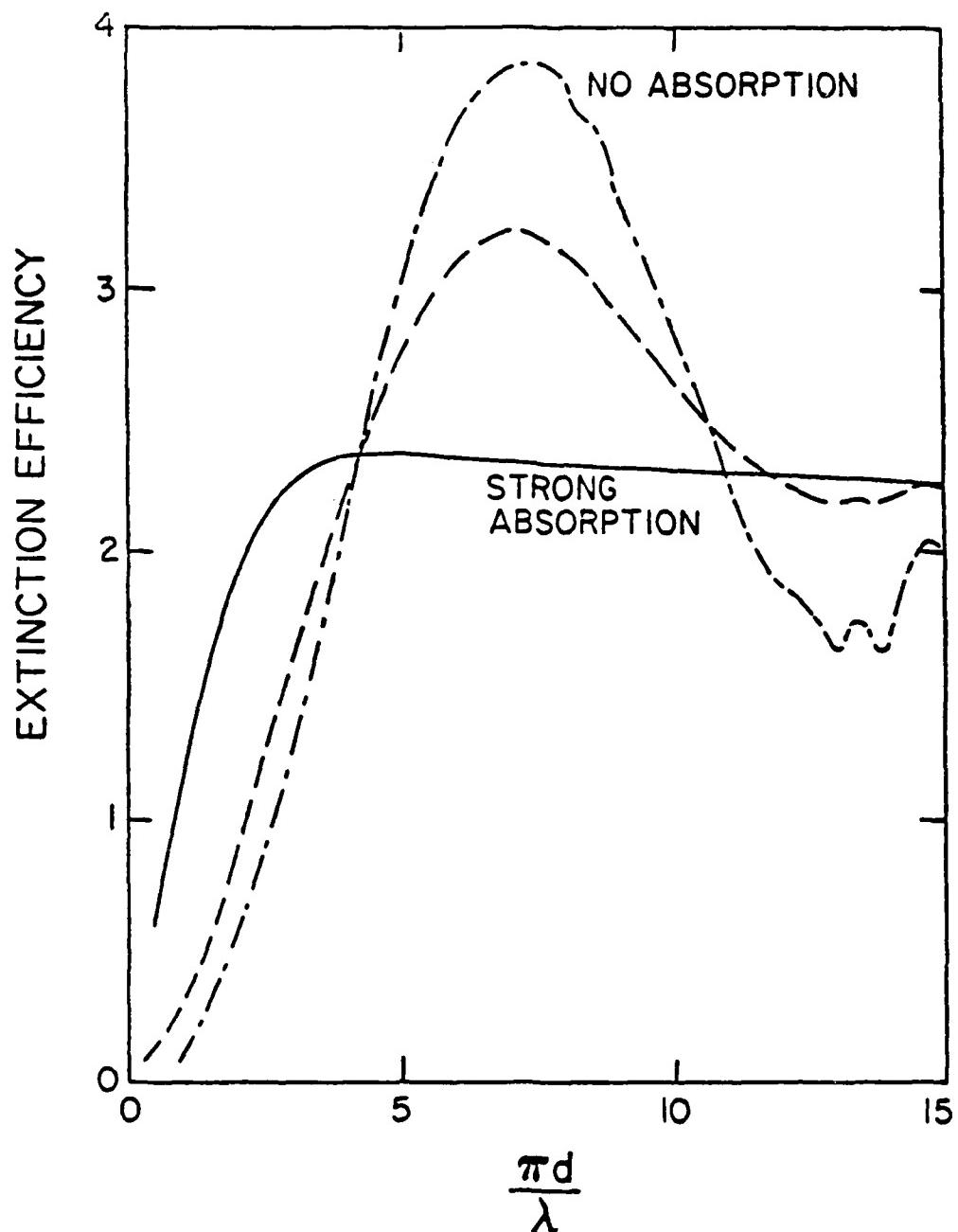
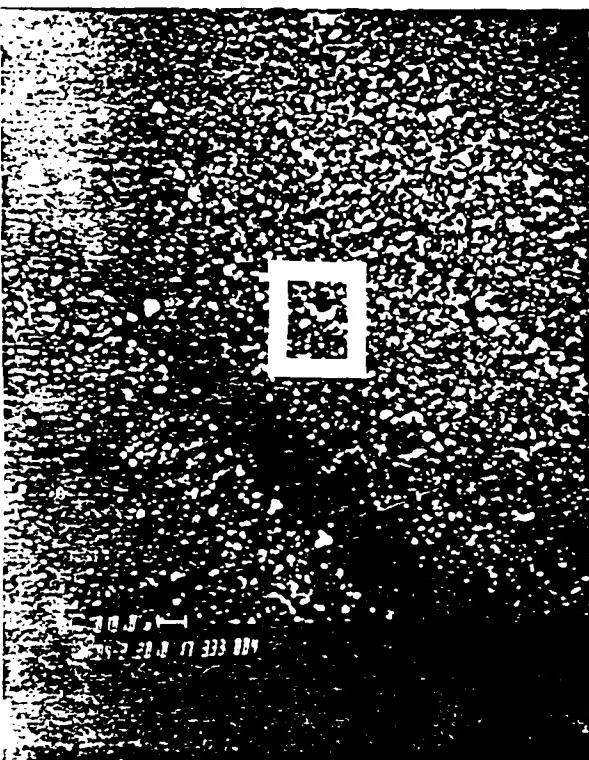
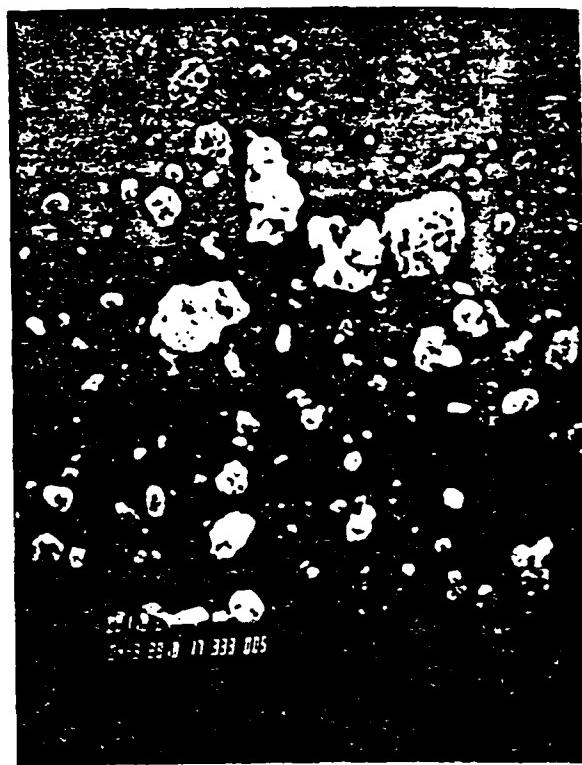


Fig. 1: Extinction efficiency for water spheres with different absorption.

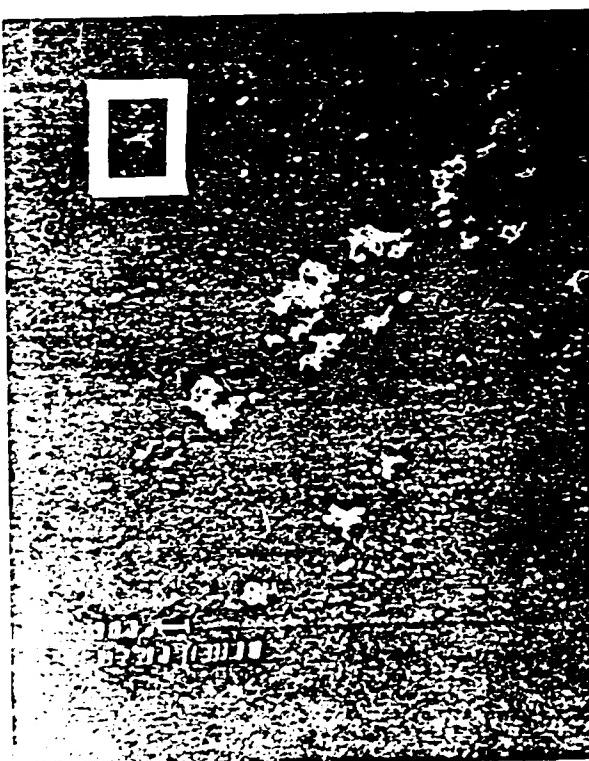


(a)



(b)

Figure 2



(c)



(d)

FIG. 2. (a)-(d) Micrographs

# CCN SPECTRA FROM SOOTY ACETYLENE

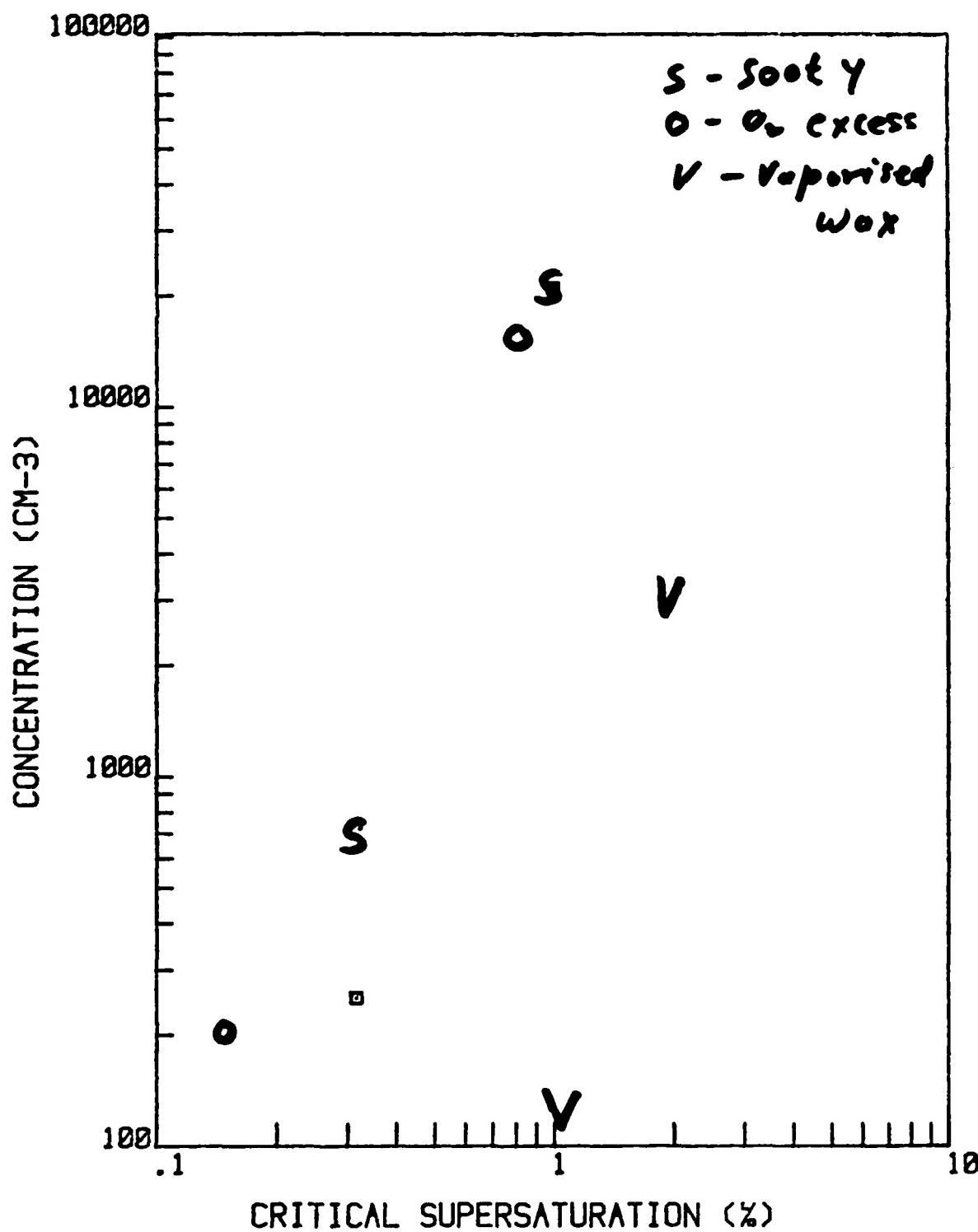


Figure 3 CCN spectra with immediate dilution.

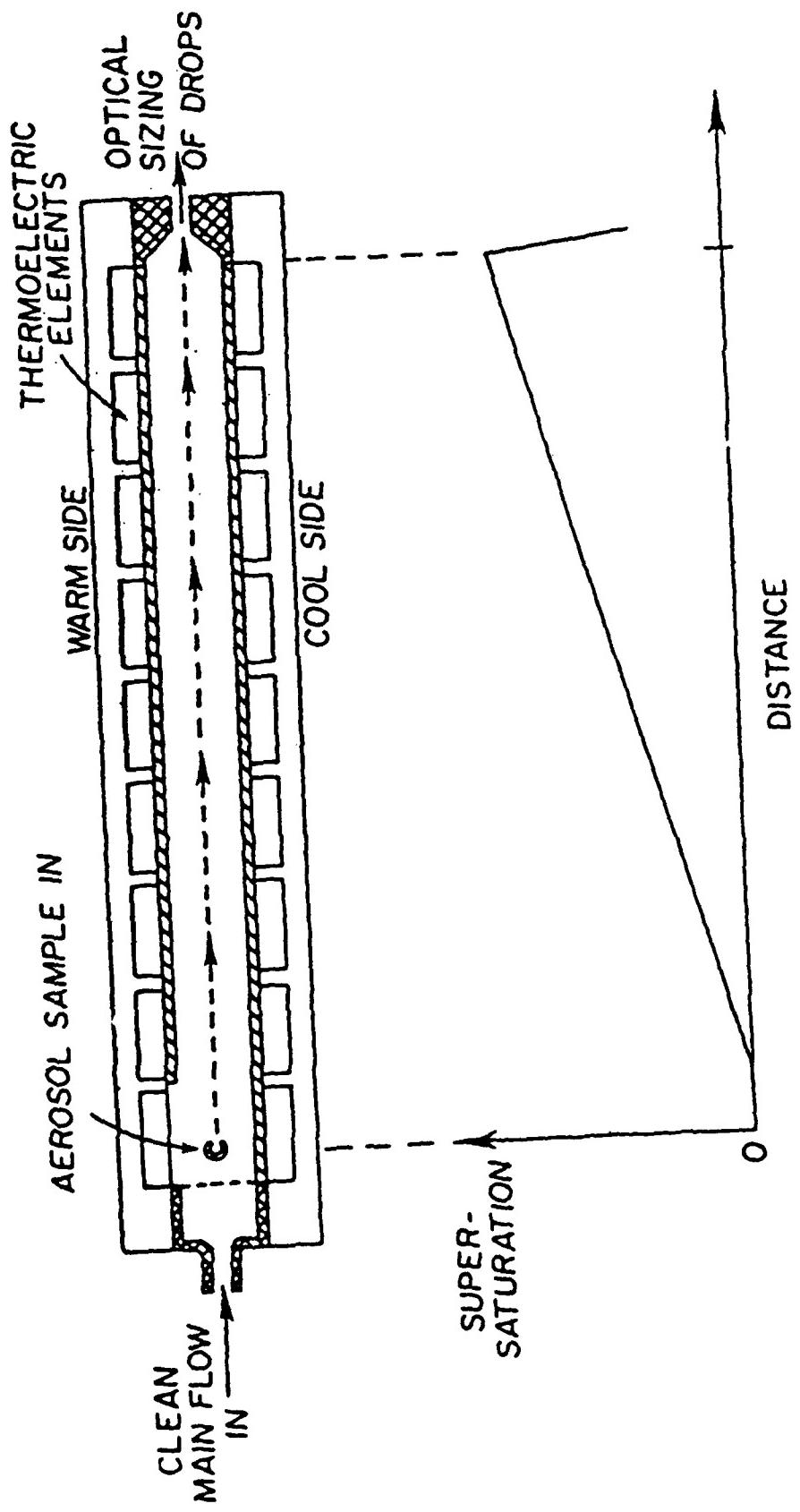


Figure 4 Diagram of the CCN thermal gradient spectrometer.

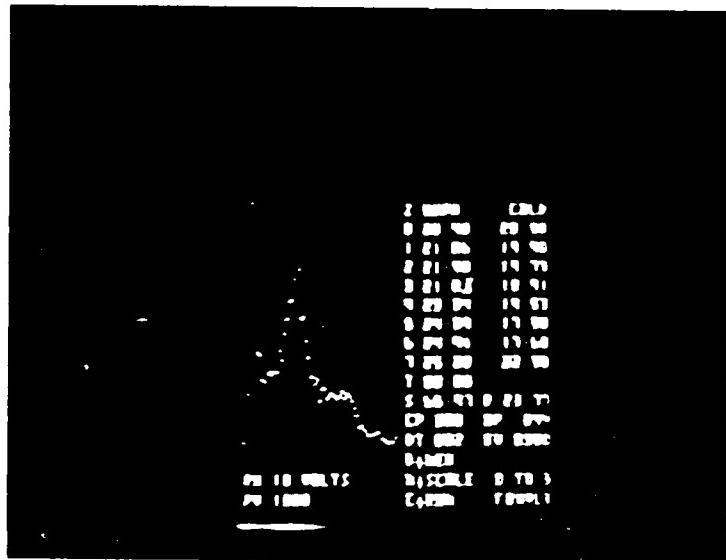
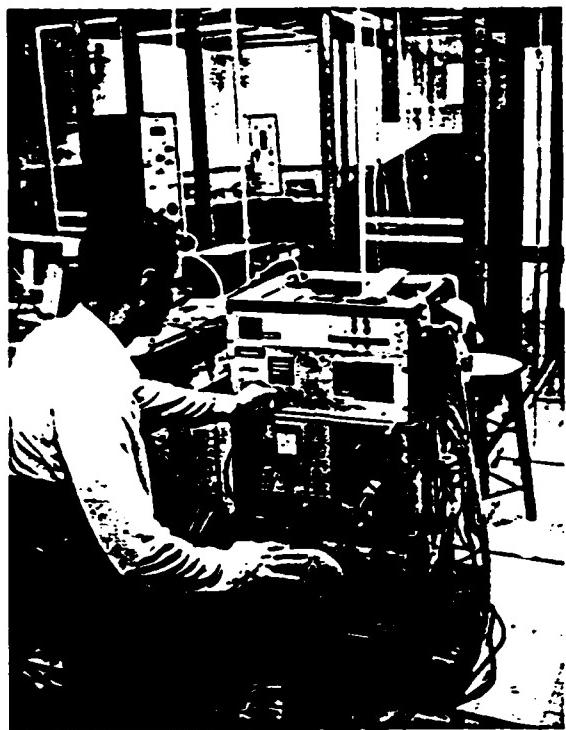


FIG 5: CCN CHAMBER AND SPECTRA

REC1HR = 02 15 62 56 36.0  
 TIME = 1145 1PF5 ST MR CP P6 P1 P2 P3 P4 P5 P6 P7 KEP  
 REC = 1145 1PF5 1070L 30 040 050 060 070 080 090 0100 0110 0120  
 HEMI = 659.16.0 62.1 579.5 268.6.12.0.92.1.83.3.81.2.85.5.67.6.85.7.94.46  
 QC = 9 82 165 94  
 RECMT = 62 15 62 56 36.0  
 RECMT

358

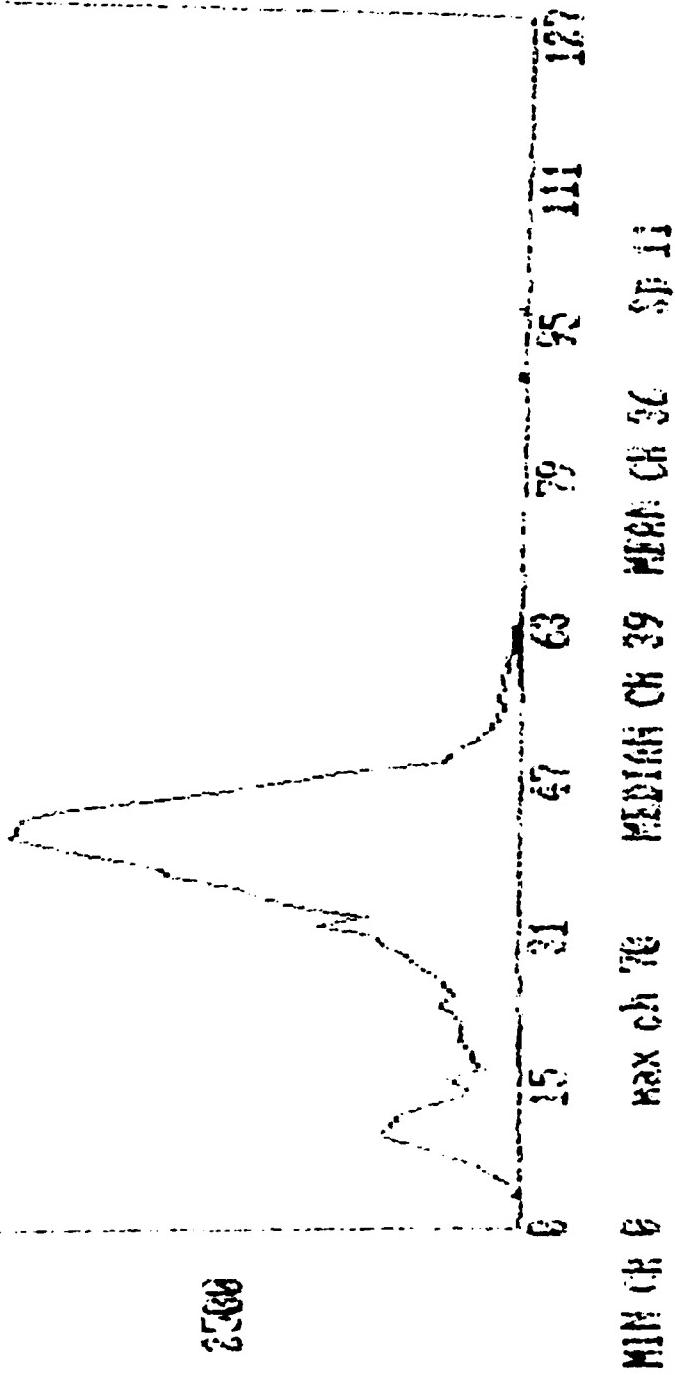


FIG. 6:

CALIBRATION AFRO SOC  
 Nell O. 80 % S.  
 RECMT

BRGUM = 382  
LIM = 15 332  
HCL = 15 332  
HCL = 15 332

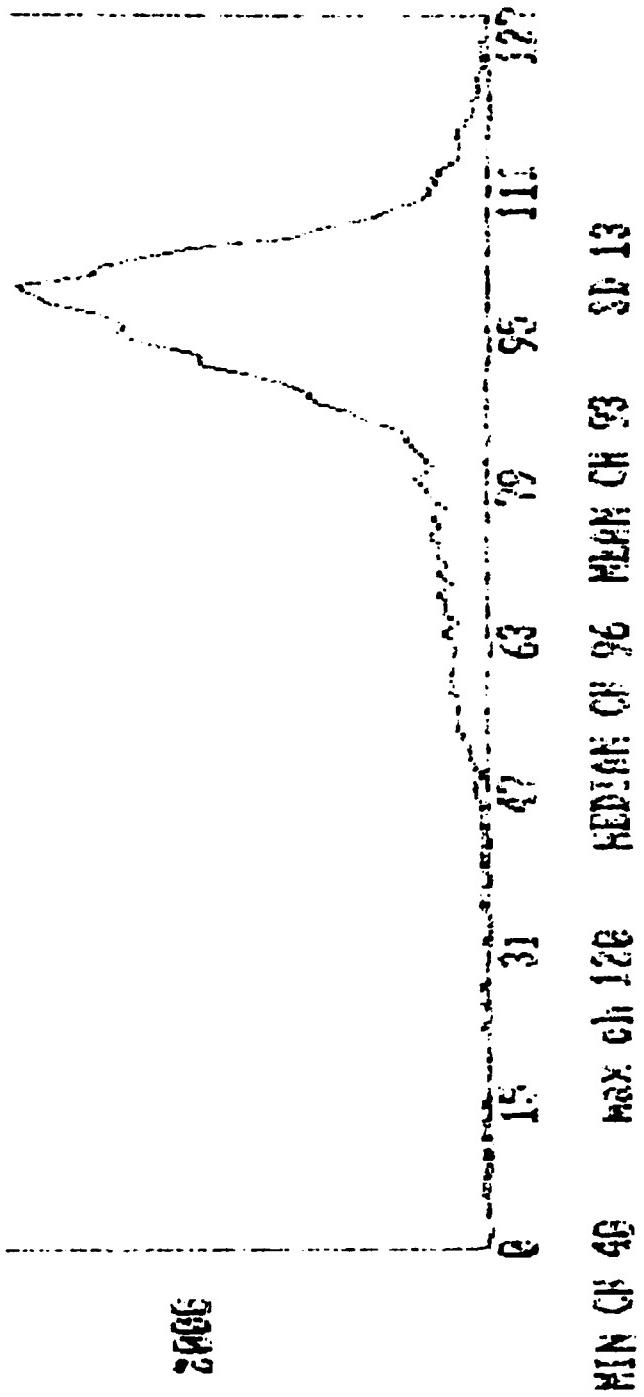


FIG. 7:  
CALIBRATION AEROSOL  
NaCl 0.04%  $S_c$

	P1	P2	P3	P4	P5	P6	P7	MED
TIME = 00:15:22:20:24.0	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00
REC	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00
MEAN	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00
TOTAL	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00
CP	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00
P0	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00
DUP	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00
CP	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00
P1	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00
P2	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00
P3	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00
P4	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00
P5	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00
P6	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00
P7	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00
MED	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00

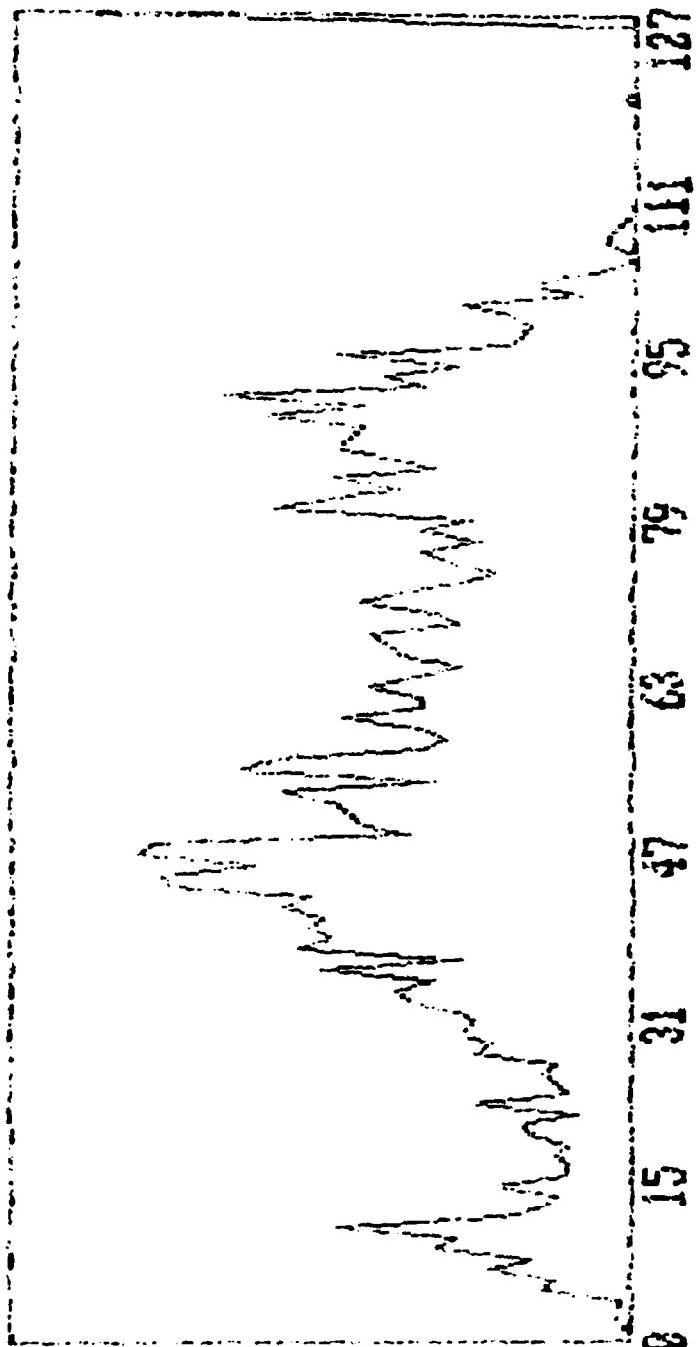


FIG. 8:  
MII CH 1 MEDIAN CH 56 MEAN CH 66

*back of neck and*

RECNUM = 366  
 TIME = 00 24 15 56 20.5  
 REC TIME FREQ CH TOTAL C36 C60 C90 C100 C110 C120  
 MEAN = 556.265 844.975 364.0 33.0 4.76 5.69 4.92 5.69 3.16 5.30 5.3  
 64.35 62.62 15.56 42.5  
 TIME = 02 24 15 56 20.5  
 MEAN = 556.00

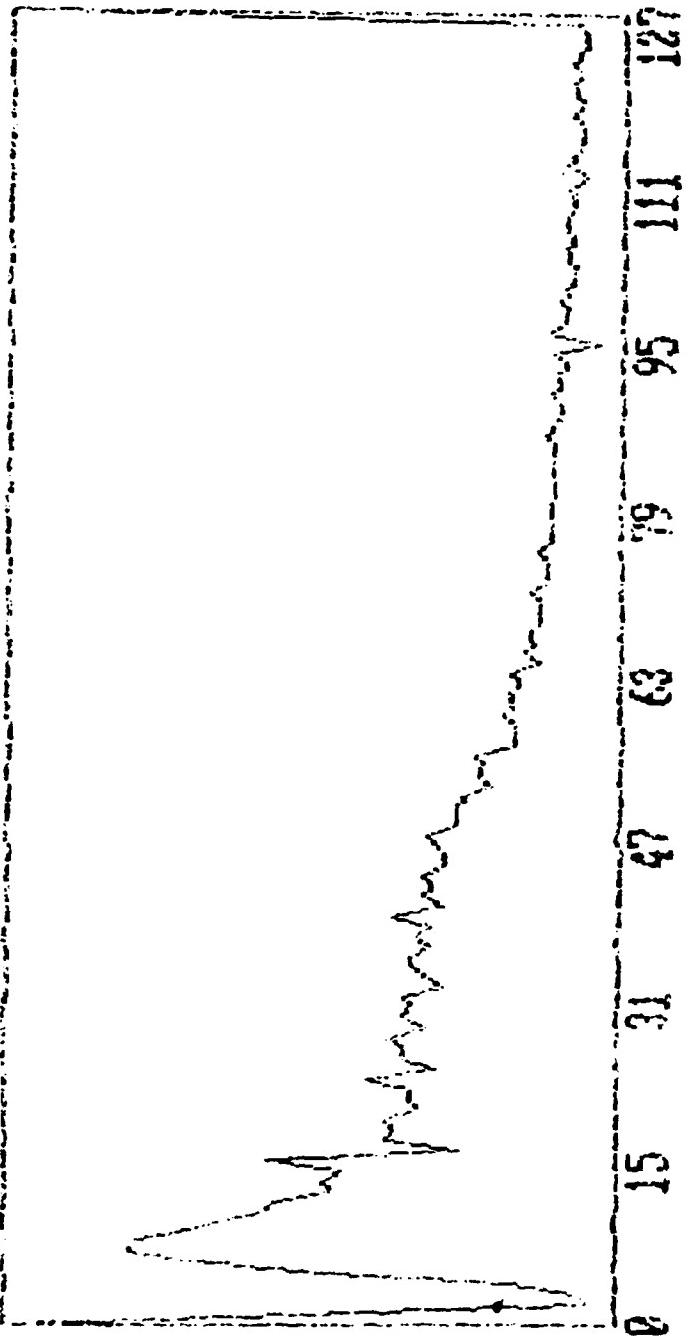


FIG. 9:

P. C. S. 1. Response - Smooth

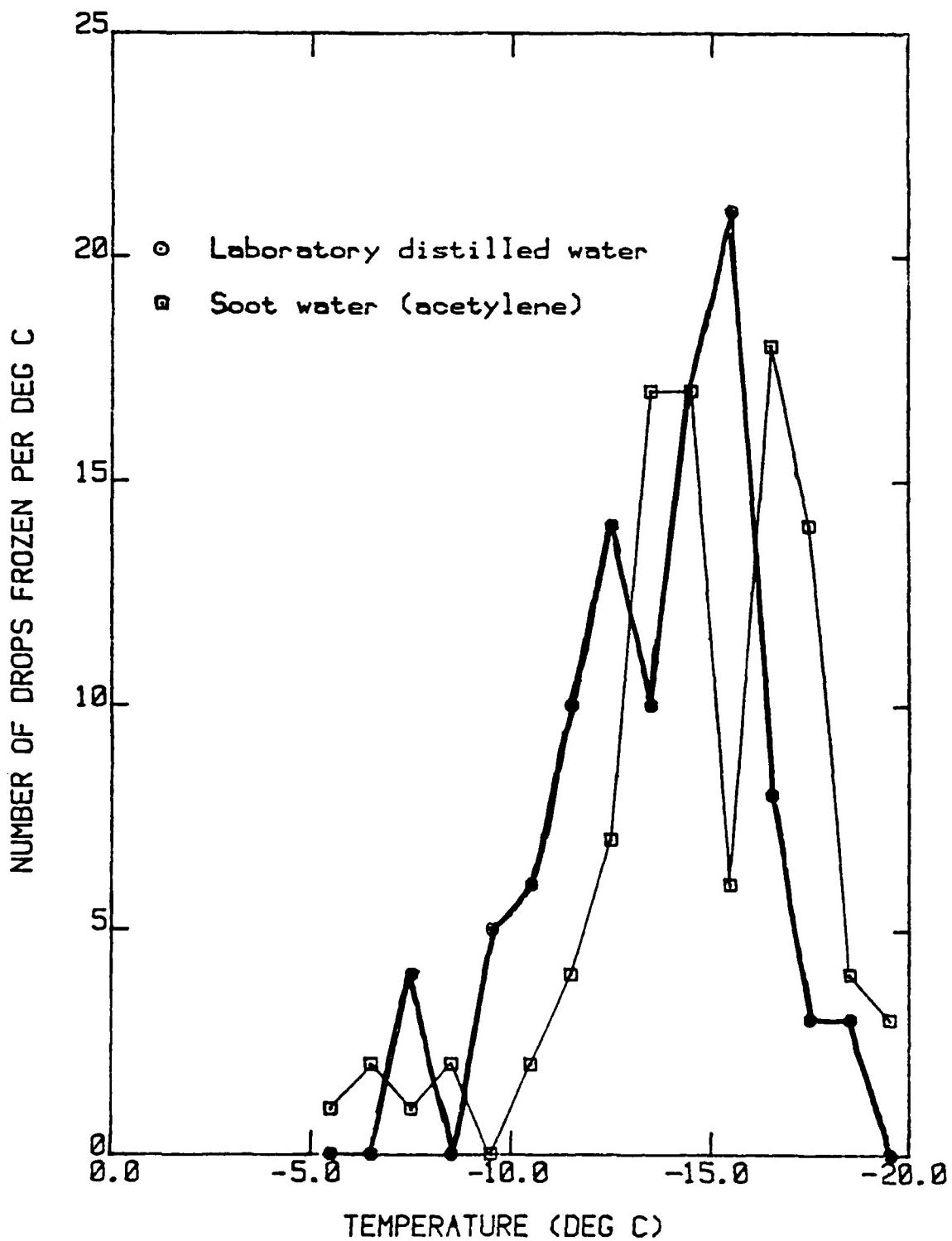
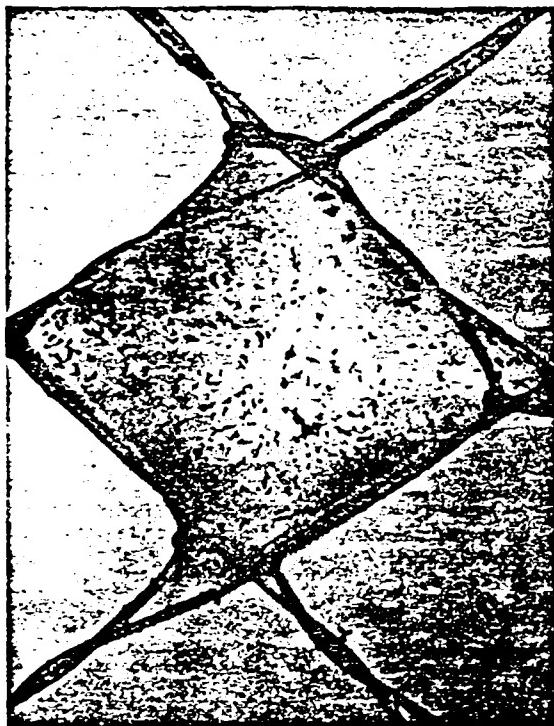
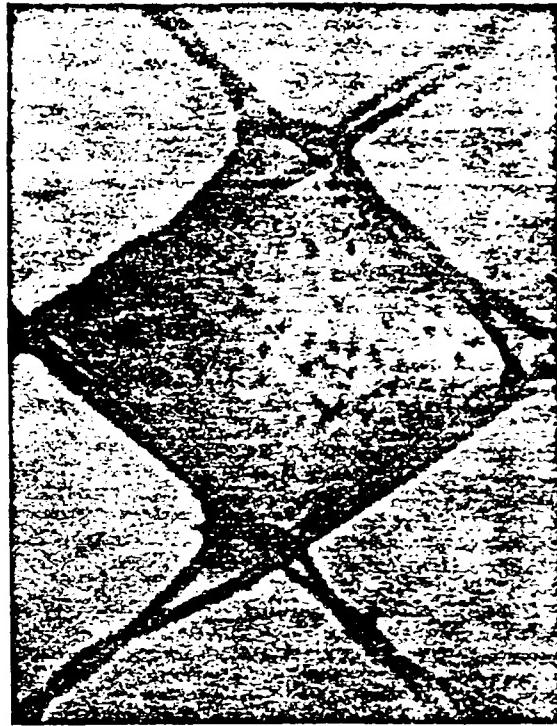


Figure 10. Number of drops frozen per degree temperature change as a function of temperature.

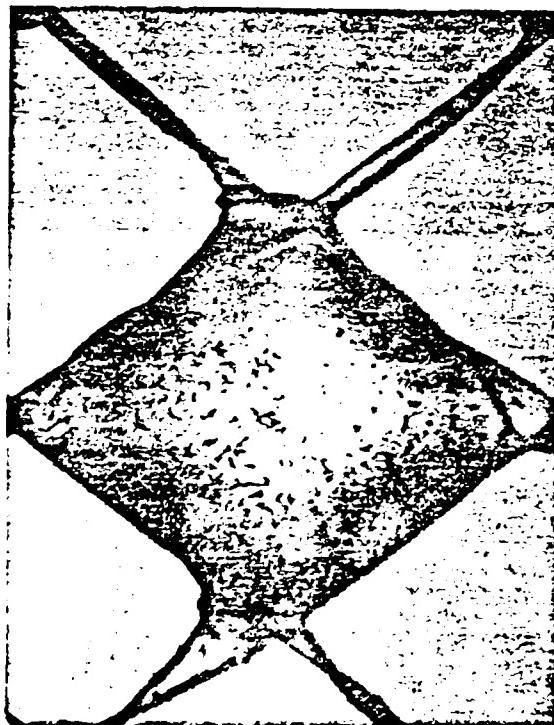
R



C



EE



E



FIG. 11: ACETYLENE SOOT ON WATER DROP SURFACE

FIG. 12: COAGULATION RATE OF AEROSOL WITH CLOUD DROPLETS, TIME CONSTANT  $\sim 1000$ s

$$\frac{dn}{n} = \frac{dt}{\frac{3\pi r^2}{kT\lambda N} \cdot \frac{1}{F}}$$

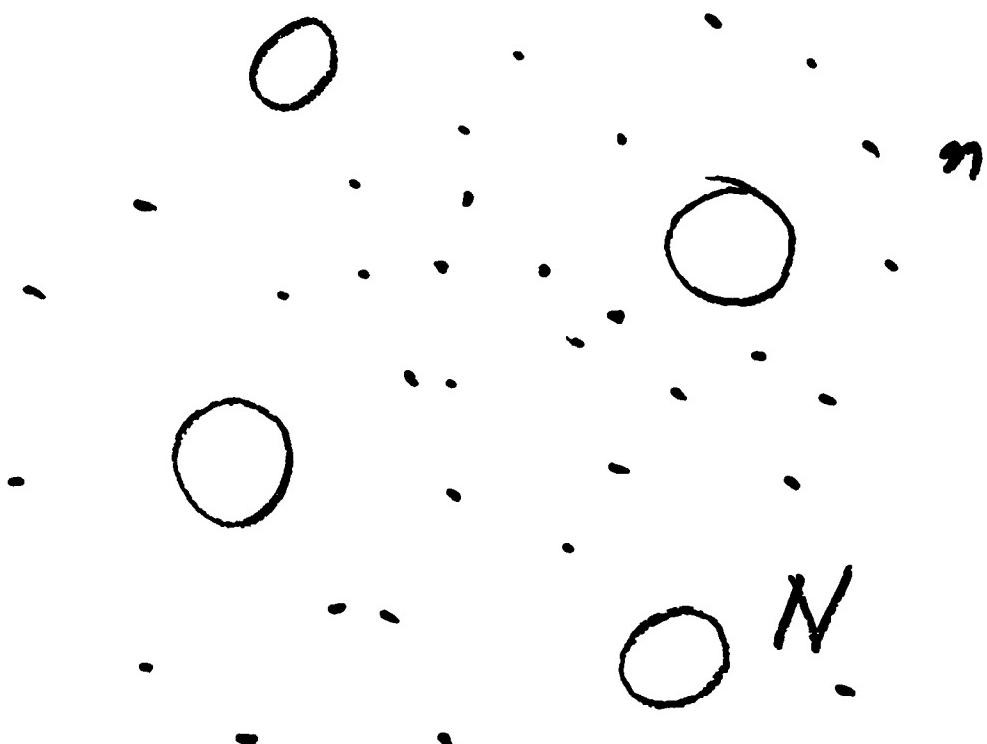


FIG. 13: SAMPLING OF RENO POOL FIRE

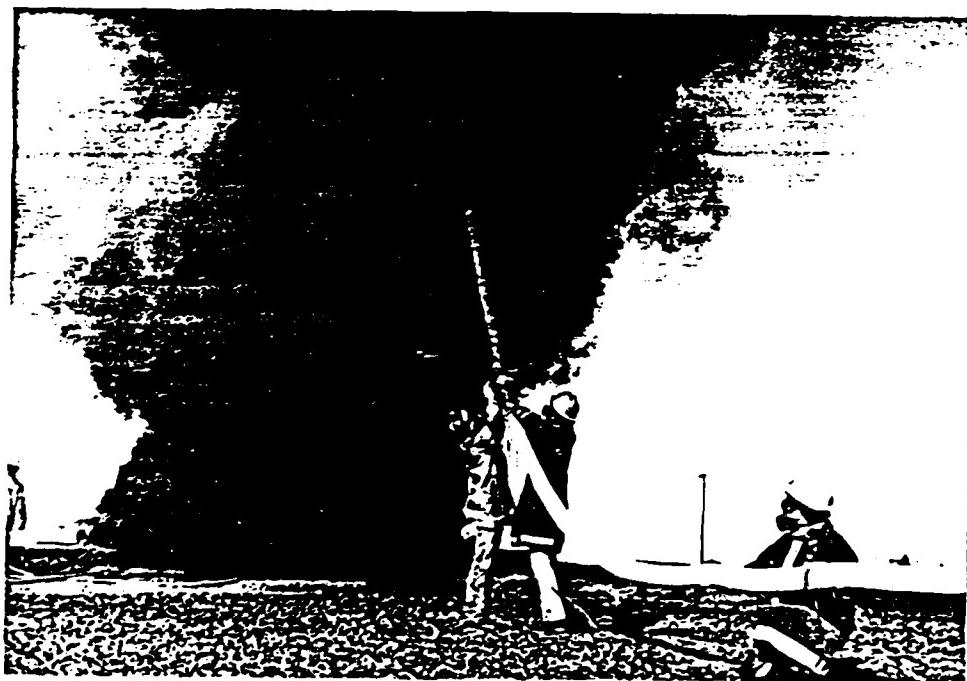




FIG. 14: SEM PICTURE OF FORMVAR REPLICATED SOOT FLAKE

Acetylene soot and/or on periphery. fig.



FIGURE 15:  
Soot particles from acetylene on the surface; (a) and protruding from the periphery; (b) of a 2 mm waterdrop.

a. Nucleate complete coverage

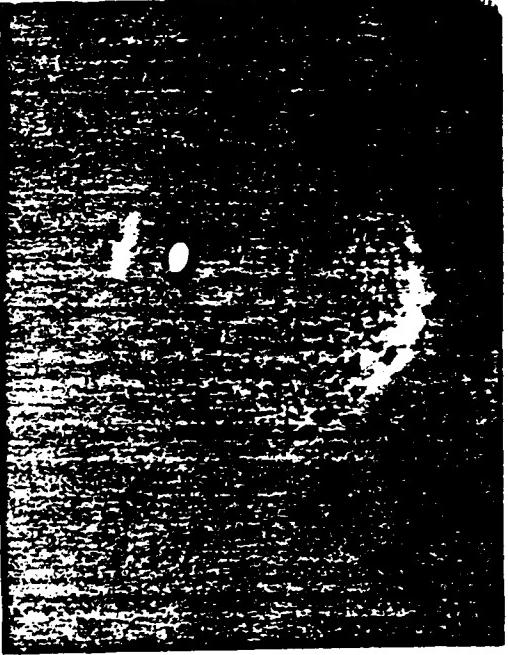
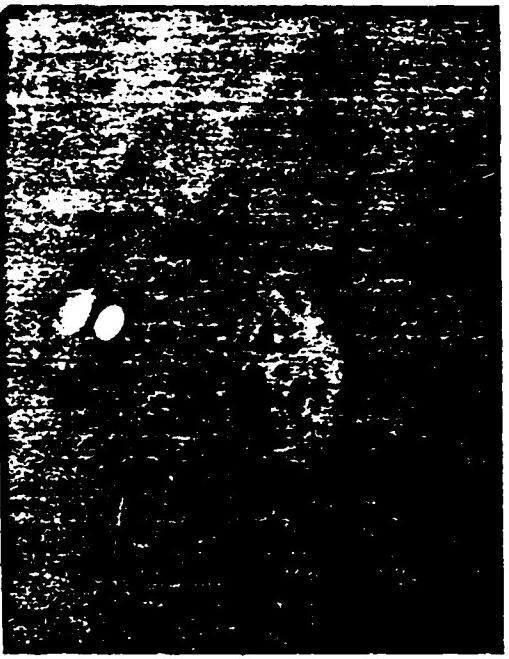
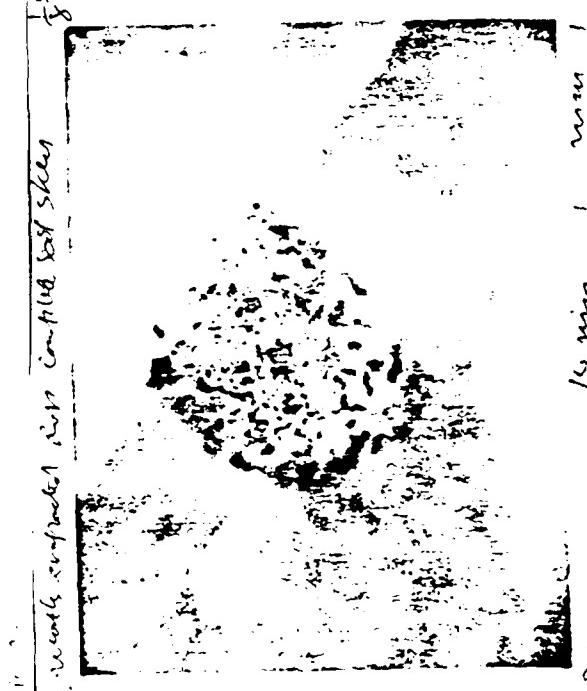


FIGURE 16:

b. Nucleate complete coverage



c. Nucleate complete coverage



c. Nucleate complete coverage

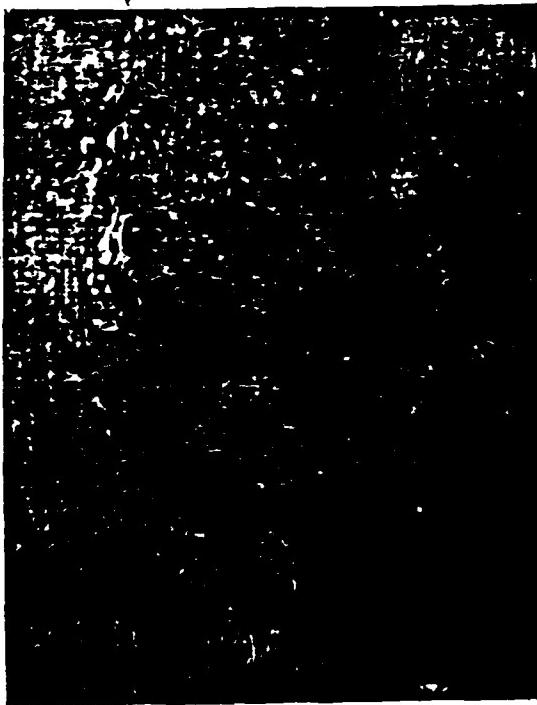


d. Drop completely surrounded by paraffin

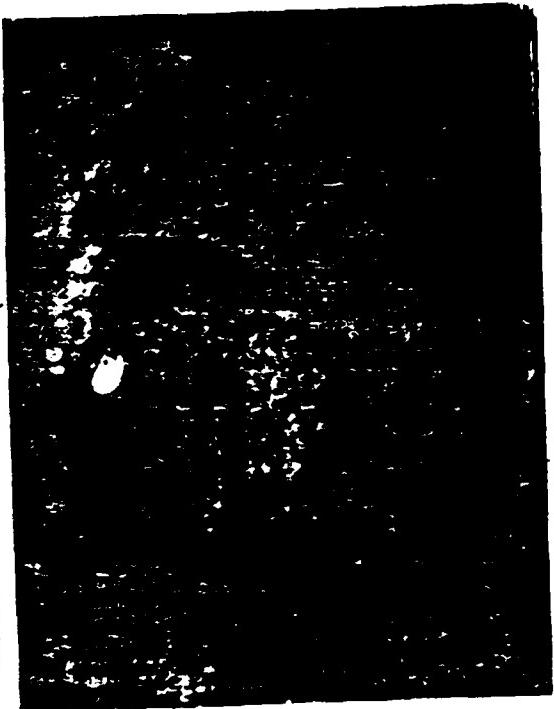
Evaporation sequence of a 2 mm soot covered waterdrop. (a) Shows an open water area (center); (b) is completely covered; (c) is beginning to shrink; (d) the shell has collapsed on the parafilm substrate.

FIGURE 17:

6. C. multi puncturing off. capsule ends



7. C. multi. no. Slight



5

C. multi. at cold skin and new drug



7

a. an open lead produced by a hypodermic needle in a soot covered waterdrop.

b. The lead persists even after the drop has dried out.

c. The dried skin is resuspended on another drop but fails to break up.

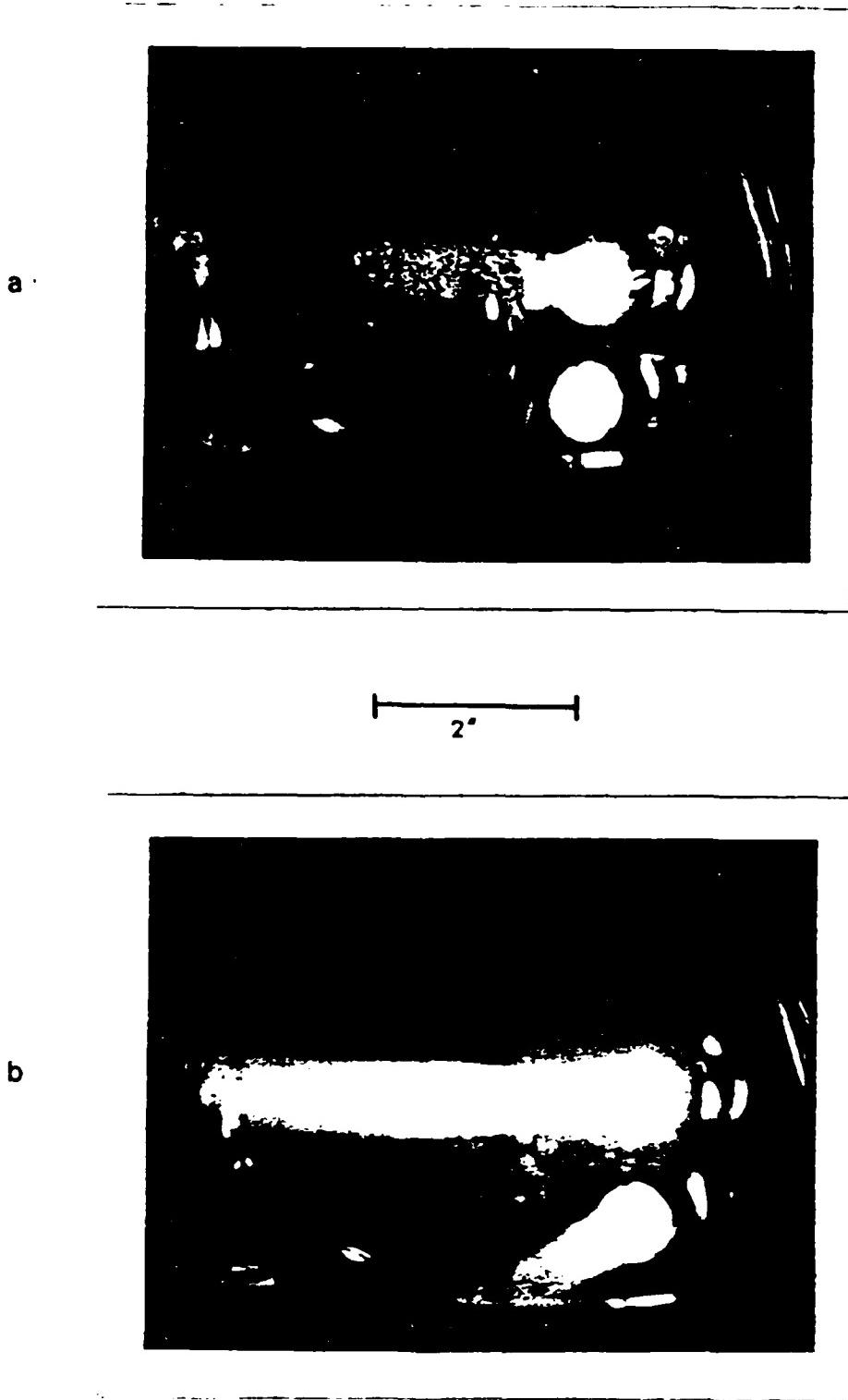


Figure 18 a) Cloud produced on natural aerosol in an expansion cloud chamber.  
b) Cloud produced on acetylene soot in an expansion cloud chamber.

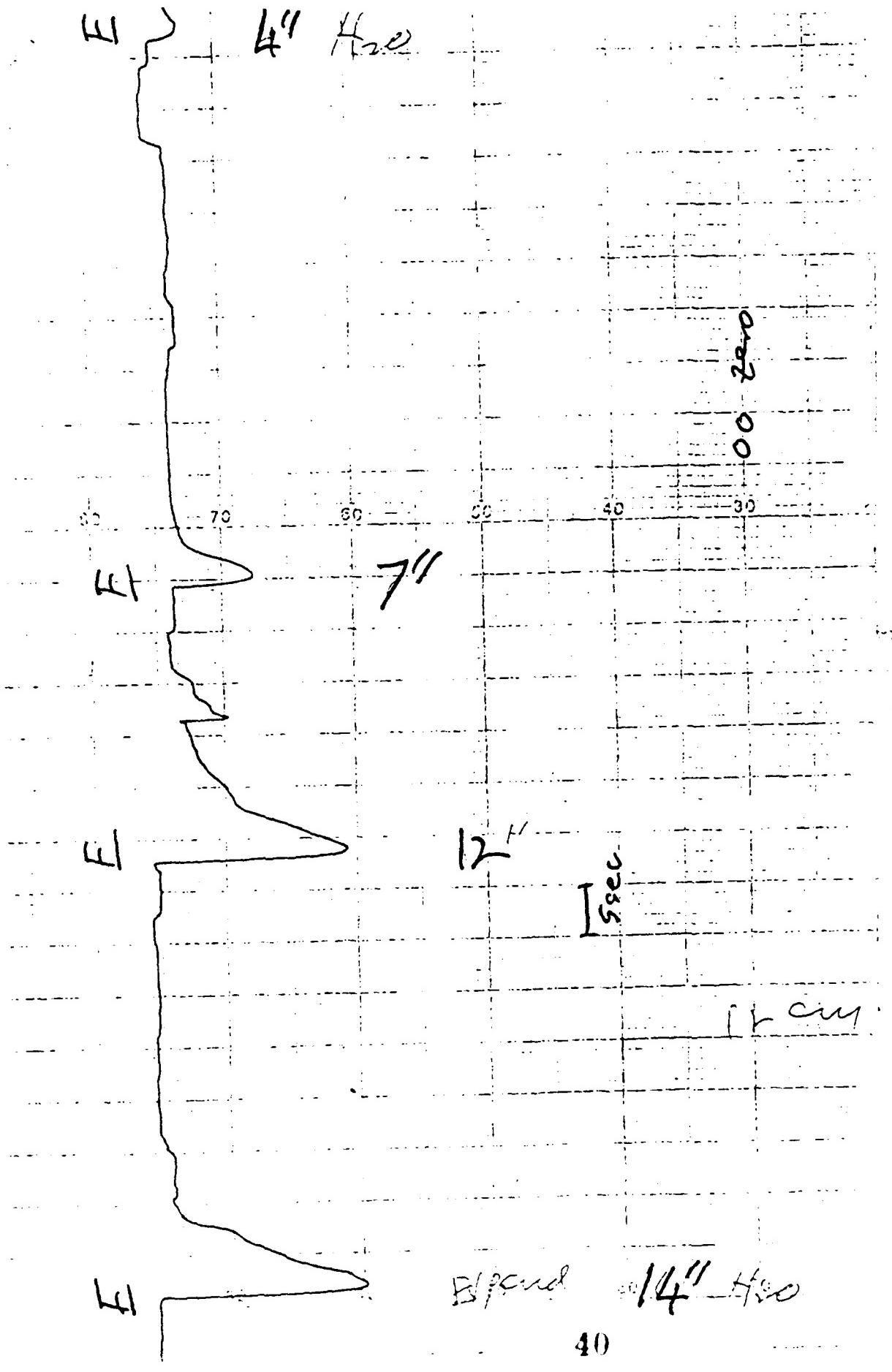


FIG. 19: EXTINCTION MEASUREMENT AFTER EXPANSION AND RECOMPRESSION FOR A ACETYLENE SOOT NUCLEATED WATER CLOUD

SOOT PARTICLE DISTRIBUTION (23-25 OCT '85)

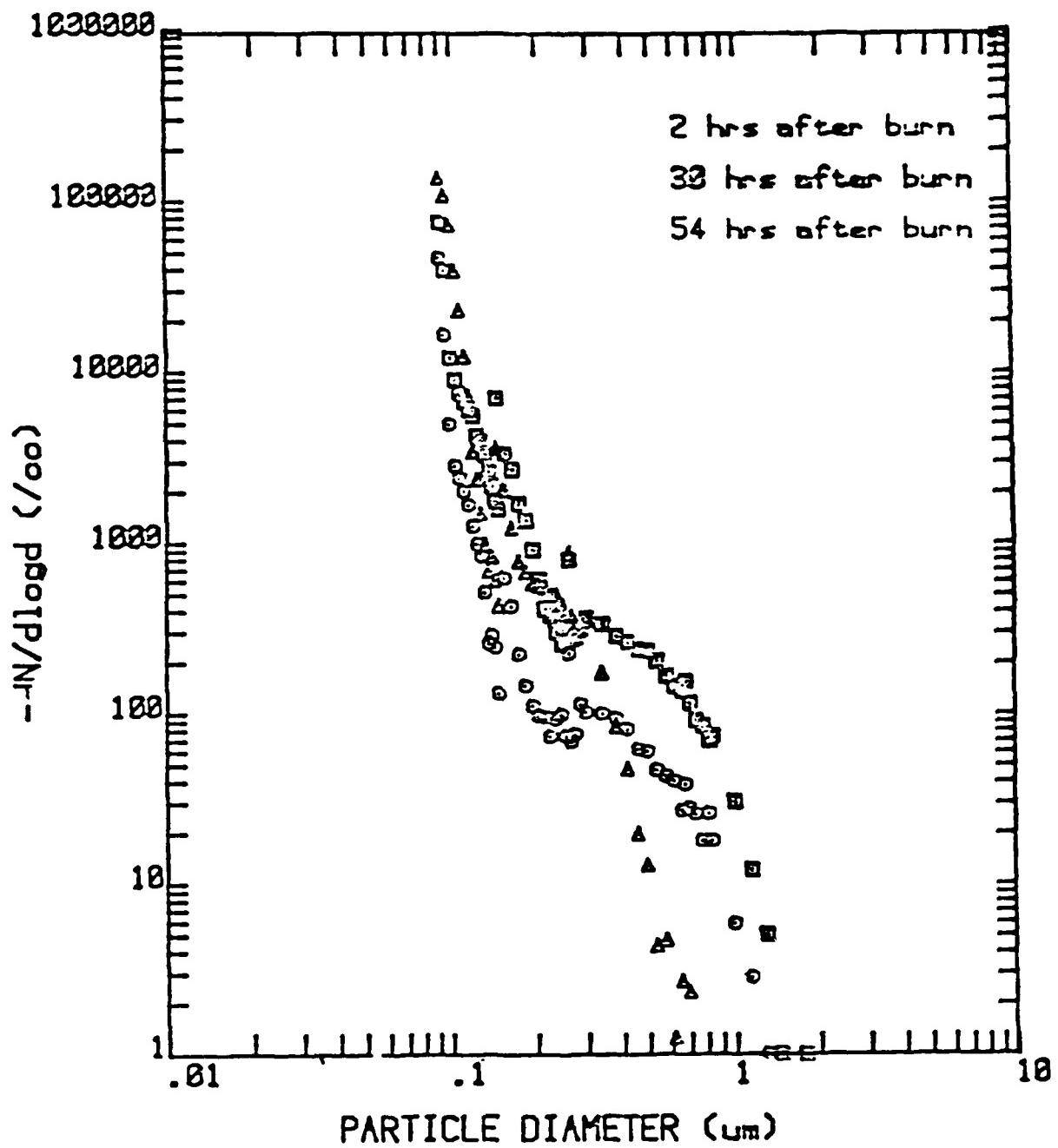
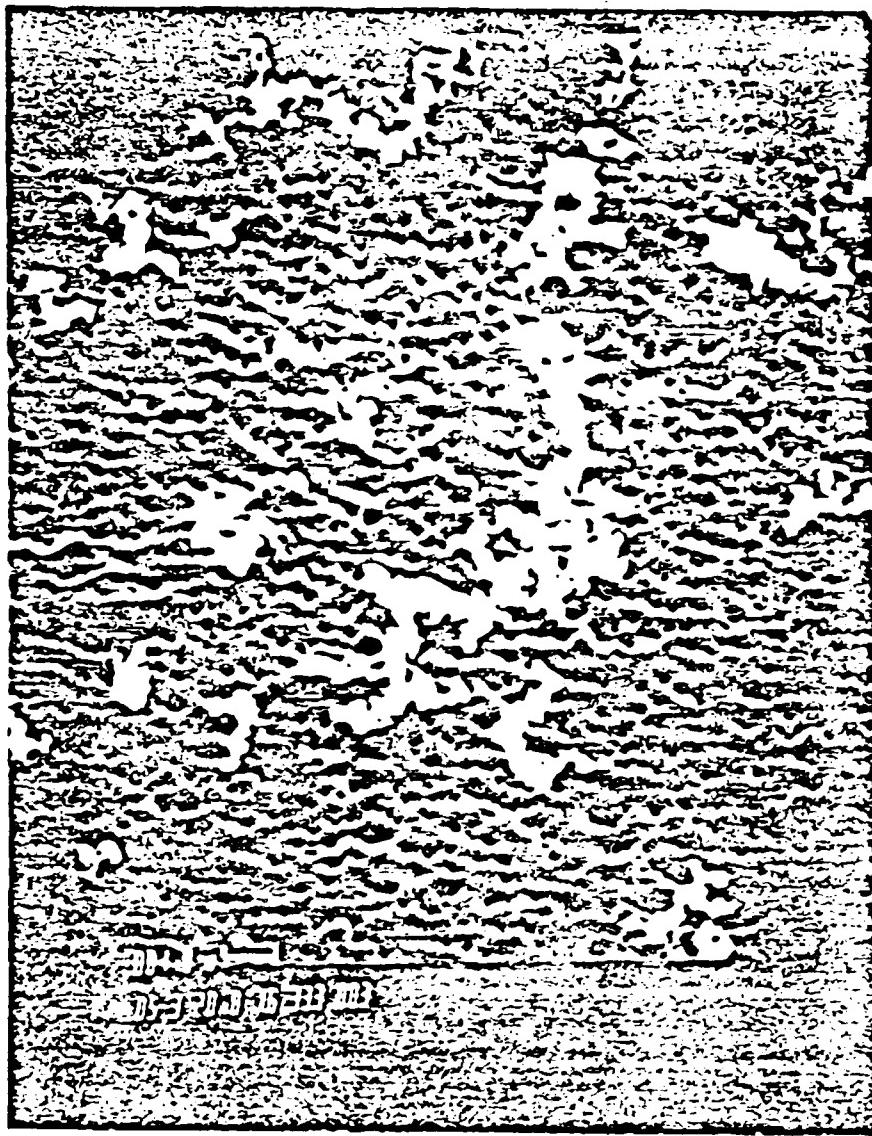


Figure 20: ASASP-X size spectra



FILTER 26

Figure 21: October 23, 1985 sample. Filter exposed  
at 5.5 hours, from 8 m<sup>3</sup> bag.

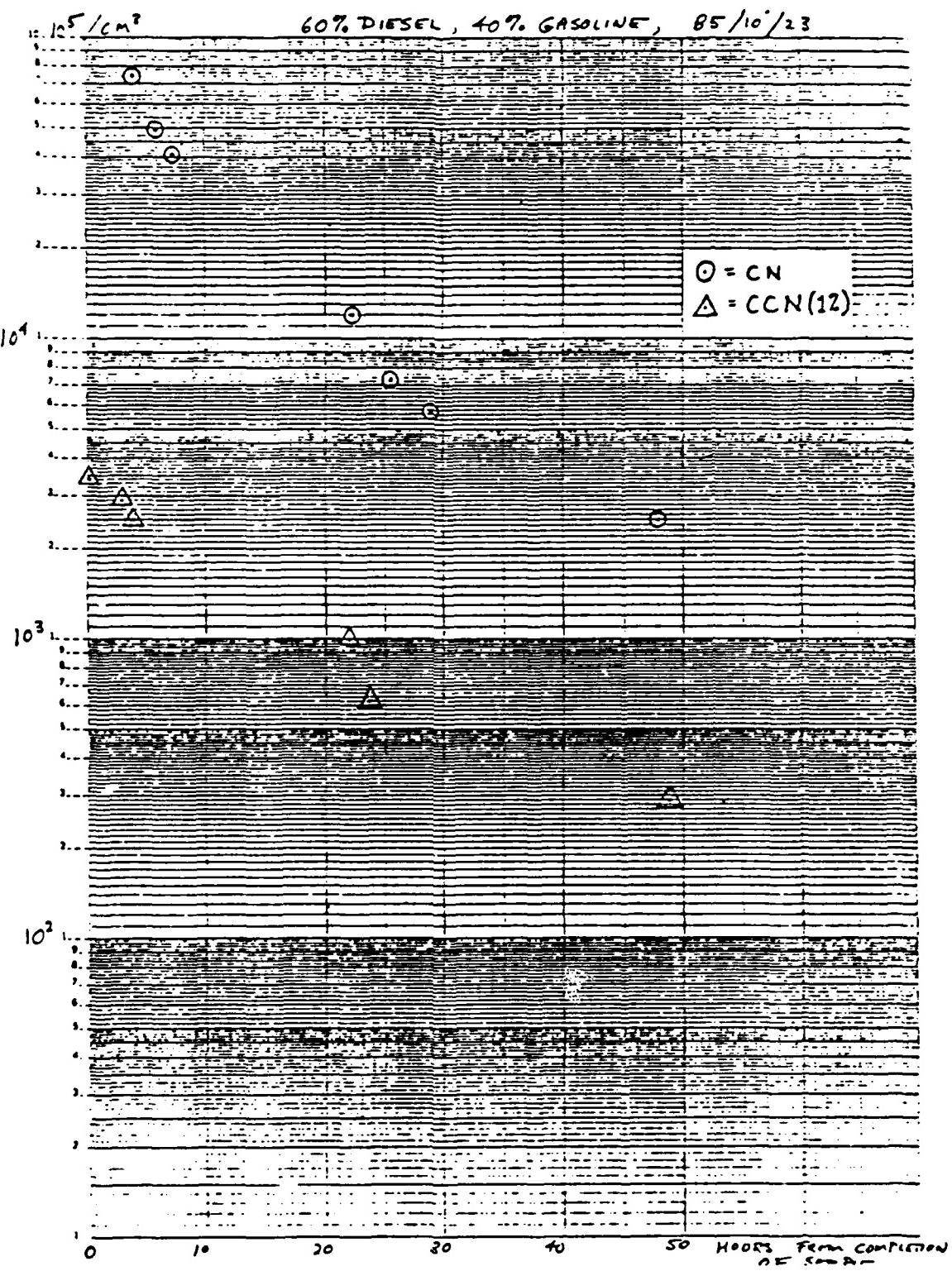
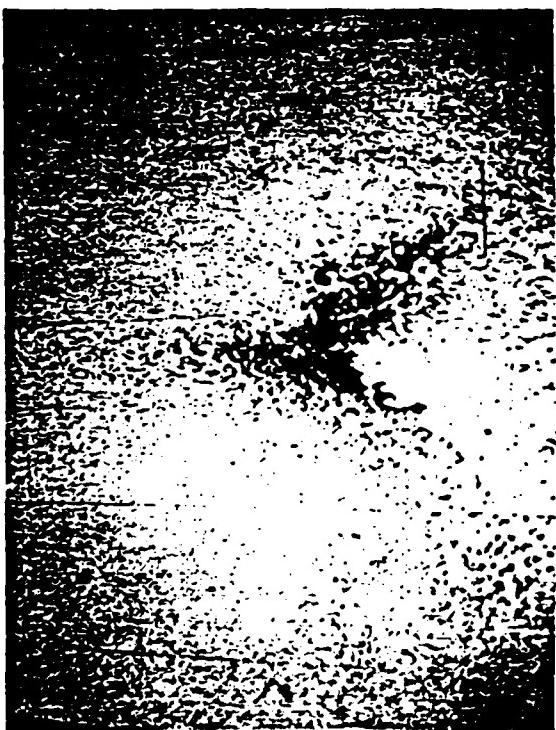
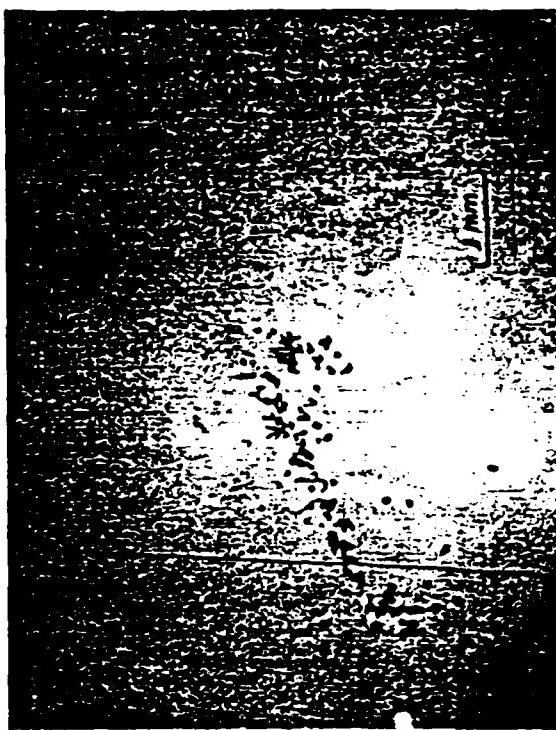


Figure 22: CN and CCN active at 1%; October 23, 1985 sample.

FIGURE 23 : SOOT FLAKES FROM ACETYLENE COLLECTED ON FORMVAR  
REPLICATOR. LABORATORY TEST.



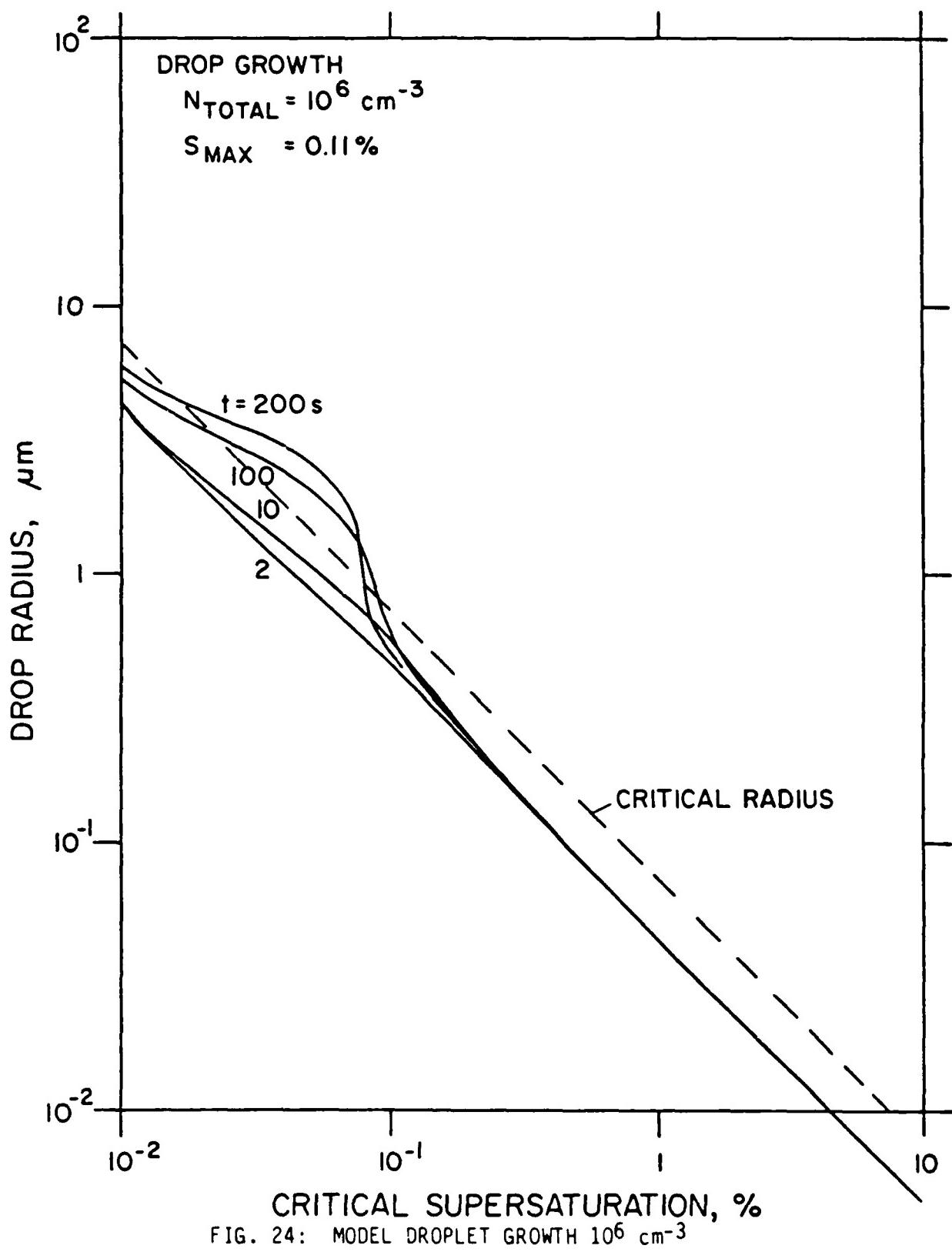


FIG. 24: MODEL DROPLET GROWTH  $10^6 \text{ cm}^{-3}$

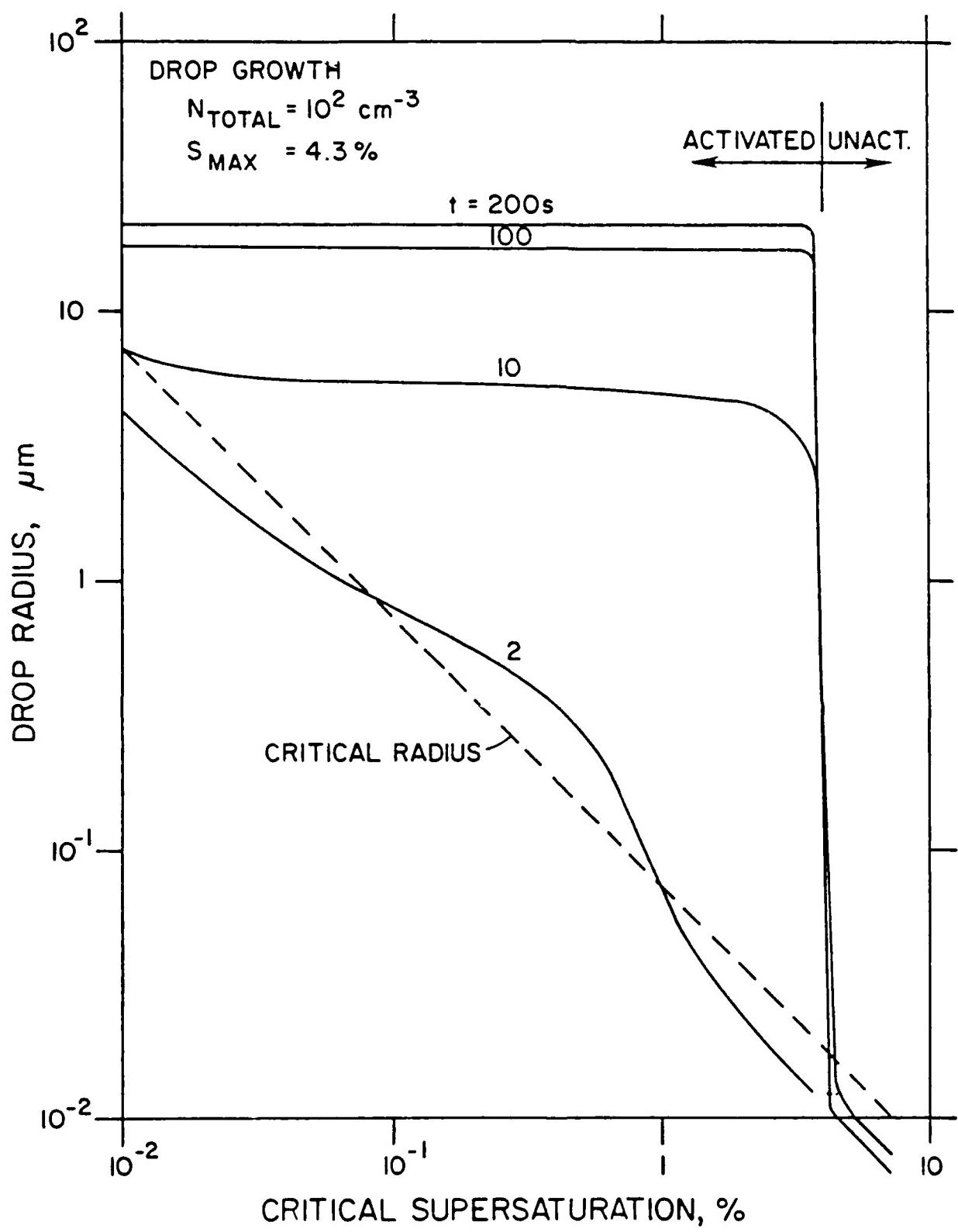


FIG. 25: MODEL DROPLET GROWTH  $100 \text{ cm}^{-3}$

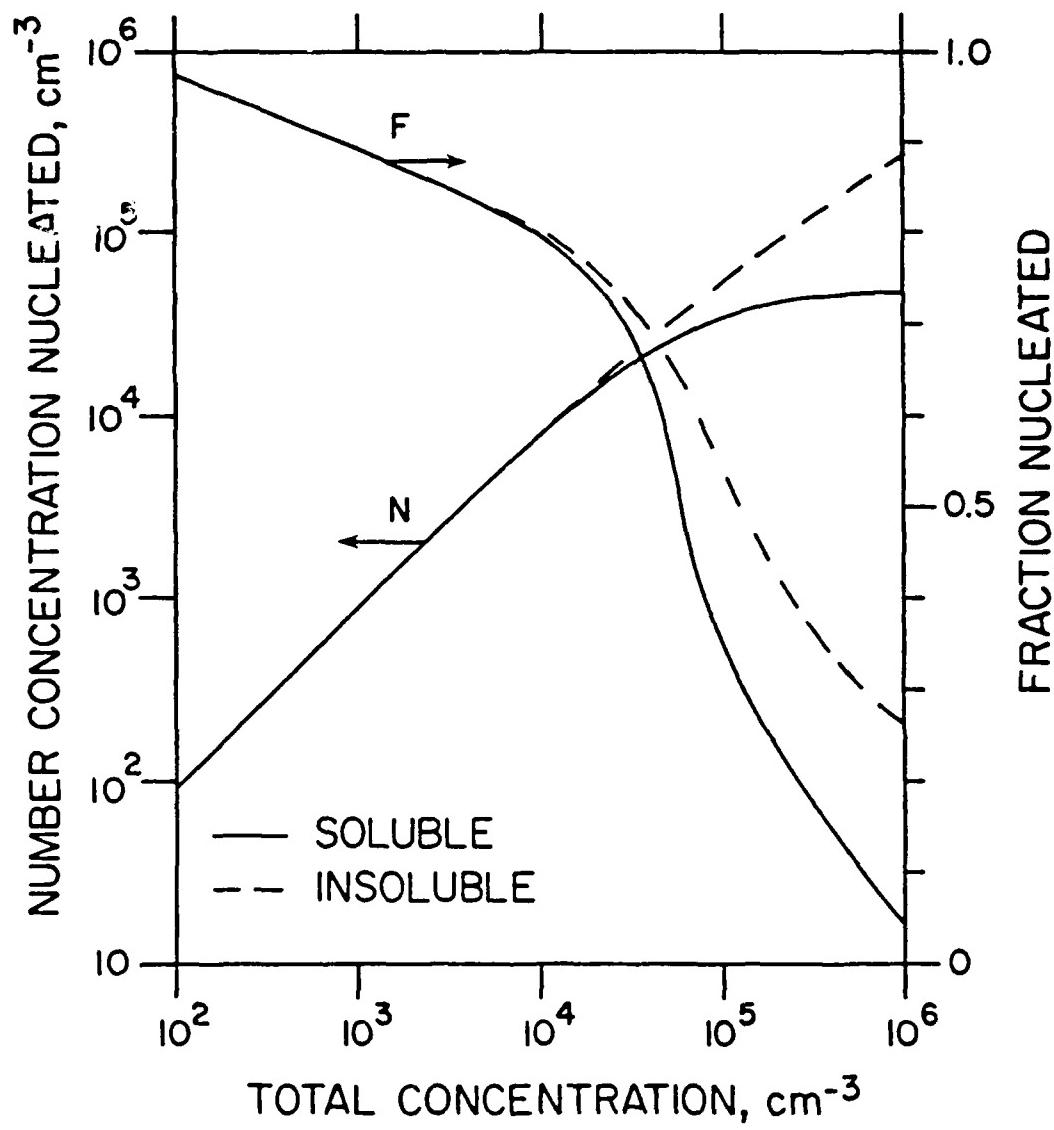


FIG. 26: MODEL NUMBER AND FRACTION ACTIVATED

## CONCLUSIONS FROM OUR STUDY:

- 1). Lab/field soot complex; 0.1 to 100  $\mu\text{m}$ .
- 2). Lab soot acetylene/crude oil  $\frac{\text{CCN}}{\text{CN}}$  ratio few to 10% increasing with time.
- 3). Ice nucleation unexciting.
- 4). Soot collects on outside of drops mm - 50  $\mu\text{m}$  as evaporation proceeds.
  - (i) Independent particles
  - (ii) lily pads
  - (iii) solidified surface
  - (iv) collapsed soccer ball
  - (v) surface resuspension
- 5). Sooty cloud doesn't change extinction over  $10^3$ s by scavenging.
- 6). CCN - cloud droplets - ice removal or coalescence; only partial pluvial constipation.

## FUTURE WORK:

- i. evaporated cloud droplets and soot coagulation.
- ii. scattering + extinction of sooty cloud.
- iii. role of growth/evaporation in scavenging process (a) droplets,  
(b) ice
- iv. soot in a supercooled cloud - rime growth.

## FIELD:

- i. soot from oil fire. CN, CCN, replica.
- ii. cloud droplets on replica, modification and scavenging in cap cloud. (Lodi?)
- iii. deeper cloud and ice role in scavenging. Ice from replica (Chapleau?)

## MAJOR PROBLEM:

Firestorm dynamics to give right microphyics for maximum scavenging.

Chemical Scavenging in the Atmosphere: Smoke-Ozone Reactions.

N. deHaas, R. M. Fristrom, M. J. Linevsky and D. M. Silver  
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Laurel, Maryland 20707

ABSTRACT

A laboratory apparatus has been constructed to examine the reaction of smoke particles with ozone. A stopped-flow experimental procedure has been developed that permits (1) a measurement of the total extinction of a He-Ne laser beam by ozone (Chappuis band) and/or by various sub-micron smoke particles; (2) a measurement of the right-angle scattering of a Nd-Yag laser by the smoke particles; and (3) a sampling of particulates for examination under the scanning electron microscope. Ozone (without smoke) is stable in the apparatus over a period of hours; smoke (without ozone) is also stable over a period of hours. Smoke-ozone mixtures begin decaying immediately and, under the experimental conditions employed, they have half-lives of a half-hour to two hours depending on the fuel source of the smoke particles. Specifically, the half-life is defined to be the time in which the optical extinction at 6328 Å is reduced to one-half its initial value. Translating these results to "nuclear winter" atmospheric conditions yields half-lives in the range of a month. These experiments provide solid qualitative evidence for the existence of atmospheric chemical scavenging of smoke and ozone.

This work was supported in part by the Naval Sea Systems Command under Contract N00024-85-C-5301.

Submitted for presentation at the DNA Global Effects Program Technical Meeting, 25-27 February 1986 by D. M. Silver, (301) 953-6265.

TABLE V - SOOT CLOUD SCAVENGING (LIMITED BY TURBULENT ENTRAINMENT WHICH  
IN TURN IS DIFFUSION LIMITED BELOW THE MINIMUM MIXING SCALE LENGTH)

Alt (Km)	$t_{1/2} (O_3)$ Scale 1 Meter sec	$t_{1/2} (O_3)$ Scale 10 Meters sec
8	1170	$1.2 \times 10^5$ ( $1\frac{1}{2}$ da)
10	500	$5 \times 10^4$ ( $\frac{1}{2}$ da)
12	160	$1.6 \times 10^4$ (5 hrs)

$t_{1/2}$  Number of C atoms in Minimal Eddy

$$t_{1/2} = \frac{\text{Minimal Eddy Area} * \text{Diffusion Velocity} * \text{Concentration} * \text{Reaction Probability}}{1/2 v_E * f * N_T}$$

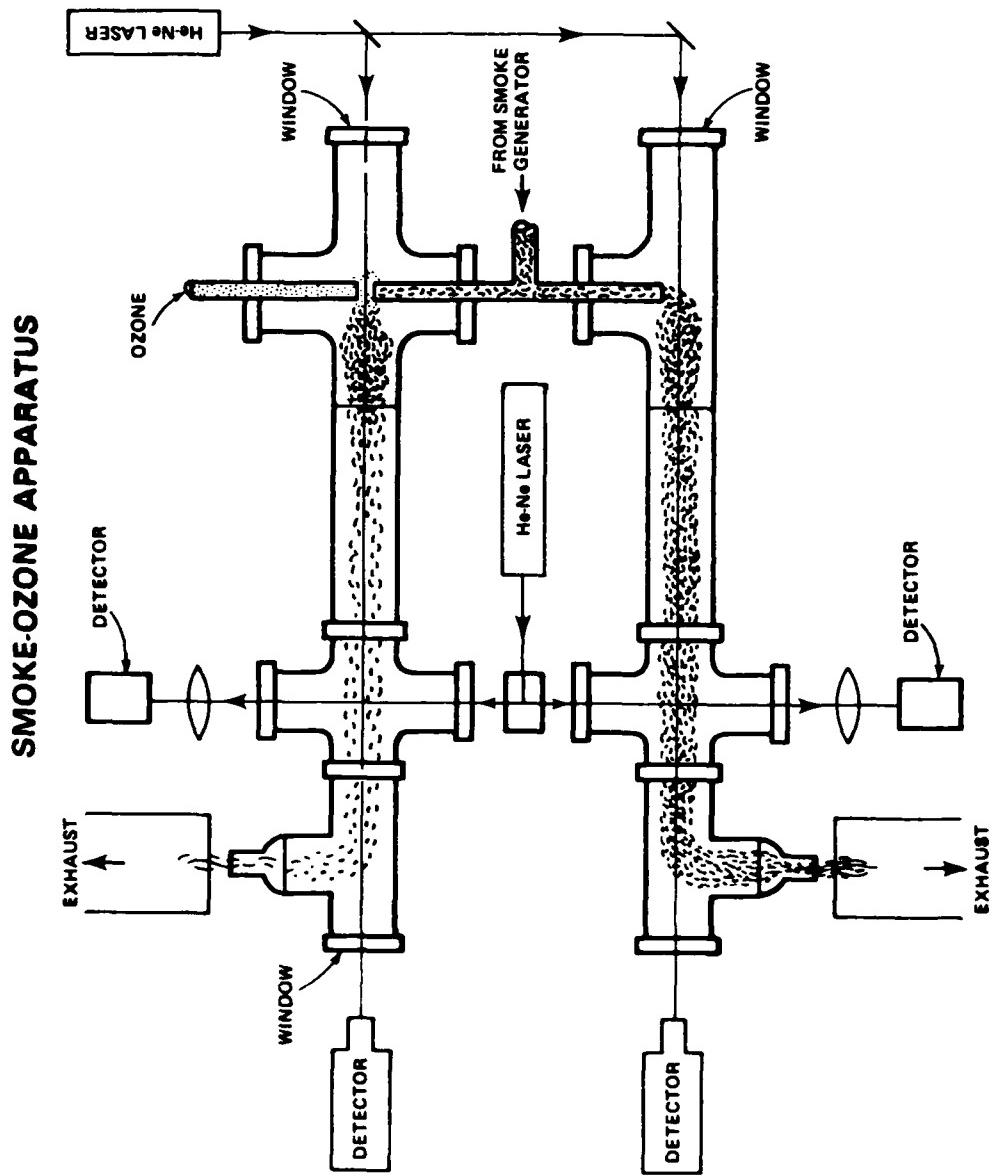
$$t_{1/2} = \frac{1/2 v_E * f * N_T}{A_E * v_D * N_i}$$

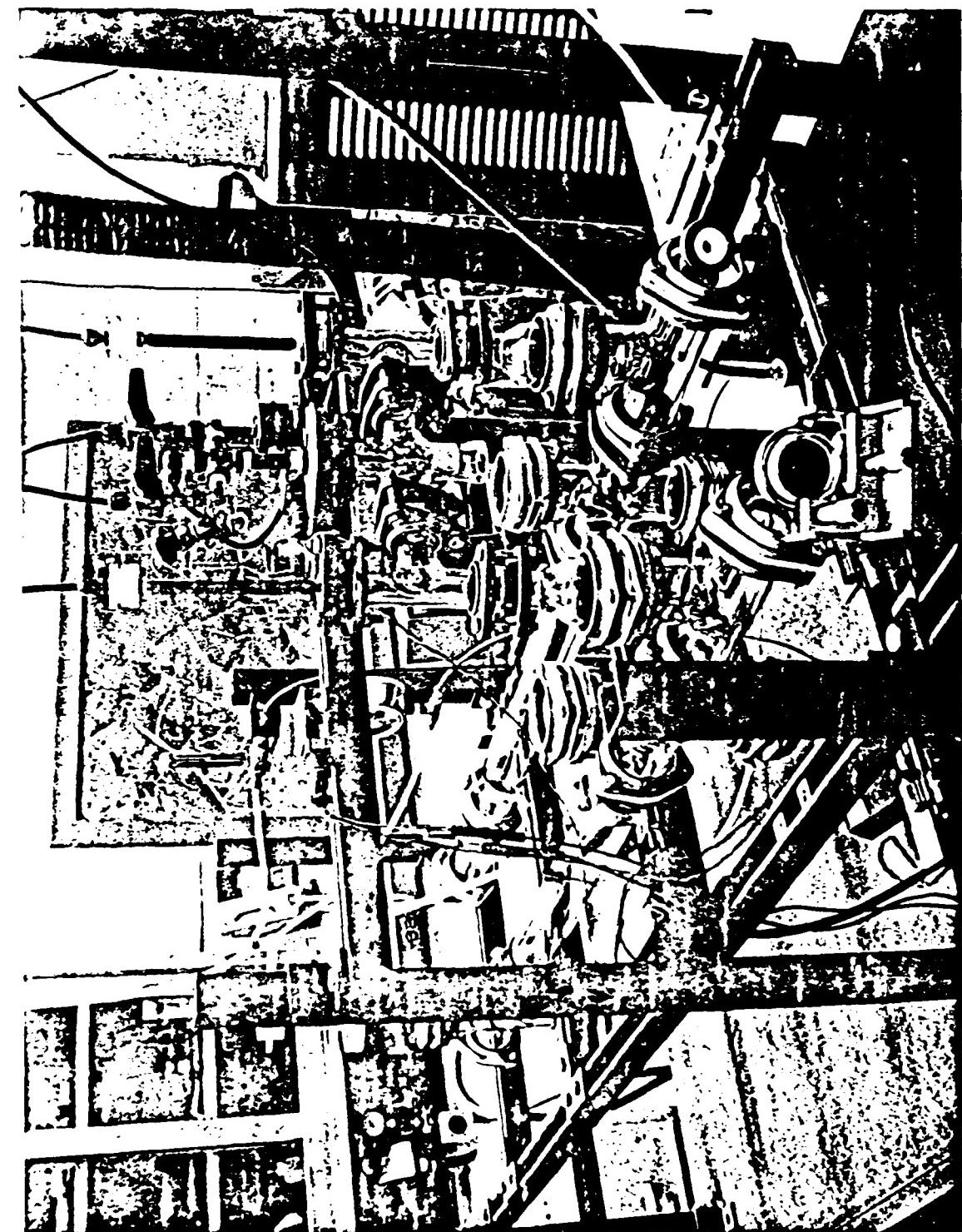
$$t_{1/2} = 1 \times 10^{-4} P * (300/T)^{1.67} * N_T / N_i * d_E^2$$

$$v_E/A_E = d_E/6$$

$$f = 4.8 \times 10^{-3} \quad (\text{Ref. C. Fenimore, Chemistry of Premixed Flames, Pergamon (1964) p. 65})$$

$$v_D = 2D_i/L_c = \left( \frac{2D_0}{PL_c} \right) \left( \frac{T}{300} \right)^{1.67}$$





Smoke - Ozone Apparatus



Cellulosic Smoke (SEM 3600 $\times$ )



Xylenic Smoke (SEM 3600 x)

Transmission of light through smoke column

$$I = I_0 \exp(-abc_s)$$

$a$  = extinction coefficient

$b$  = length of smoke column

$c_s$  = concentration of smoke

$\ln(I_0/I)$  is linearly proportional to  
smoke concentration

$$\ln(I_0/I) = abc_s$$

Rate of disappearance of smoke

$$-\frac{dc_s}{dt} = k c_0 c_s$$

$k$  is second-order rate constant

$c_0$  is concentration of ozone

$c_s$  is concentration of smoke

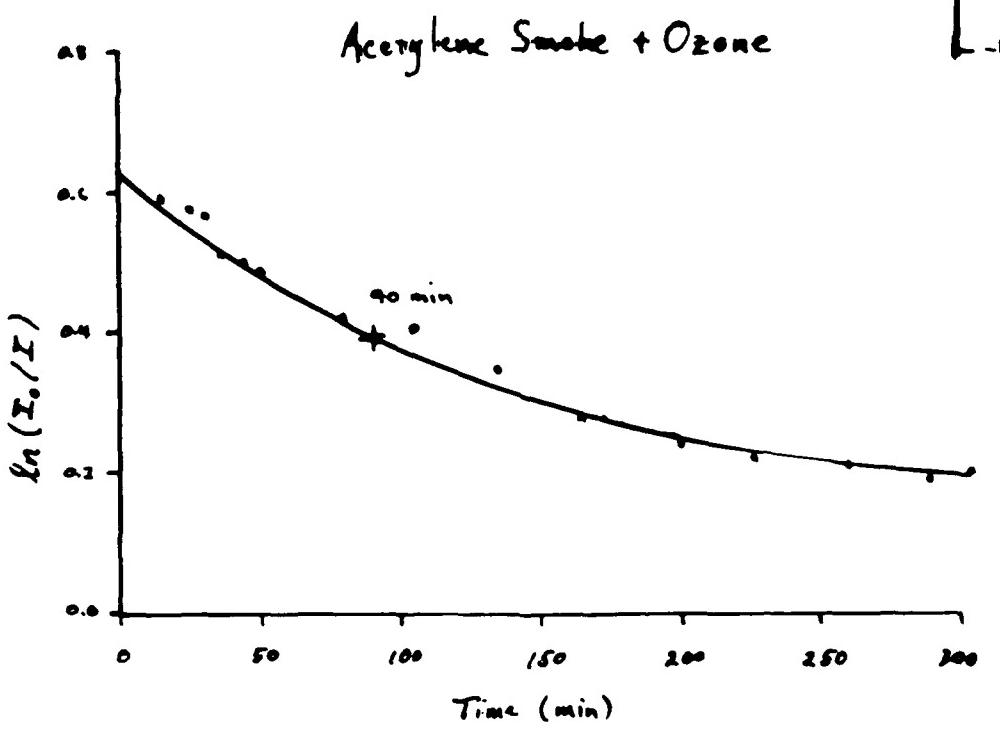
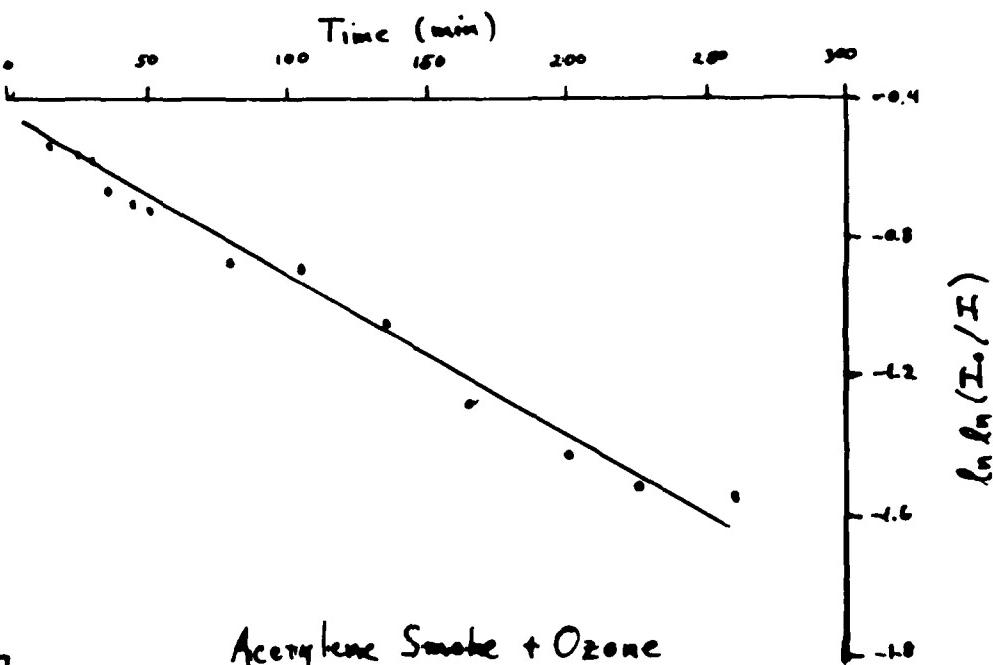
With excess ozone, rate is pseudo first-order in smoke concentration: ( $c_0$  assumed constant)

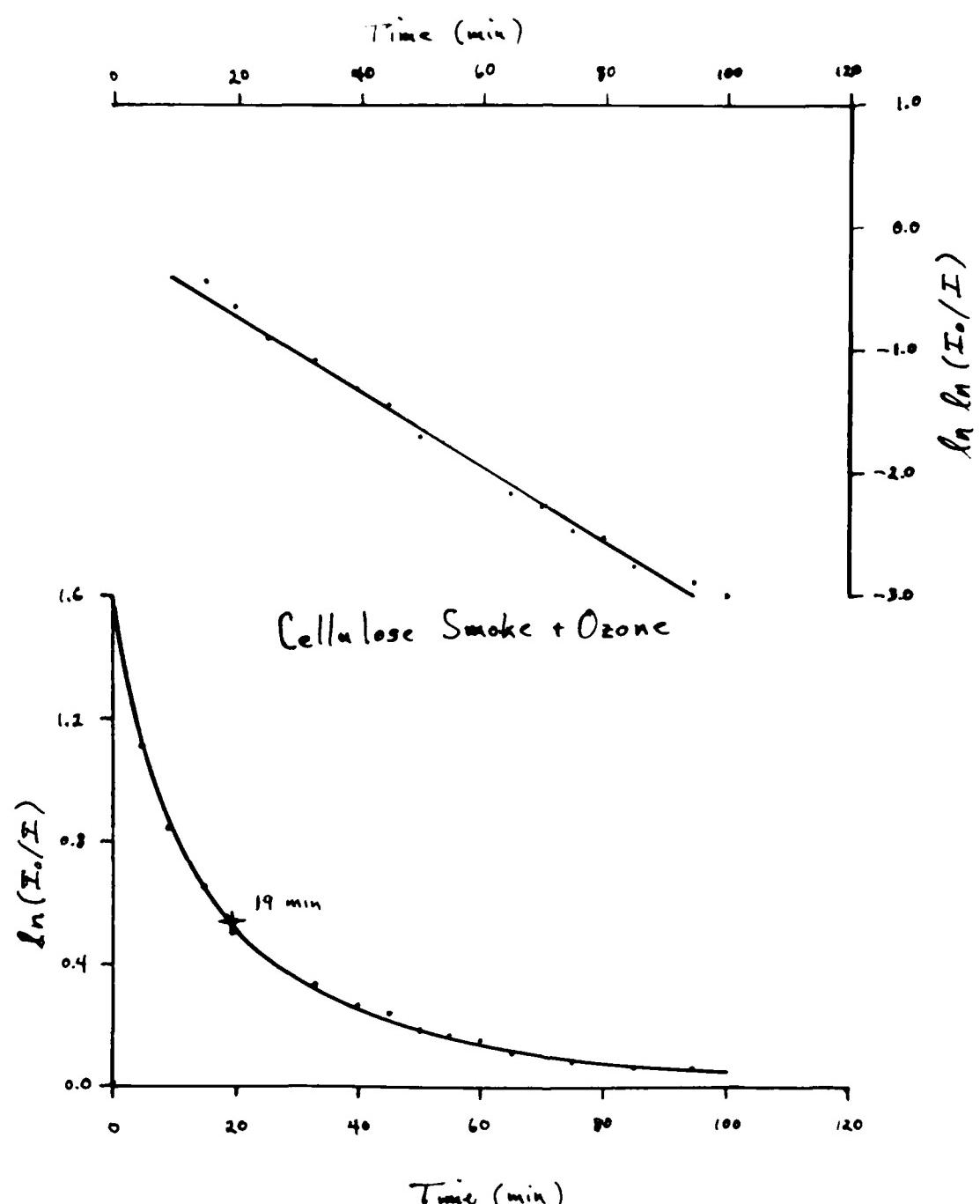
$$\therefore c_s = A \exp(-k c_0 t)$$

But  $\ln(I_0/I)$  is proportional to  $c_s$

$$\therefore \ln \ln(I_0/I) = A' - k c_0 t$$

i.e., linearly proportional to time  
if the reaction is pseudo  
first-order in smoke concentration





Ozone (without smoke) is stable in apparatus over a period of hours (no change in absorption in the Chappuis bands).

Smoke (without ozone) is stable in apparatus over a period of hours (no change in optical extinction at 6328 Å).

Smoke-ozone mixtures show a time-dependent decay in optical extinction at 6328 Å

- Half-life depends on source of smoke
- Kinetics appear to have first-order dependence on smoke concentration

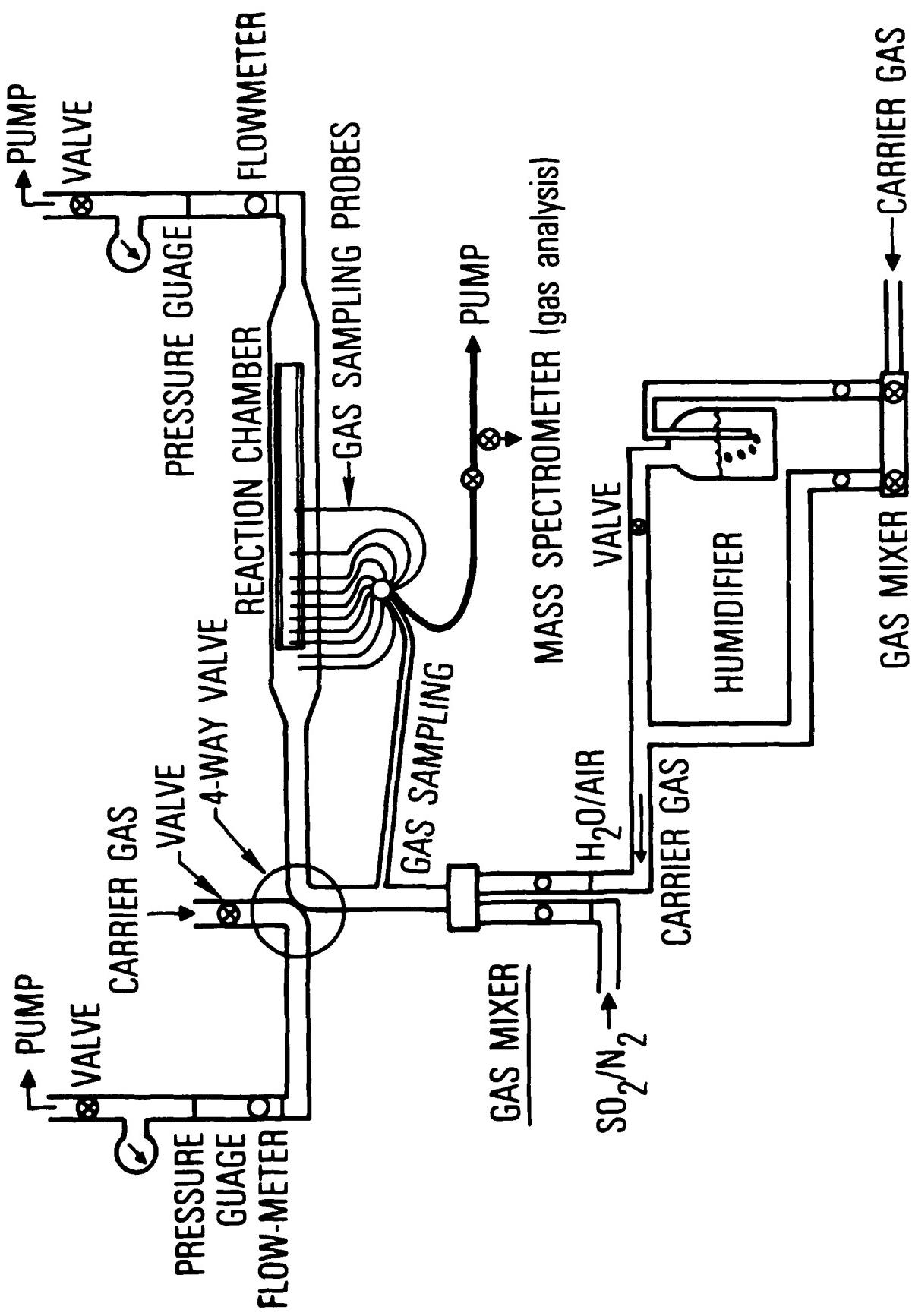
Translated to "nuclear winter" atmospheric conditions, optical extinction is estimated to be reduced to one-half its initial value in a time range of a month.

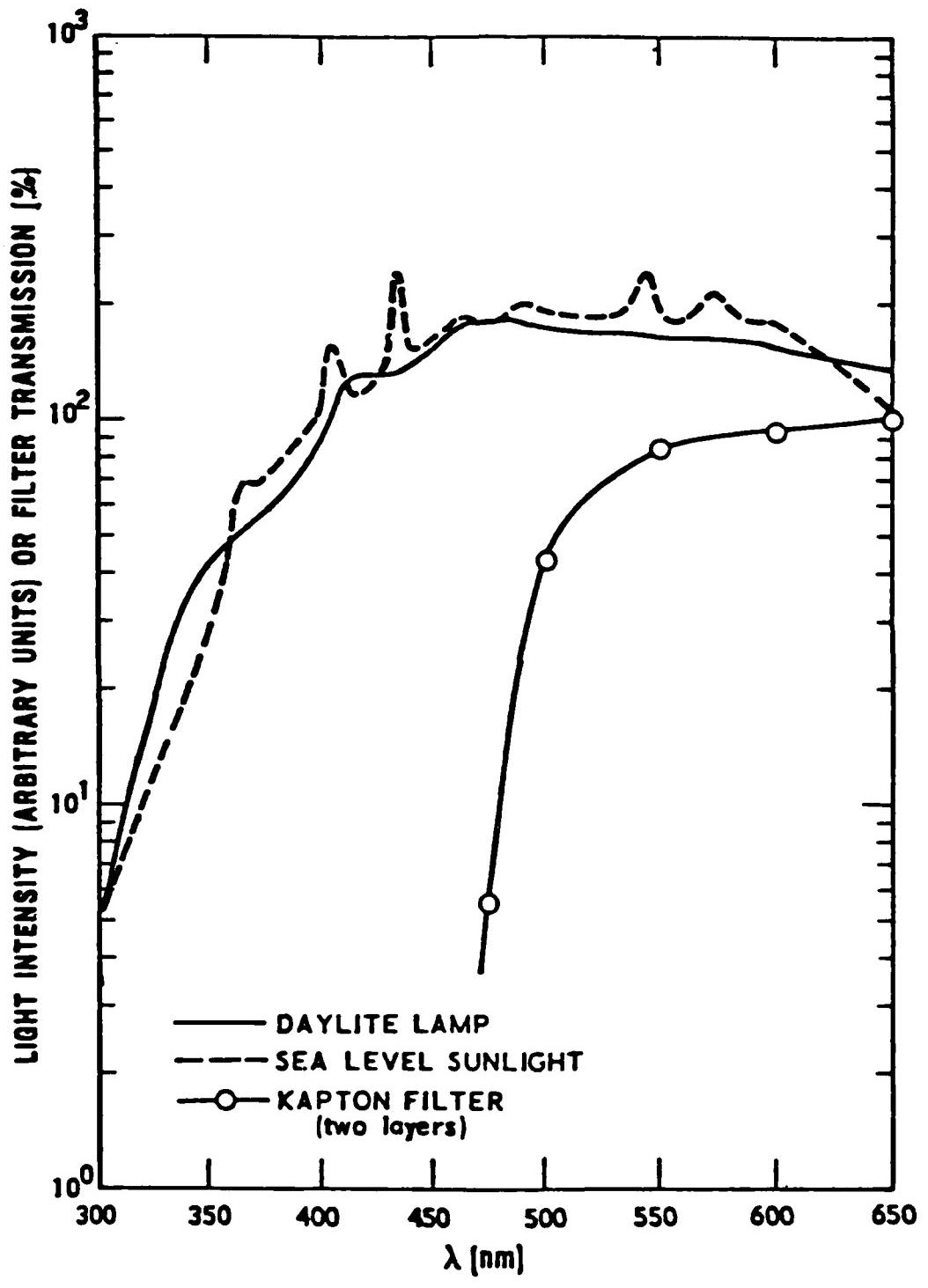
REACTION OF STRATOSPHERIC CARBON AEROSOL WITH OZONE

L. R. MARTIN, H. S. JUDEIKIS, R. B. COHEN

THE AEROSPACE CORPORATION







## CHEMICAL ANALYSES

- GAS PHASE
  - ON-LINE MASS SPEC
  - CO, CO<sub>2</sub>, OTHERS
- ABSORBED PRODUCTS
  - DESORPTION AND MASS SPEC
  - DISSOLUTION AND HPLC

$$D \left( \frac{\partial^2 C}{\partial R^2} + \frac{1}{R} \frac{\partial C}{\partial R} + \frac{\partial^2 C}{\partial Z^2} \right) - V_z \frac{\partial C}{\partial Z} = 0$$

WHERE

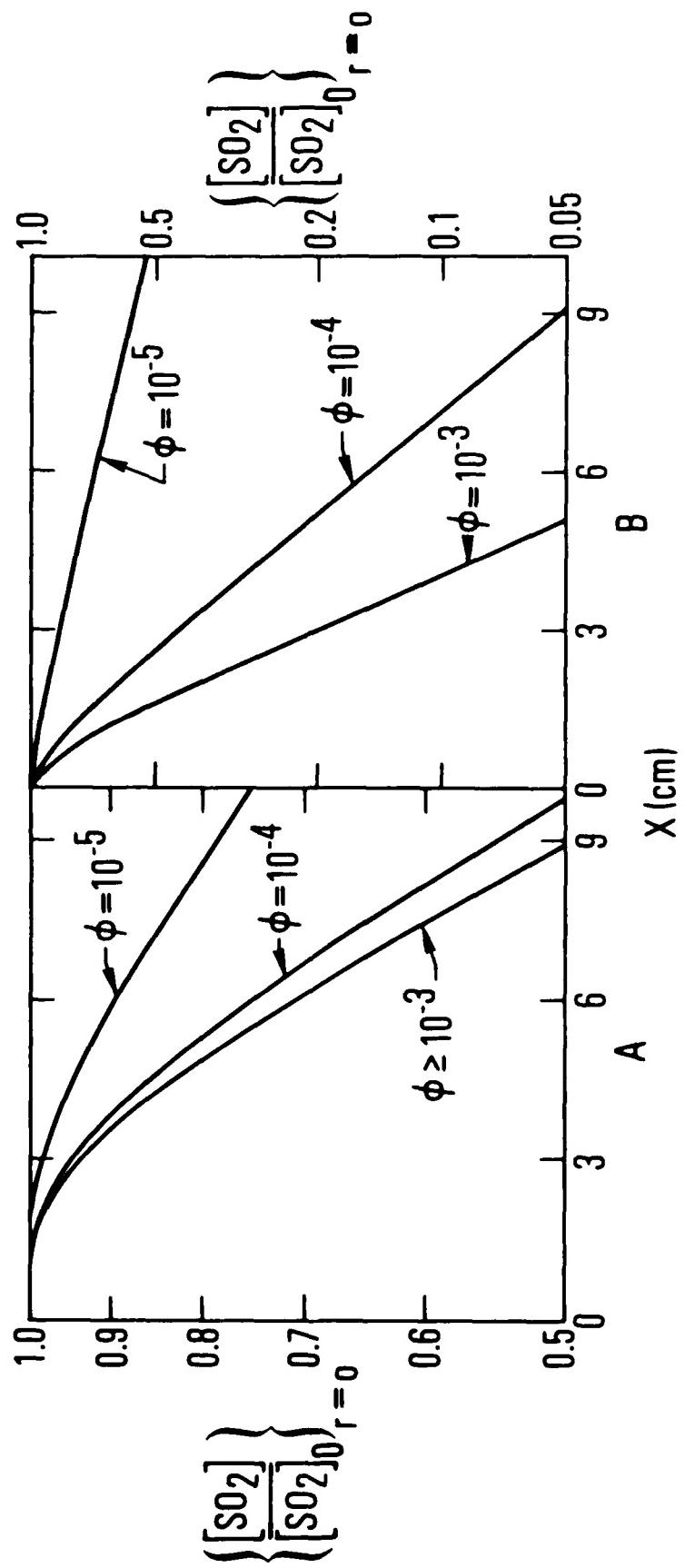
$$C = \frac{[SO_2]}{[SO_2]_0}$$

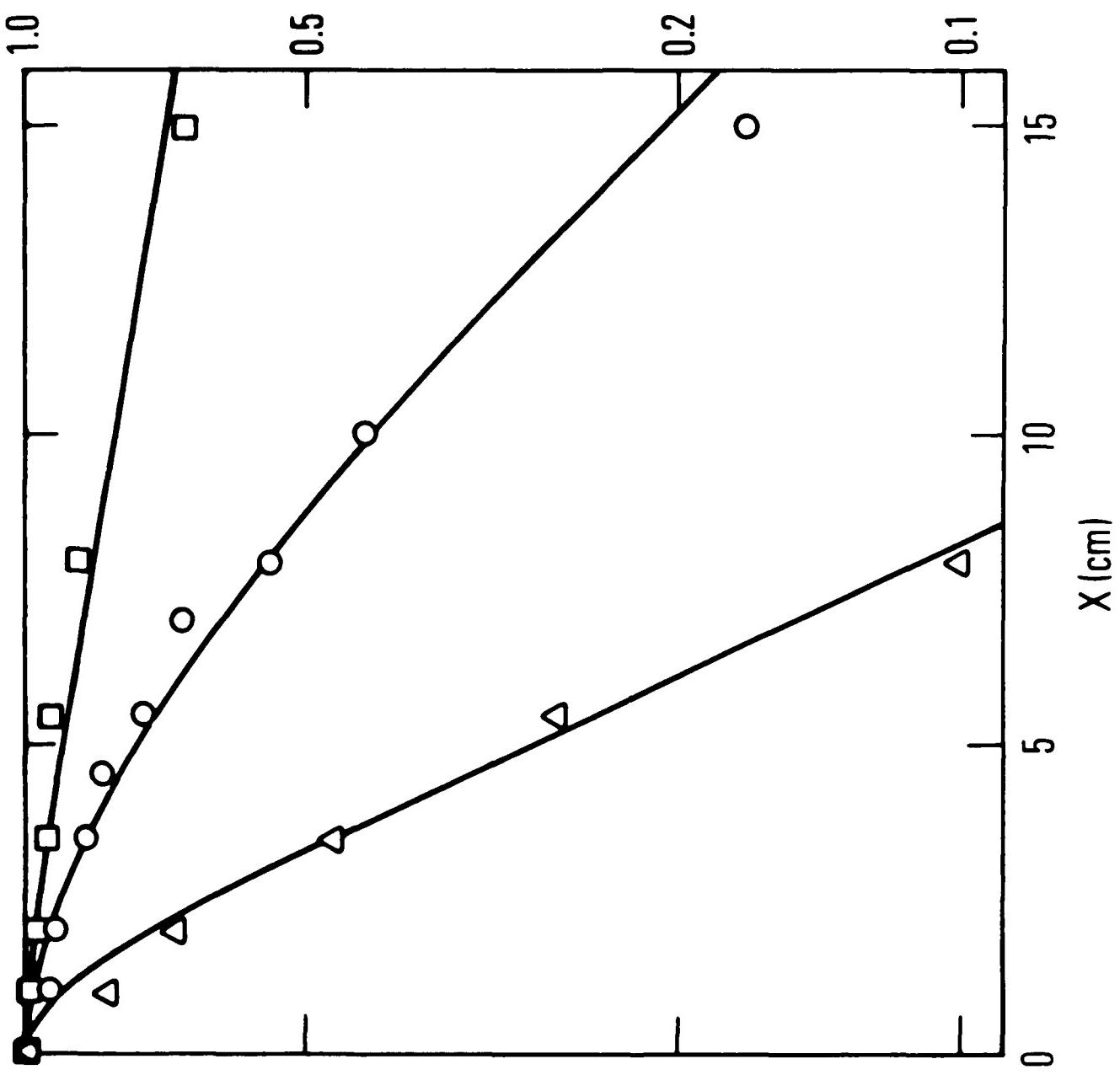
### BOUNDARY CONDITIONS

$$\begin{aligned} C &= 1 \text{ AT } Z = 0 \\ C &= 0 \text{ AT } Z = \infty \\ \frac{\partial C}{\partial R} &= 0 \text{ AT } R = 0 \\ D \frac{\partial C}{\partial R} &= \phi K_R C \text{ AT } R = R \end{aligned}$$

### GENERAL SOLUTION

$$C = \sum I F_I(R) E^{-\beta_I Z}$$





$$\left\{ \frac{[SO_2]}{[SO_2]_0} \right\}_{r=0}$$

ATMOSPHERIC MODEL

$$\text{AEROSOL REACTION RATE} = \frac{3}{R^2} \frac{\text{AD}}{6} C_{03}$$

$$G = 1 \left[ + K_N \left( \frac{1.33 + 0.7 K_N^{-1}}{1 + K_N^{-1}} + \frac{4}{3} \frac{1-\gamma}{\gamma} \right) \right]^{-1}$$

FOR  $K_N \gg 1$

$$\frac{C}{C_0} = (1 - 0.009 \frac{\gamma}{R_0} \tau)^3$$

$$\frac{R}{R_0} = (1 - 0.009 \frac{\gamma}{R_0} \tau)$$

AEROSOL DESTRUCTION

$\frac{\gamma}{10^{-3}}$	$\frac{0.1}{3 \text{ HR}}$	$\frac{0.2}{1 \text{ DAY}}$	$\frac{0.5}{13 \text{ DAYS}}$
$10^{-4}$	6 HR	2 DAYS	5 DAYS
$10^{-5}$	13 DAYS	26 DAYS	65 DAYS
$10^{-6}$	4 MOS.	8 MOS.	20 MOS.
$10^{-7}$	4 YRS.	8 YRS.	20 YRS.

## AEROSOL DESTRUCTION

$$\begin{aligned}
 D_0 &= 0.3\mu \\
 Y &= 1 \times 10^{-5} \\
 O_3 &= 3 \times 10^{12}/\text{cc}
 \end{aligned}$$

<u>T (DAYS)</u>	<u>AEROSOL RADIUS (<math>\mu</math>)</u>	<u>% AEROSOL MASS REMAINING</u>
0	0.30	100.00
3	0.28	78.44
6	0.25	60.22
9	0.23	45.07
12	0.21	32.70
15	0.18	22.83
18	0.16	15.18
21	0.14	9.46
24	0.11	5.40
27	0.09	2.70
30	0.07	1.10
33	0.04	0.30
36	0.02	0.03
39	0.00	0.00

**HETEROGENEOUS REACTIONS  
IN A SOOT LADEN ATMOSPHERE**

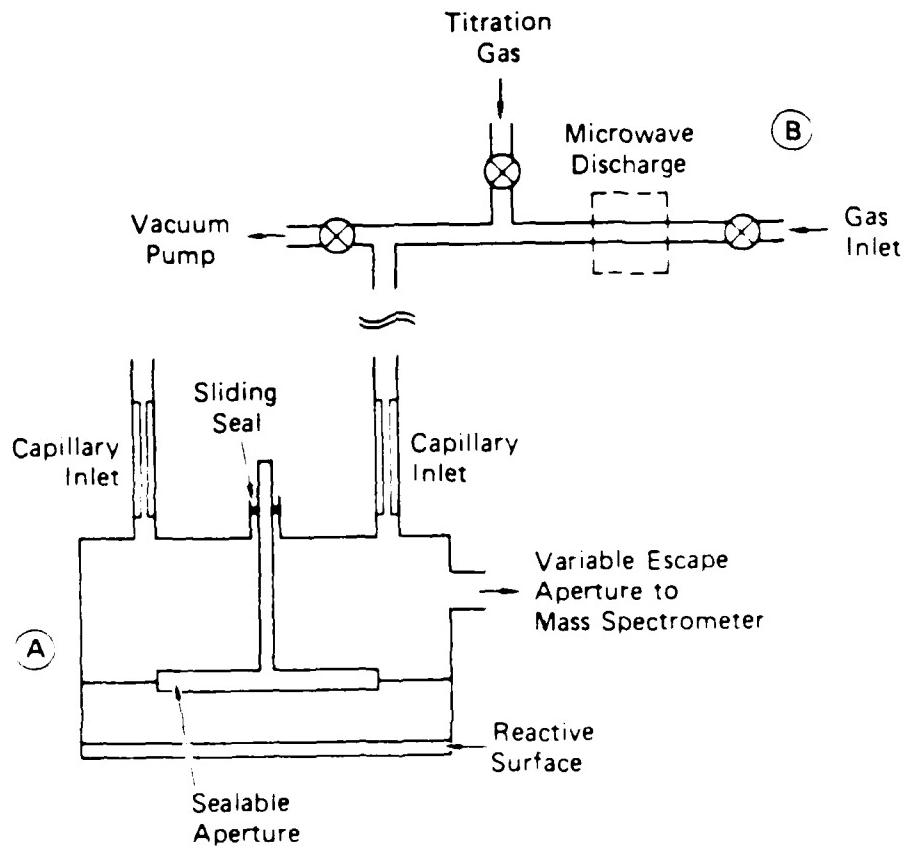
**DAVID M. GOLDEN**

**MICHEL J. ROSSI**

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**SHERRY STEPHENS**

**DEPARTMENT OF CHEMICAL KINETICS  
CHEMICAL PHYSICS LABORATORY  
SRI INTERNATIONAL**



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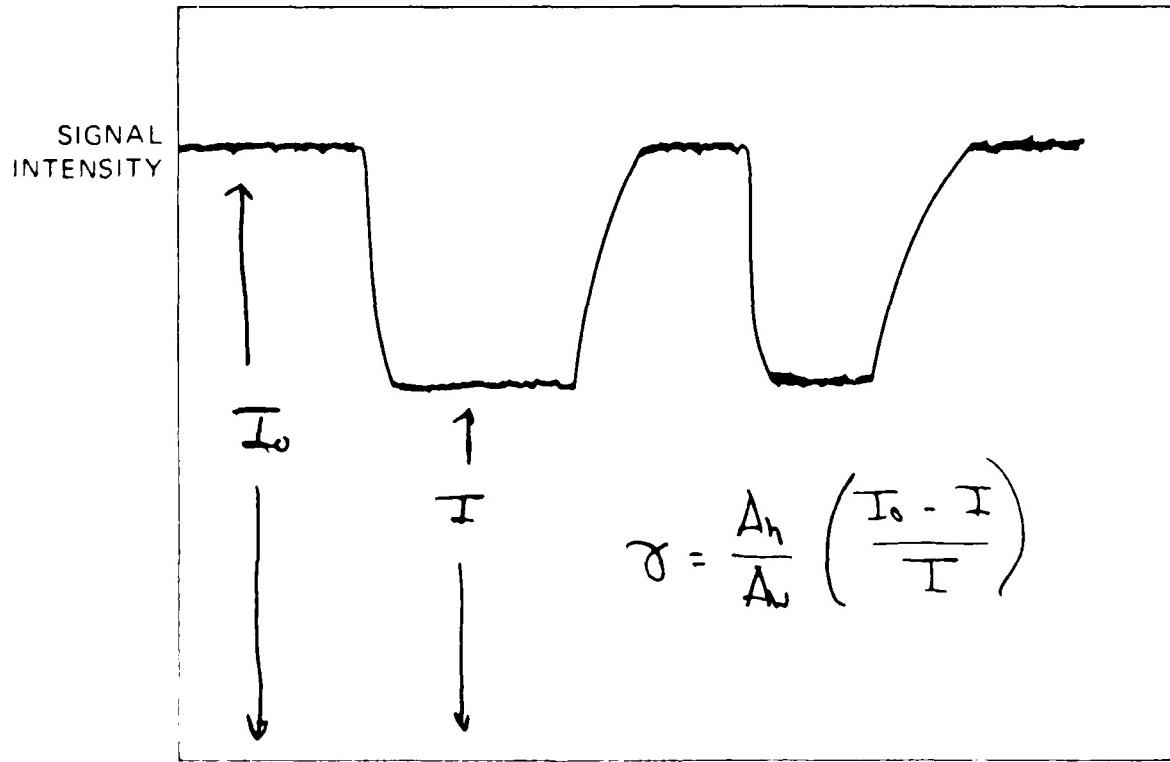
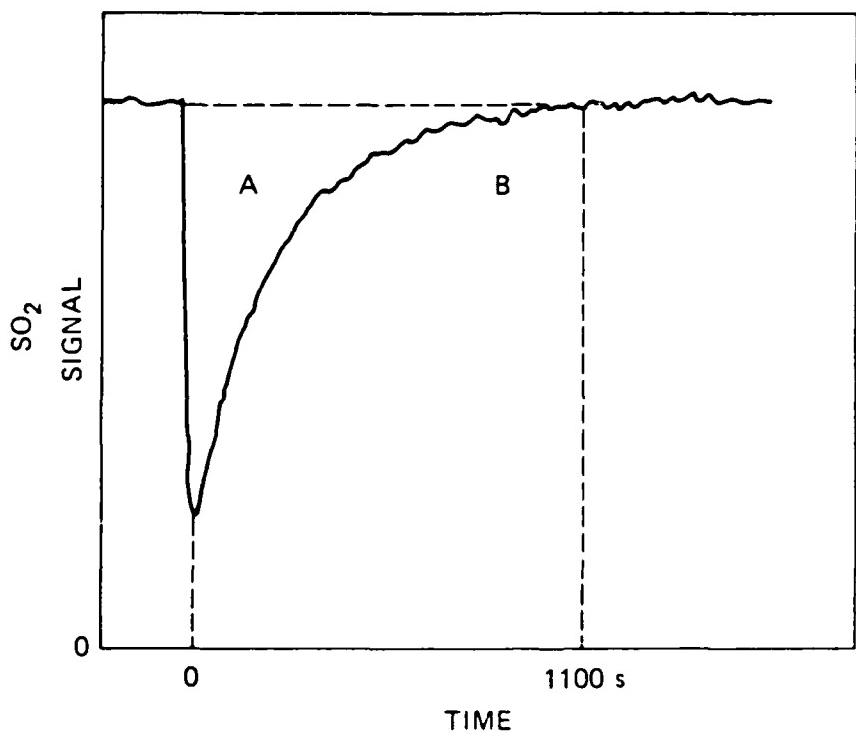


FIGURE TYPICAL EXPERIMENTAL DATA

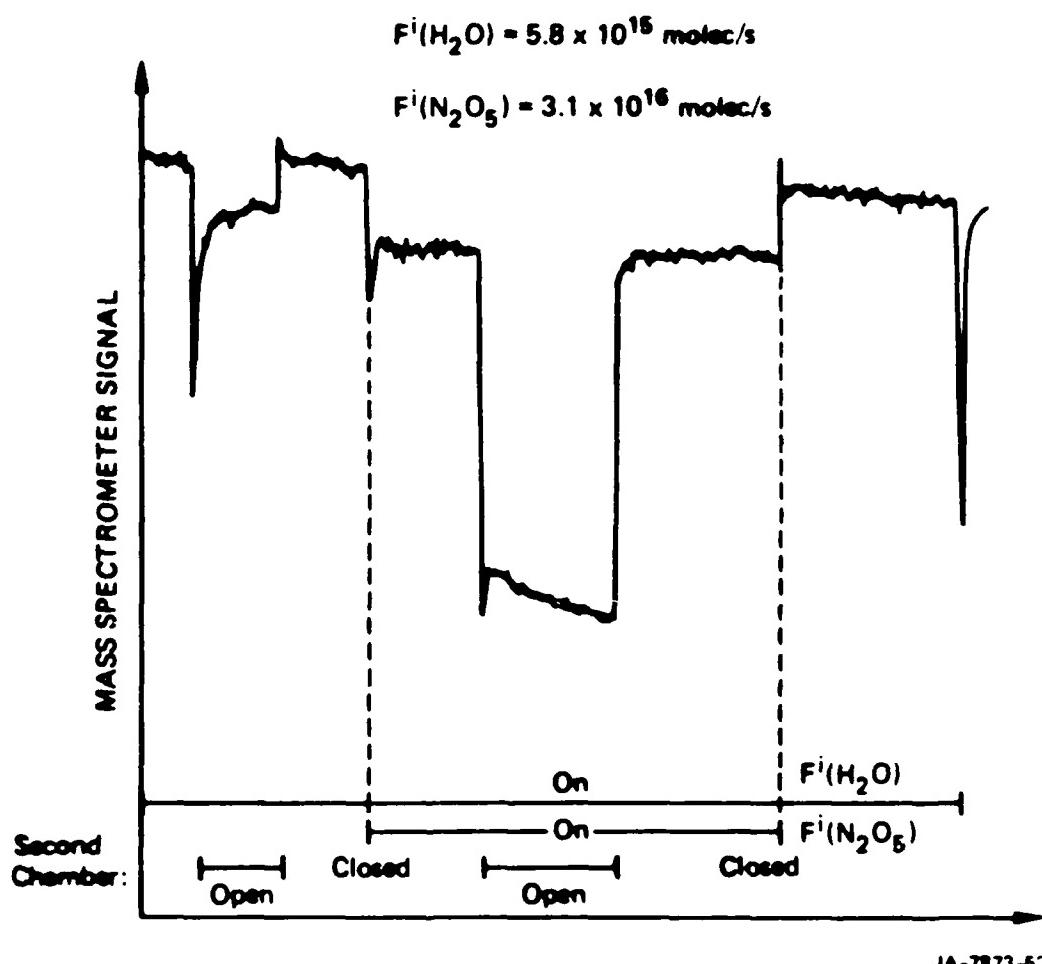
Table 1. Collisional reaction probabilities on a H<sub>2</sub>SO<sub>4</sub> surface at 300 K.

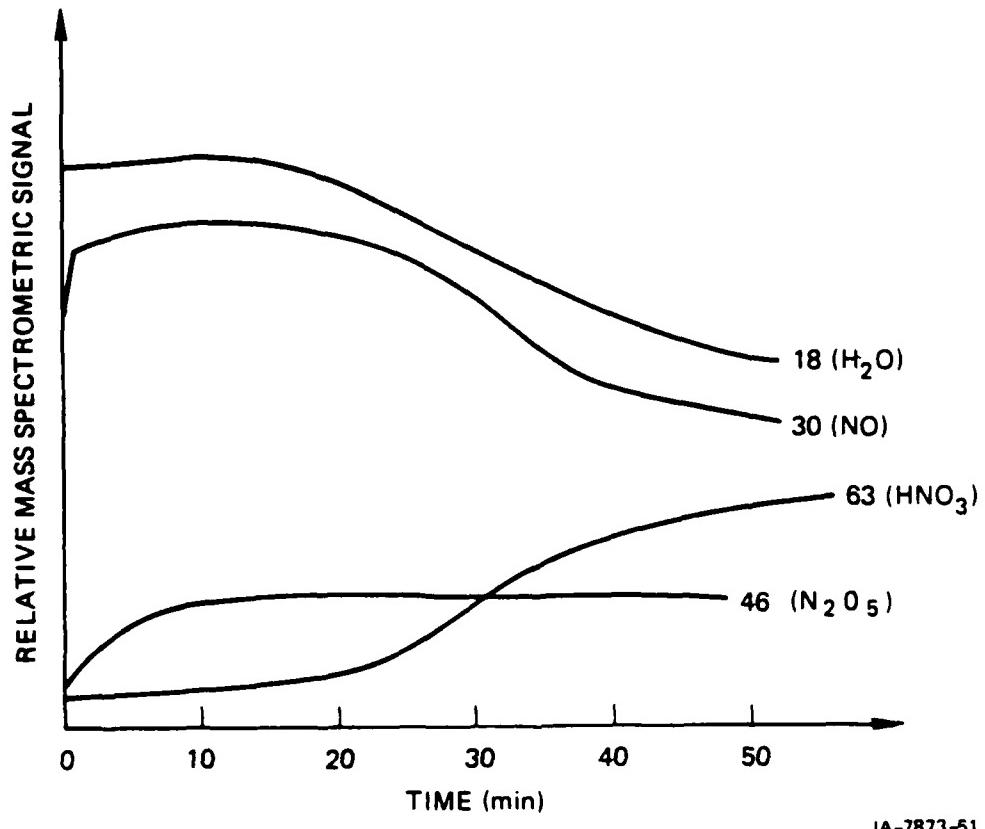
Species	Collisional reaction probability (γ)
H <sub>2</sub> O <sub>2</sub>	7.8 × 10 <sup>-4</sup>
HNO <sub>3</sub>	≥ 2.4 × 10 <sup>-4</sup>
HO <sub>2</sub> NO <sub>2</sub>	2.7 × 10 <sup>-5</sup>
ClONO <sub>2</sub>	1.0 × 10 <sup>-5</sup>
N <sub>2</sub> O <sub>5</sub>	≥ 3.8 × 10 <sup>-5</sup>
H <sub>2</sub> O	~ 2.0 × 10 <sup>-3</sup>
NH <sub>3</sub>	> 1.0 × 10 <sup>-3</sup>
OH	4.9 × 10 <sup>-4</sup>
O <sub>3</sub>	< 1.0 × 10 <sup>-6</sup>
NO	< 1.0 × 10 <sup>-6</sup>
NO <sub>2</sub>	< 1.0 × 10 <sup>-6</sup>
SO <sub>2</sub>	< 1.0 × 10 <sup>-6</sup>
Alkenes	< 1.0 × 10 <sup>-6</sup>
Alkanes	< 1.0 × 10 <sup>-6</sup>
CF <sub>4</sub>	< 1.0 × 10 <sup>-6</sup>
CCl <sub>2</sub> F <sub>2</sub>	< 1.0 × 10 <sup>-6</sup>
O( <sup>3</sup> P)	< 1.0 × 10 <sup>-6</sup>
N	< 1.0 × 10 <sup>-6</sup>

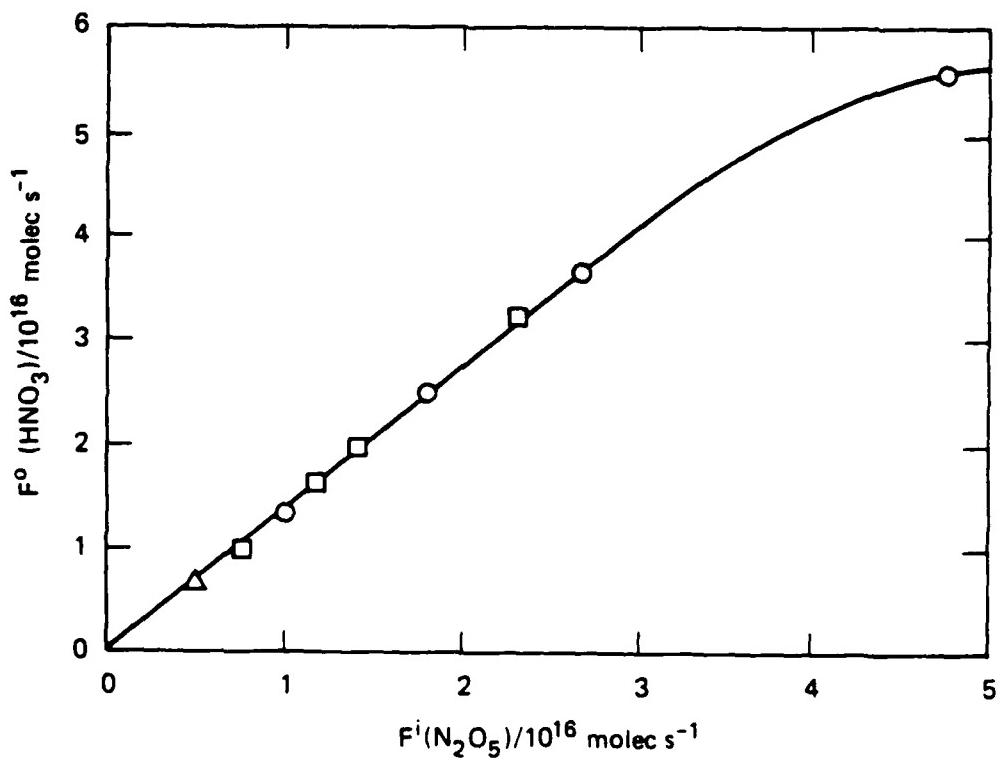


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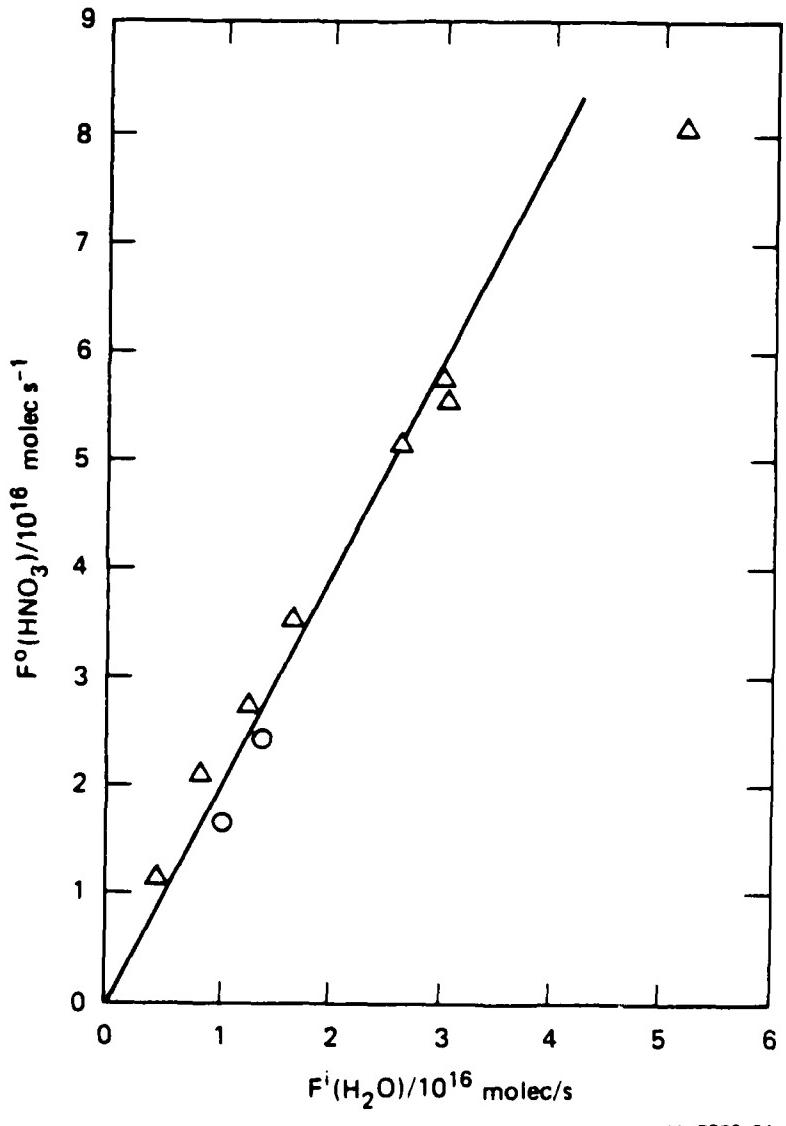
$\text{H}_2\text{O}$  MASS SPECTROMETER SIGNAL ( $m/e = 18$ ) IN THE SYSTEM  
 $\text{N}_2\text{O}_5/\text{H}_2\text{O}$  ON CARBON



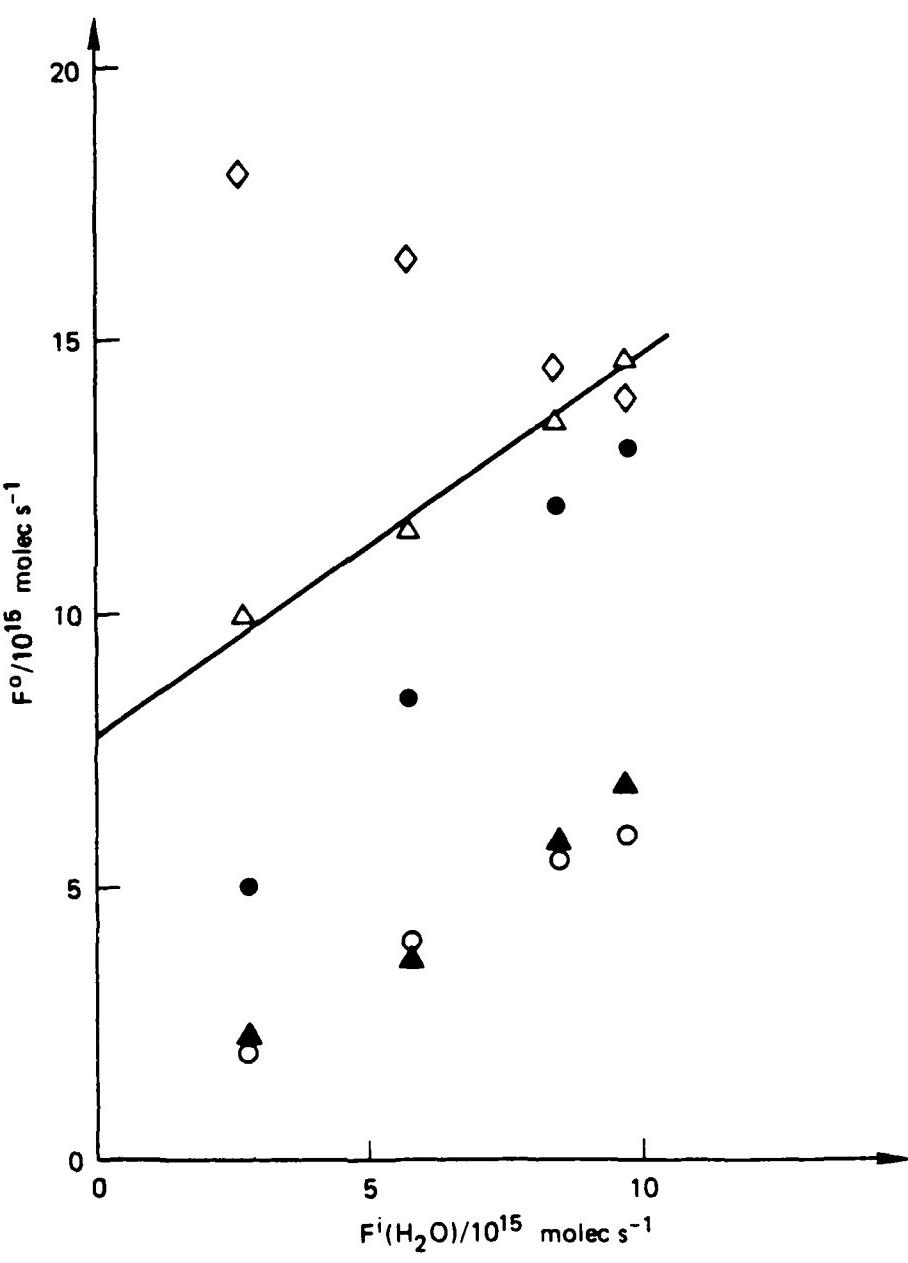




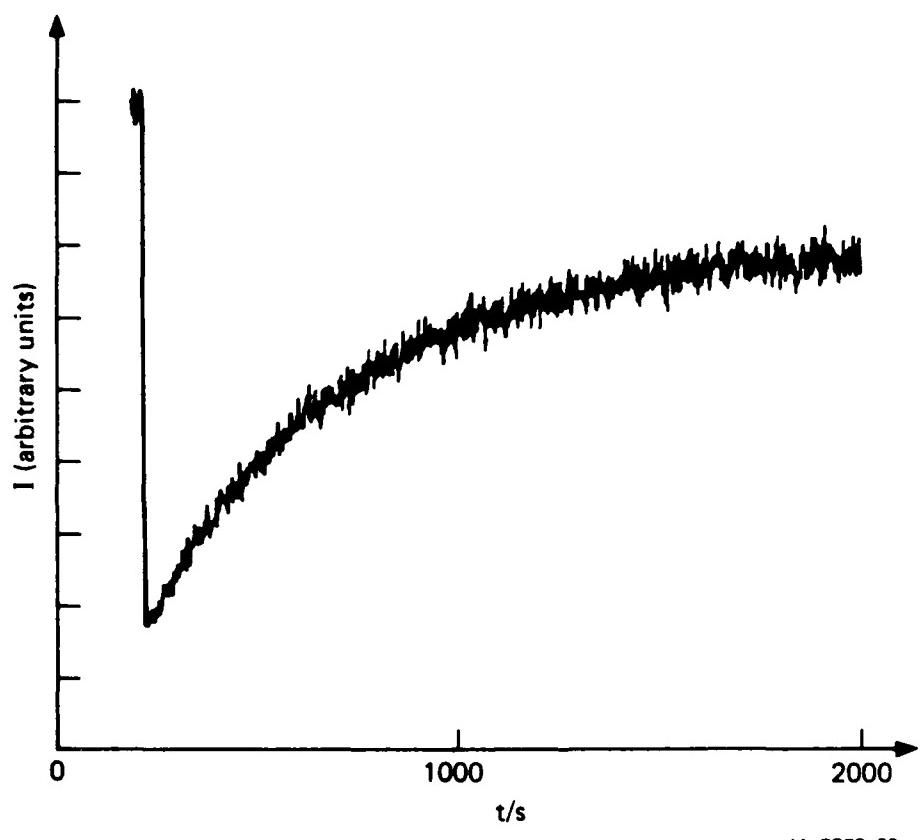
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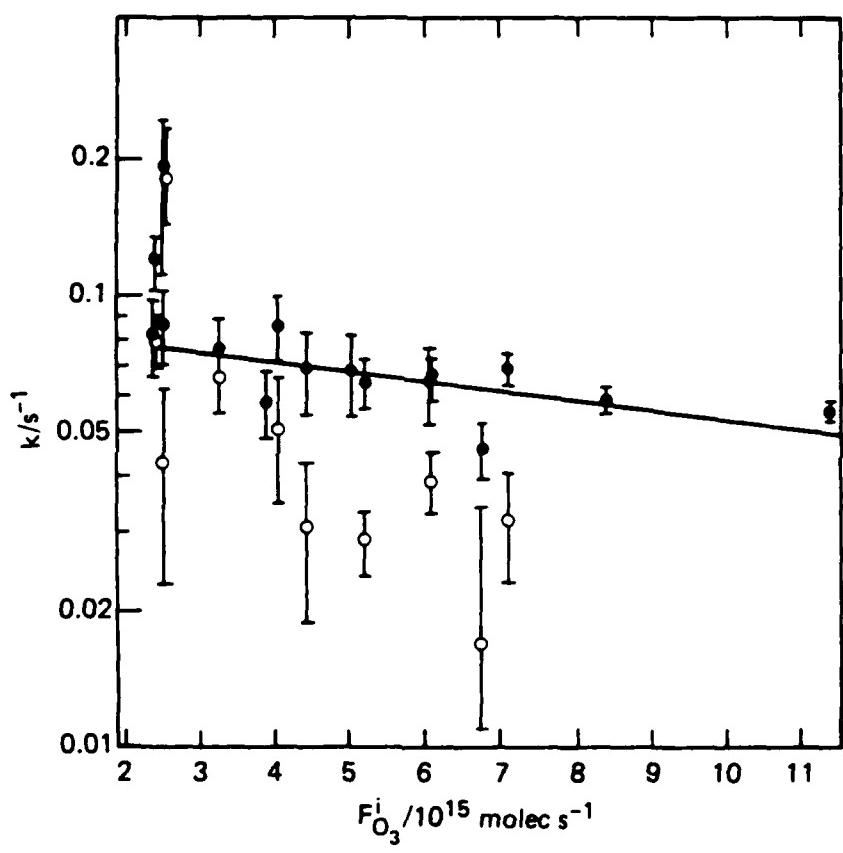
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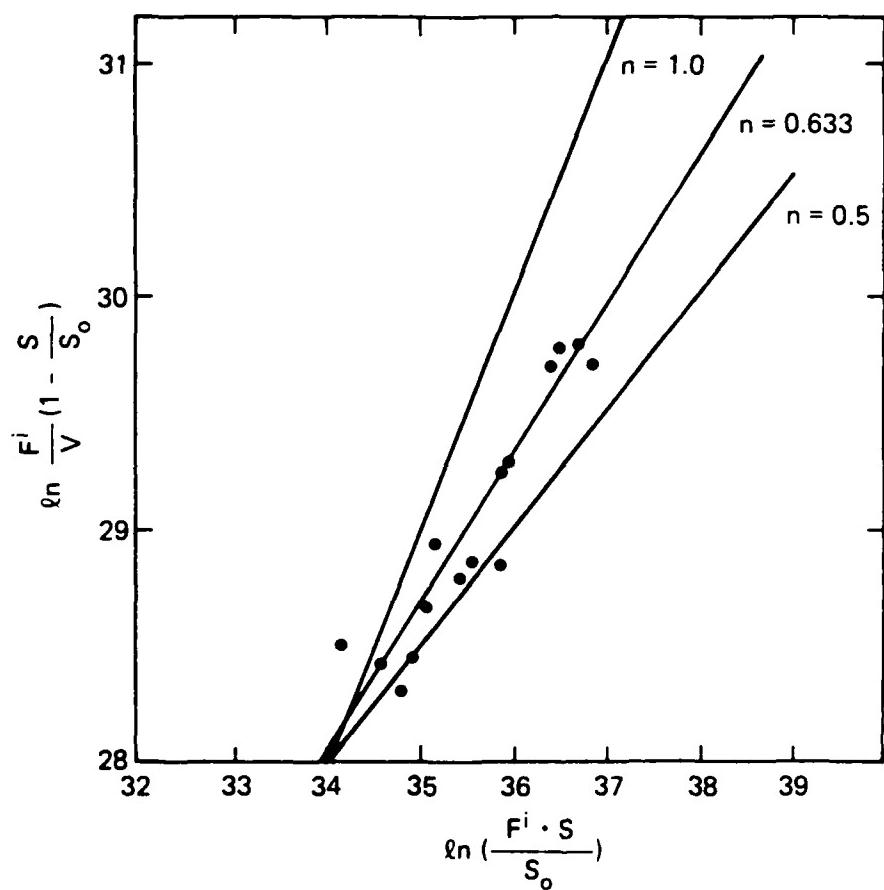
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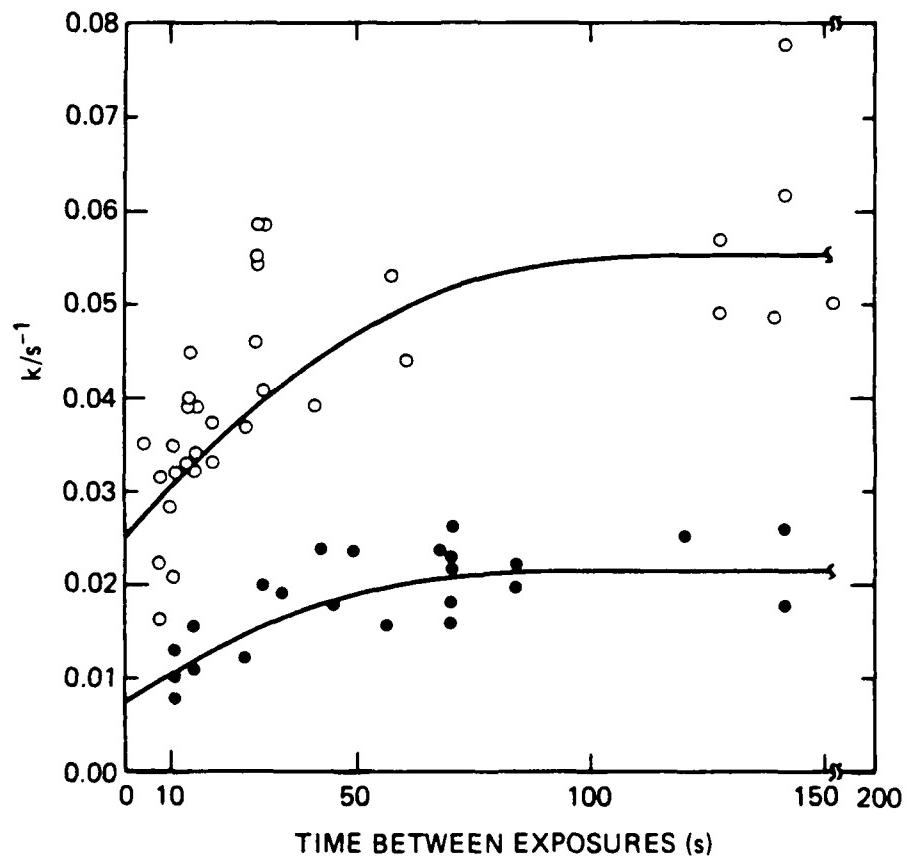
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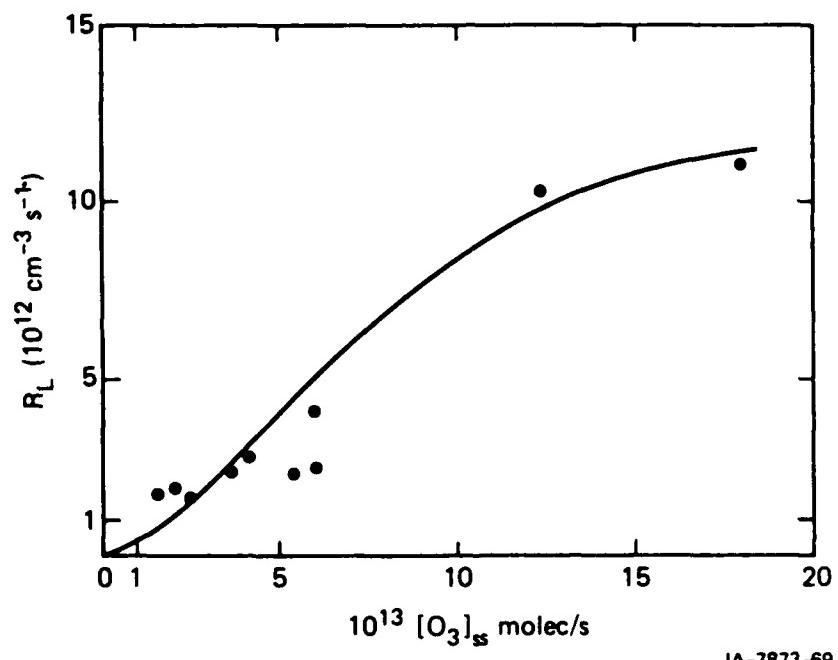
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JA-7873-63



JA-7873-66



JA-7873-69

PROPOSED MODELING STUDIES ON REMOVAL BY CLOUD PROCESS  
OF MASSIVE EMISSIONS OF PARTICLES

In Young Lee

Atmospheric Physics Program  
Environmental Research Division  
Argonne National Laboratory, Argonne, IL 60439

ABSTRACT

In general, atmospheric aerosols originate by several processes: (1) condensation and sublimation of vapors and the formation of smokes in combustion, (2) reactions among trace gases, (3) mechanical disruption and dispersal of matter at the earth's surface as mineral dusts, and (4) coagulation of nuclei to produce larger particles of mixed constitution. The particles participate actively during cloud formation either as cloud condensation nuclei or ice-forming nuclei, or passively by being removed by cloud droplets and raindrops through processes of collision and coalescence. The evolution of dust and smoke particles is additionally controlled by dynamic processes such as transport, turbulent mixing and deposition. These natural processes of aerosol evolution would be greatly affected if extraordinary amounts of dust and smoke particles were injected into the atmosphere by blasts or large fires.

The scavenging efficiency for aerosols produced by blasts and fires can be examined in numerical simulations which describe coagulatory processes involving aerosols, cloud droplets and raindrops. The investigation of the evolution

within clouds of dust and smoke particles of very high concentrations is important for clear understanding of the resulting global budget of aerosols and for better assessment of the potential climatic impact. The collision efficiency between particles can be formulated from a combination of theoretical analyses and available experimental data. A numerical module for wet removal can be used to evaluate scavenging efficiency, modification of cloud properties, and applicability of cloud microphysics parameterizations currently in use.

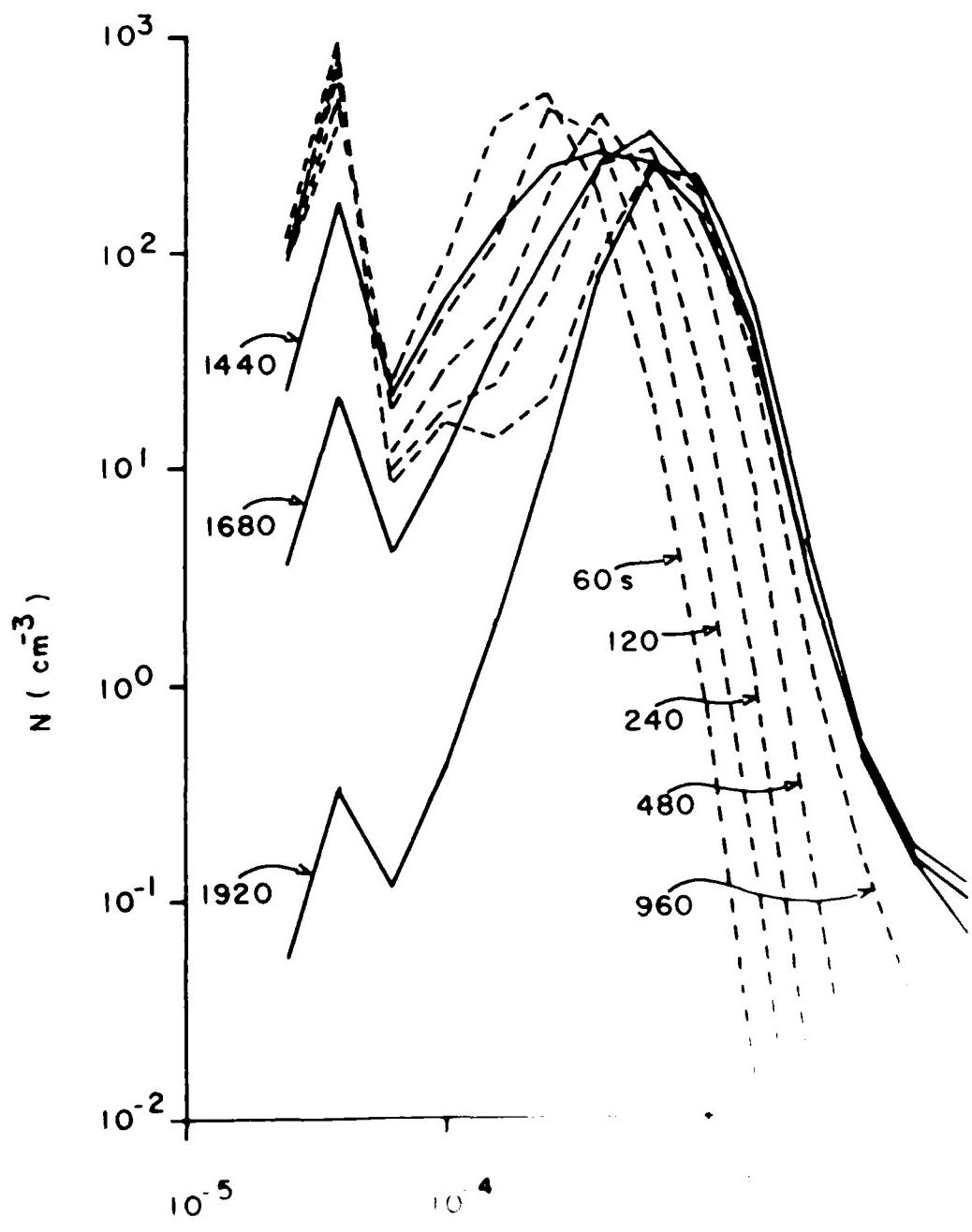
A cloud model originally developed at Argonne to study cloud microphysics parameterizations has been modified to examine the behavior of submicron aerosols in clouds. Lagrangian computations following a rising parcel of air were made in order to understand and identify mechanisms that control the fate of smoke and dust particles in clouds. The air parcel was assumed to be saturated initially and to contain aerosols of high number density ( $1650 \text{ cm}^{-3}$ ). Some portion of the aerosols in the parcel could be activated to form cloud droplets and the rest could remain inactivated due mainly to the Kelvin effect. Thereafter, the spectrum was allowed to evolve and form raindrops. The model treated particle sizes ranging from  $0.04 \mu\text{m}$  to  $164 \mu\text{m}$  in radius.

Computations have been made for 40 min of simulated time. The spectral evolution and the role of each microphysical mechanism in the spectral evolution are presented in Figs. 1 and

2, respectively. Figure 1 shows that the spectrum evolves quickly during the first five minutes of cloud development due mainly to condensation, which affects the growth of particle sizes approximately up to the mode radius near  $10 \mu\text{m}$ . The spectral evolution starts to produce raindrops, associated with sizes greater than about  $50 \mu\text{m}$  in radius, after 10 min of simulated time. The submicron size particles are not removed or activated until simulated time exceeds 20 min, but thereafter they are quickly scavenged as the cloud starts to accumulate drops greater than about  $100 \mu\text{m}$  in radius. In Fig. 2, we see clearly that the combined processes of condensation and coagulation can remove submicron particles much more effectively than either process alone. We also found that the removal of submicron particles is augmented if precipitation is included.

Figure 1. Number distributions of aerosols, cloud droplets and raindrops after various periods of simulation.

Figure 2. The role of cloud microphysical processes of condensation and coagulation in the spectral evolution of particles.



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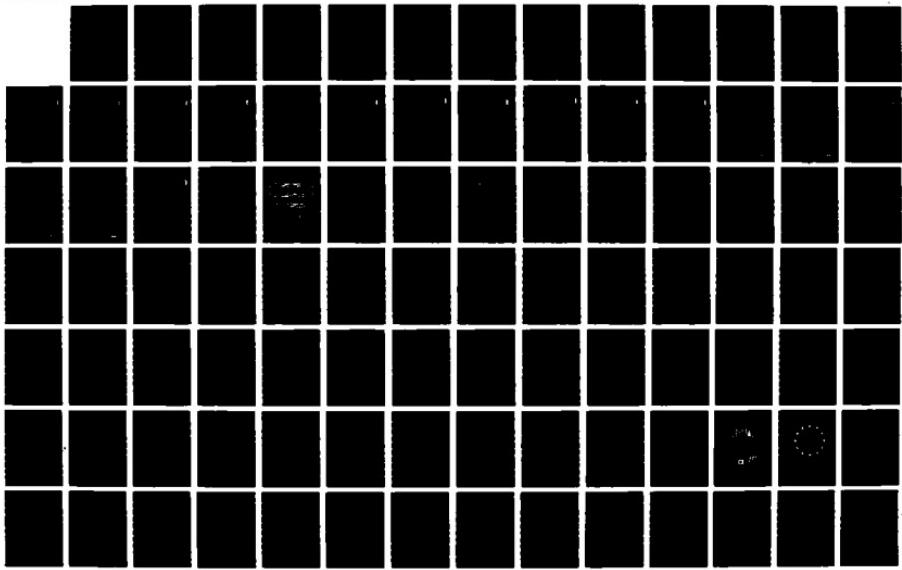
2/4

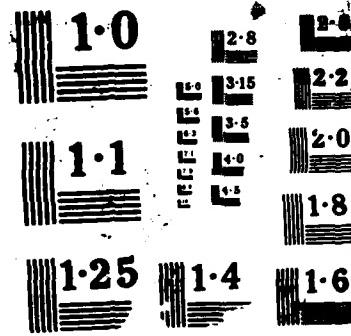
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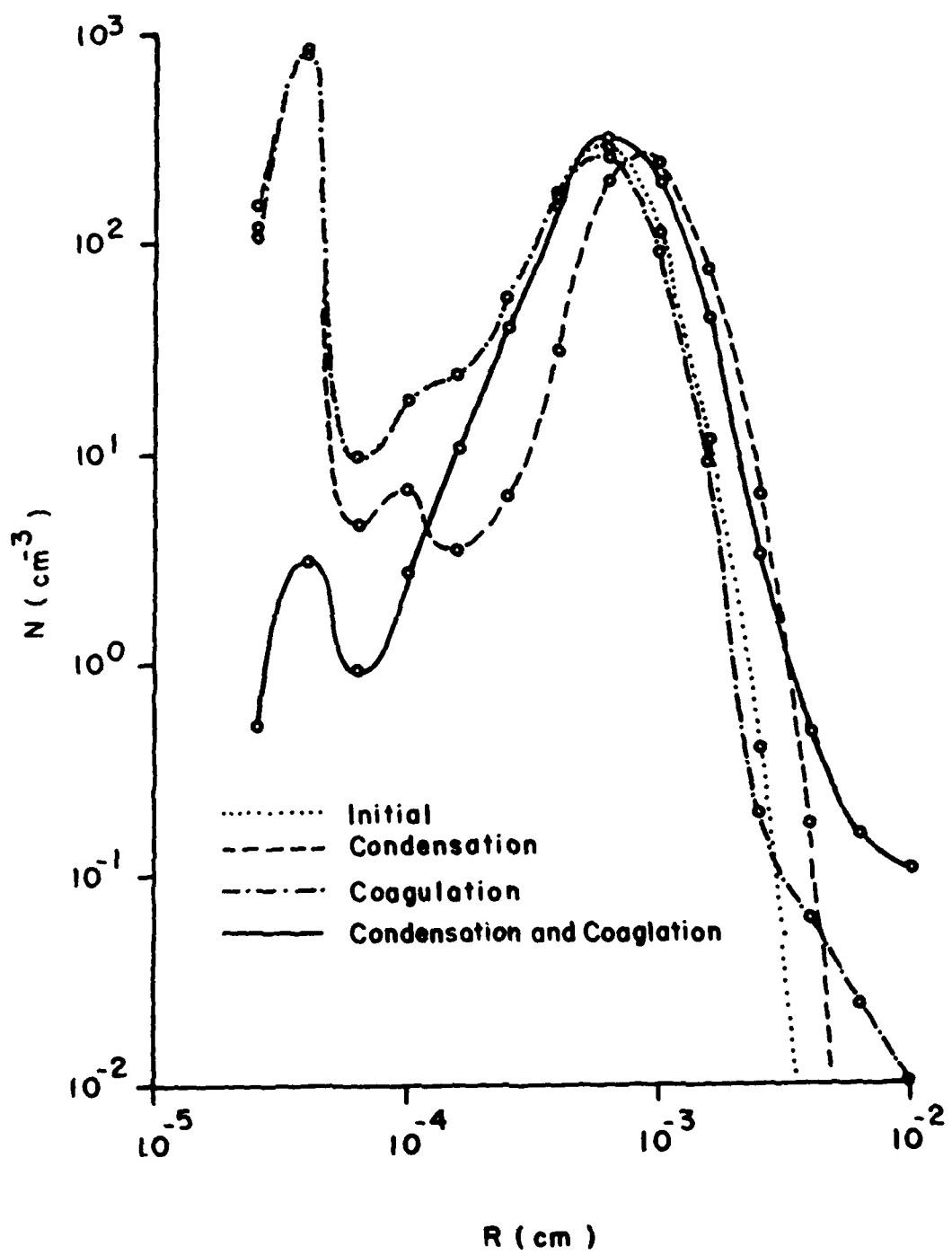
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NL







**The University of Missouri-Rolla Cloud Simulation  
Chambers**

**Daniel R. White  
Donald E. Hagan  
John C. Carstens**

**Graduate Center for Cloud Physics Research  
University of Missouri-Rolla**

## THE UMR CLOUD SIMULATION CHAMBERS

The UMR cloud simulation program is based on two cooled-wall expansion chambers designed to produce uniform cooling comparable to that produced in cloud updrafts. The smaller chamber, called Proto II, is currently producing cloud, and the larger--more sophisticated--Romulus chamber is scheduled for completion in the summer of 1985.

These chambers feature thermoelectric modules to cool the walls at the same rates as the expansion cools the gas. Thermal wall effects are thus effectively removed and a closed parcel thermodynamic environment similar to that found in the core of a convective cloud in the absence of turbulent mixing can be simulated.

After a sample of moist aerosol-laden air has been introduced into the chamber, the sensitive volume undergoes a controlled expansion which cools the gas and produces the increased relative humidity required for cloud formation. The chamber's initial condition can be set at any desired state of relative humidity less than 100%. The cooling capability is sufficient to allow ice nucleation and crystal growth studies to be carried out.

The creation of a viable cloud simulation program has necessitated developmental work on a broad front so that the required science and technology would come together to support the experimental effort at about the same time. Much of this effort has gone into the development of the following peripheral and supporting equipment:

1. Air preconditioning system for drying and filtering large volumes of air (see Fig. A2) for flushing the chamber and for use by aerosol generation equipment.
2. Precision saturator columns and thermal regulation columns for relative humidity (see Fig. A3).
3. A variety of aerosol generation techniques have been developed for the production of artificial nuclei. Much effort has gone into characterizing aerosols and rendering them monodispersed. Natural nuclei can be employed.
4. Precision continuous flow thermal diffusion chamber (CFDs) for characterizing the critical activation supersaturation spectrum of the condensation nuclei have been developed. Three vertical plate CFDs are currently operational. Three vertical plate CFDs are currently under construction. These will operate simultaneously, two in the CFD mode and one in the isothermal mode, to cover the entire CCN spectrum with high resolution and will provide automatic data acquisition in a format for immediate assimilation by the simulation chamber minicomputer system to aid in establishing last minute parameters for experiment control.
5. Optical and electrical aerosol characterization equipment for measuring the size distribution of both wet and dry aerosol particles. A GE CNC counter and a UMR Aitken nucleus counter are also available for measuring total particle concentrations and sensing the Aitken nuclei content of sample air. This makes possible scavenging experiments.

6. Laser light scattering equipment enables one to determine the Mie scattering intensity at one angle. The total attenuation of a light beam (from a light emitting diode operating at a different wavelength from the argon ion laser) is also measured with a precision of three parts in  $10^4$ . A photograph is also taken with flash illumination in order to verify the droplet concentration at specified intervals. These optical systems are described in more detail in Appendix C.
7. The digital data acquisition and control system is built around a NOVA 840 and a NOVA 3 minicomputer. When both are operational, a greater real time computational capacity is available for experiments. State-of-the-art techniques are employed extensively through the system, particularly in the measurement and control of temperature and pressure.
8. A wide variety of simulation chamber control options have been developed. These can be readily configured to best suit the demands of individual experiments. It is possible to run various microphysical models in real time on one of the minicomputers and use the output in the control of the chamber or to compare with the output of the various optical sensors. In this area we envision a continuous evolution of capability as new experiments are developed.

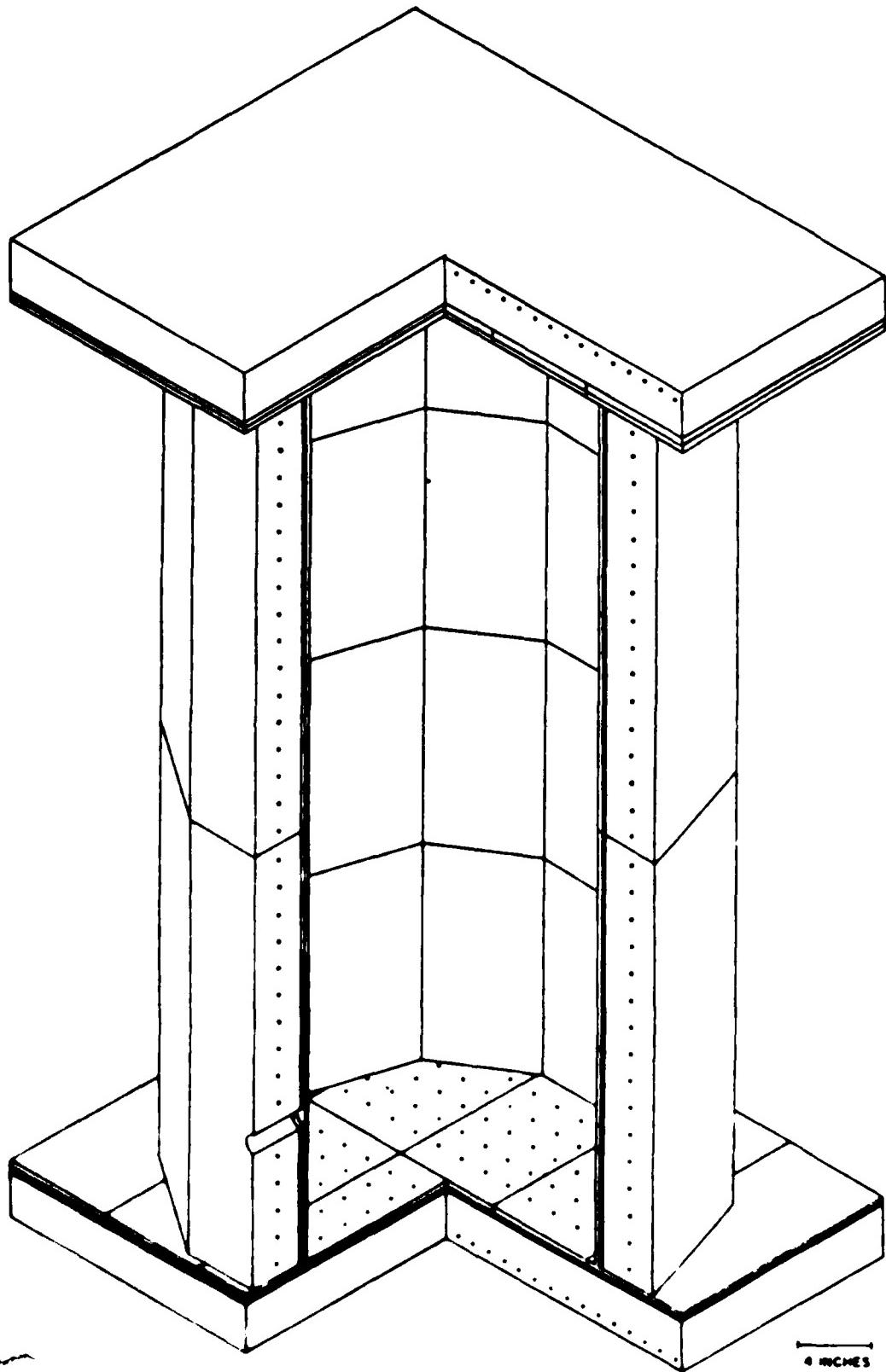


Fig. A1 Prot. Cloud Simulation Chamber

UMR CLOUD SIMULATION FACILITY

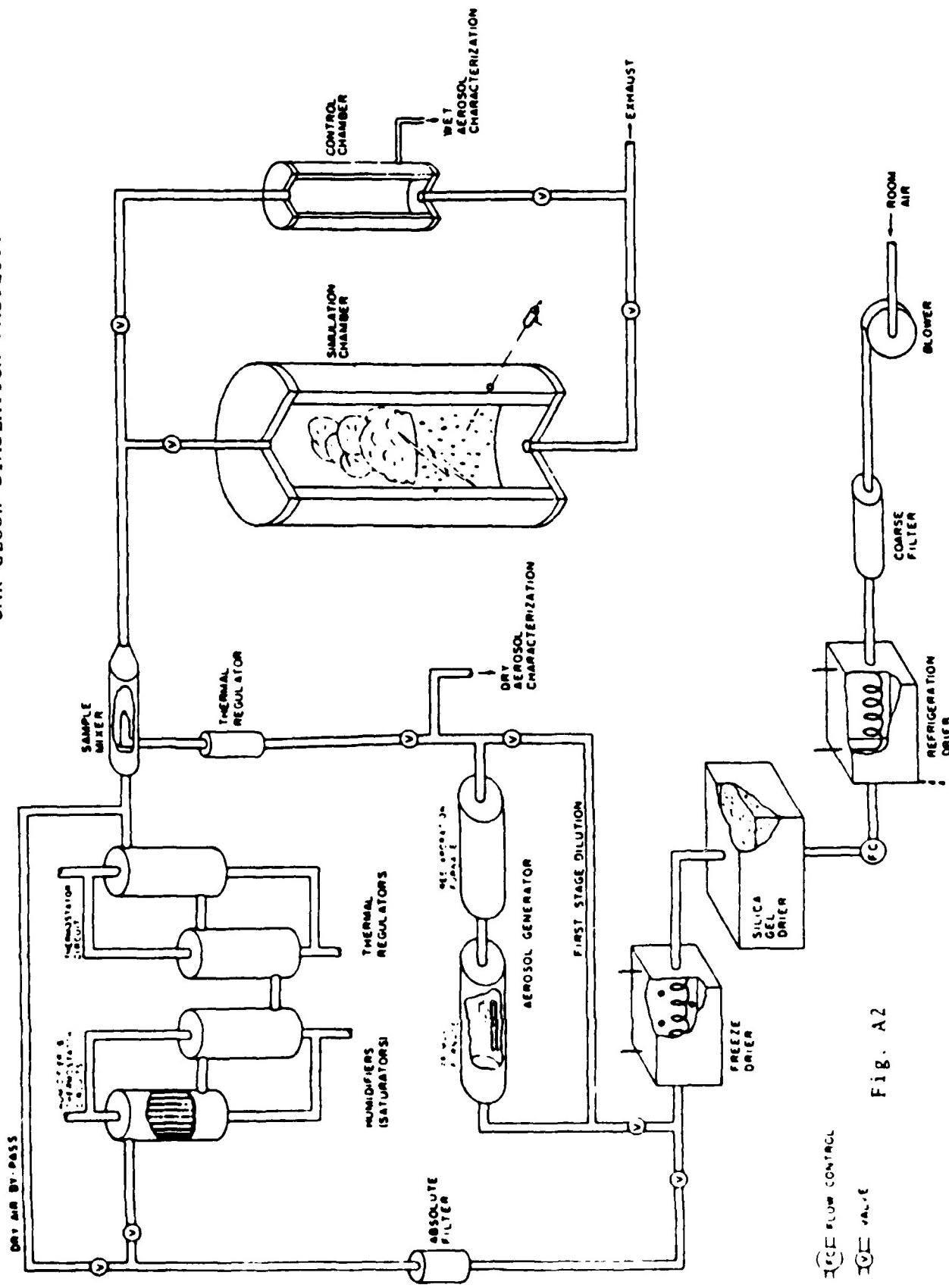


Fig. A2

CF = CLOSURE CONTACT  
OC = OPEN CONTACT

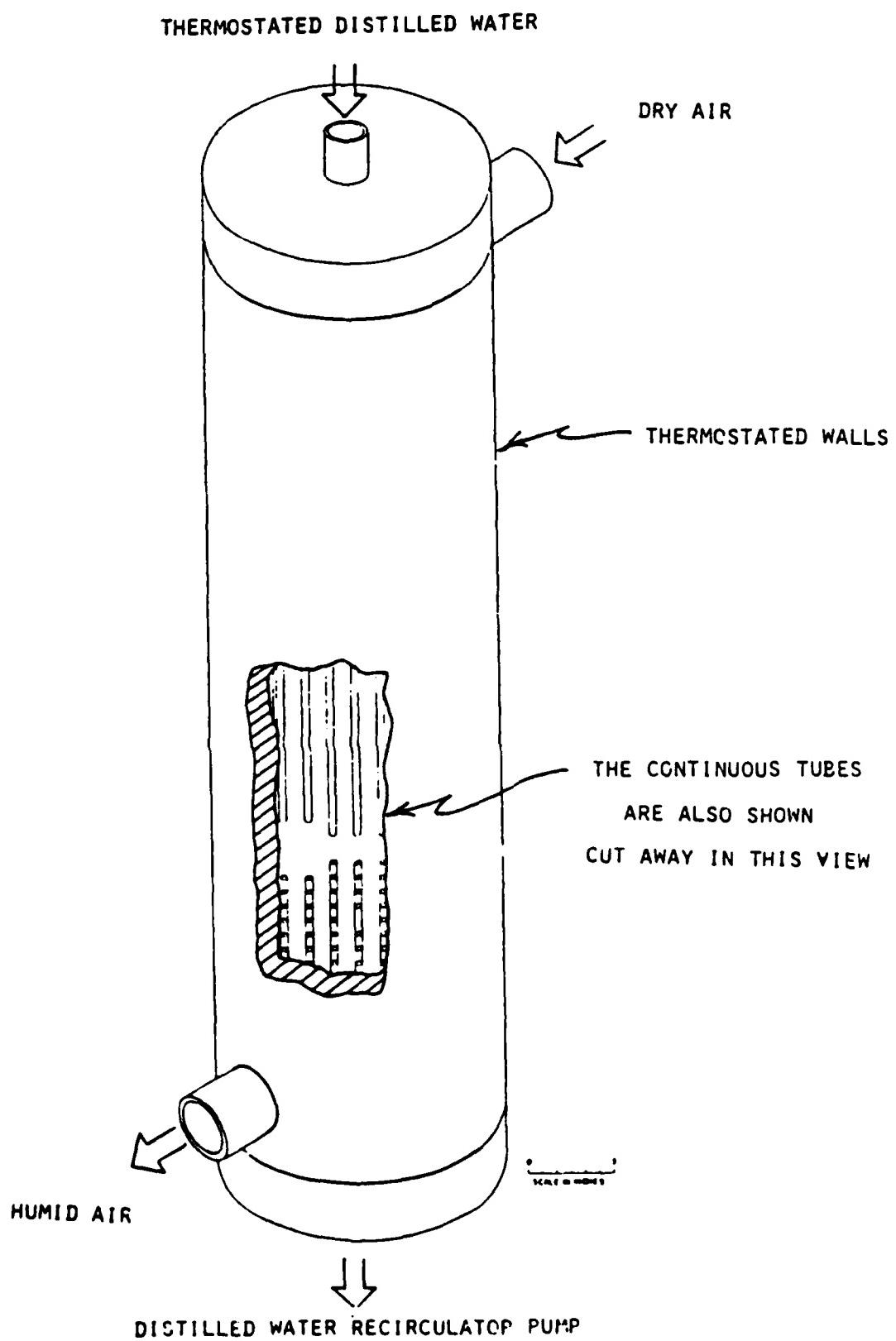


Fig. A3 Humidifier

TABLE I  
PROTO II CHAMBER SPECIFICATIONS

<u>Temperature</u>	<u>Prototype Chamber</u>
Range Control	40° to -40°C
Holding: At fixed temperature RMS deviation from spatial average, 40 sensors	0.030°C
Cooling Mode: Maximum rate, at least Wall temperature lag from control point RMS deviation from spatial average	10°C/min. 0.01°C/(°C/min.)* 0.080°C @ 6°C/min. 0.040°C @ 3°C/min.
Wall Measurement Precision	<u>±0.002°C</u>
Pressure	
Reference pressure absolute accuracy	±0.01%, dead weight pressure gauge
Dynamic sensor range	0 to 100 Torr differential
Resolution	<u>±0.05 Torr</u>
Control:	
Holding: Standard deviation (time)	0.5 Torr
Expansion: Maximum rate	3.9 Torr/sec 12 (C/min. wet adiabatic)
Offset from command valve Standard deviation with (time)	0.5 Torr 0.36 Torr
Wall Cooling Method	Thermoelectric modules plus fluid heat exchange in heat sink
Volume	48 cm dia x 61 cm ht OR 48 cm dia x 122 cm ht

\*Cooling Rate

SECTION 2

PLUME DYNAMICS AND CLOUD PHYSICS

**Scavenging of submicron aerosols by clouds.**

**R.F.Pueschel, NASA Ames Research Center, Moffett Field, CA  
94040.**

Scavenging by clouds and precipitation is the most effective mechanism by which aerosols are removed from the atmosphere. Data on the relationships between aerosol and cloud spectra and concentrations before, during and after cloud formation show the extent and time scales with which the aerosol-cloud interaction occurs. Specifically, the following results will be discussed:

(1) The accumulation mode, the longest lived portion of the atmospheric aerosol, provides the cloud condensation nuclei around which drops form. Particles as small as 0.1 um diameter are ingested into cloud drops during the cloud's formation; the nucleation scavenging of submicron particles is the more efficient the larger their size.

(2) In polluted airmasses, there are more submicron particles removed from the aerosol than there are cloud drops formed. This is different from clean air situations where a 1:1 relationship between the loss of submicron particles and the formation of cloud drops exists. Thus, in polluted air a mechanism in addition to nucleation scavenging has to be invoked to account for the aerosol losses. Stephan flow, a hydrodynamical movement of a vapor-gas mixture normal to the surface of a condensing liquid, to compensate for gas diffusion away from the surface, is a possible explanation.

(3) Nucleation scavenging during cloud formation is 180 times faster than is in-cloud attachment of the residual aerosol after the cloud has formed.

*THE ROLE OF BUOYANCY INDUCED TURBULENCE  
IN SCAVENGING PARTICULATES FROM  
LARGE SCALE FIRES*

R. B. Edelman

Science Applications International Corporation

Combustion Science and Advanced Technology Department

Chatsworth, CA 91311

Presented at the Global Effects Program Technical Meeting,

February 25 - 27, 1986

NASA - Ames Research Center

Moffet Field, CA



## *Topics*

- Area of Interest
- Problem Definition
- Objective
- Method of Approach
- Summary



## AREA OF INTEREST

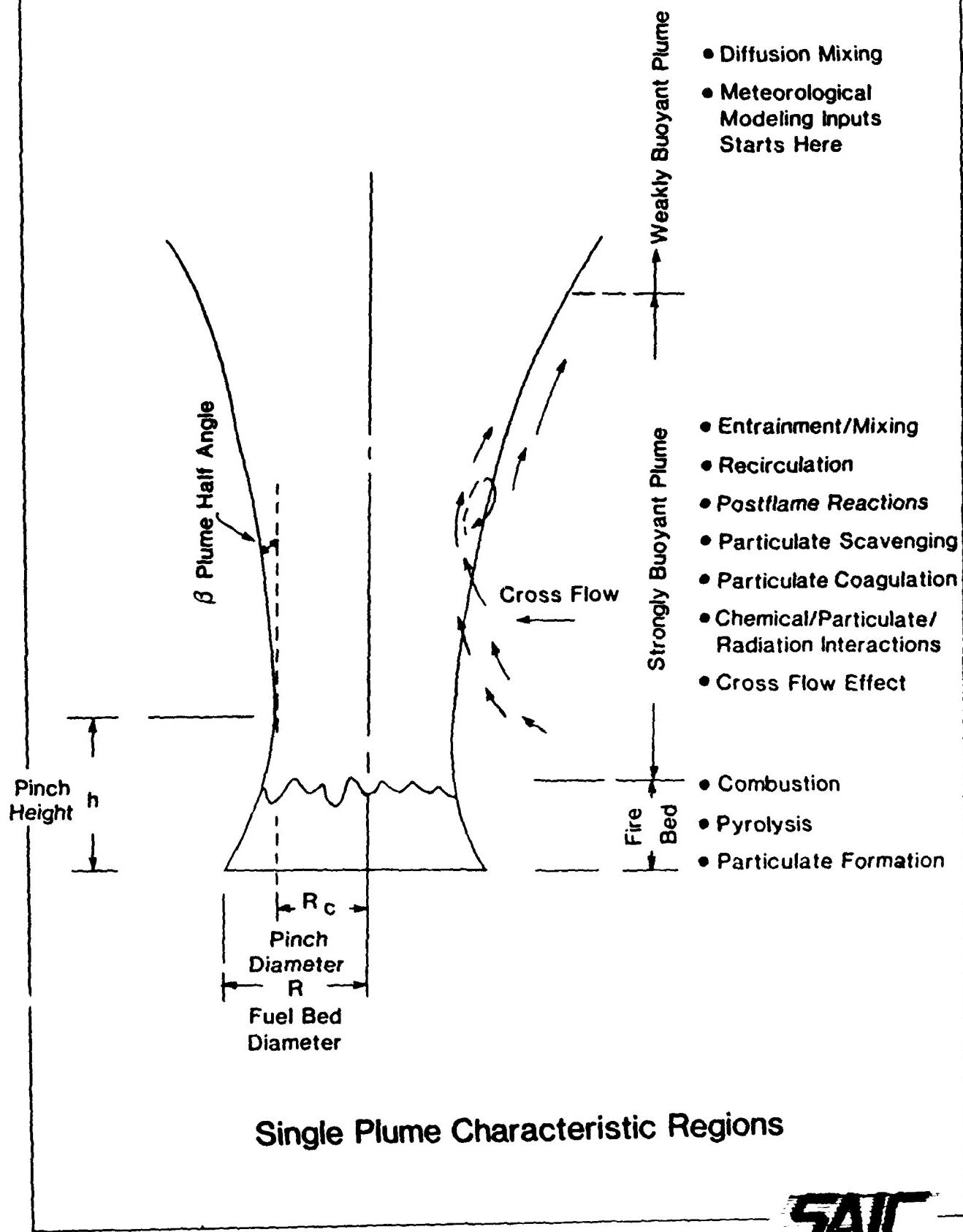
- Uncertainty in Predicting Amount and Structure of Soot and other Particulates Reaching High Altitudes
- Scavenging and Coagulation can have a Strong Effect

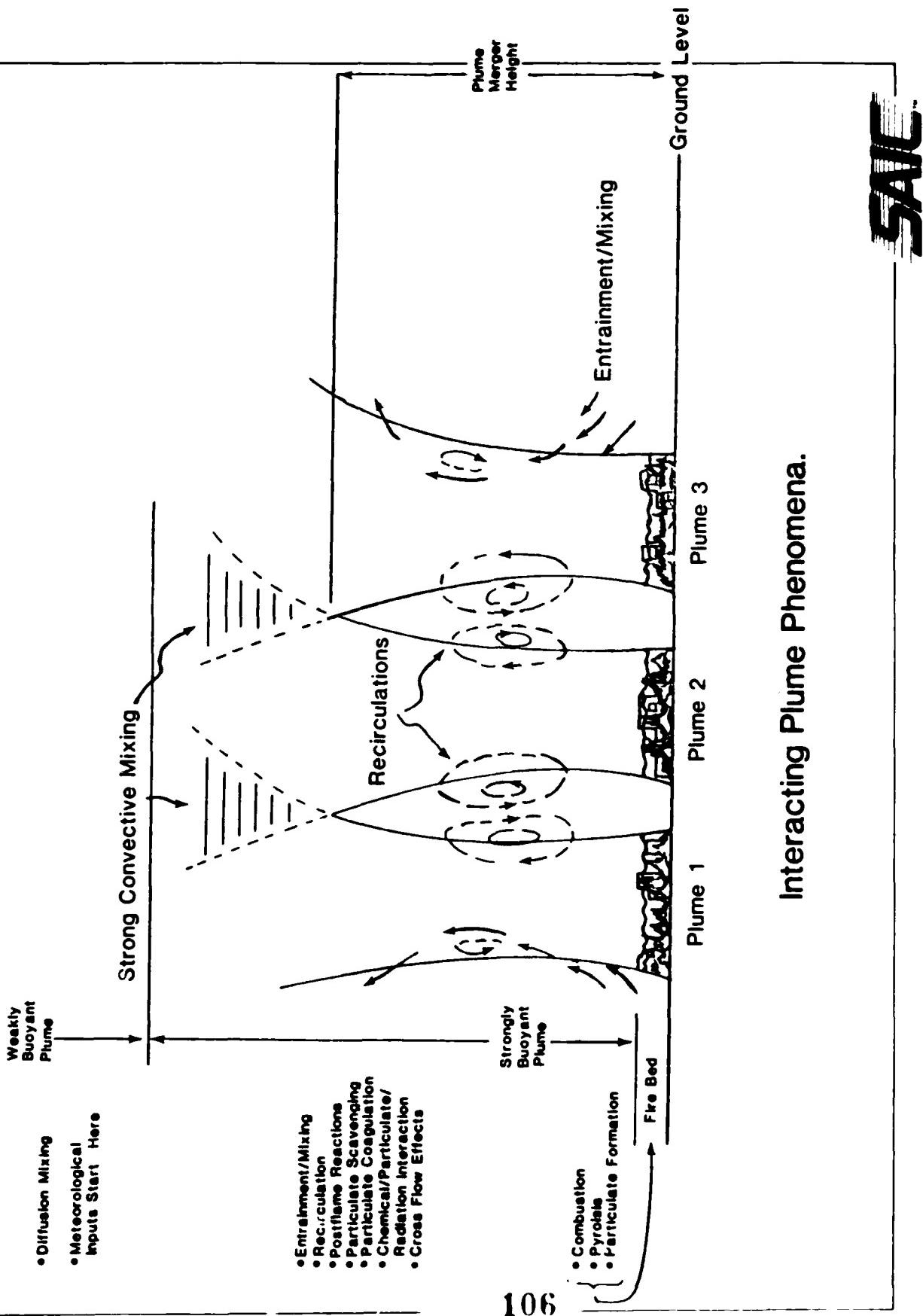
## PROBLEM DEFINITION

- Local Plume Properties are Required
- Spatial and Temporal Variations in Temperature, Concentration and Velocity are Important in Plume Flow Fields
- Properties Depend upon Mixing Rates which, in Turn, Depend Upon Buoyancy-Induced Turbulence and Entrainment Rates

## OBJECTIVES

- To Describe Local Properties in the Near and Intermediate Plume Flow Fields with Emphasis on Buoyancy – Induced Turbulence and Entrainment Processes





### Interacting Plume Phenomena.

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## SUMMARY OF OBSERVATIONS

- Complex Interplay of Time Dependent Entrainment and Mixing Processes
- Combustion Completion and Secondary Rate Processes are Relatively Slow and therefore are Non-Equilibrium Processes
- Mass Fire is Likely Comprised of Small Multiple Fire Plumes Interacting with One Another , Merging to Form a Single Large Scale Plume

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## METHOD OF APPROACH

- Single Plume Modeling
  - Axisymmetric
  - Steady
  - Transient
- Multiple Plume Modeling
  - 3-D
  - Steady
  - Transient
- Plume Structure
  - Local Properties
  - Scaling

## *PRELIMINARY RESULTS ON TURBULENT PLUMES*

- Buoyancy Induced Mean Flow Acceleration
- Buoyancy Induced Turbulence

## BUOYANCY GENERATED TURBULENCE

$$\underbrace{\frac{\partial k}{\partial t} + u_j \frac{\partial k}{\partial x_j}}_{\text{rate of change}} = - \frac{\partial}{\partial x_j} \left[ u_j \left( \overline{u_i u_j} + \frac{\rho}{\rho} \right) - v \frac{\partial k}{\partial x_j} \right]$$

convection

diffusive transport

$$\underbrace{- \overline{u_i u_j} \frac{\partial u_i}{\partial x_j}}_{\text{production by shear}} - \beta g_1 \overline{u_i T'} \underbrace{- v \frac{\partial u_i}{\partial x_j} \frac{\partial u_i}{\partial x_j}}_{\text{dissipation by buoyancy forces}}$$

where,

$$k = \frac{1}{2} \overline{u_i u_i}$$

# $k - \epsilon - q$ TURBULENCE MODEL

## 1. TURBULENT KINETIC ENERGY

$$\rho u \frac{\partial k}{\partial x} + \rho v \frac{\partial k}{\partial r} = \frac{1}{r} \frac{\partial}{\partial r} \left( \frac{r \mu_T}{Pr_k} \frac{\partial k}{\partial r} \right) + \mu_T \left( \frac{\partial u}{\partial r} \right)^2 - \rho \epsilon + \rho g \frac{\beta}{2} (qk)^2$$

## 2. TKE DISSIPATION RATE

$$\rho u \frac{\partial \epsilon}{\partial x} + \rho v \frac{\partial \epsilon}{\partial r} = \frac{1}{r} \frac{\partial}{\partial r} \left( r \frac{\mu_T}{Pr_\epsilon} \frac{\partial \epsilon}{\partial r} \right) + 1.44 \frac{\epsilon}{k} \mu_T \left( \frac{\partial k}{\partial r} \right)^2 - 1.92 \frac{\rho \epsilon^2}{k} + \rho g \beta \epsilon \left( \frac{q}{k} \right)^2$$

## 3. TEMPERATURE [ $q \equiv (T'2)^{1/2}$ ]

$$\rho u \frac{\partial q}{\partial x} + \rho v \frac{\partial q}{\partial r} = \frac{1}{r} \frac{\partial}{\partial r} \left( r \frac{\mu_T}{Pr_q} \frac{\partial q}{\partial r} \right) + 2.7 \mu_T \left( \frac{\partial T}{\partial r} \right)^2 - 1.8 \frac{\rho \epsilon}{k} q - 2k \left( \frac{\partial \epsilon}{\partial r} \right)^2$$

$$Pr_k = 0.6$$

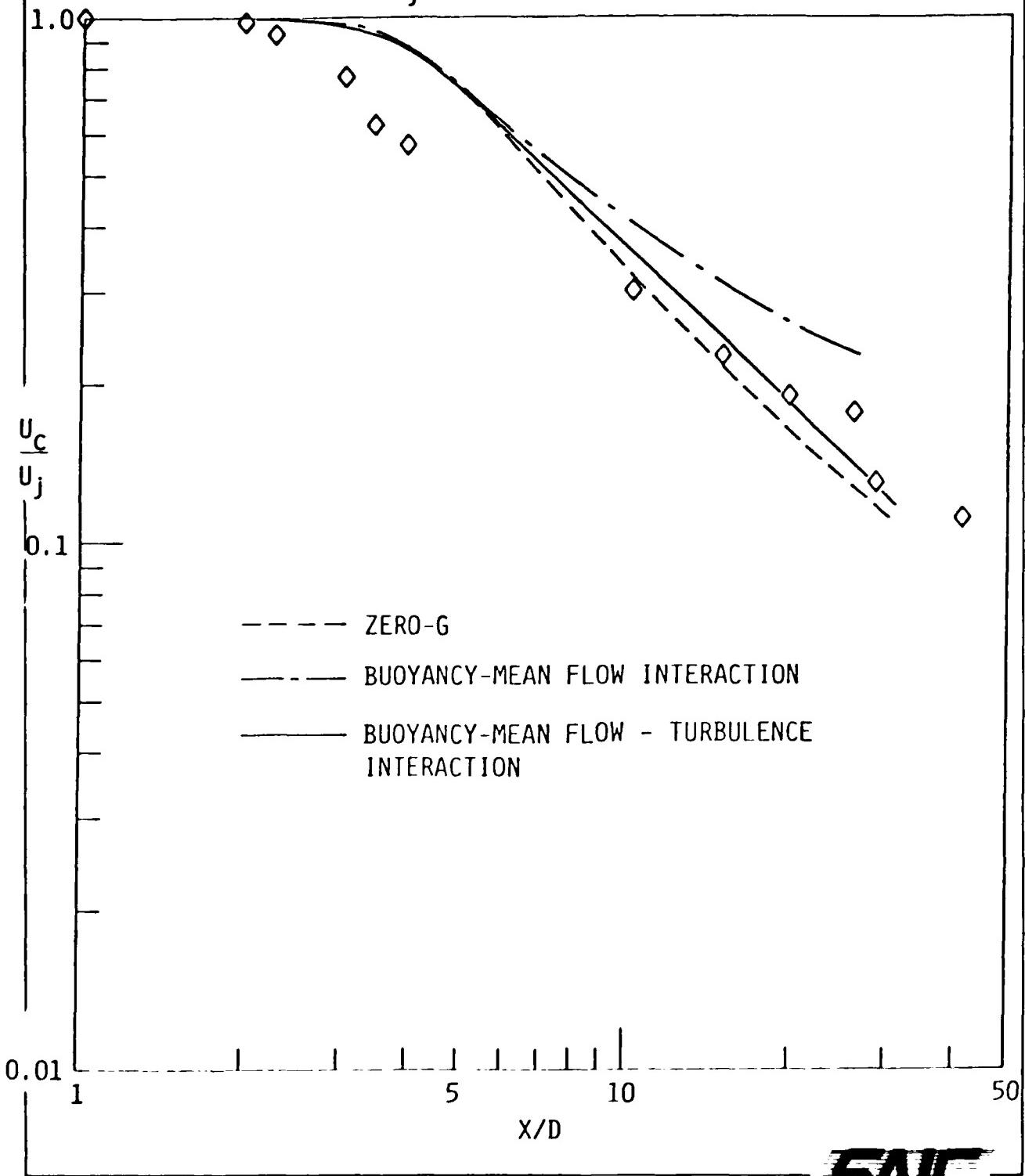
$$Pr = 0.6$$

$$Pr_q = 0.7$$

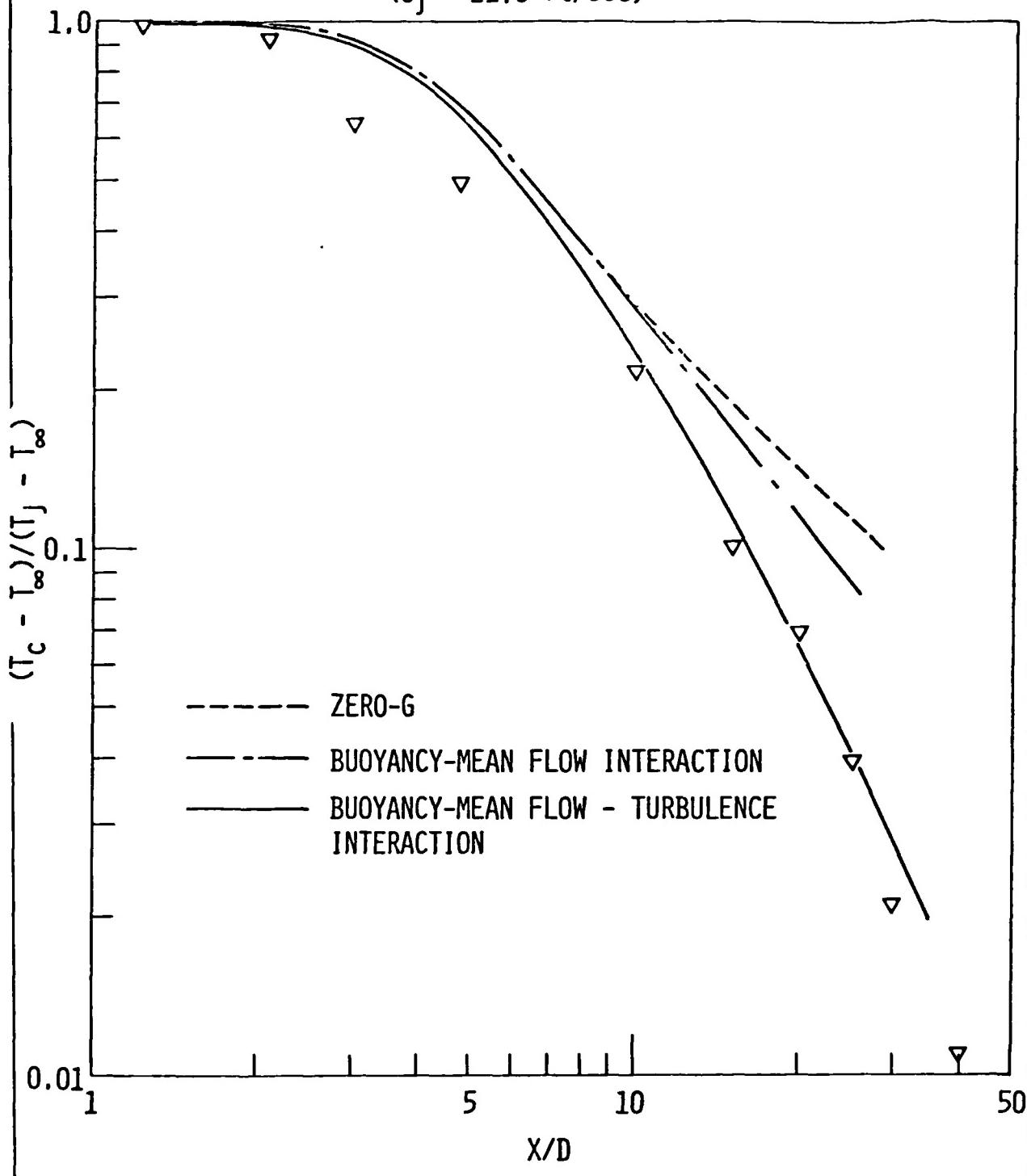
$$\mu_T = 0.09 \rho \frac{k^2}{\epsilon}$$

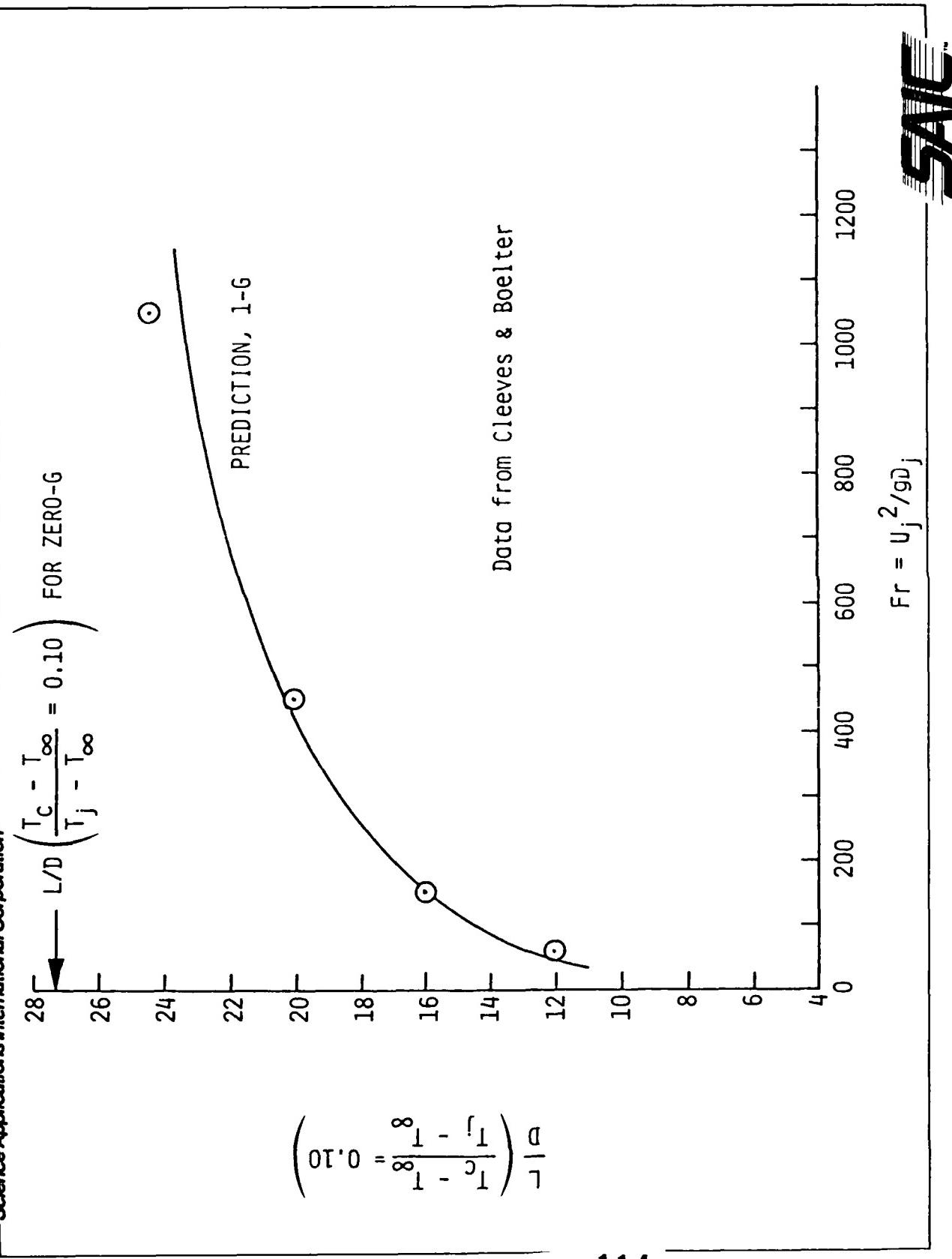
EFFECT OF MODELING ASSUMPTIONS ON PREDICTION OF  
CENTERLINE VELOCITY, HOT BUOYANT JET

( $U_j = 21.6$  ft/sec)

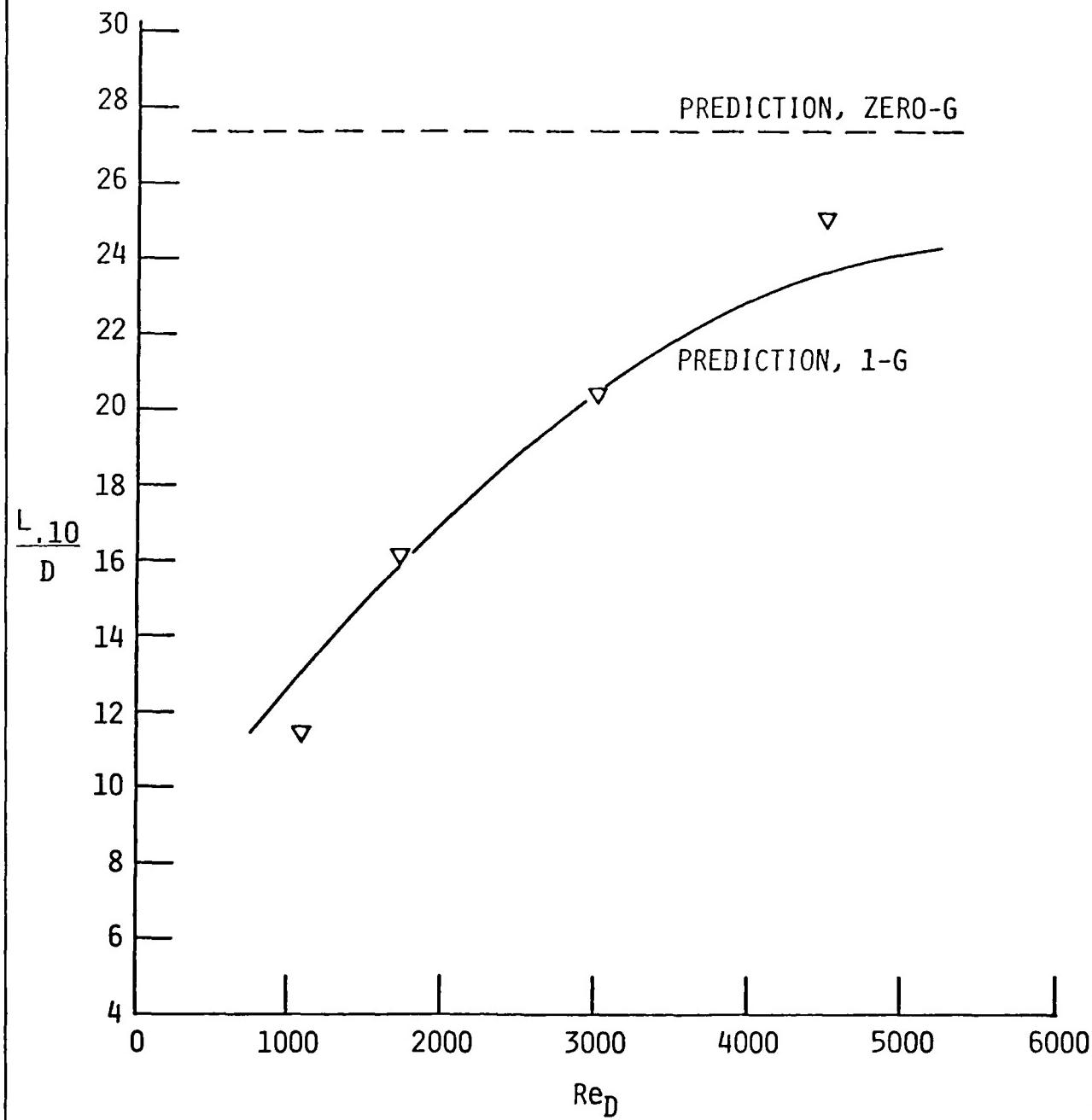


EFFECT OF MODELING ASSUMPTIONS ON PREDICTION OF  
CENTERLINE TEMPERATURE, HOT BUOYANT JET  
( $U_j = 21.6$  ft/sec)

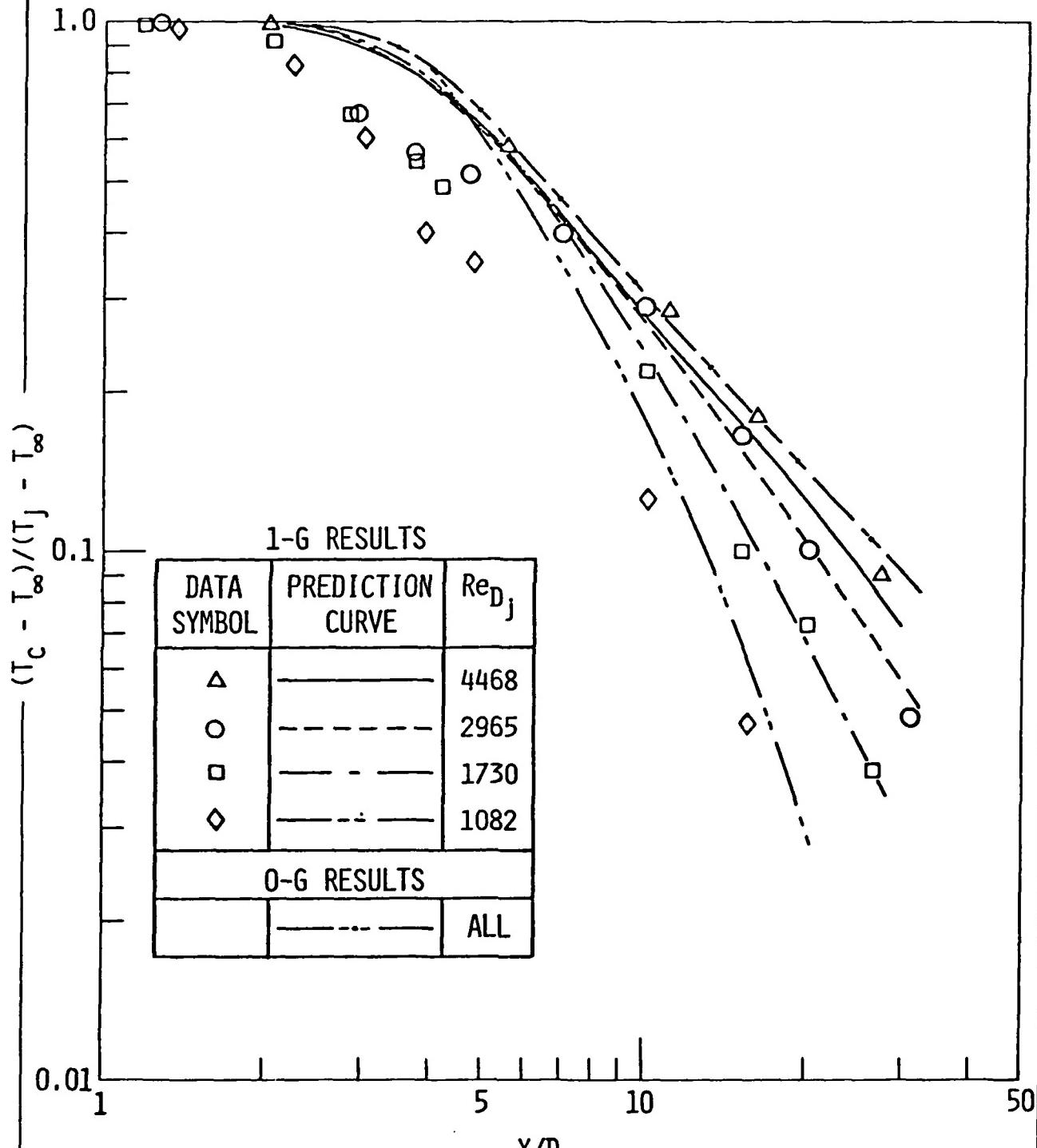




COMPARISON OF PREDICTED AND MEASURED  
BUOYANT PLUME LENGTHS



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## SUMMARY OF REQUIREMENTS

- Modeling
  - Application of Existing Multidimensional Analyses and Computer Codes Developed through the Incorporation of Appropriate Buoyancy-Induced Turbulence and Entrainment Submodels
    - ... Axisymmetric
    - ... 3-D
    - ... Chemical and Physical Processes
- Data
  - Laboratory Scales (cms)
  - Intermediate Scale (meters)
  - Large Scale (up to 100 meter)
- Various Programs of Differing Levels can be Defined to Address
  - Scaling
  - Multiple Plume Interactions
  - Fuel Effects



# On the Scavenging of Aerosol Particles by Snow Crystals

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## Aerosol Removal

Dry Removal

Wet Removal

In-Cloud  
Scavenging

Below-Cloud  
Scavenging

Rain  
Scavenging

Snow  
Scavenging

# Previous Research In Precipitation Scavenging

## 1. Rain Scavenging

Wang, P.K., and H.R. Pruppacher, 1977b

"An experimental determination of the efficiency with which aerosol particles are collected by water drops in subsaturated air"  
J. Atmos. Sci., 34, 1664-1669.

Wang, P.K., S.N. Grover, and H.R. Pruppacher, 1978

"On the effect of electric charges on the scavenging of aerosol particles by clouds and small raindrops"  
J. Atmos. Sci., 35, 1735-1743.

Wang, P.K., 1979a

"Particular solutions to the steady-state diffusion equation and their application to aerosol scavenging problems"  
Pap. Meteor. Res. (Taipei), 2, 37-42.

Wang, P.K., and H.R. Pruppacher, 1980a

"The effect of an external electric field on the scavenging of aerosol particles by cloud and small raindrops"  
J. Coll. Interf. Sci., 75, 286-297.

Walcek, C., P.K. Wang, J.H. Topalian, S.K. Mitra, and H.R. Pruppacher, 1981

"An experimental test of a theoretical model designed to determinate the rate at which freely falling water drops scavenge SO<sub>2</sub> in air"  
J. Atmos. Sci., 38, 871-876.

Wang, P. K., 1983b

"Collection of aerosol particles by a conducting sphere in an external electric field-continuum regime approximation"  
J. Coll. Interf. Sci., 94, 301-318.

## 2. Snow Scavenging

Wang, P. K., and H. R. Pruppacher, 1980b

"On the efficiency with which aerosol particles of radius less than 1 micron are collected by columnar ice crystals"

Pure Appl. Geophys., 118, 1090-1108.

Martin, J.J., P.K. Wang, and H.R. Pruppacher, 1980a

"On the efficiency with which aerosol particles of radius larger than 1 micron are collected by columnar ice crystals"  
Pure Appl. Geophys., 118, 1109-1129.

## 2. Snow Scavenging (cont'd)

Martin, J.J., P.K. Wang, and H.R. Pruppacher, 1980b  
"A theoretical determination of the efficiency with which aerosol particles are collected by simple ice plates"  
J. Atmos. Sci., 37, 1628-1638.

Martin, J.J., P.K. Wang, and H.R. Pruppacher, 1980c  
"A theoretical study of the effect of electric charges on the efficiency with which aerosol particles are collected by ice crystal plates"  
J. Coll. Interf. Sci., 78, 44-56.

Wang, P.K., and S.M. Denzer, 1983  
"Mathematical description of the shape of plane hexagonal snow crystals"  
J. Atmos. Sci., 40, 1024-1028.

Wang, P. K., 1983a  
"On the definition of collision efficiency of atmospheric particles"  
J. Atmos. Sci., 40, 1051-1052.

Wang, P.K., 1985  
"A potential and stream function analysis of two-dimensional steady state convective diffusion equations involving Laplace fields"  
Int. J. Heat Mass Transfer, 28, 1089-1095.

Wang, P. K., C. H. Chuang, and N. L. Miller, 1985  
"Electrostatic, temperature, and vapor density fields surrounding stationary columnar ice crystals"  
J. Atmos. Sci., 42, 2371-2379.

Wang, P. K., 1985  
"A convective diffusion model for the scavenging of submicron aerosol particles by snow crystals of arbitrary shapes"  
accepted by J. de Rech. Atmos.

Wang, P. K., 1985  
"Brownian diffusion of charged fine particles surrounding a conducting cylinder in the presence of an external electric field"  
accepted by J. Aerosol Sci.

Wang, P. K., and D. Sauter  
"Experimental measurements of the snow scavenging efficiencies of aerosol particles"  
(to be submitted to J. Atmos. Sci.)



\*

\*

Shutter

Front Optics

Detector

AP

AP Chamber

Mass

Aeroco

Vector

Chamber

& Fred.



Front

Shutter

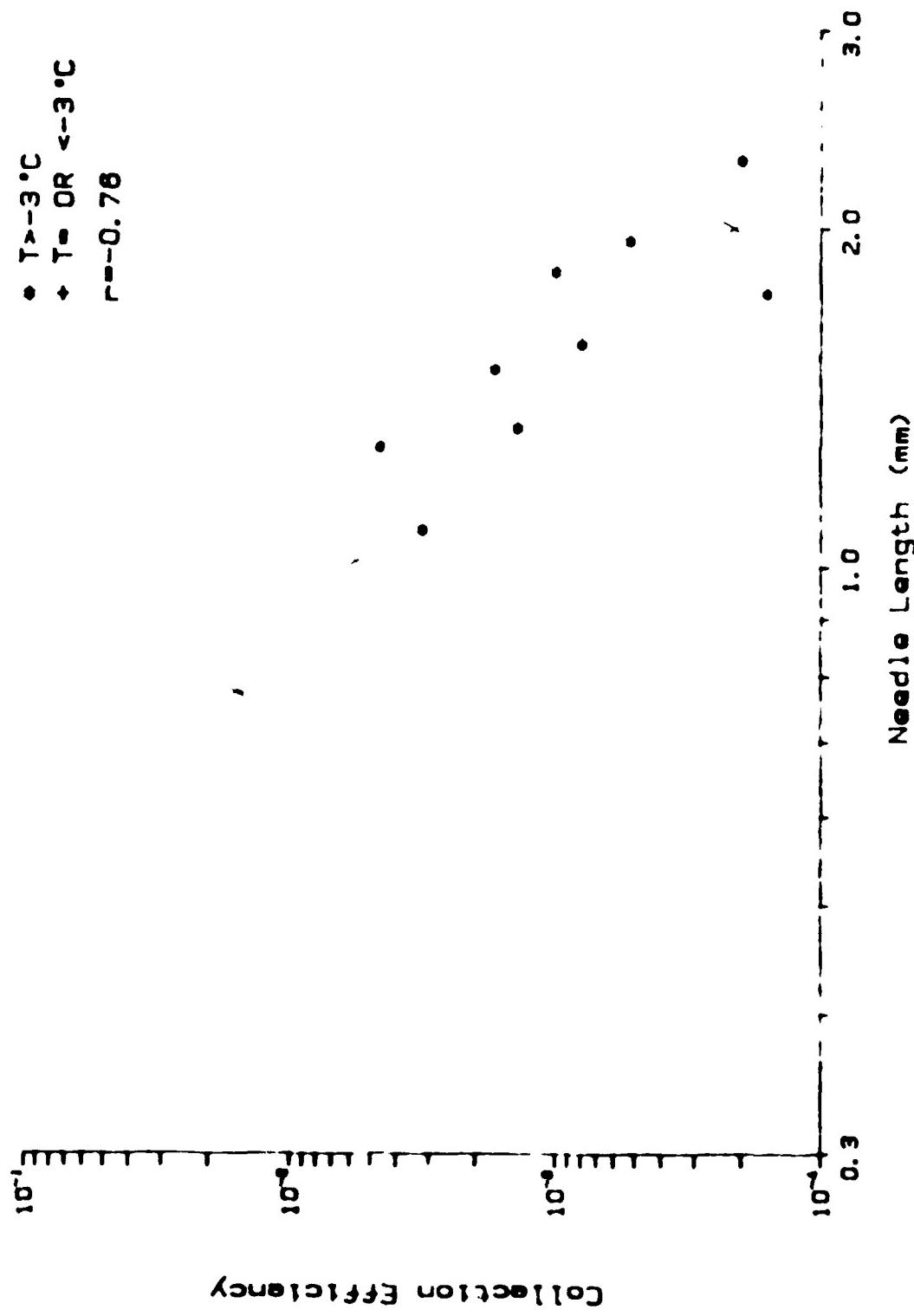
Filter

Collimator

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Snow-AP Shutter Vector

NEEDLES



STELLAR CRYSTALS

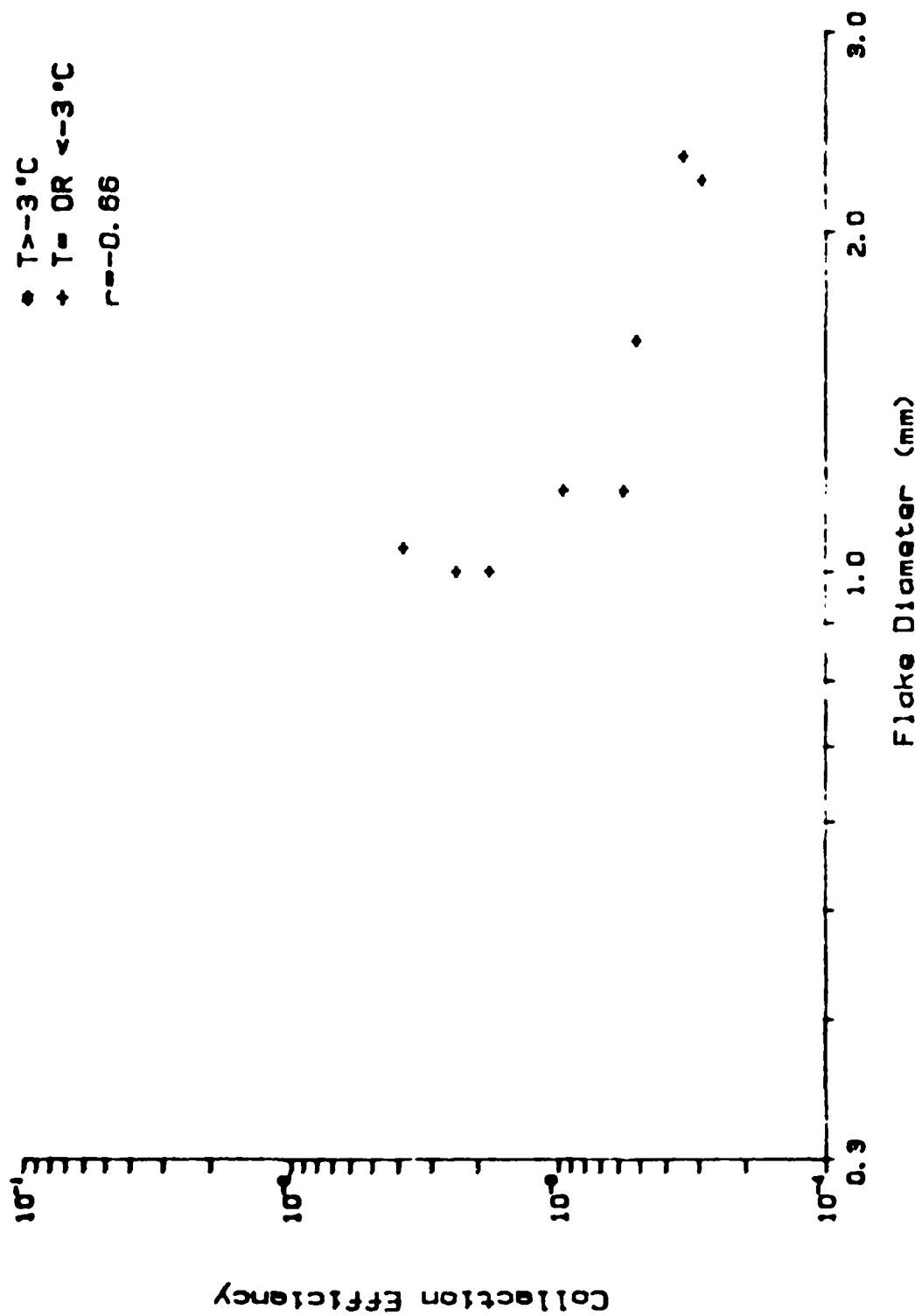


Fig. 4.11  
124

HEXAGONAL PLATES

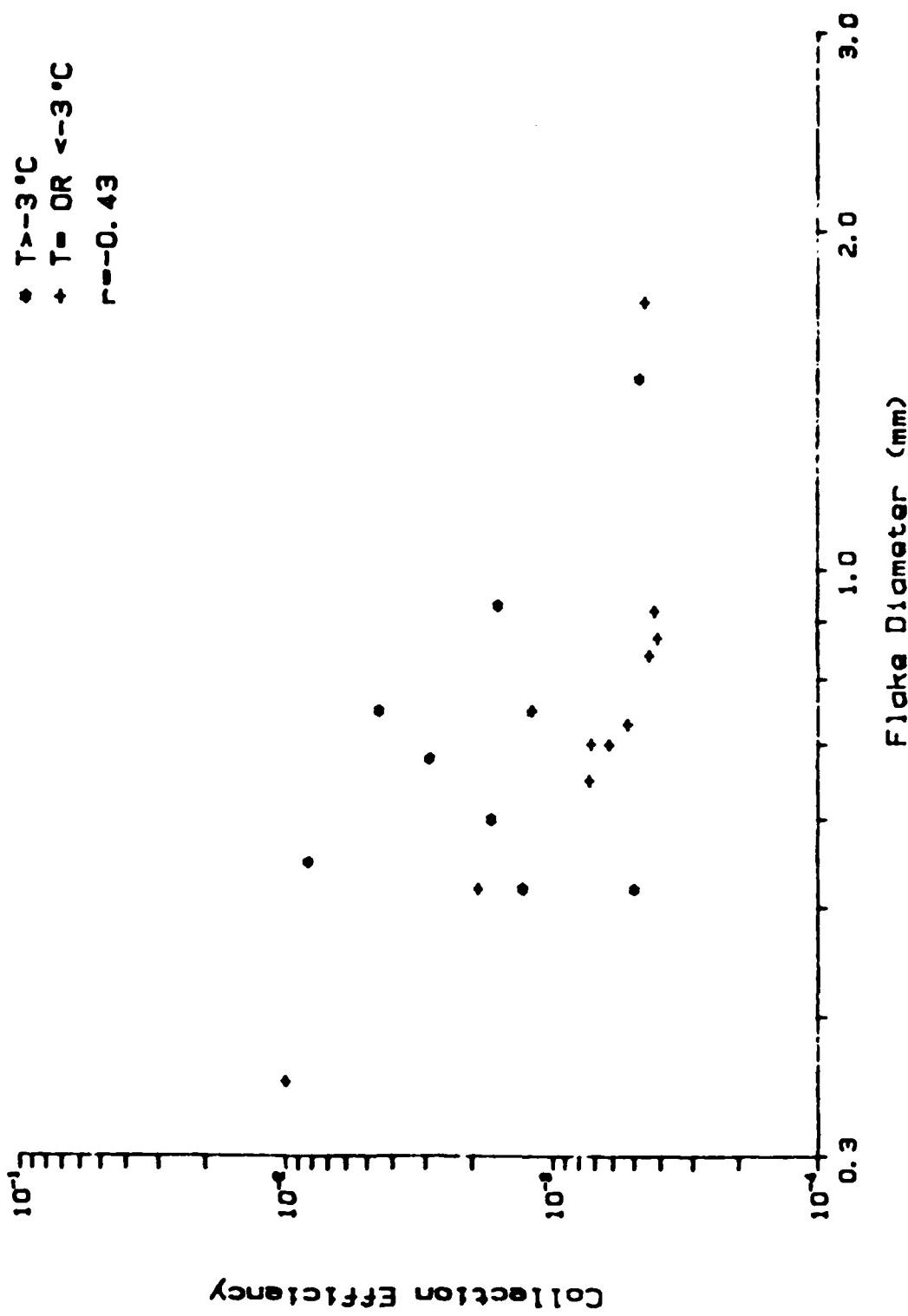


Fig 4.21

BROAD BRANCHED CRYSTALS

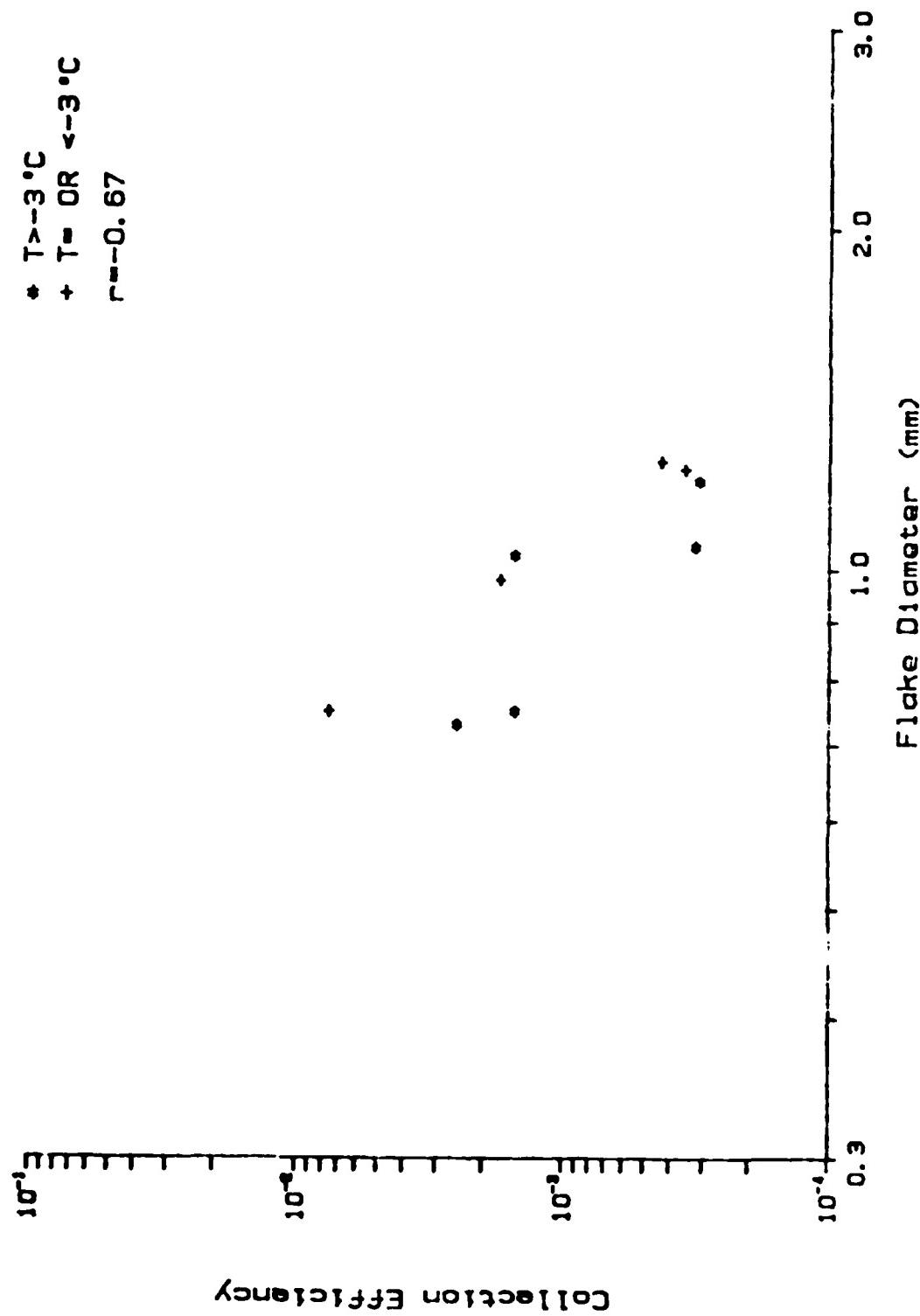


Fig. 4.16  
126

ALL PLATELIKE FLAKES

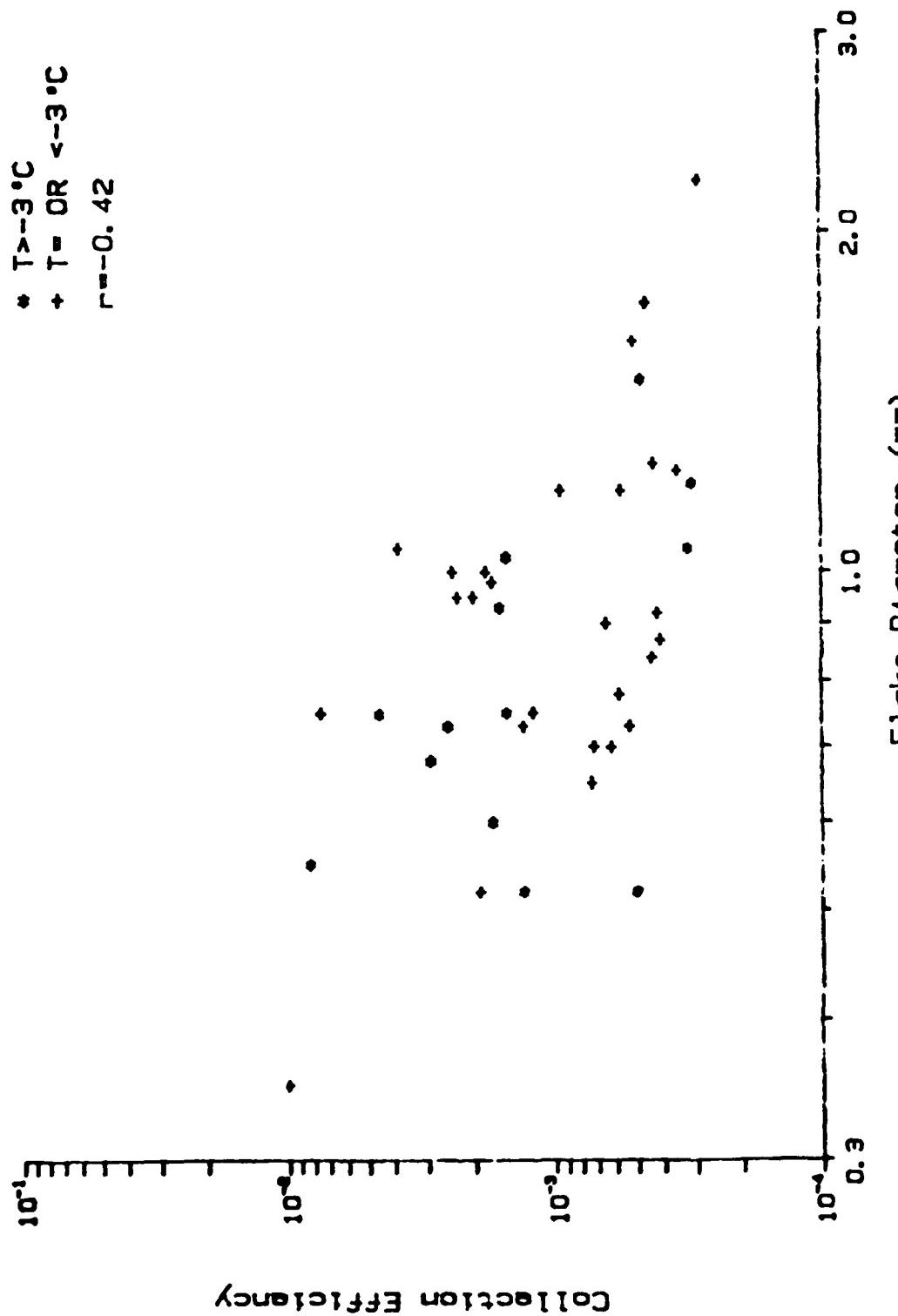
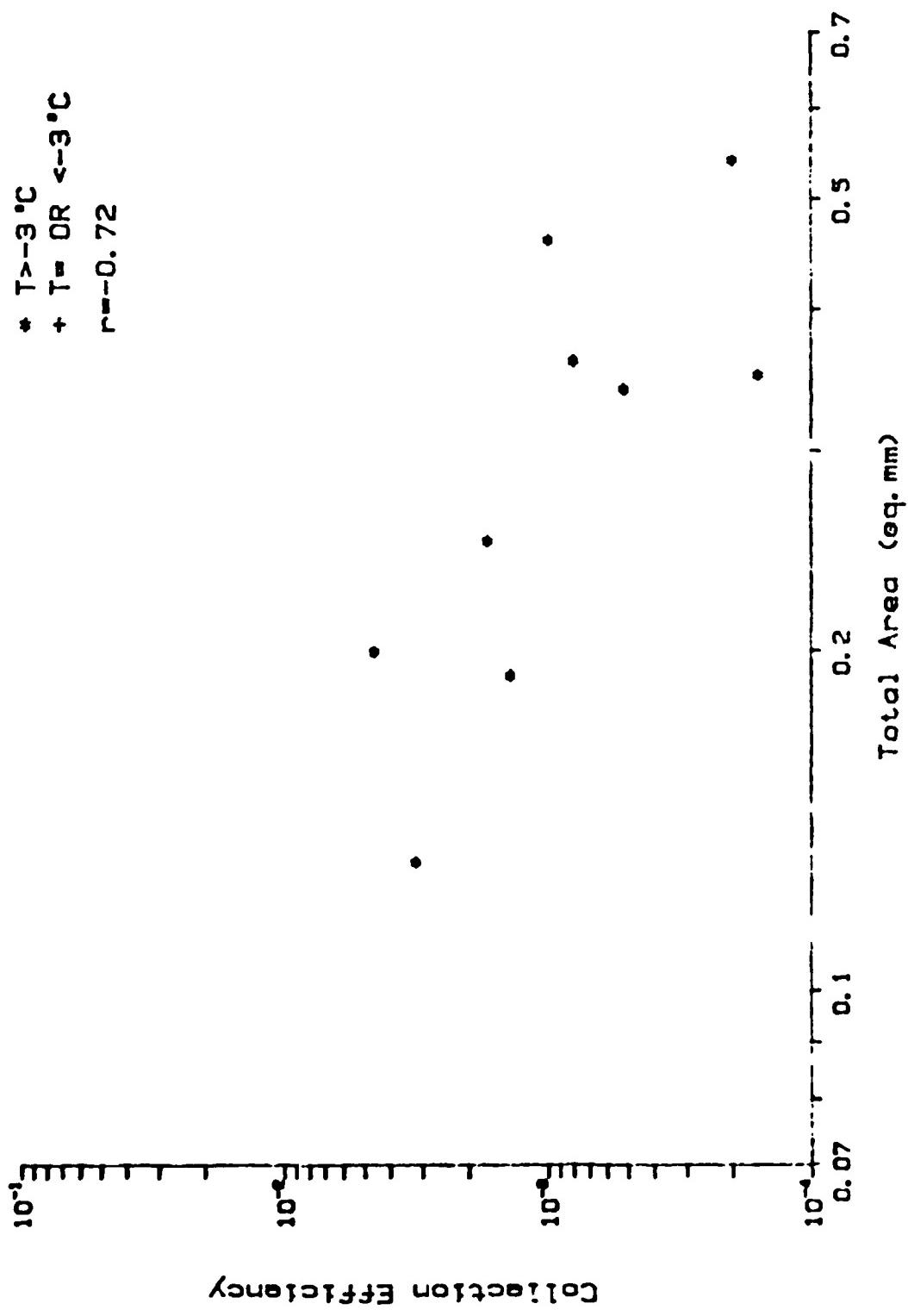


Fig 4.1  
127

NEEDLES



Collection Efficiency

~~Fig 83~~

128

STELLAR CRYSTALS

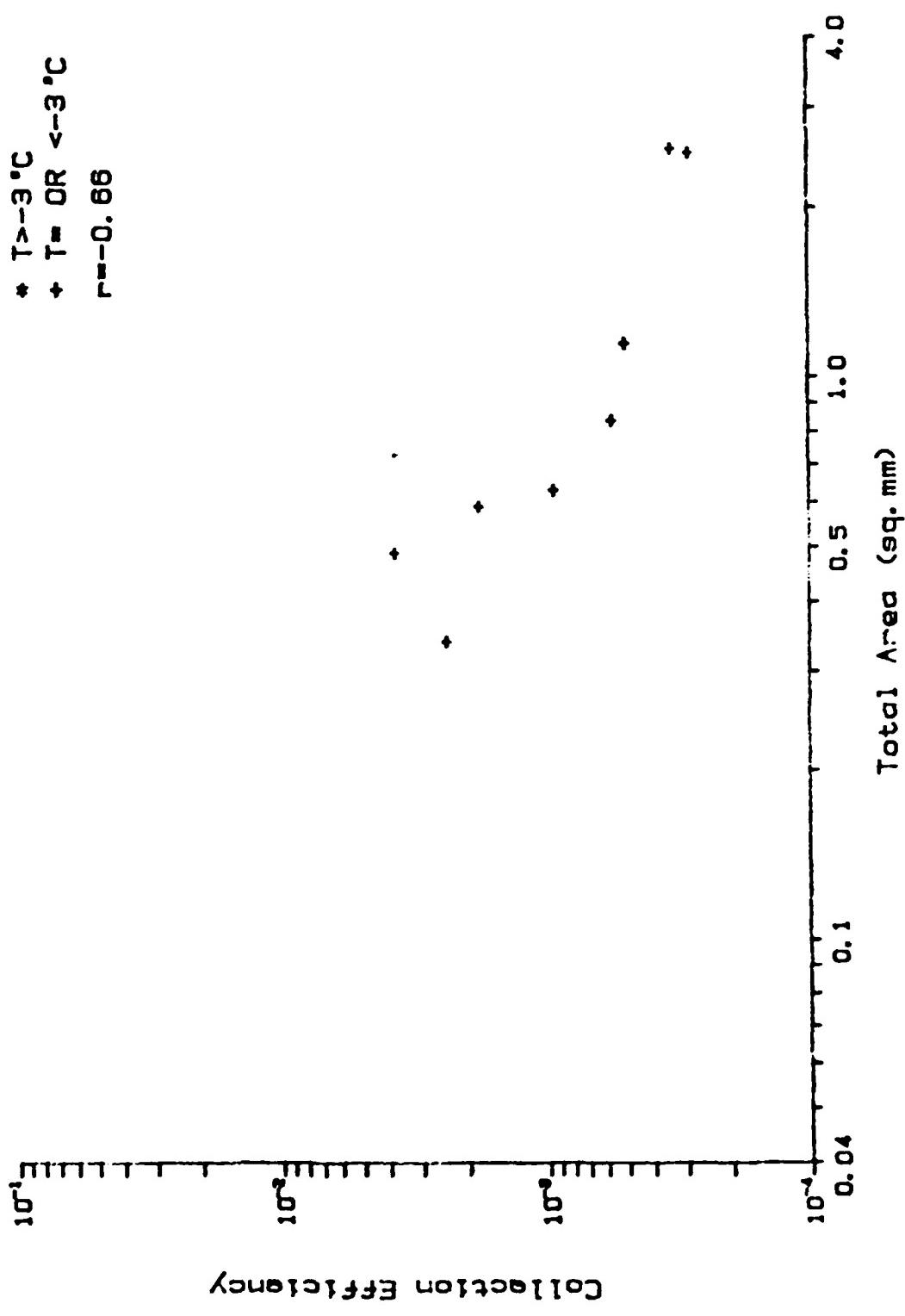
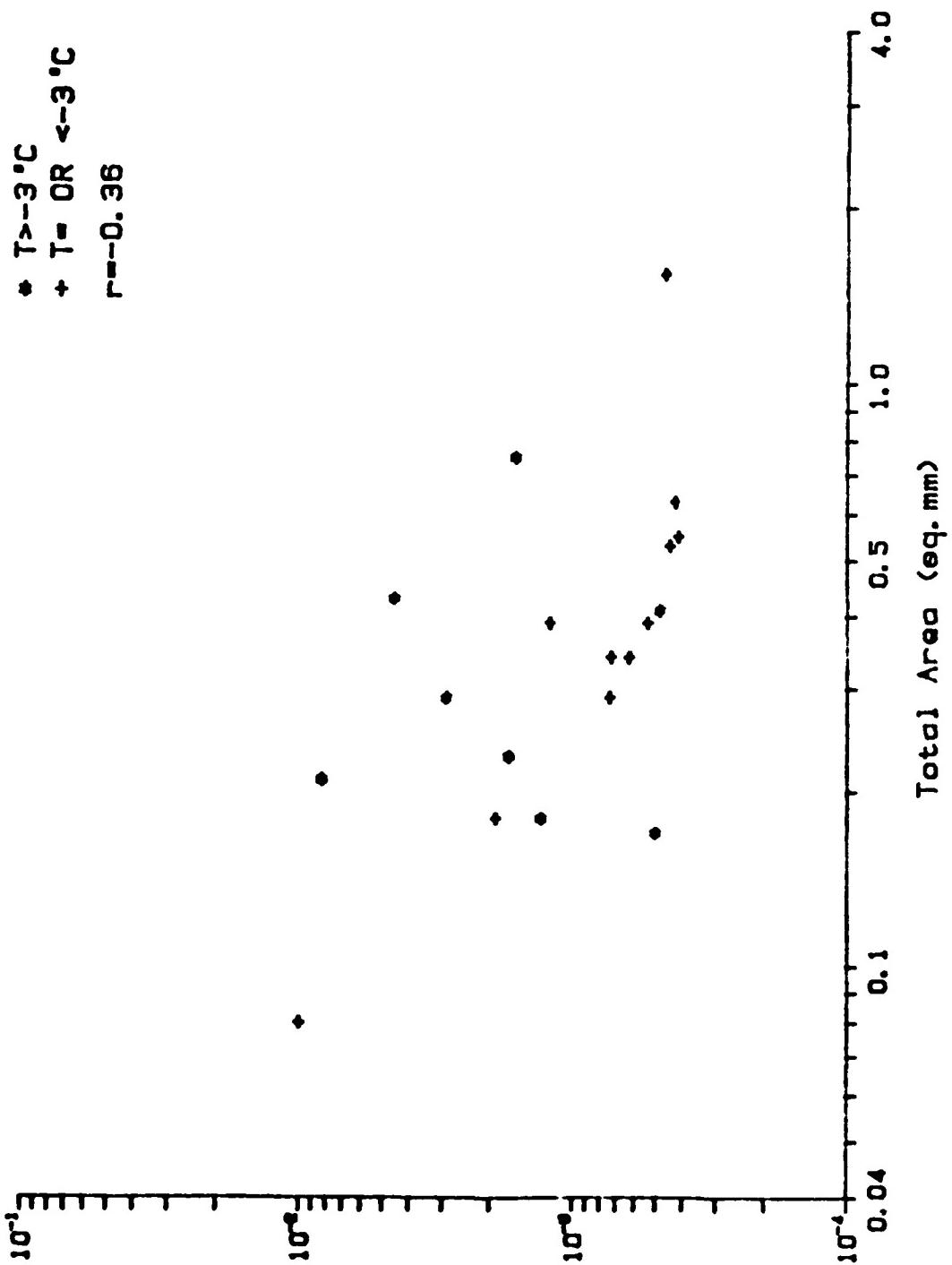


Fig. 4.14

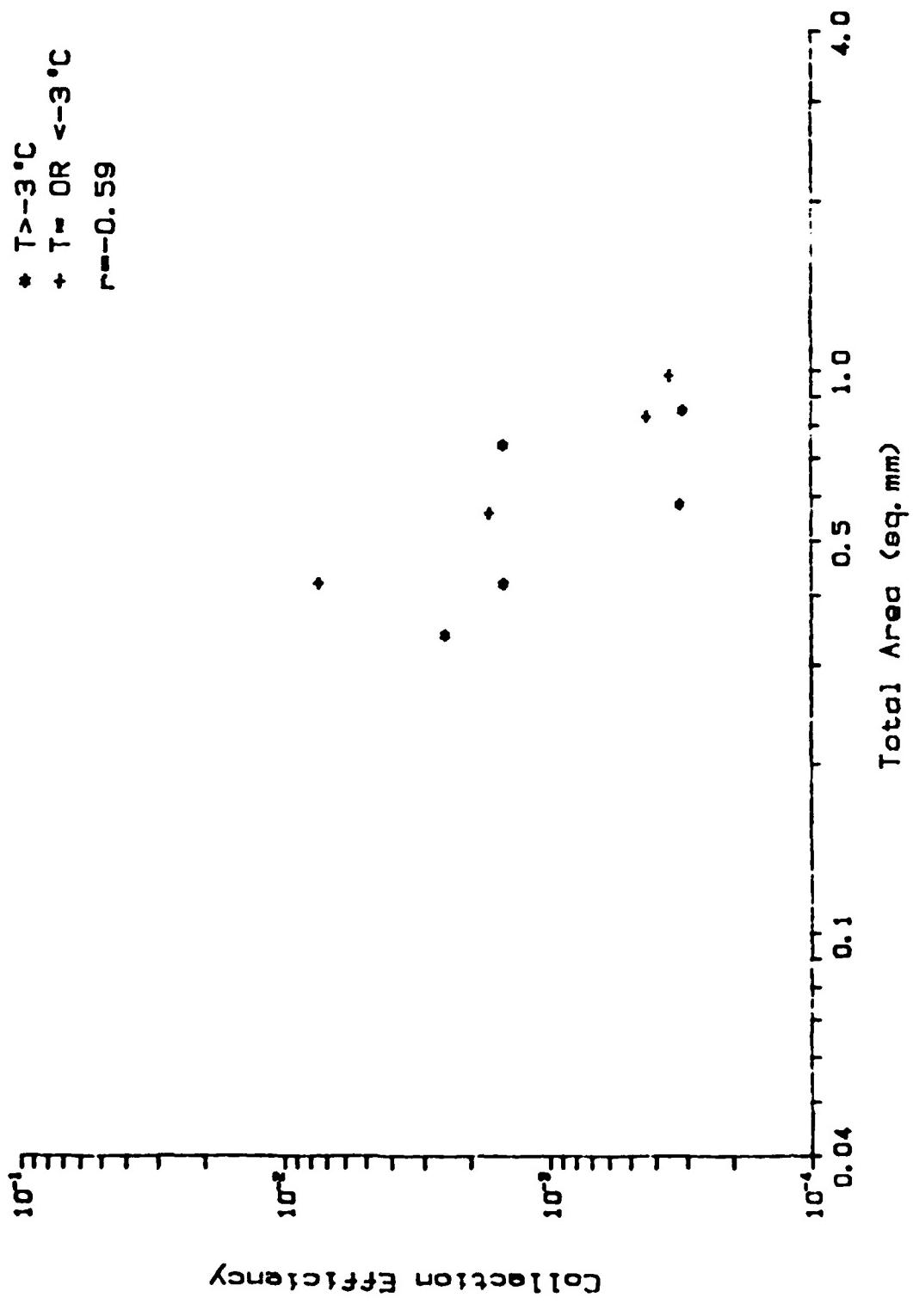
HEXAGONAL PLATES



COLLECTION EFFICIENCY

Fig. 4.24

BROAD BRANCHED CRYSTALS



ALL PLATELIKE FLAKES

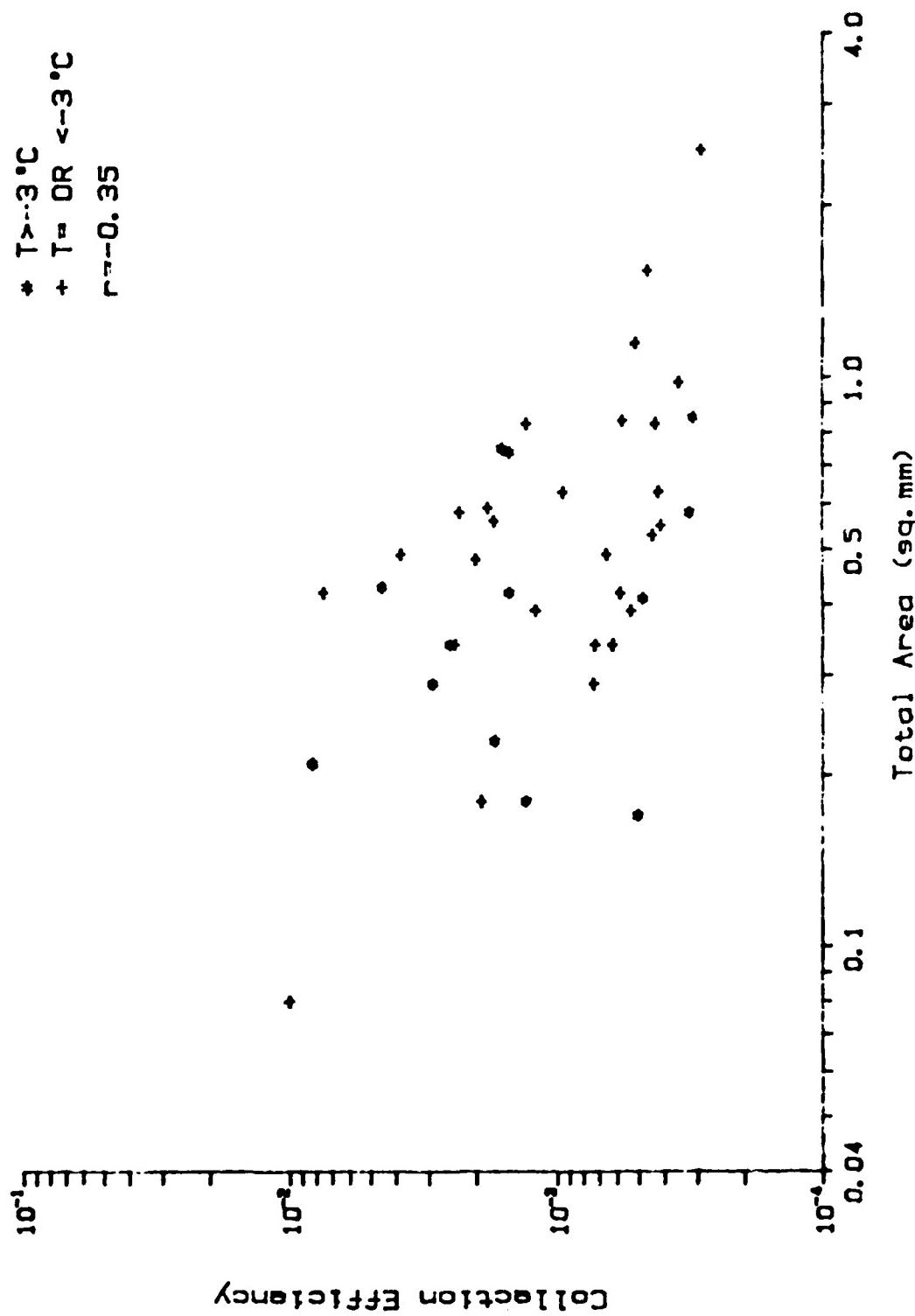


Fig 4.4

## Theoretical Consideration

Basically a convective diffusion problem

The mechanisms involved are:

1. Inertial impaction
2. Brownian diffusion
3. Thermophoresis
4. Diffusiophoresis
5. Electric forces
  - Static charges
  - Extrenal electric field
6. Turbulence

## Models:

1. Trajectory Model – Brownian diffusion not treated  
valid for particle radius  $> 0.5$  micron.
2. Flux Model – inertial impaction not treated  
valid for particle radius  $\leq 0.5$  micron.

**Flux Model:** (Good for  $r \leq 0.5 \mu\text{m}$ )

④

Convective Diffusion Equation (CDE)

$$D \nabla^2 n - B \vec{F} \cdot \nabla n = 0 \quad \dots (1)$$

Ventillation-Enhanced CDE

$$\bar{f}_p D \nabla^2 n - B \sum \bar{f}_i \vec{F}_i \cdot \nabla n = 0 \quad \dots (2)$$

where  $\bar{f}_i$ 's are overall ventilation factors.

B.C.  $\begin{cases} n=0 & \text{at } \Phi=\Phi_0 \text{ (Surface)} \\ n=n_\infty & \text{at } R \rightarrow \infty \end{cases} \dots (3)$

where  $\Phi$  is the force potential.

Solution of (2) :

$$n = n_\infty \left[ e^{\frac{B}{D\bar{f}_p}(\Phi-\Phi_0)} - 1 \right] / \left( e^{\frac{B\Phi_0}{D\bar{f}_p}} - 1 \right) \dots (4)$$

(5)

Resulting collection kernel :

$$K = -\frac{1}{n_\infty} \frac{\partial N}{\partial t} = -\frac{1}{n_\infty} \oint (n_B \vec{F} - D \vec{f}_p \nabla n) \cdot d\vec{S}$$

$$= \frac{B}{e^{\frac{D \vec{f}_p \Phi}{B}} - 1} \oint \vec{F} \cdot d\vec{S} \quad (\vec{F} = \sum_i \bar{f}_i \vec{F}_i) \quad \dots (5)$$

Collection Efficiency

$$E = \frac{K}{K^*} \quad \dots \dots \dots (6)$$

Consider the following external forces :

$$\vec{F} = \vec{F}_e + \vec{F}_{th} + \vec{F}_{df} \quad \dots \dots \dots (7)$$

$\vec{F}_e$  : Electric force ,  $\vec{F}_{th}$  : Thermophoretic force

$\vec{F}_{df}$  : Diffusiophoretic force.

(6)

Applying electrostatic theory,

$$\oint \vec{F}_e \cdot d\vec{s} = 4\pi C g (V_s - V_\infty) = 4\pi Q g$$

$$\oint \vec{F}_{th} \cdot d\vec{s} = 4\pi C \bar{f}_h Z_{th} (T_s - T_\infty)$$

$$\oint \vec{F}_{df} \cdot d\vec{s} = 4\pi C \bar{f}_h Z_{df} (\rho_{v,s} - \rho_{v,\infty})$$

V: Electric potential, T: Temp.,  $\rho_v$ : Vapor density

C: Capacitance (different for different shapes + sizes)

$$Z_{th} = \frac{12\pi\eta_a r (k_a + 2.5 k_p N_{thn}) k_a}{5(1+3N_{thn})(k_p + 2k_a + 5k_p N_{thn}) P}$$

$$Z_{df} = \frac{6\pi\eta_a r (0.74 D M_a)}{(1+\alpha N_{thn}) M_w \rho_a}$$

(7)

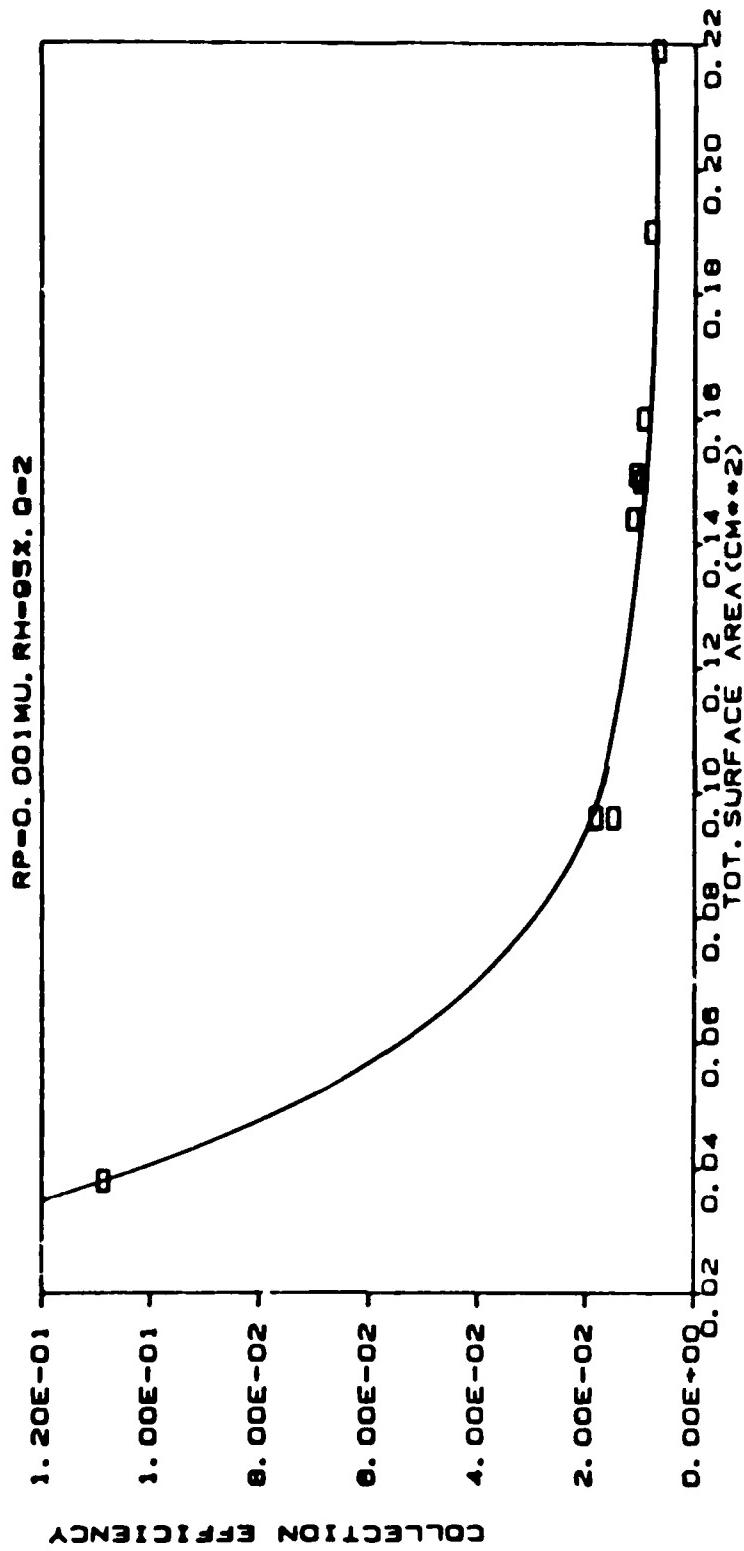
Therefore the collection kernel is

$$K = - \frac{4\pi B [Q_f + C \bar{f}_h \bar{\epsilon}_{th} (T_s - T_a) + C \bar{f}_h \bar{\epsilon}_{af} (\rho_{v,s} - \rho_{v,a})]}{e^{\frac{Q}{D_f p} \phi_0} - 1}$$

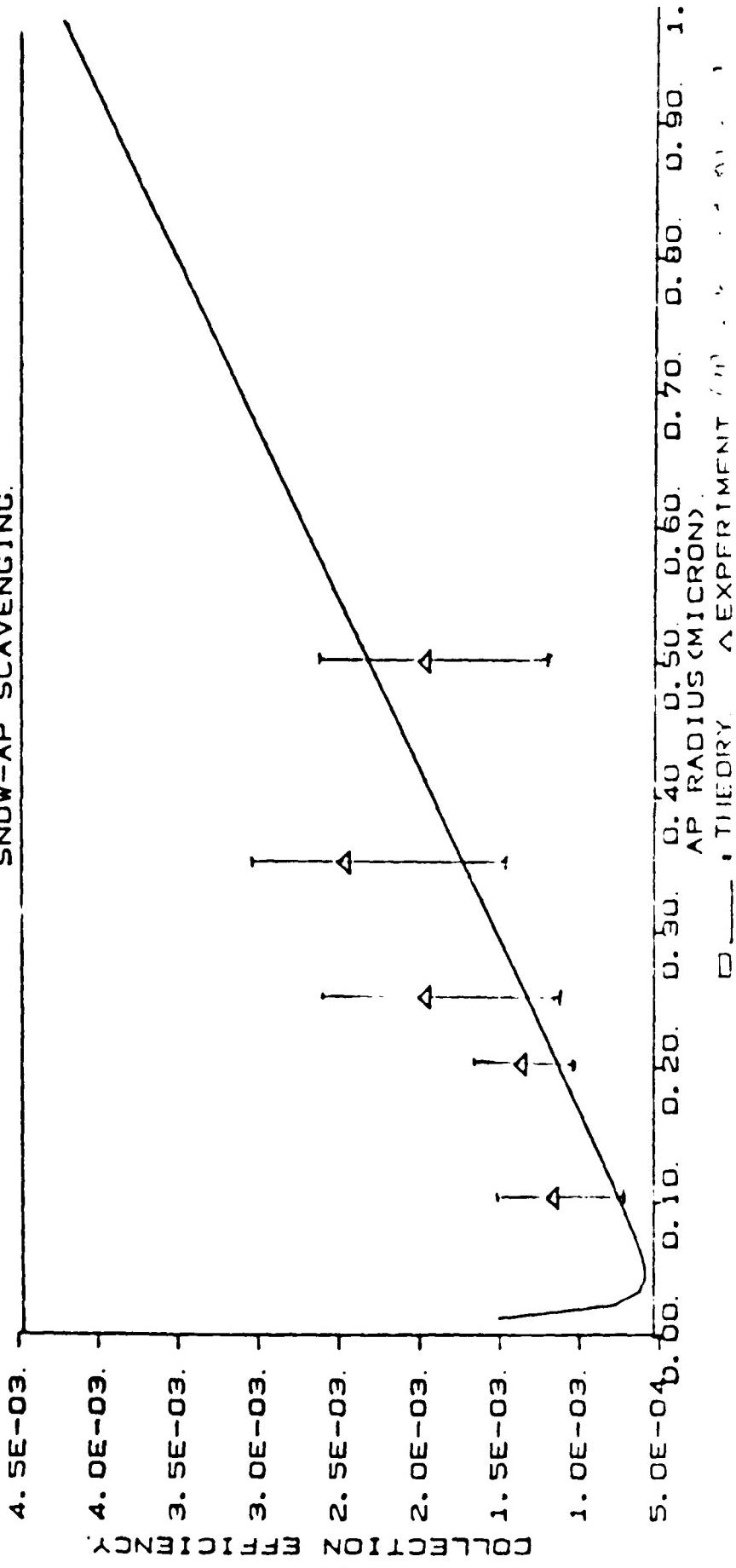
$$\epsilon = - \frac{4\pi B \phi_0 C}{e^{\frac{Q}{D_f p} \phi_0} - 1}$$

For sphere :  $C = a$  (radius)

For complicated shapes such as snow flakes,  
 $C$  can be measured experimentally.

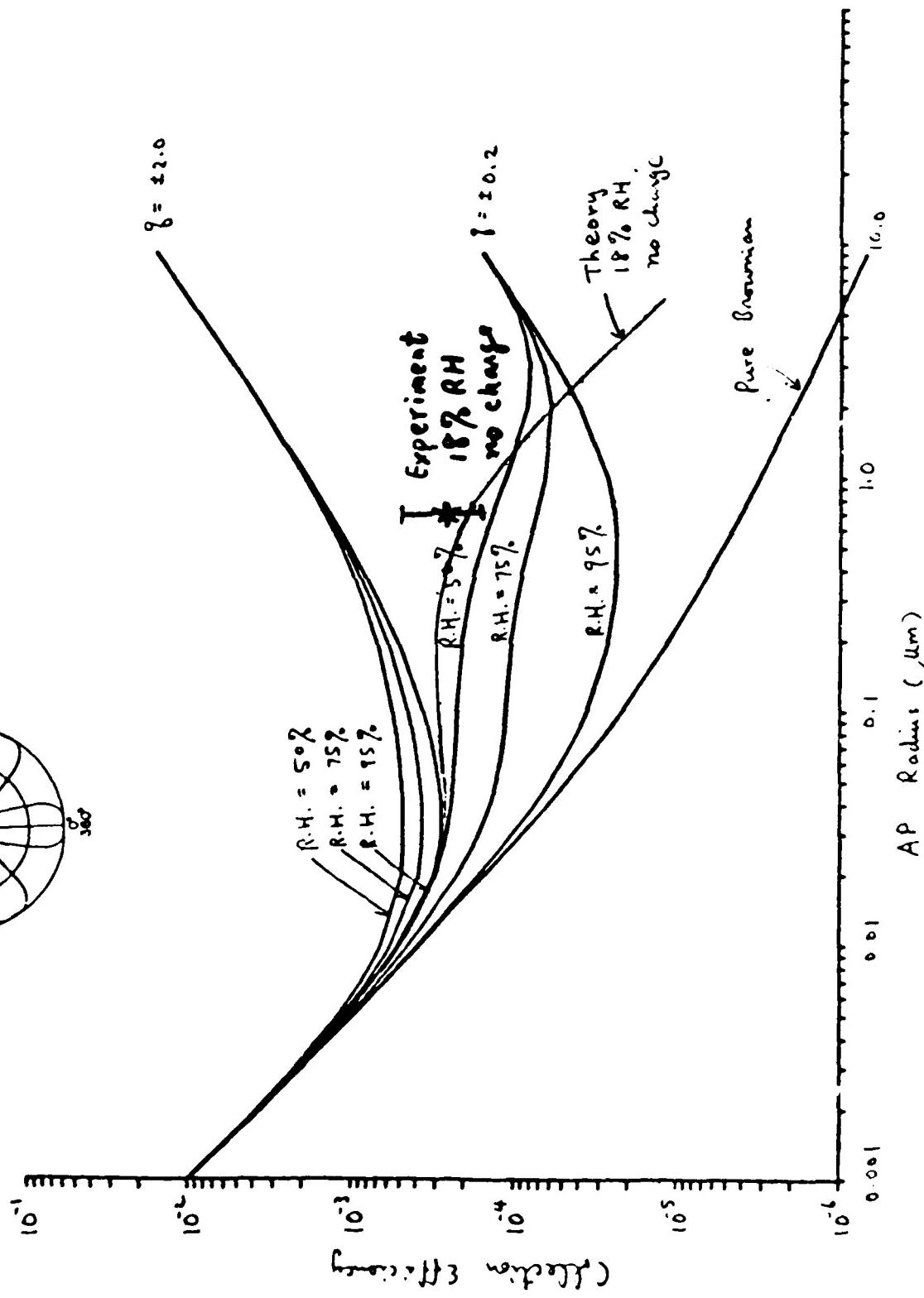
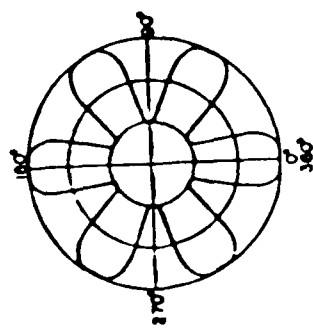


SNOW-AP SCAVENGING.



COLLECTION EFFICIENCY

Mar. Dimension = 0.2 cm



CLOUD MODEL SENSITIVITY STUDIES  
ON THE VERTICAL PENETRATION OF SOOT

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## OVERVIEW

- Review 3-D cloud model results  
dependence of upward soot penetration on  
atmospheric conditions
- Offers hope for simple parameterisation of vertical  
soot profiles
- Atmospheric stability indices - distributions  
by season  
by region
- Sample procedure: how distribution approach  
might be used to estimate amount of soot initially  
entering stratosphere

- Forest fire (and other) smoke plume heights — sensitive to atmospheric stability, moisture content, in addition to fire intensity — (NRC, '85)
- If urban fires have larger heat fluxes ( $10x$ ), are those larger fires also sensitive to atmospheric structure?
- 3-D numerical cloud simulations indicate that they are.

### 3-D Cloud Model Simulations

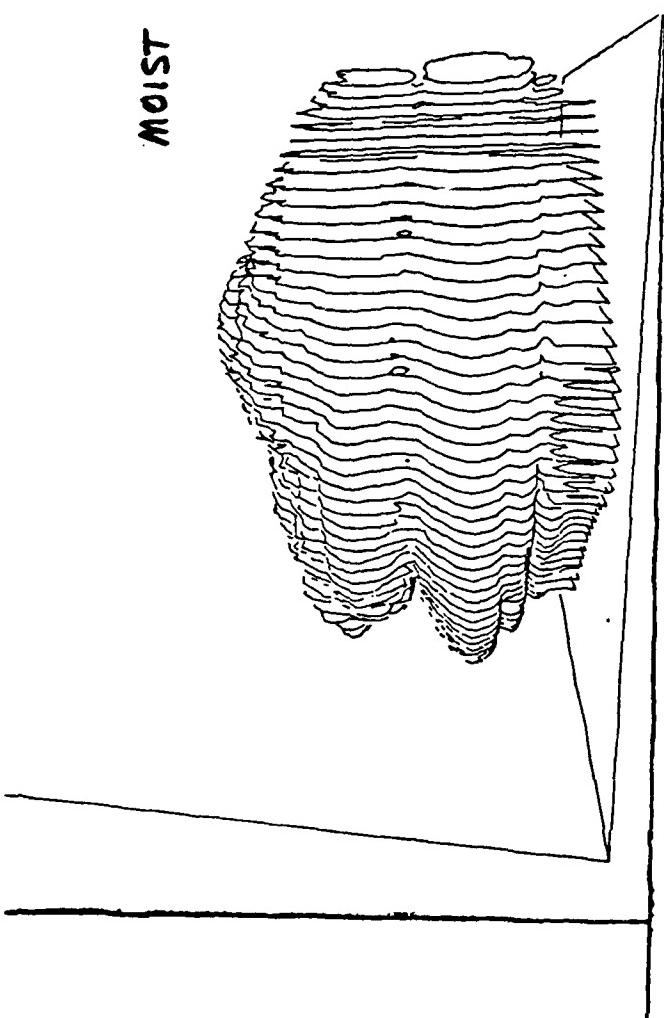
- standard atmosphere + moisture, winds
- to 30 min (60 min w/ reduced heat flux)
- cloud microphysics (cloud, rain, ice, "snow", graupel/hi!)
- soot tracer

see descriptions by Tripoli & Cotton

## SAMPLE OUTPUT

- 3-D pictures
- 2-D cross sections
- numerical output fields

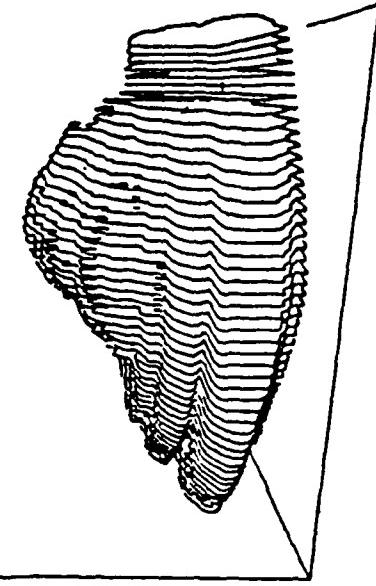
MOIST



DRY

131  
TIME = 1000. SECONDS

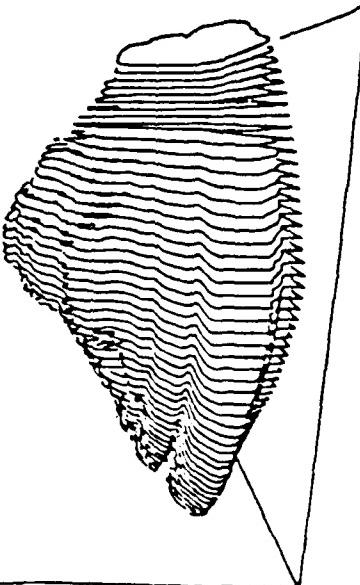
SHORE

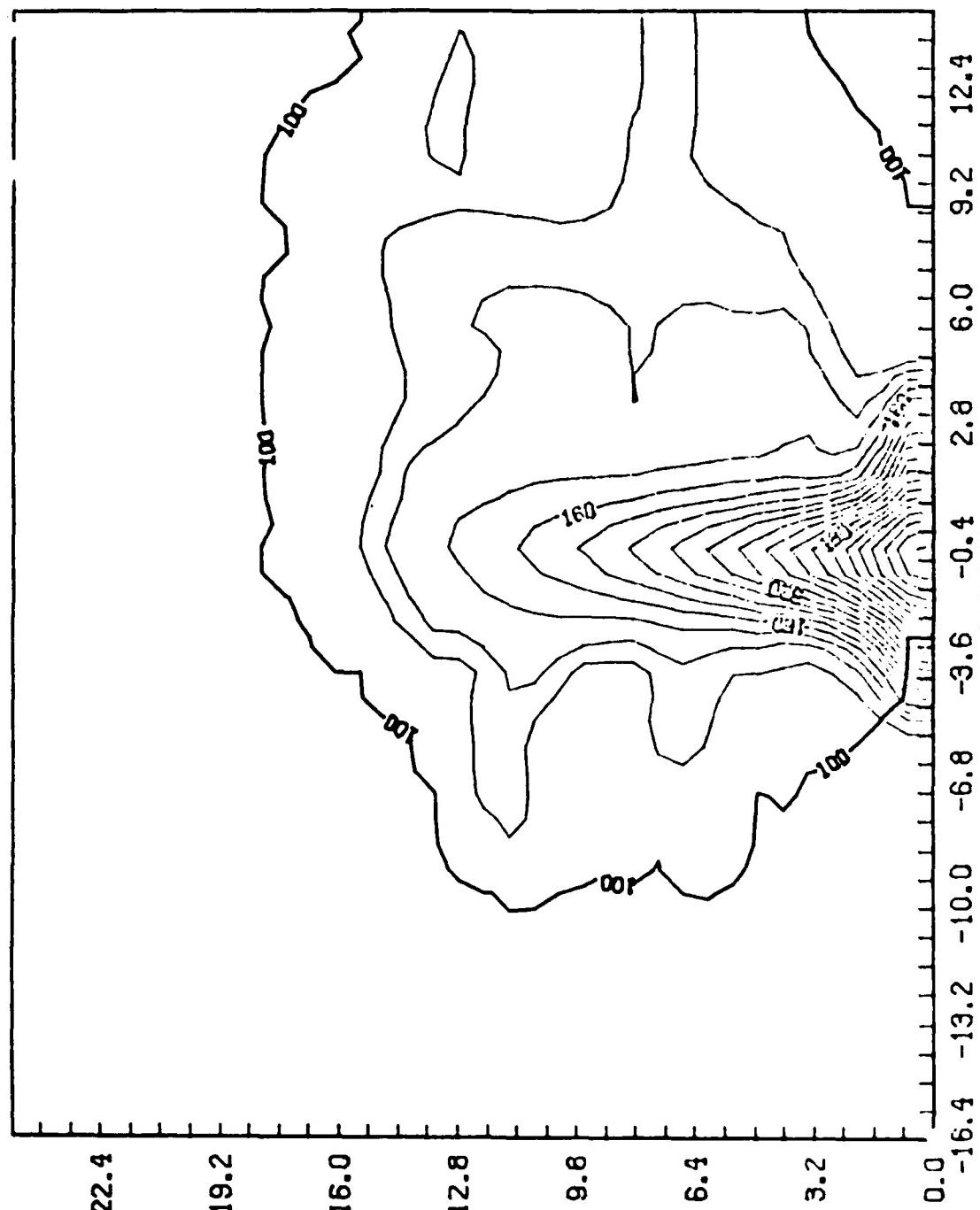


DRY  
ATMOSPHERE

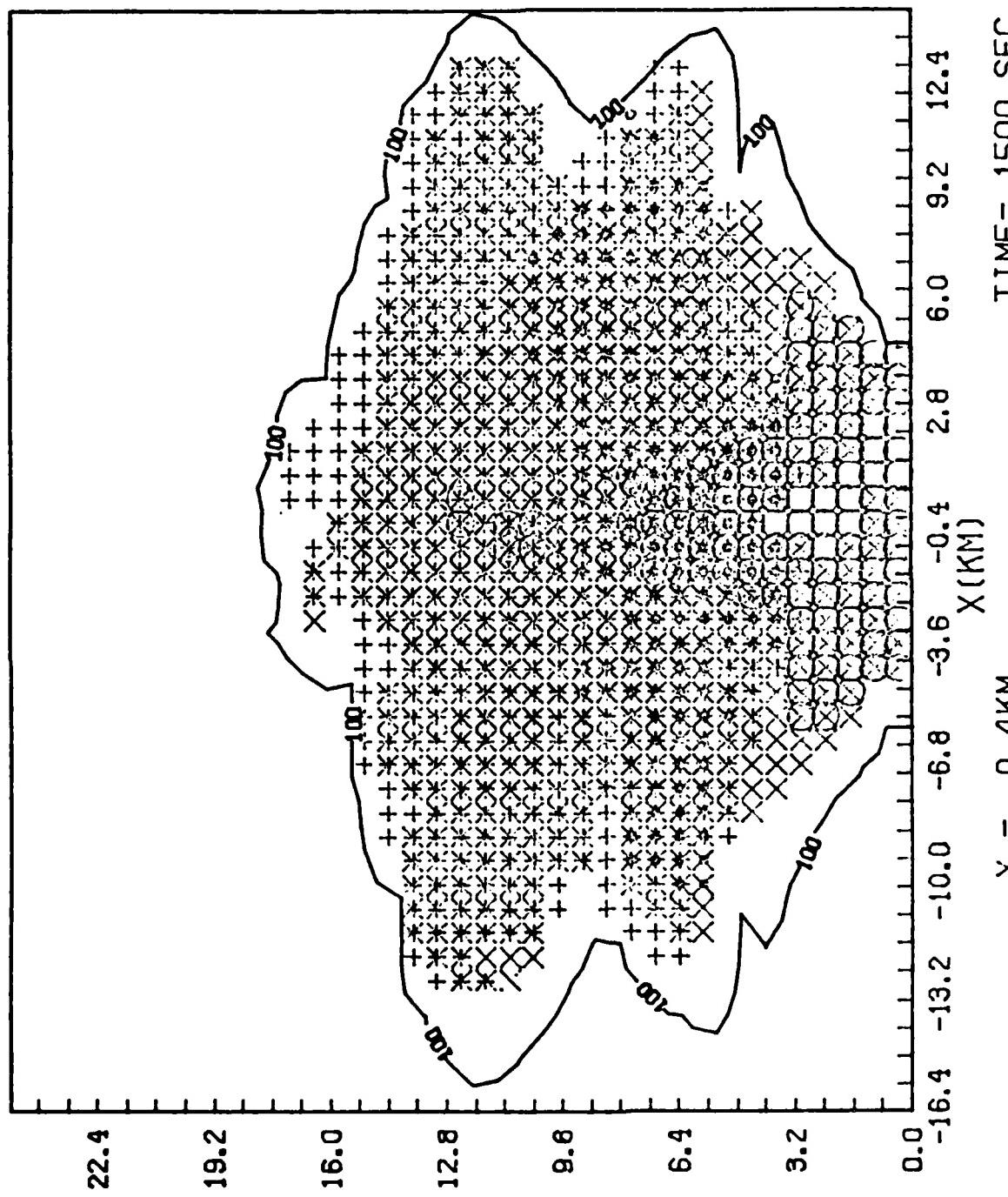
131.  
TIME = 1000. SECONDS

SHORE





X = 0.4KM SMOKE (G/M3) TIME= 1500. SEC



## SUMMARY OF MODEL-RUN COMPARISONS

### Stability

unstable (Cotton & Tripoli): 44% into stratosphere  
stable (std. atmos): 22% " "  
" "

### Moisture (std atmos)

with 40-50% RH in troposphere: 22% into stratosphere  
with 0 moisture < 1% " "

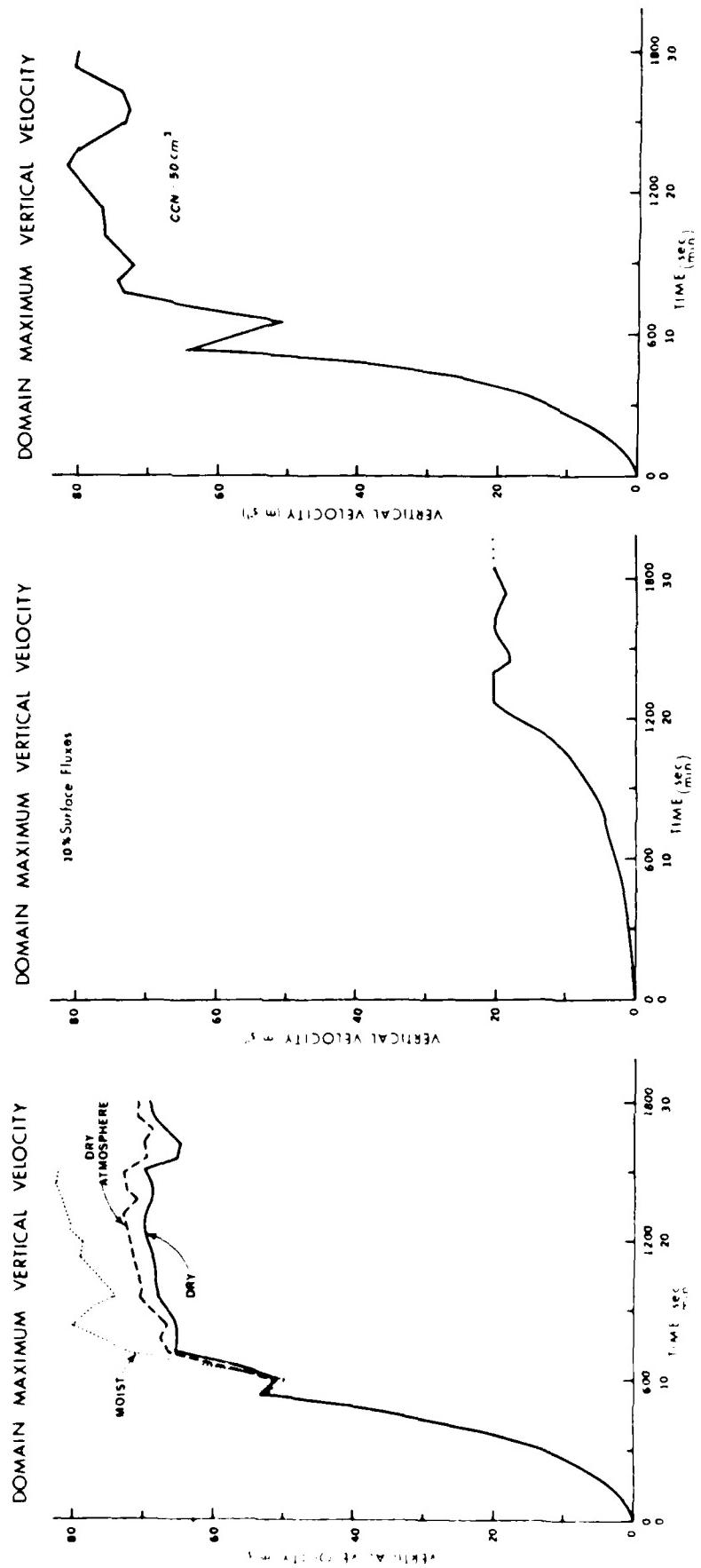
### Heat flux (std atmos)

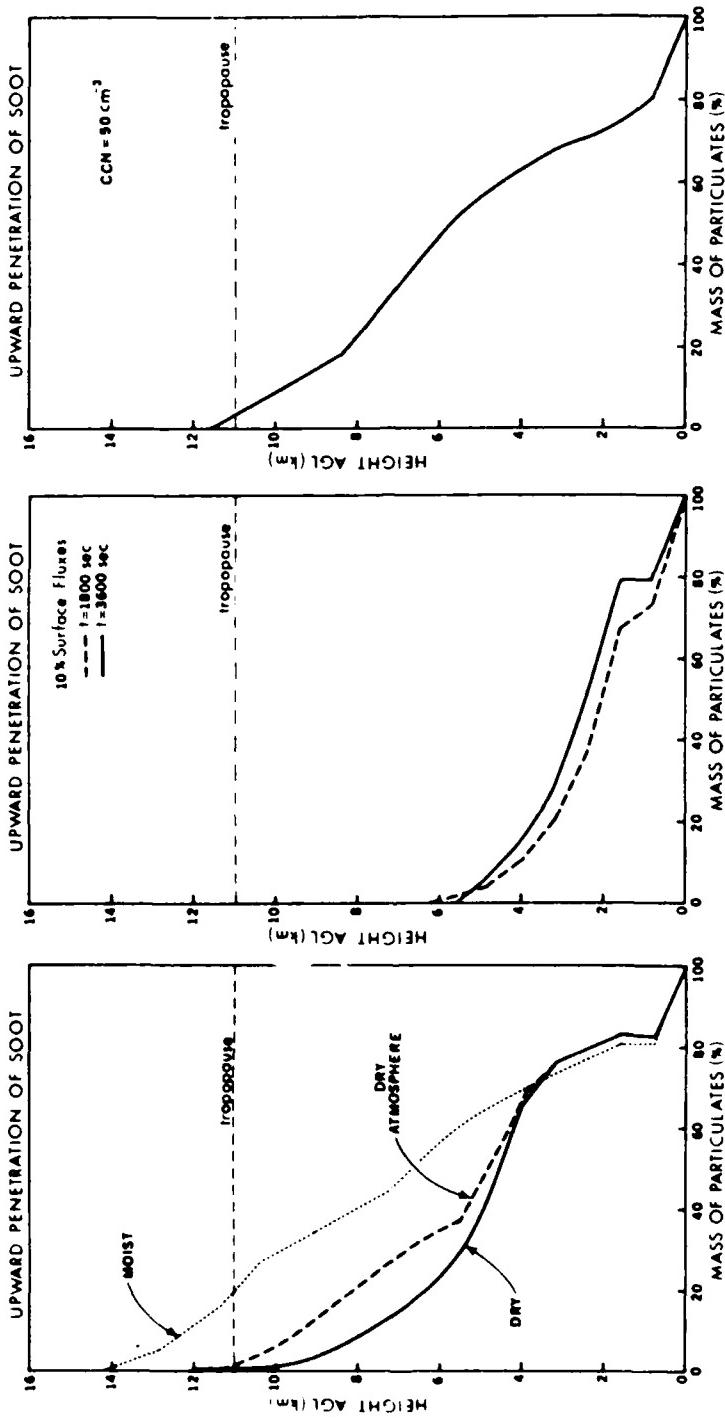
100% - smoke penetrates to 14 km.  
10% - no smoke above 6 km.

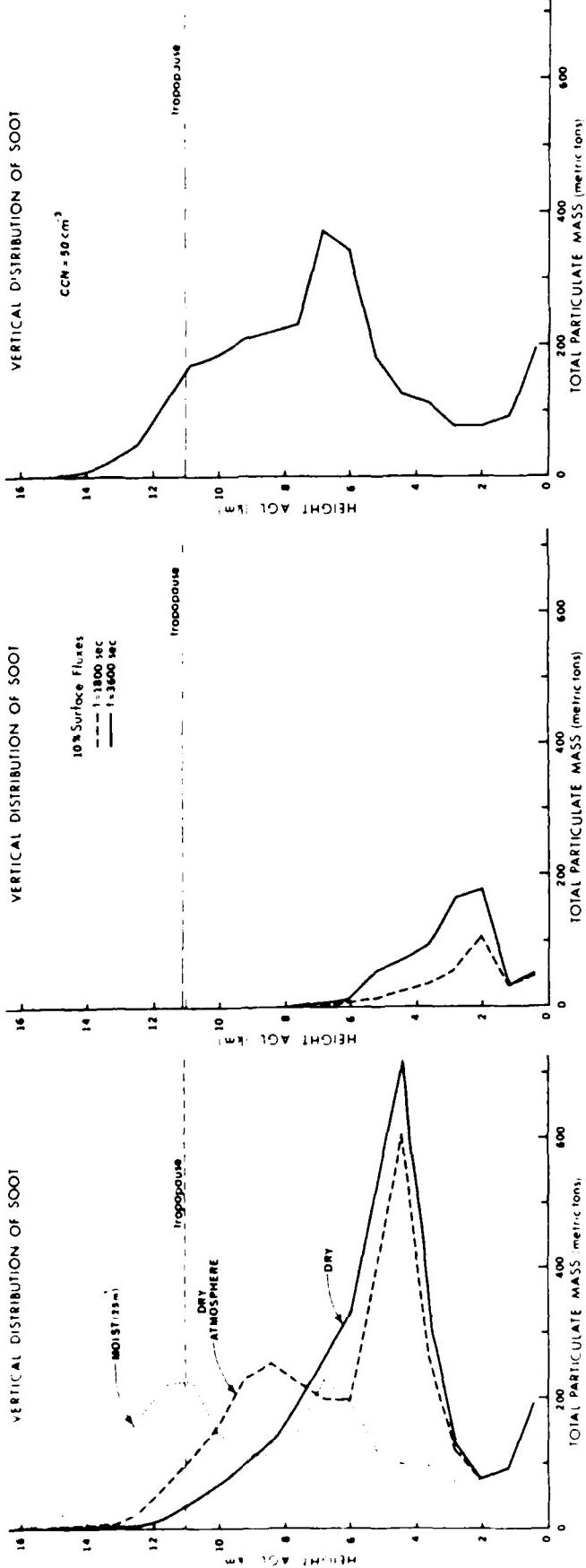
Large heat-flux lines - sensitive to atmosphere

## MODEL - RUN COMPARISONS

One more issue - CCN concentrations  
5000  $\text{cm}^{-3}$  22% into stratosphere  
50  $\text{cm}^{-3}$  11% into stratosphere  
(attempt to represent some large nuclei)  
less latent heat of fusion  
but with the phoretic effects modeled,  
 $< 2\%$  scavenging in both cases







## PARAMETERIZATION OF VERTICAL INJECTION PROFILE

Previous results suggest that vertical injection profiles can be parameterized in terms of some stability (etc.) parameter as input into larger-scale models. Need

- profile shape
  - uniform
  - detrainment curve
- height of plume
  - fire heat flux
  - atmospheric stability
  - atmospheric moisture (low-level precipitable water)
  - wind profile
  - tropopause height (normalize?)
  - etc.

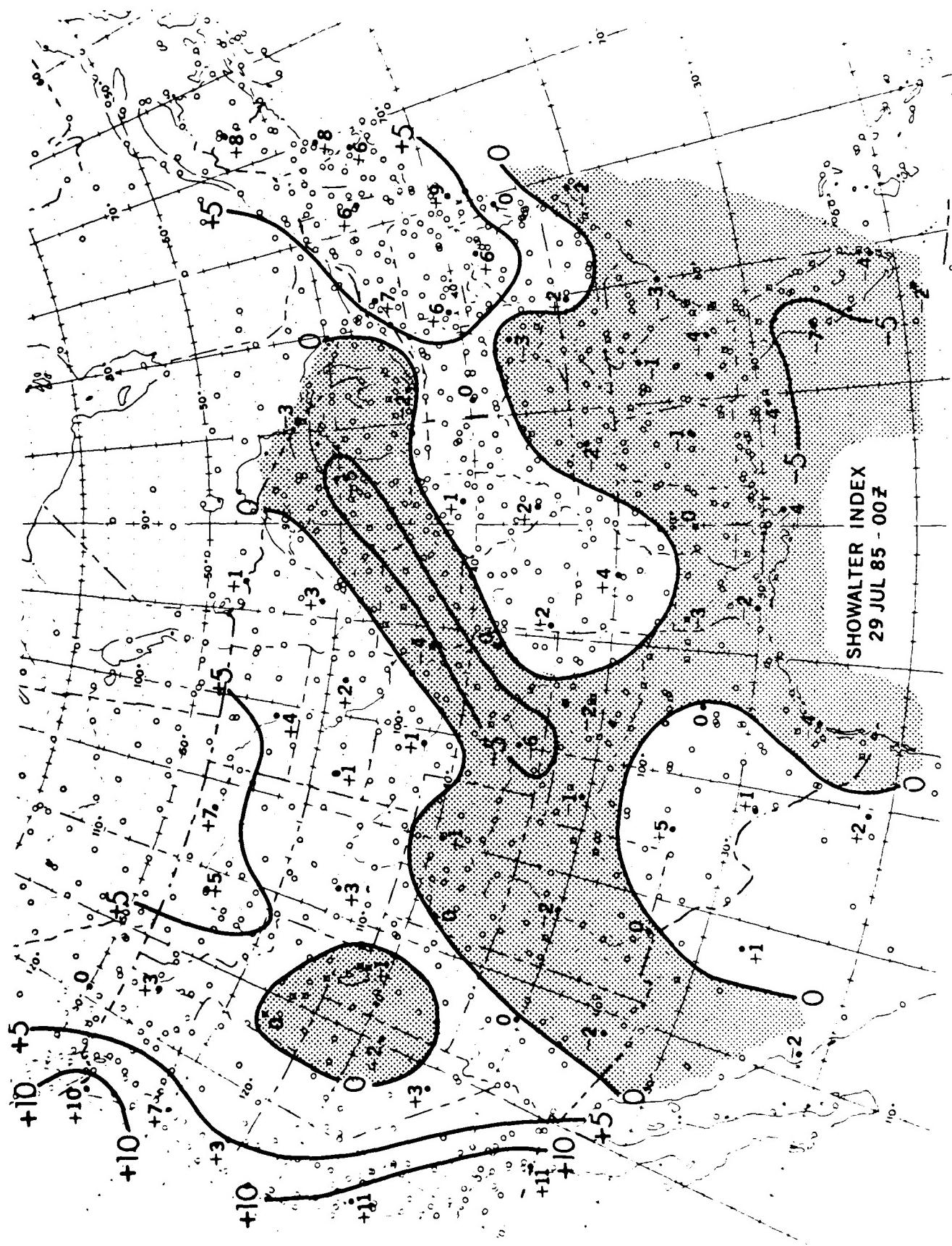
## VARIABILITY OF STABILITY INDICES BY SEASON, BY REGION

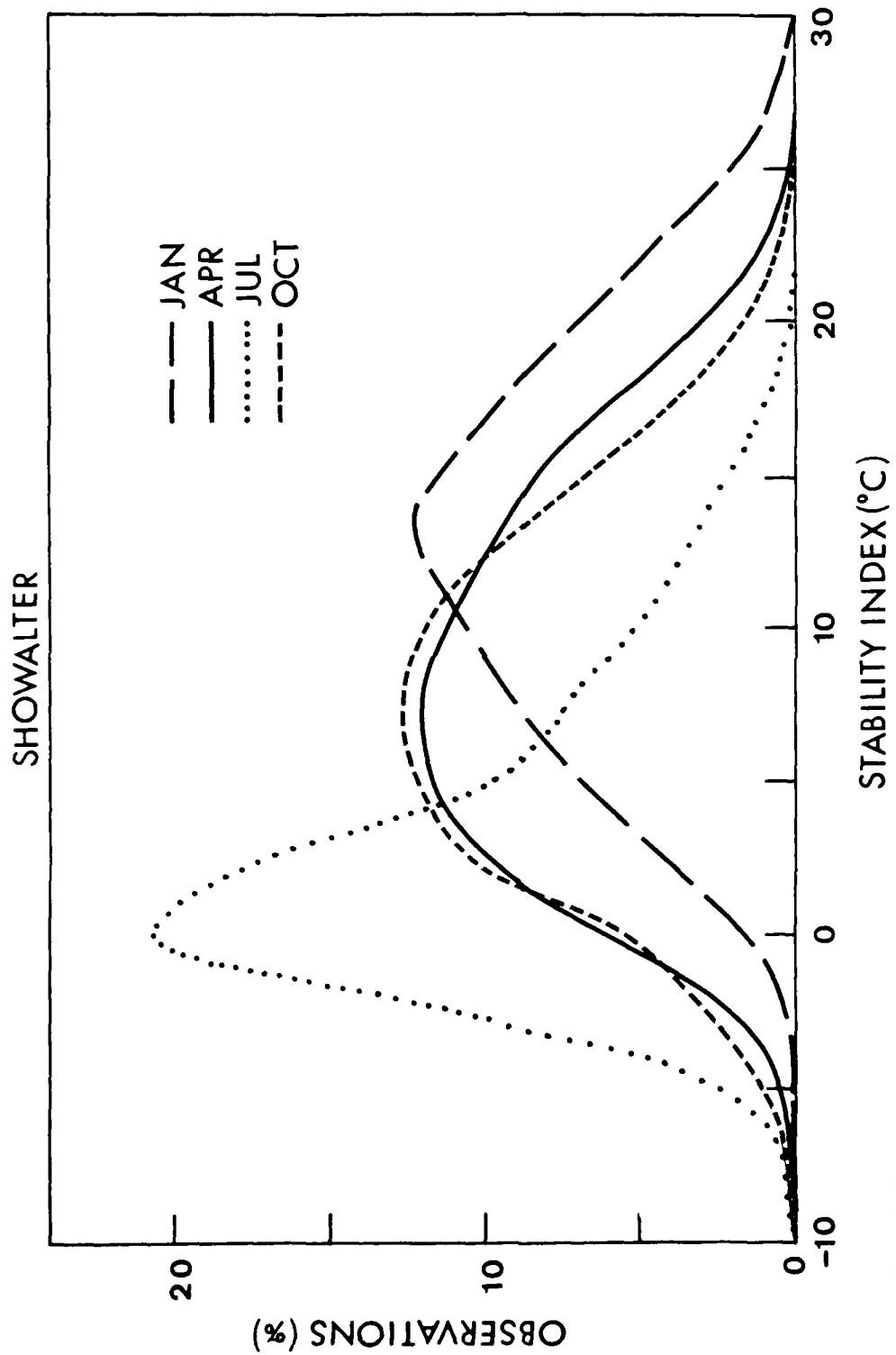
Analog of the desired parameter

Stability Indices, e.g.

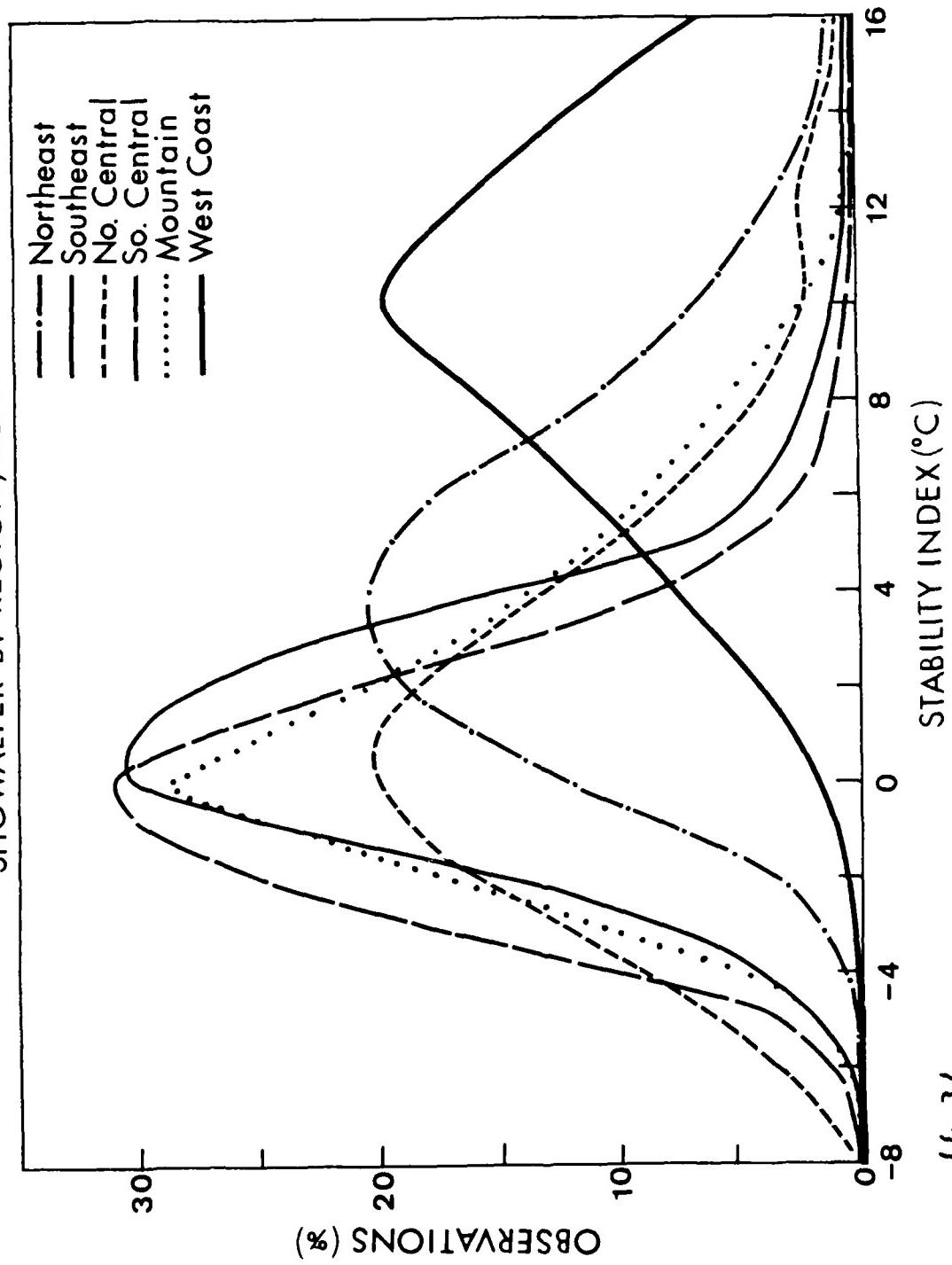
- simple: Showalter Stability Index (SSI)  
850mb temp, humidity; 500mb temp.
- complex: SWEAT Severe Weather Threat Index  
mixture of temp, moisture, wind shear values  
at various levels

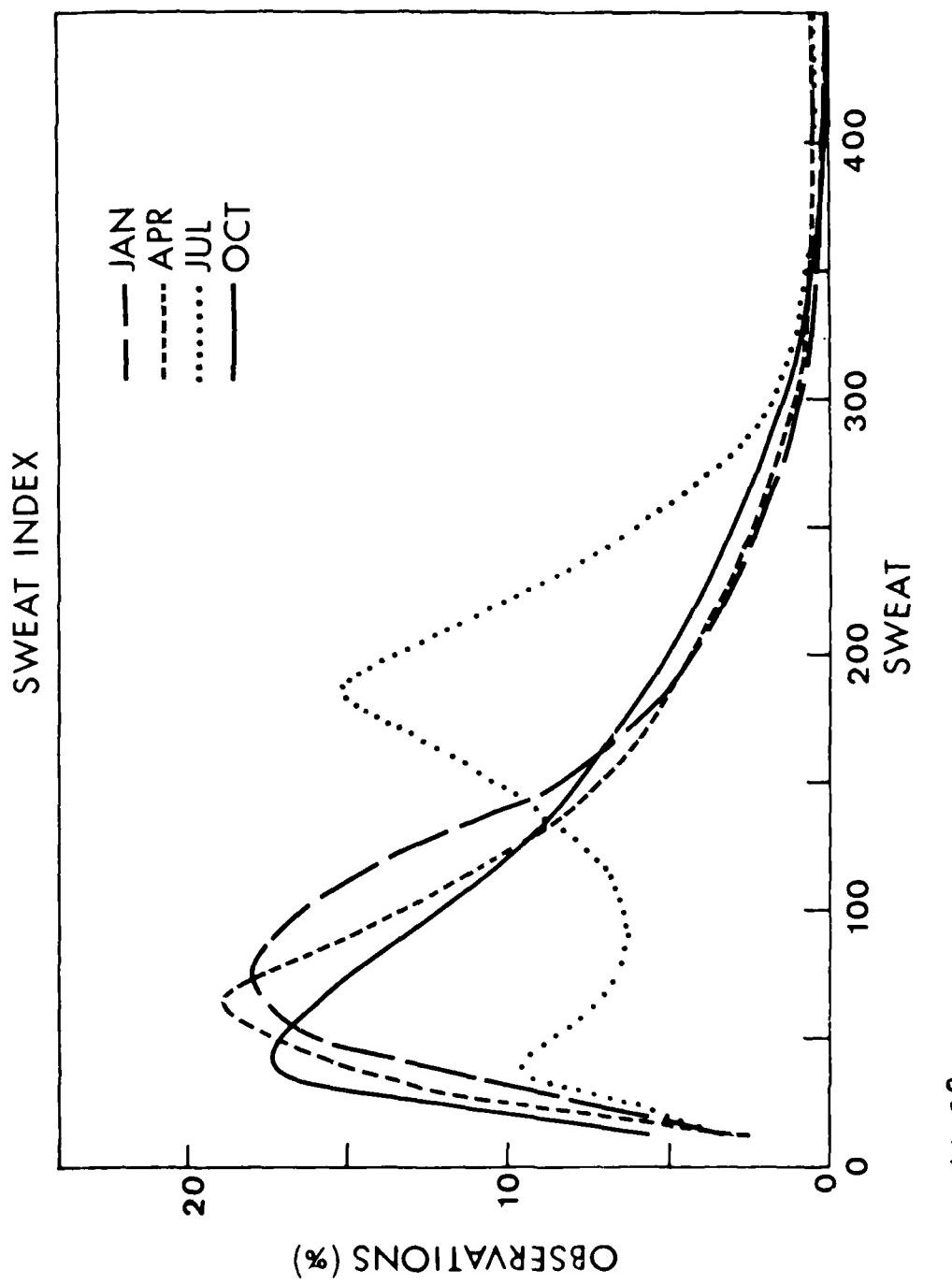
Both compiled by U.S. Air Force's Environmental  
Technical Applications Center (ETAC)  
12-year period ('73-'84)  
~ 8,000 obs per month (12-yr sample)



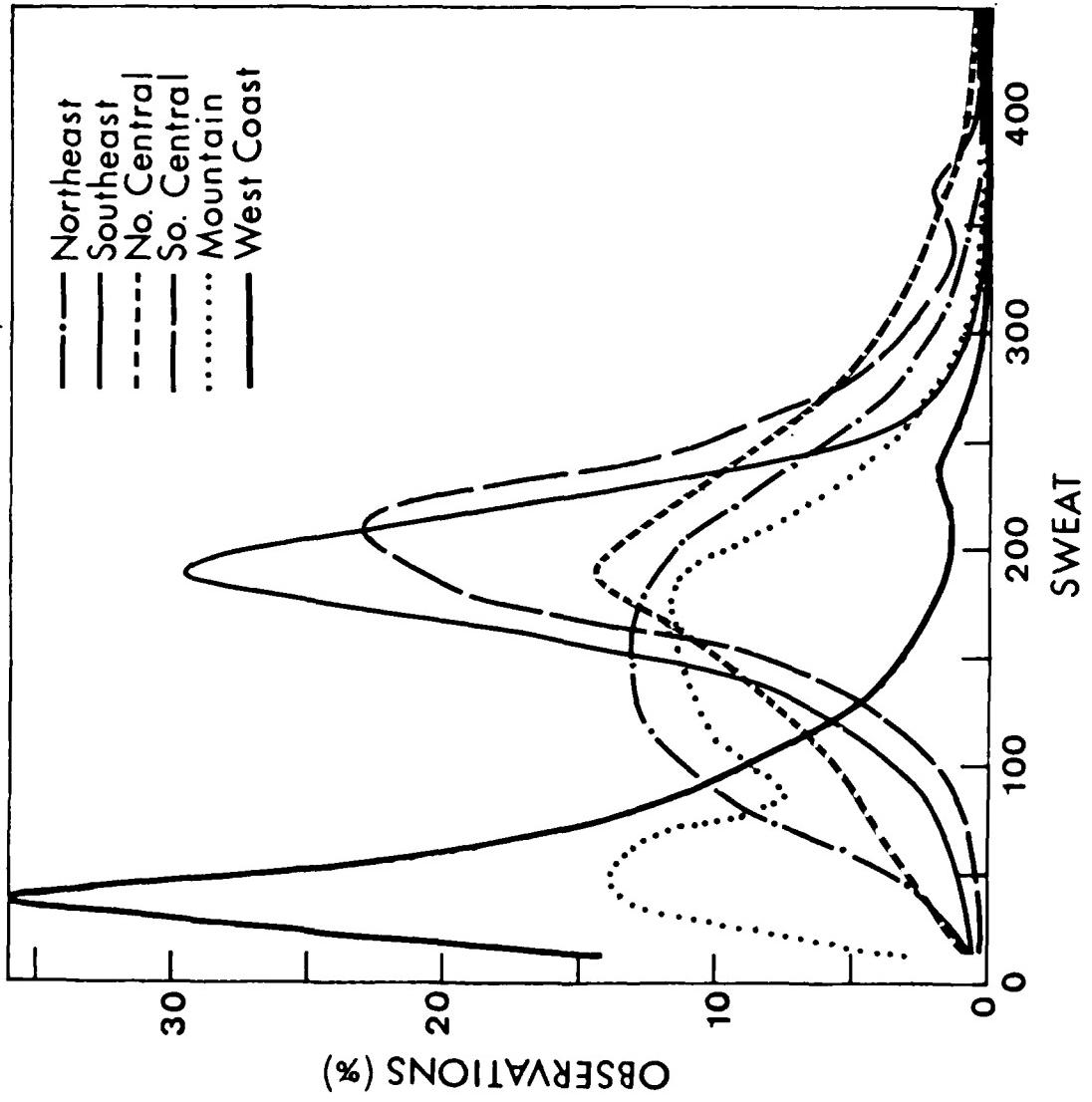


SHOWALTER BY REGION, JULY





SWEAT BY REGION, JULY



## APPROACH: CRUDE EXAMPLE (using SSI)

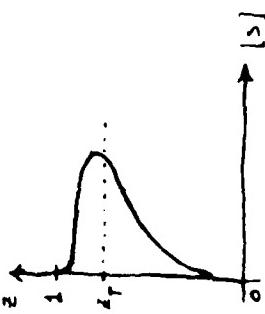
Assume vertical smoke distribution can be described by

$$p(z) = 30(z^4 - z^5) \quad 0 \leq z \leq 1 \quad \sim B(5, 2)$$

cumulative distribution

$$S(z) = \begin{cases} 0 & z \leq 0 \\ 6z^5 - 5z^6 & 0 < z < 1 \\ 1 & z \geq 1 \end{cases}$$

1.



If  $z_T$  is taken to be the tropopause height, then  
 $S(z_T)$  represents the amount of smoke in troposphere  
 $1 - S(z_T)$  " amount of smoke in stratosphere

## DISTRIBUTION APPROACH

Where does tropopause fall in the distribution of smoke?

②

$$\text{let } z_T = z_T (\text{SSI})$$

$$\text{e.g. } z_T = 0.9 + \text{SSI}/50$$

Then:

SSI	$z_T$	$S_{\bar{z}}(z_T)$
+5	1.00	1.00
0	0.9	0.89
-5	0.8	0.66

Now use observed SSI distributions to estimate stratospheric injection ~~of~~ of smoke.

## RESULTS: STRATOSPHERIC SMOKE INJECTION

	$\frac{T_g \text{ (nrc)}}{\text{Optical depth } \frac{1}{2} \text{ hemisphere}}$
JULY	7.5
JAN	0.6

## STABILITY DISTRIBUTION APPROACH

- ① Vertical soot profile - determine here - assumed a B detrainment-like profile
- ② Relate stability (etc.) parameter to level of tropopause in this distribution  
here - simple linear  $z_T = 0.9 + \frac{SSI}{60}$
- ③a Climatological: use observed stability-parameter distributions to determine amount of soot at levels of interest  
e.g. above tropopause
- ③b Model input (larger-scale): use model-derived values to calculate stability (etc.) parameter, then obtain geometric altitudes by multiplying distribution heights by a tropopause height /  $z_T$

## SUMMARY

- Vertical soot profile, max. heights related to fire, atmosphere properties
- Should be able to determine a parameter which takes all those effects into account (?)
- How would such a parameter behave  
Clues from stability indices - by season, region
- How could such a parameter be used to estimate vertical smoke profile info?  
Crude example using SSI,  
assuming B dist'n of smoke in vertical  
Estimated 7.5 % of smoke would enter stratosphere  
in July case. Corresponds to optical depth of 0.5  
using NRC baseline data.

## WHY BOTHER?

Framework, guidance in interpreting cloud observations and cloud model results.

Use model, obs:

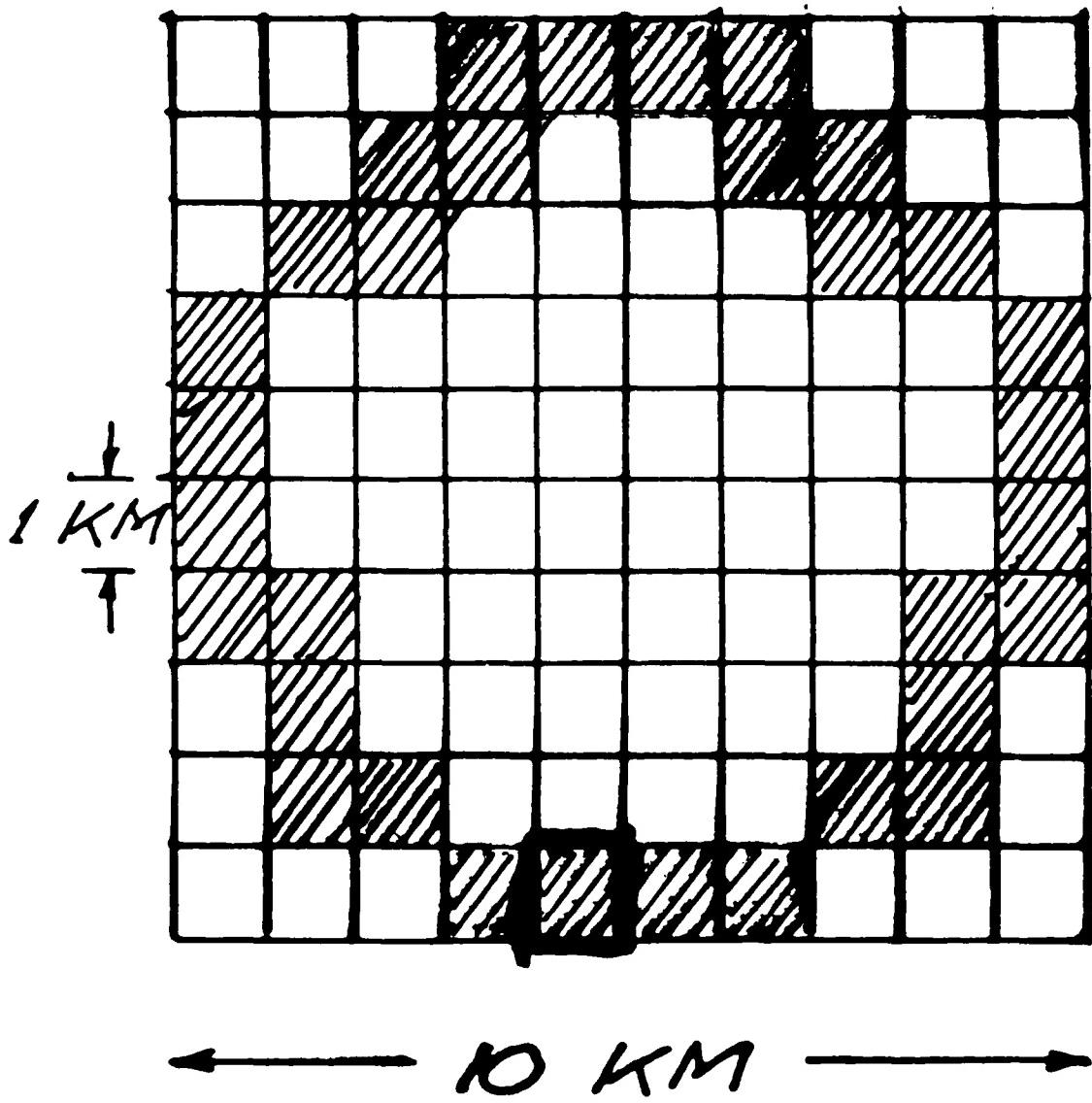
- to find a suitable parameter
- to find an appropriate vertical soot profile
- to relate the two
- to evaluate the scheme.

Then use albedo data to determine distributions of the parameter.

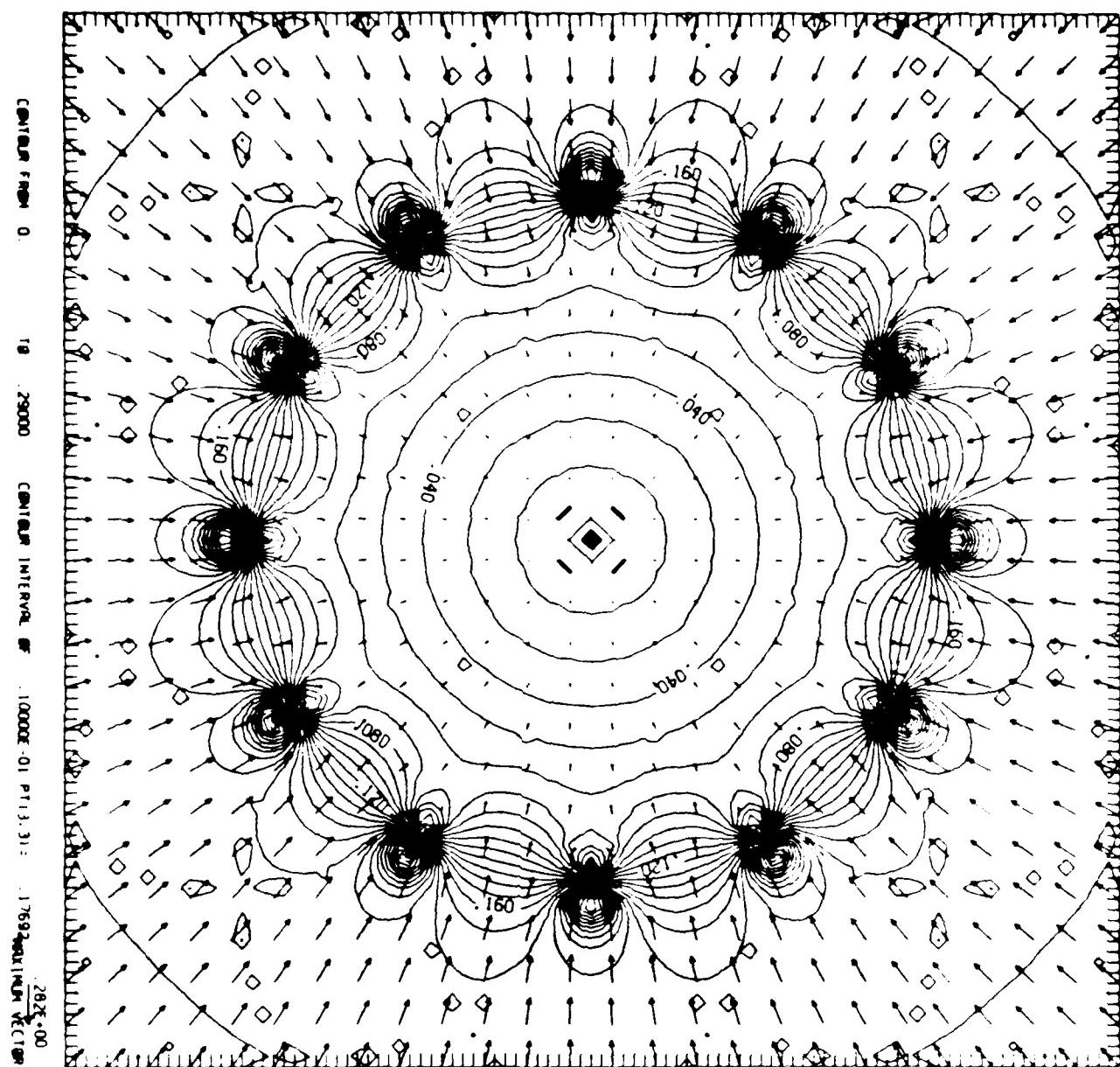
# MULTIPLE PLUME FLOW FIELDS

HOWARD R. BAUM  
NATIONAL BUREAU  
OF  
STANDARDS  
GAIITHERSBURG, MD. 20899

# MASS FIRE BURN GEOMETRY



CROSS HATCH INDICATES  
BURNING AREA AT 1 KM  
RESOLUTION



## VELOCITY DECOMPOSITION

$$\vec{U} = \nabla \phi + \vec{V}$$

$$\phi = \sum_i \phi_i$$

$$\vec{V} = \sum_i \vec{V}_i$$

$$\nabla^2 \phi_i = \frac{\rho}{\rho_0} Q_i$$

$$\nabla \cdot \vec{V}_i = 0$$

$$\nabla \times \vec{V}_i = \vec{\omega}_i$$

NOTE THAT ALTHOUGH EACH COMPONENT OF THE FLOW IS AXIALLY SYMMETRIC WITH RESPECT TO A LOCAL ORIGIN, THE NET RESULTANT FLOW IS THREE DIMENSIONAL.

## ROLE OF KINEMATICS

CONSERVATION OF MASS & ENERGY  
TOGETHER WITH PERFECT GAS LAW  
YIELD

$$\nabla \cdot \vec{U} = \frac{\gamma - 1}{\gamma P_0} \sum_i Q_i$$

$Q_i$  = LOCAL HEAT RELEASE RATE  
IN "i" TH PLUME

$P_0$  = AMBIENT PRESSURE  
 $\gamma$  = SPECIFIC HEAT RATIO

THE VORTICITY FIELD  $\vec{\omega}$  IS DEFINED  
IN TERMS OF THE VELOCITY  $\vec{U}$  BY

$$\nabla \times \vec{U} = \vec{\omega}$$

GIVEN A REPRESENTATION FOR  
 $Q_i$  AND  $\vec{\omega}$ , THE VELOCITY FIELD  
IS DETERMINED UNIQUELY EVERY-  
WHERE

## REPRESENTATION OF $\vec{\omega}$

VORTICITY PRIMARILY IN PLUMES

$$\vec{\omega} = \sum_i \vec{\omega}_i$$

$\vec{\omega}_i$  = VORTICITY OF "i"TH PLUME

EACH PLUME IS AXIALLY SYMMETRIC

$$\vec{\omega}_i \cong \omega_\theta \hat{\theta}_i$$

$\hat{\theta}_i$  = UNIT VECTOR IN LOCAL AZIMUTHAL DIRECTION

$$\begin{aligned}\omega_\theta &\cong \frac{\partial u_z}{\partial r} \\ &= -\frac{2U_0(z)}{R(z)} \left(\frac{r}{R}\right) \exp\left\{-\left(\frac{r}{R}\right)^2\right\}\end{aligned}$$

$U_0(z)$  = CENTRILINE VELOCITY OF PLUME

$R(z)$  = LOCAL PLUME WIDTH

## SCALING LAWS

$$D^* = \left\{ Q_0 / \rho_0 c_p T_0 V g \right\}^{2/5}$$

$$\vec{U} = (g D^*)^{1/2} \vec{U}^*(r^*, z^*)$$

$$r^* = r/D^* \quad ; \quad z^* = z/D^*$$

$$\omega_\phi = (g/D^*)^{1/2} \omega^*(r^*, z^*)$$

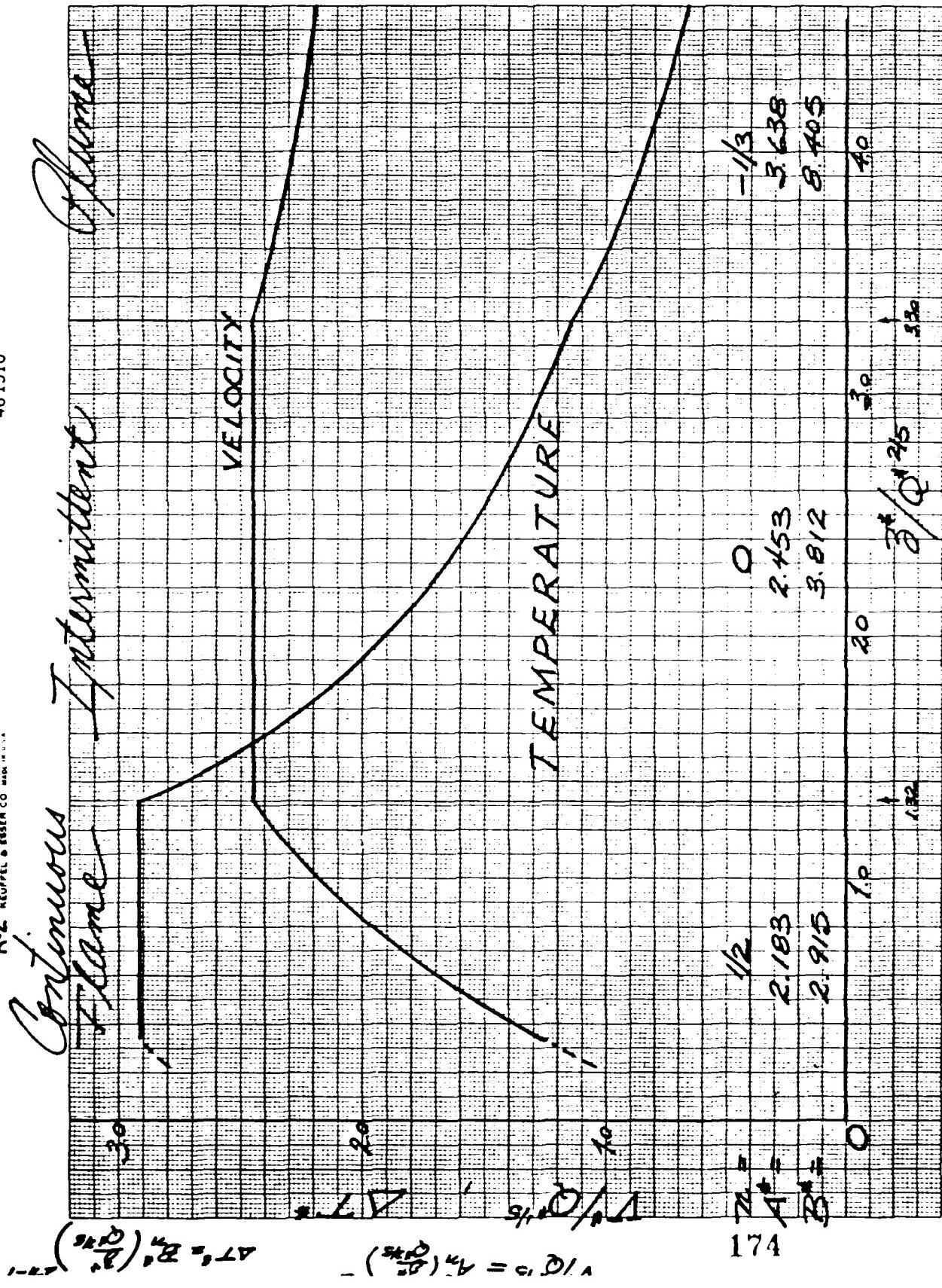
$$\phi = (g D^*)^{1/2} D^* \bar{\phi}^*(r^*, z^*)$$

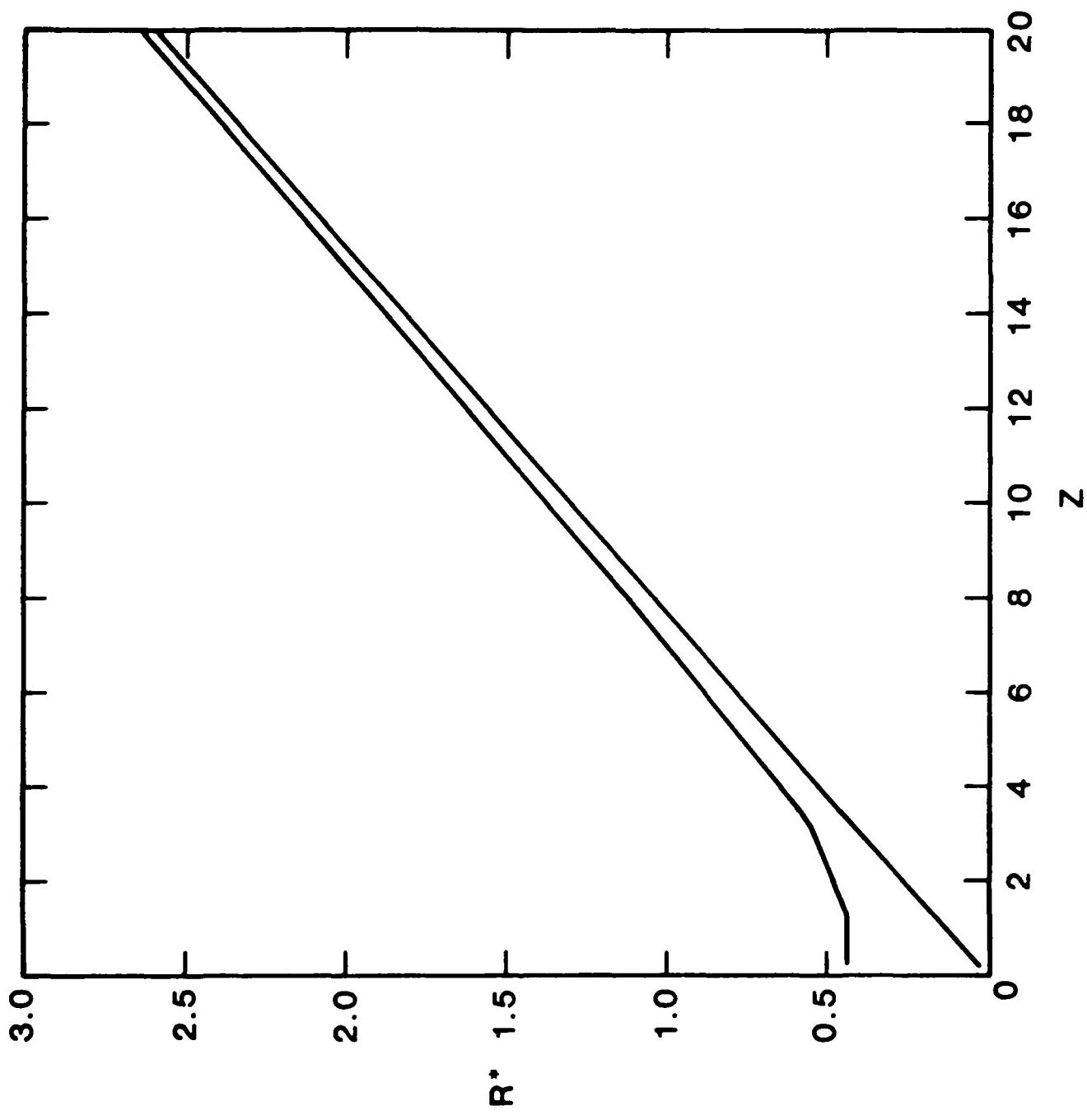
$$\psi = (g D^*)^{1/2} D^* \bar{\psi}^*(r^*, z^*)$$

THIS SCALING LEADS TO A  
 "UNIVERSAL" FLOW FIELD WHICH  
 DEPENDS ONLY ON THE RADIATED  
 FRACTION OF THE CHEMICAL HEAT  
 RELEASE

THE CENTIMETER IN THE METRIC SYSTEM

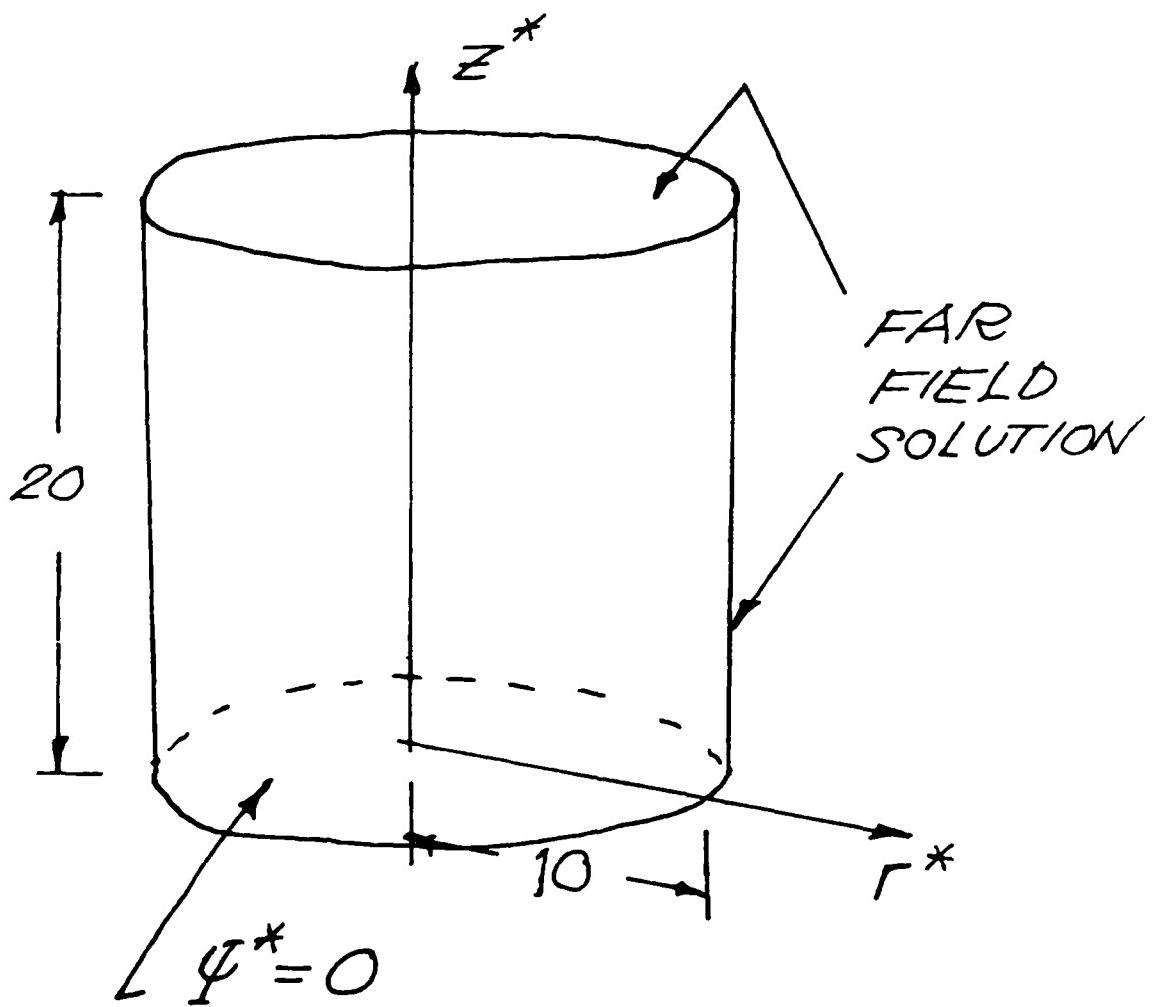
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## VECTOR POTENTIAL EQ.

$$\frac{\partial^2 \psi^*}{\partial z^*^2} + \frac{\partial^2 \psi^*}{\partial r^*^2} - \frac{1}{r^*} \frac{\partial \psi^*}{\partial r^*} = -r^* \omega^*$$



## POINT SOURCE PLUME

$$\psi^* = \rho^{5/3} F(\mu)$$

$$\omega^* = \rho^{-4/3} \Omega(\mu)$$

$$\rho^2 = r^2 + z^2$$

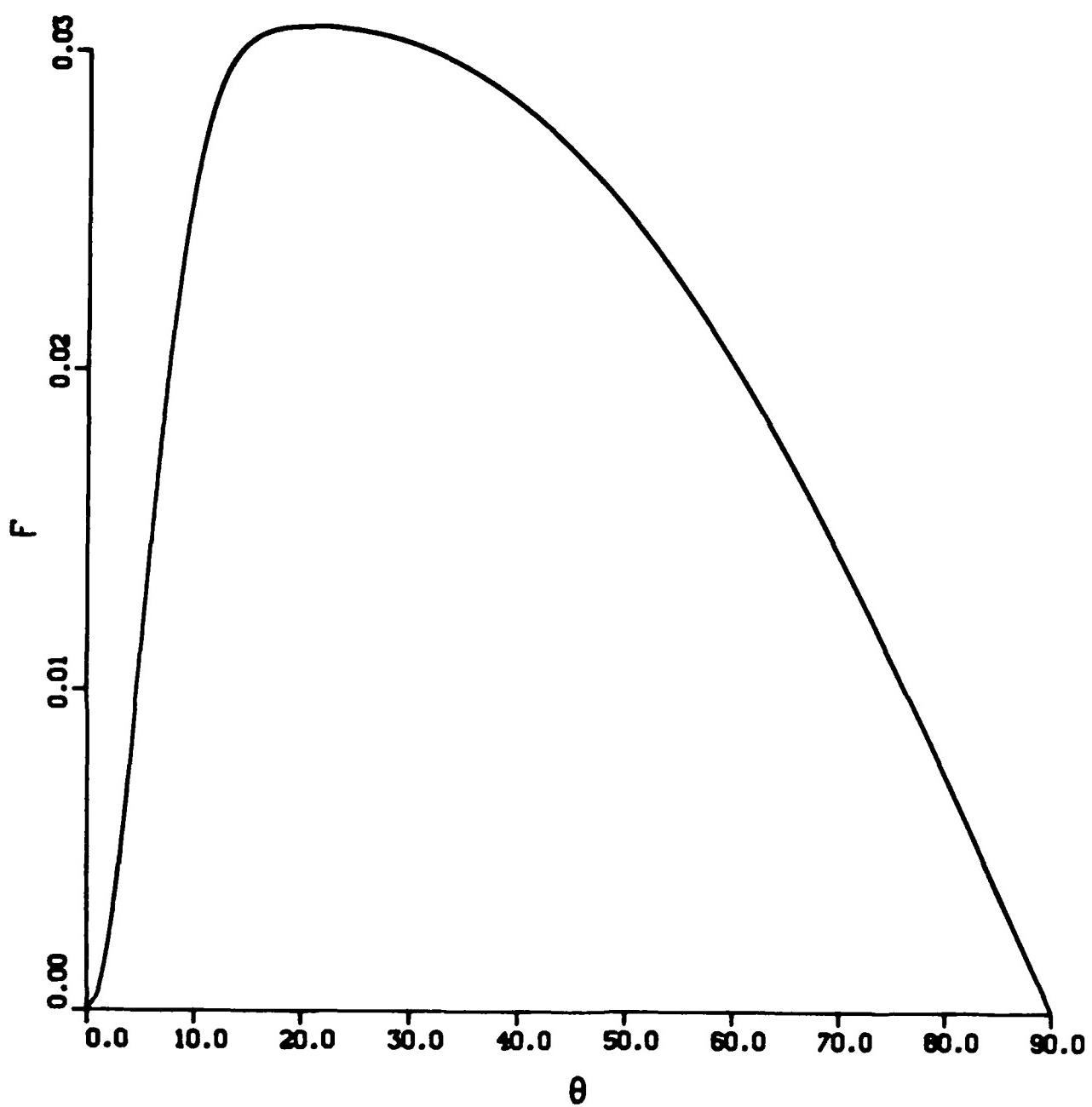
$$\mu = \cos \theta$$

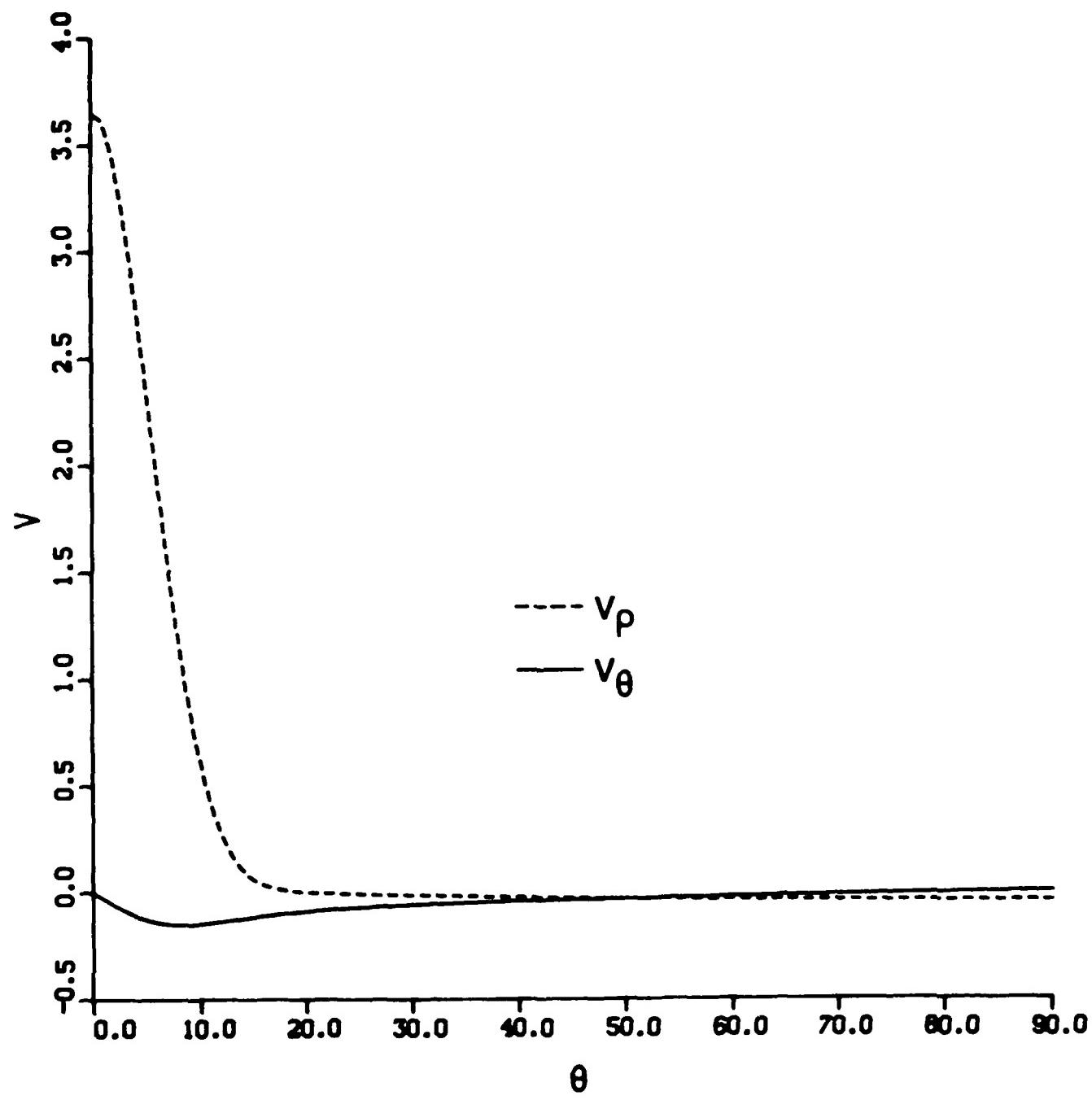
$$\frac{d^2 F}{d\mu^2} + \frac{10}{9(1-\mu^2)} F = -\frac{\Omega(\mu)}{\sqrt{1-\mu^2}}$$

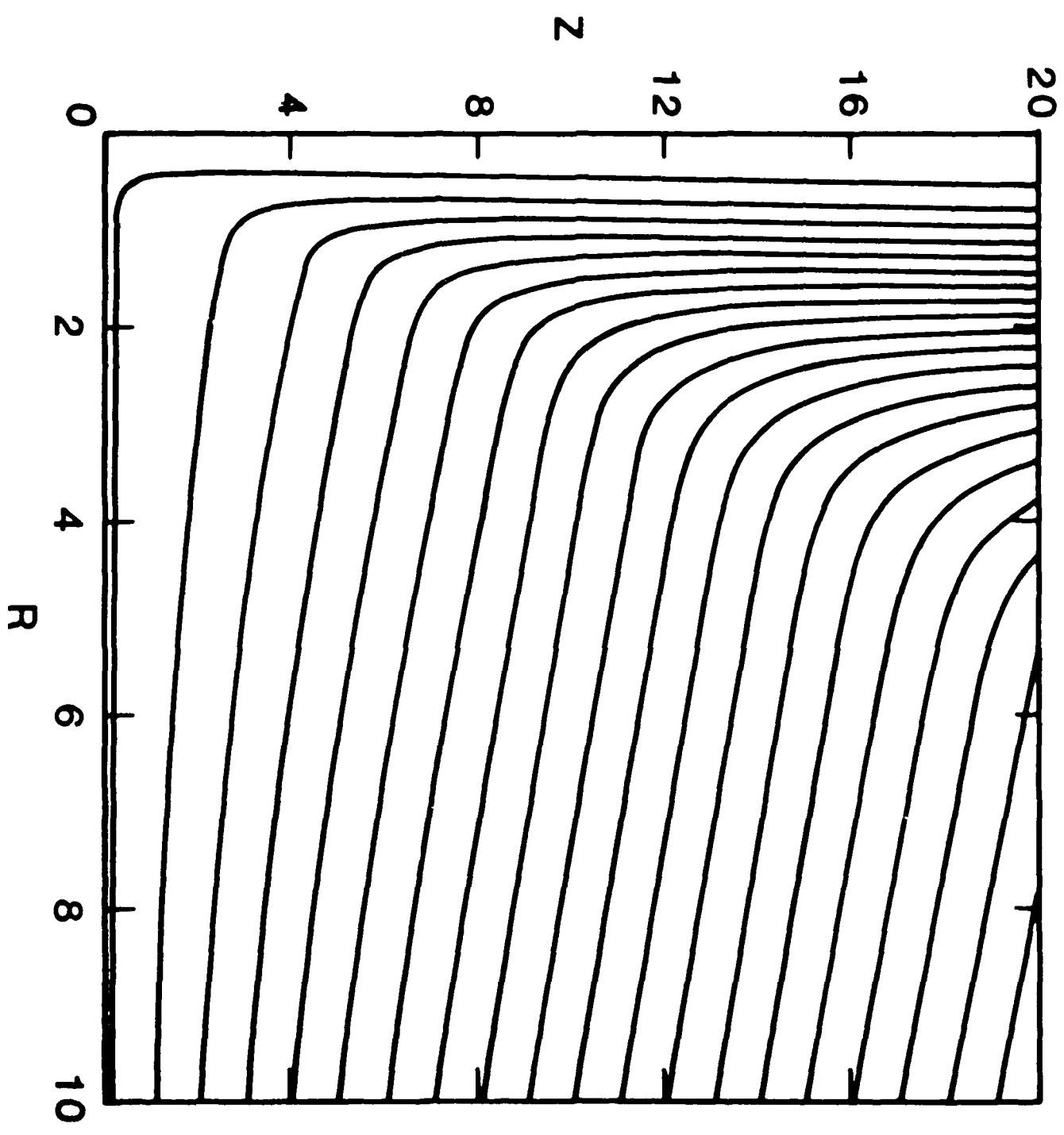
$$r = \rho \sin \theta ; z = \rho \cos \theta$$

$$F(0) = F(1) = 0$$

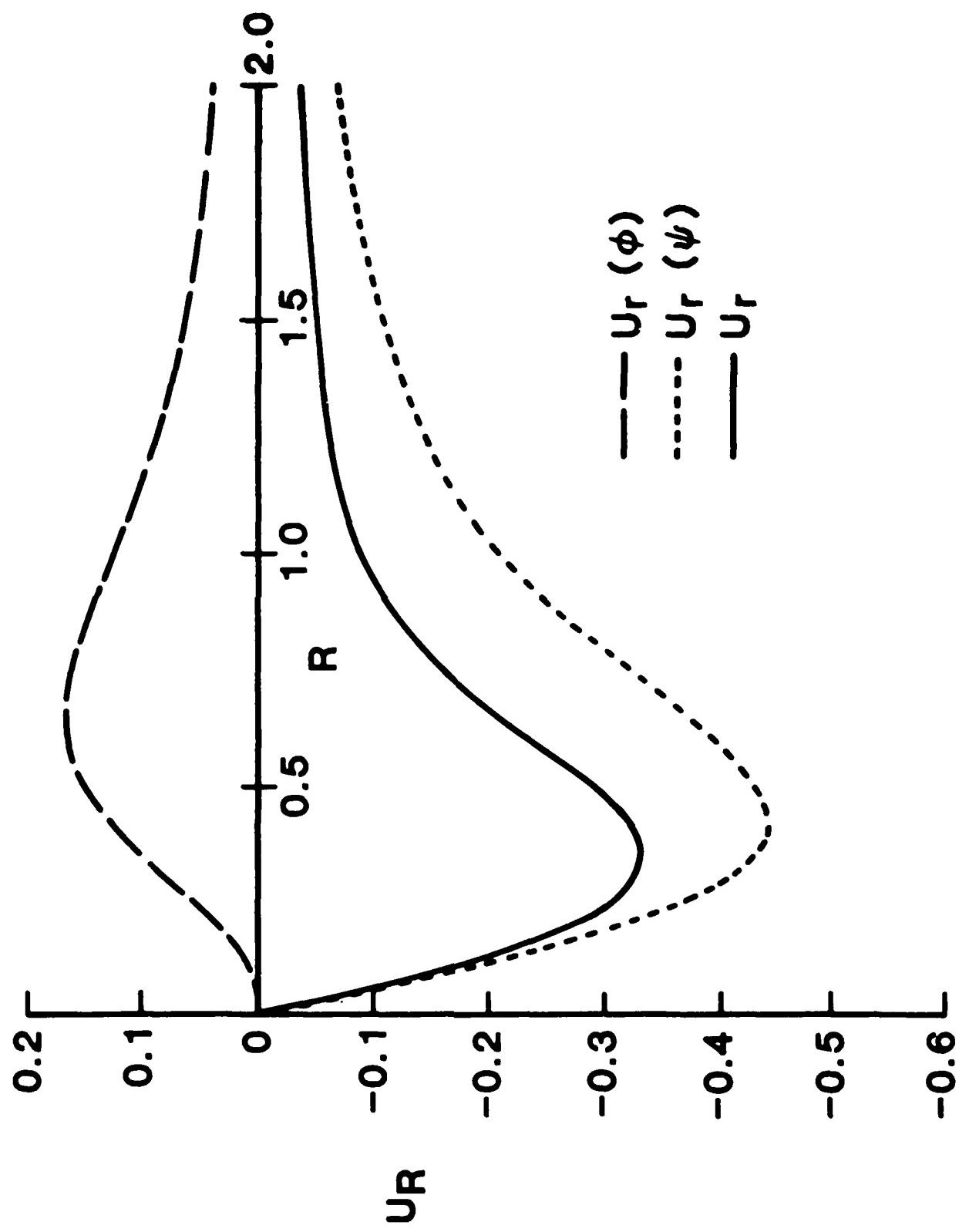
THE SOLUTION TO THIS PROBLEM  
GENERATES THE FAR FIELD FOR  
EACH PLUME

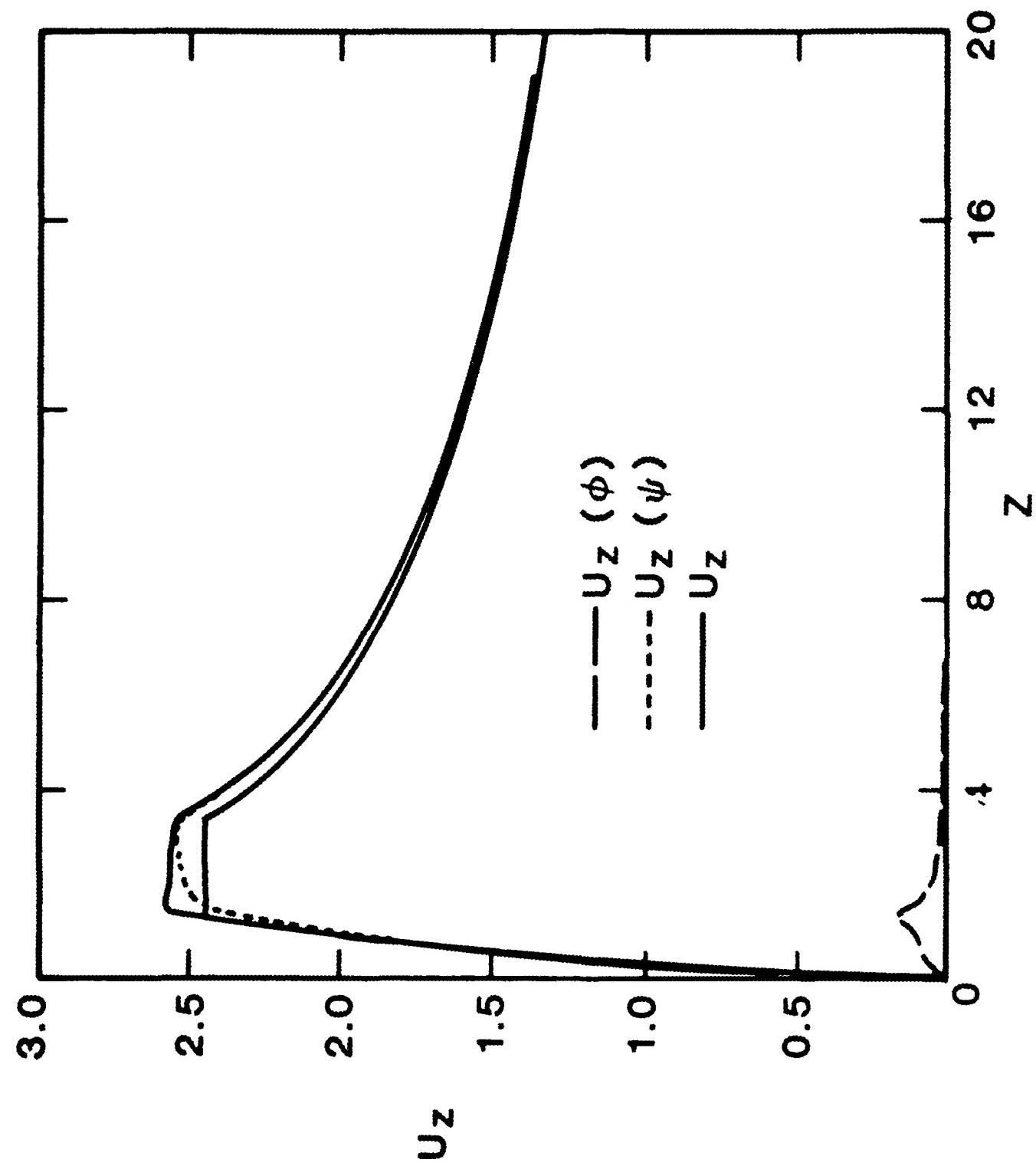


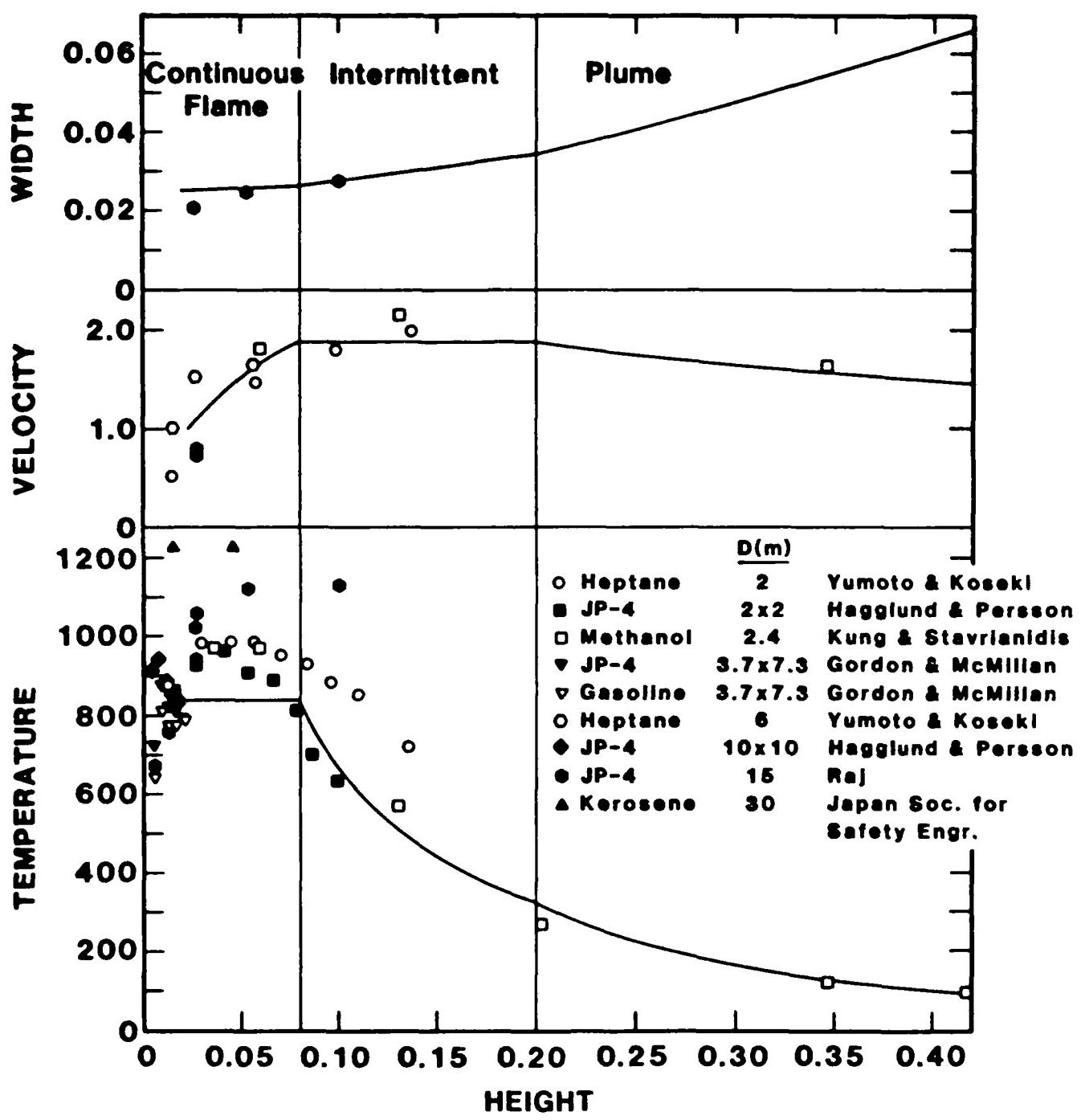


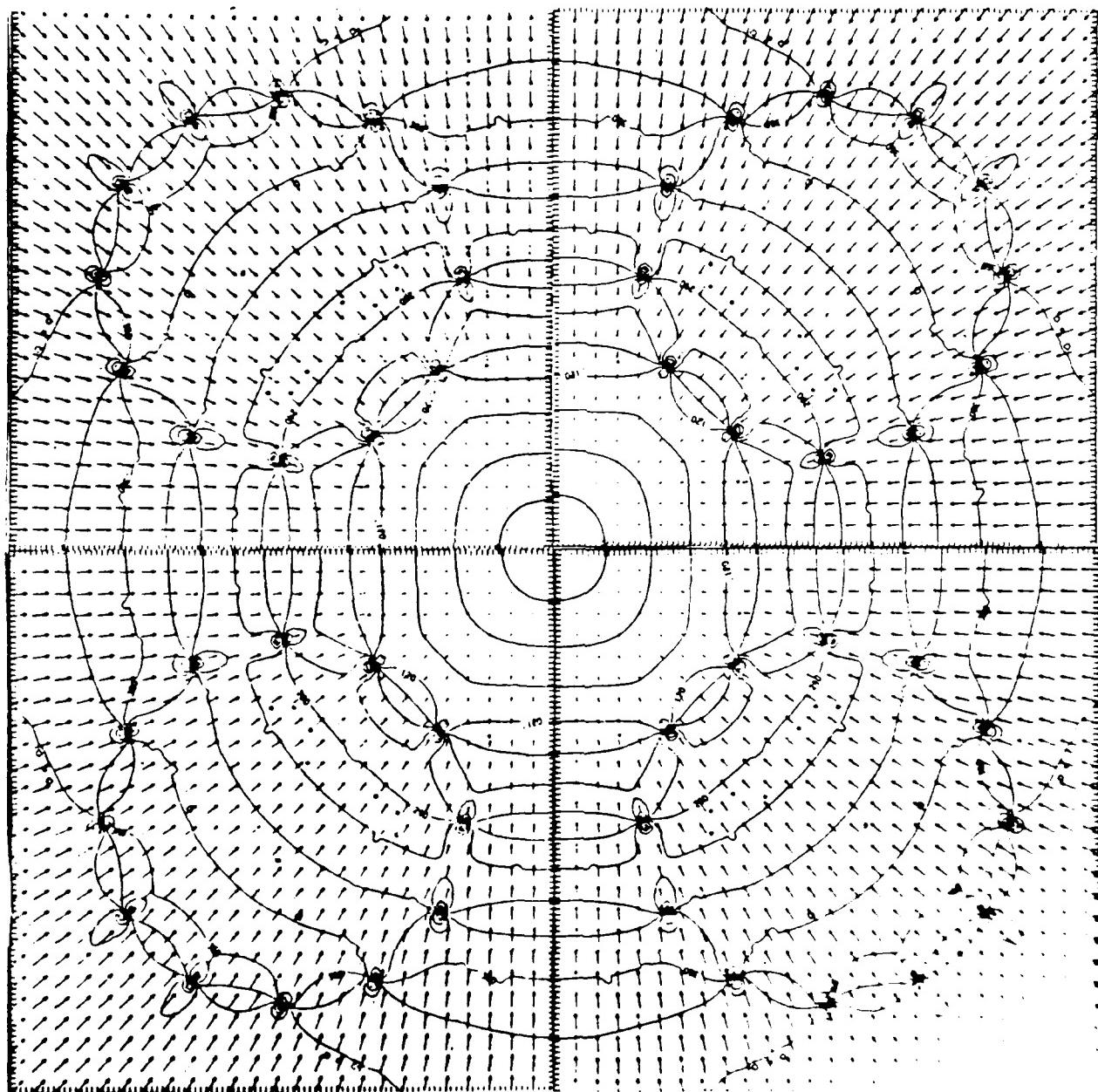


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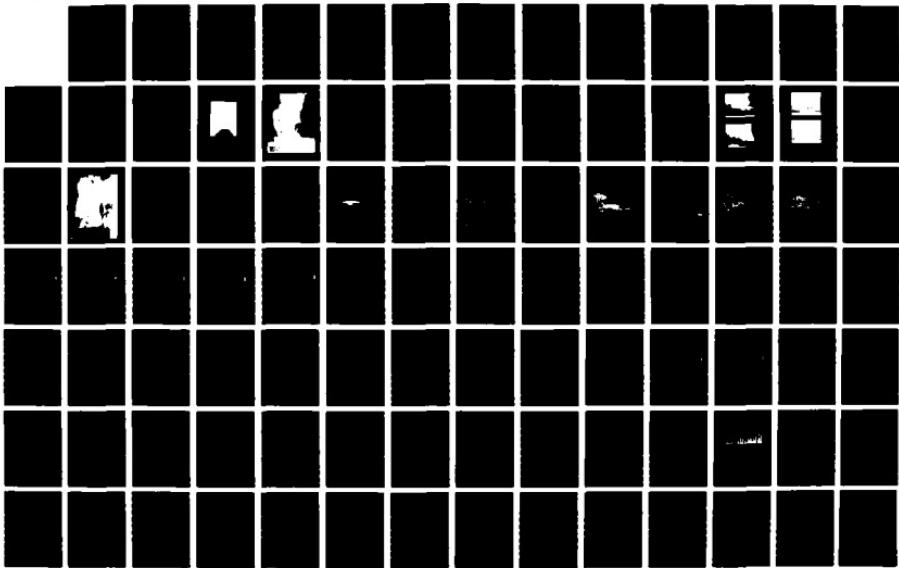
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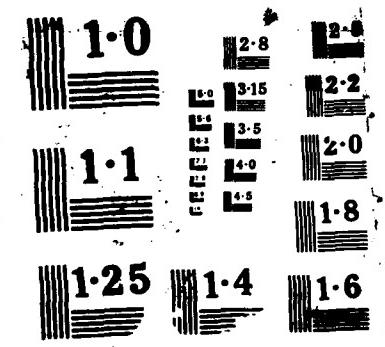
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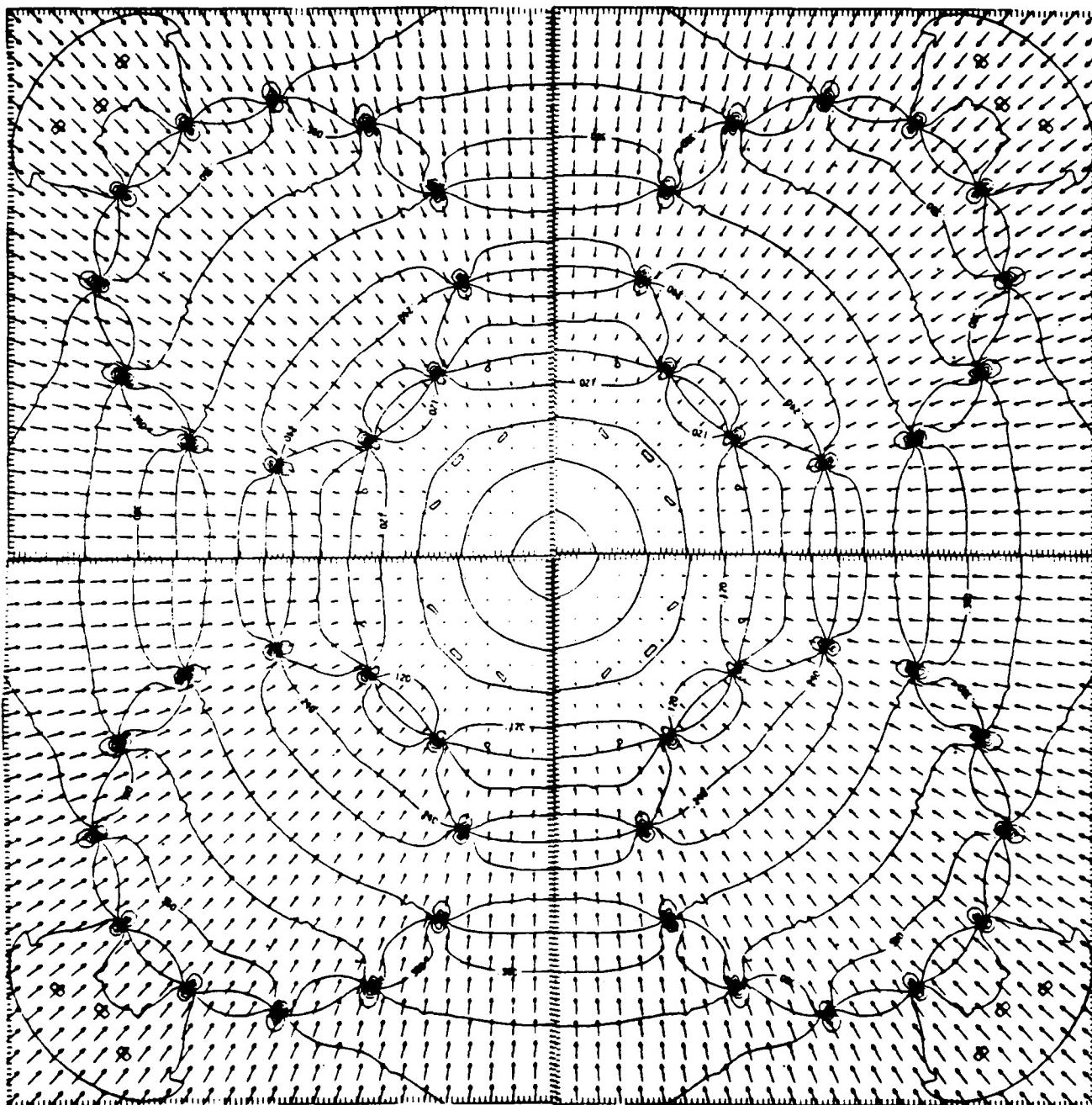
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# CAPABILITIES

- 1) VERY GENERAL FIRE PLANFORM
- 2) PERMITS MODELING FROM  
ESTABLISHED DATABASE
- 3) RANDOM AS WELL AS DETERMINISTIC  
INPUT IN SIZE AND LOCATION
- 4) HIGH SPATIAL RESOLUTION FOR  
MODEST COMPUTER RESOURCES

**PHYSICAL SIMULATION OF  
LARGE-AREA FIRE PLUMES\***

by

M. Poreh, J. E. Cermak and J. A. Peterka

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CSU Project 2-95660

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March 1986

CEP85-86MP-JEC-JAP19

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# PHYSICAL SIMULATION OF LARGE-AREA FIRE PLUMES

by

M. Poreh, J. E. Cermak and J. A. Peterka

## ABSTRACT

The dynamics of large-area fire plumes has been investigated by physical simulation using small-scale models. The models use heat or helium to generate the buoyancy flux. They neglect the radiative heat transfer and assume a relatively shallow combustion layer, an assumption which is found to be consistent with the measurements in the model, but maintain the correct dimensionless parameters which determine the dynamics of the plume outside the combustion layer.

Four regions of distinct flow characteristics have been identified. A combustion layer whose height is small compared to the horizontal dimension of the fire. A radial convective boundary layer, where radially-inward induced winds carry the rising plume toward the center. A central core and a convective plume above it, which rises to large heights.

Vigorous mixing with the ambient air occurs in the radial convective boundary layer, where the hot burned fuel rises in the form of thermals. Such thermals are formed for both continuous combustion layers and for large fires which are made of many distinct, close, small fires. It is found that these two cases produce similar plumes.

The simulations show that the plume above the burning area contracts considerably as it rises, and that the upper convective plume behaves as a weak plume with a virtual origin below the

surface, at  $Z/R$  of the order of -1.5. The subsequent plume rise in an upper stably-stratified atmosphere would thus be much smaller than that of a plume originating from a point source at the ground, with the same buoyancy flux.

The simulations also show that the induced radial velocities of a large-area fire in a calm neutral atmosphere are not as large as those estimated by previous numerical and analytical models. In addition, the study suggests that the kinetics of a single fire, which is surrounded by numerous other fires would not be much different from that of an individual fire in a calm atmosphere, as oxygen depletion is found to decrease with the size of the fire.

## SUMMARY OF THE PRESENTATION

### Simulation Rationale

- The simulation is based on the conventional approach to fluid modelling. All the important dimensionless parameters and boundary conditions must be the same in the model and prototype ( $\pi_m = \pi_p$ ).
- The Reynolds number in the model cannot be matched with that of the prototype. However, it is shown to be large, so that 'Reynolds number independence' is expected.
- The models neglect radiative effects.
- It is assumed that the relative thickness of the combustion layer (CL) in large-area fires is small compared to the size of the fire. Thus, the model simulates the flow above the relatively shallow CL.
- The flow above the CL is assumed to be governed by
  - σ - The buoyancy flux per unit area produced in the CL, which is proportional to the burning ratio (J/sec) [ $\sigma = L^2/T^3$ ].

R - The size (radius) of the (assumed circular) CL.

$\pi_i$  - Dimensionless boundary conditions describing the ambient wind, ambient circulation and atmospheric stratification.

The presentation focuses on Large Area Fires (LAF) in a calm, neutrally stratified atmosphere. It has been shown, however, that the physical simulation can include the effects of ambient stratification and circulation (Poreh et al., 1986).

The above assumptions imply that the velocity field  $u(\vec{x})$  and the effective gravitational field  $g'(\vec{x}) = g\Delta\rho(x)/\rho_a$  are given by

$$\frac{u(\vec{x})}{(\sigma R)^{1/3}} = f(\vec{x}/R) \quad (1)$$

and

$$\frac{g'(\vec{x})}{\sigma^{2/3} R^{-1/3}} = g(\vec{x}/R) \quad (2)$$

Buoyancy flux in the models have been generated by:

- (1) Heat flux, as suggested by R. R. Long.
- (2) Flow of lighter-than-air gas (helium) from a porous circular plate.

It is noted that the value of  $\Delta\rho/\rho_a$  of helium is approximately equal to that of burned fuel at stoichiometric proportions (BFSP). Namely, ideal burning of the fuel without mixing with surplus air. Thus, the model simulates an area source of buoyancy produced by a flux of BFSP before it mixes with additional air.

Theoretical considerations and previous experience suggest that  $\Delta\rho/\rho_a$  of the buoyant gases at the source need not be the same in the model and the prototype, unless a strong whirl is expected. This assumption is confirmed by the experimental results (Fig. 6). The value of  $\Delta\rho/\rho_a$  in one of the model ( $R = 50$  cm) was matched with the expected value in a  $R = 2$  km urban fire.

### Experimental Configuration (Figure 1)

- Two circular helium models with  $R = 27$  cm and  $R = 50$  cm were used.
- A two-dimensional area source ( $2\text{ m} \times 2\text{ m}$ ) was placed in a  $2\text{ m}$  wide test section of a wind tunnel.

### Summary of the Experimental Results

- The shape of the plume based on flow visualization and helium concentration measurements is shown in Figures 2 and 3. Considerable necking of the plume is observed, similar to the necking of a 14 acre fire in Project Flambeau (Figure 4).
- Four regions of distinct flow were identified (Figure 3):
  - (1) A relatively shallow CL.
  - (2) A radial convective boundary layer.
  - (3) A central core.
  - (4) A convective plume. The convective plume which can be described as a plume from a virtual point source at approximately  $z_v = -1.5 R$ .

The assumption of a relatively small thickness of the combustion layer is based on the flame height correlation of Thomas (1963) shown in Figure 5. The value of  $m''/(\rho_0 \sqrt{gD})$  for large urban fires is expected to be of the order of  $10^{-5}$ , so that the ratio of the flame height to the size of the fuel bed ( $L/D$ ) is expected to be, according to this correlation, of the order of  $10^{-2}$ .

The model does not have a CL. However, a 'pseudo combustion layer' (PCL) has been defined by the height where the injected helium mixes with 6 times air by mass. This is the order of the observed mixing at the edge of turbulent combustion layers. The thickness of this PCL in the model was smaller than 1 mm, as shown in Figure 6. This figure, which shows the value of the dimensionless gravity coefficient  $g'(z)$  in the two circular models also confirms the validity of Eq. 2. Near the ground; however, the value of  $g'$  may be affected by its initial value at the source.

Since  $g'$  is proportional to  $\Delta\rho/\rho_a$ , which is approximately proportional to  $\Delta T/T_a$ , the order of magnitude of  $\Delta T/T_a$  in large-area fires was estimated (Figure 7).

The conclusion that  $\Delta T/T_a$  in the fire plume decreases with the size of the fire appears surprising, but it is a direct consequence of Eq. 2. One may also conclude that the average oxygen depletion in large-area fire plumes is smaller than in small-area fires. This conclusion does not imply that large oxygen depletion cannot exist within individual fires burning within a LAF.

#### Induced Radial Velocities

The induced radially-inward velocities at the edge of the fire were found to be

$$U_R = k(\sigma R)^{1/3}$$

where  $k = 0.24 - 0.5$ , much smaller than in numerical models and in reports on large-area fires. (It should be stressed that the

above estimate is for fires in a calm, dry, neutrally stratified atmosphere.)

Figure 8 compares the shape of the radial velocity profile for a point source of buoyancy (a) and for an area source of buoyancy (b), which can be described, at  $z > R$ , as a convective plume from a virtual source at  $z_v$ . The figure demonstrates why, for the same buoyancy flux, the induced radial velocities increase when  $z_v$  is decreased. The conclusion is that area sources produce large inward-radial velocities which depend on the effective necking of the plume.

#### Estimated Plume Rise in a Stably-Stratified Atmosphere

Briggs equation or the analysis of Morton, Turner and Taylor are for a buoyant plume from a point source. The initial mixing in the radial convective boundary layer (CBL) of an area source would decrease the final plume rise.

Figure 9 shows a rough estimate of the plume rise for an area source with a given total buoyancy flux as a function of the position of the virtual origin. The estimate is based on the assumption that the behavior of the plume at  $z > 0$  may be calculated using the width of the virtual plume at  $z = 0$ . The data is for a buoyancy flux produced by a  $100 \text{ km}^2$  fire.

The average plume rise for  $z_v = 0$  is identical with the estimate given in the NCR report (1985, p. 74). One clearly sees from Figure 10 that the plume rise is reduced drastically with the lowering of the virtual origin. At  $z_v/R = -1.5$  one gets an average plume rise of the order of 5 km.

This drastic reduction is initially surprising. However, if one examines the theoretical solution of Morton et al. for the loss of buoyancy flux from a point source with height (Figure 10), which shows that most of the buoyancy flux is lost when the plume achieves a large width, the reason for the large difference in plume rise between point sources and area sources becomes clear.

It is interesting to note that while lowering of  $z_v$  decreases the plume rise, it increases the induced radial velocities.

#### The Nature of the Radial Convective Boundary Layer (CBL)

The CBL is the largest region above the combustion layer. Our study indicates that it is very similar to atmospheric CBL:

- The Monin-Obukhov length is very small  $L/R = 0(10^{-4})$ .
- The effect of the horizontal shear is small and the induced horizontal velocities do not cause mixing.
- Buoyancy flux is in the form of thermals with large upward velocities,  $w_z > u_R$ .
- To satisfy continuity fresh cool air descends to the ground between the thermals (or the individual fire plumes).
- Rapid dilution of individual fire plumes is expected.

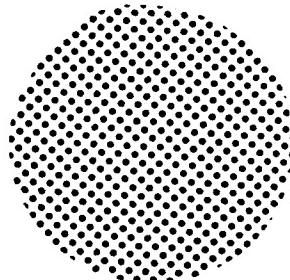
The flow visualization in Figure 11 shows the radial CBL in the two-dimensional fire model. Figure 12 shows, as well, a smoke filament introduced outside the fire.

Figure 13 shows smoke filaments introduced from outside the fire as they enter the CBL. A rapid descent of the fresh air is

observed, as in cases of fumigation. The same pattern was observed in the circular models and in Project Flambeau. Many numerical and analytical models neglect this unique form of mixing.

#### The Effect of Urban Streets

The effect was simulated by covering the porous source with perforated plates. The porosity of the plates was 12 percent and 30 percent. The gross plume shape appeared in the flow visualization to be unchanged.



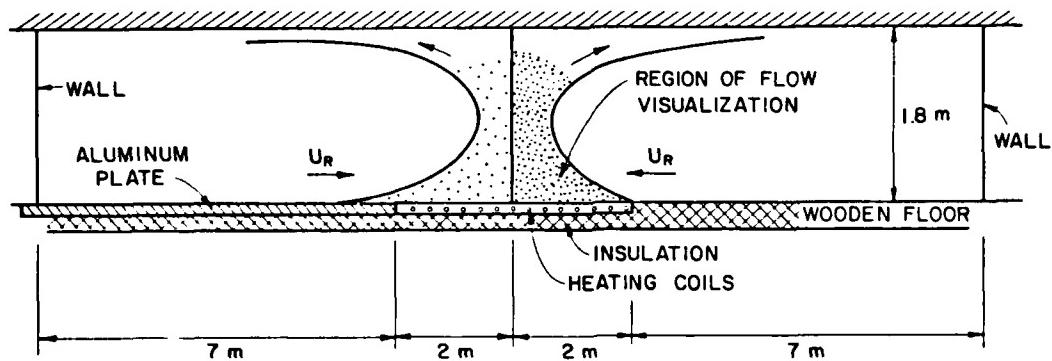
#### Conclusions

- Physical models can simulate important features of large area fire plumes.
- The rise, in a stably stratified atmosphere, of a plume with a given buoyancy flux decreases with the area of the source.
- Area sources of buoyancy induce larger radial, inward, ground-level velocities than point sources, but the magnitude of these velocities is not as high as previously reported.
- The interaction between simultaneously burning fires over a large area is not expected to be large. This conclusion is based on two observations: the average oxygen depletion in plumes of LAF is not expected to be large. The upward velocities of the individual plumes

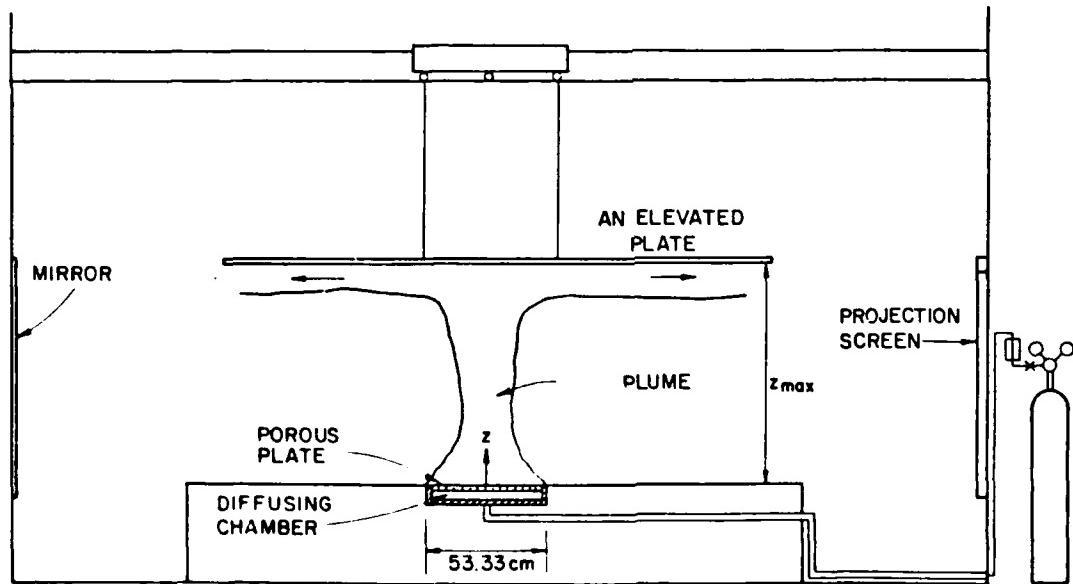
within a LAF are expected to be larger than the induced horizontal velocities.

Proposed Future Simulations

- Measure, in small-scale models, the instantaneous temperature (oxygen depletion) LAF and in an individual fire surrounded by other fires.
- Study the detailed effect of streets and noncombustible areas on the flow and temperature field of a LAF.
- Study the behavior of LAF plumes in a stably stratified atmosphere.
- Study the merger of fire plumes from close area fires.



The two-dimensional fire model in the Meteorological Wind Tunnel.



The circular fire model.

**Figure 1.** Schematic description of the experimental configurations.

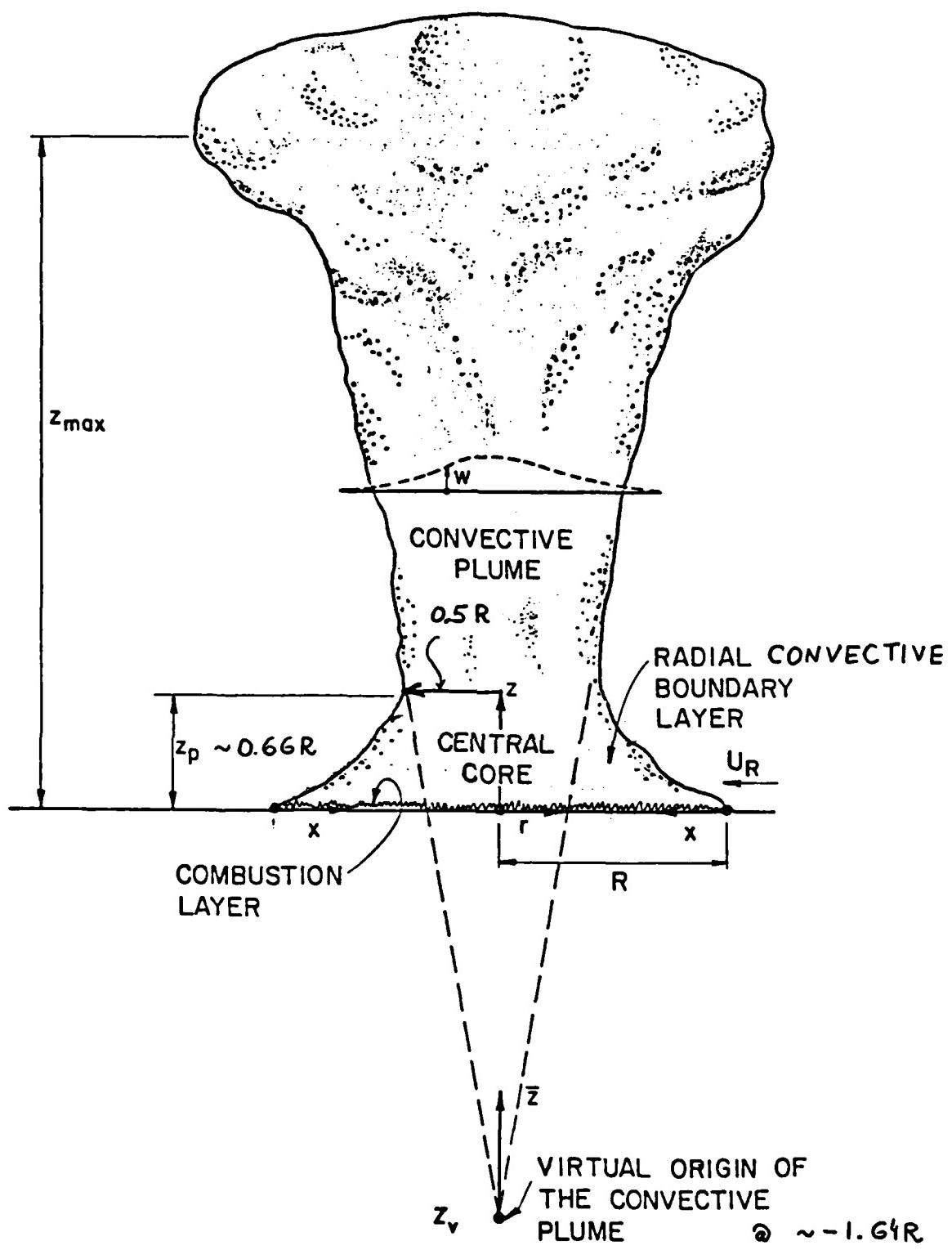
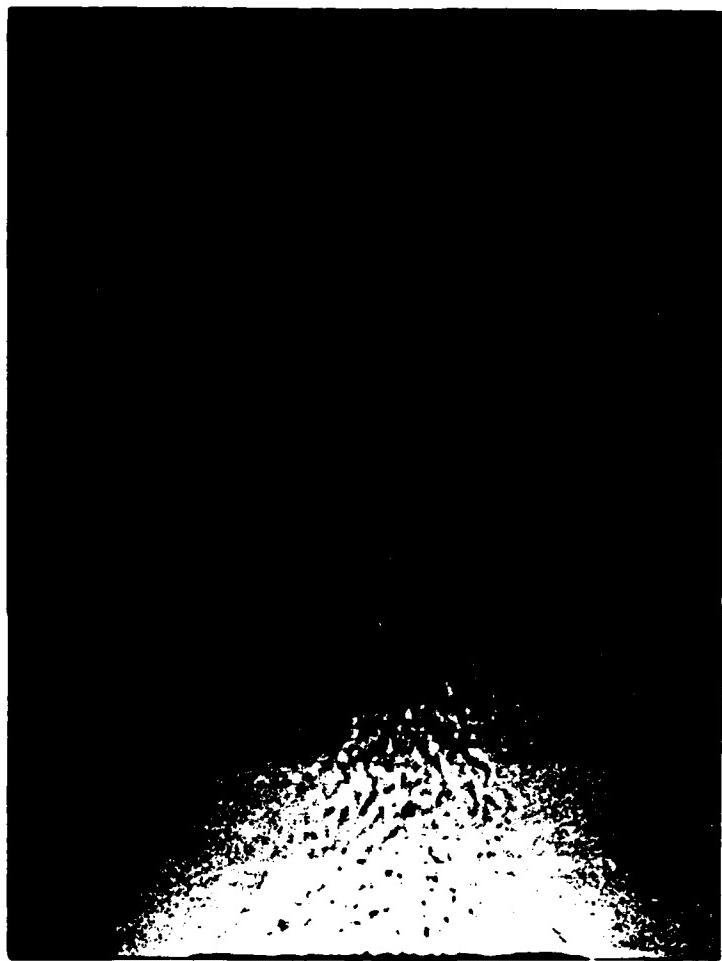


Figure 2. Schematic description of a large fire.



**Figure 3.** A photograph of the plume shadowgraph.



Figure 4. A 15-acre fire in the Flambeau Project. Trees  
in foreground are 12 ft high.

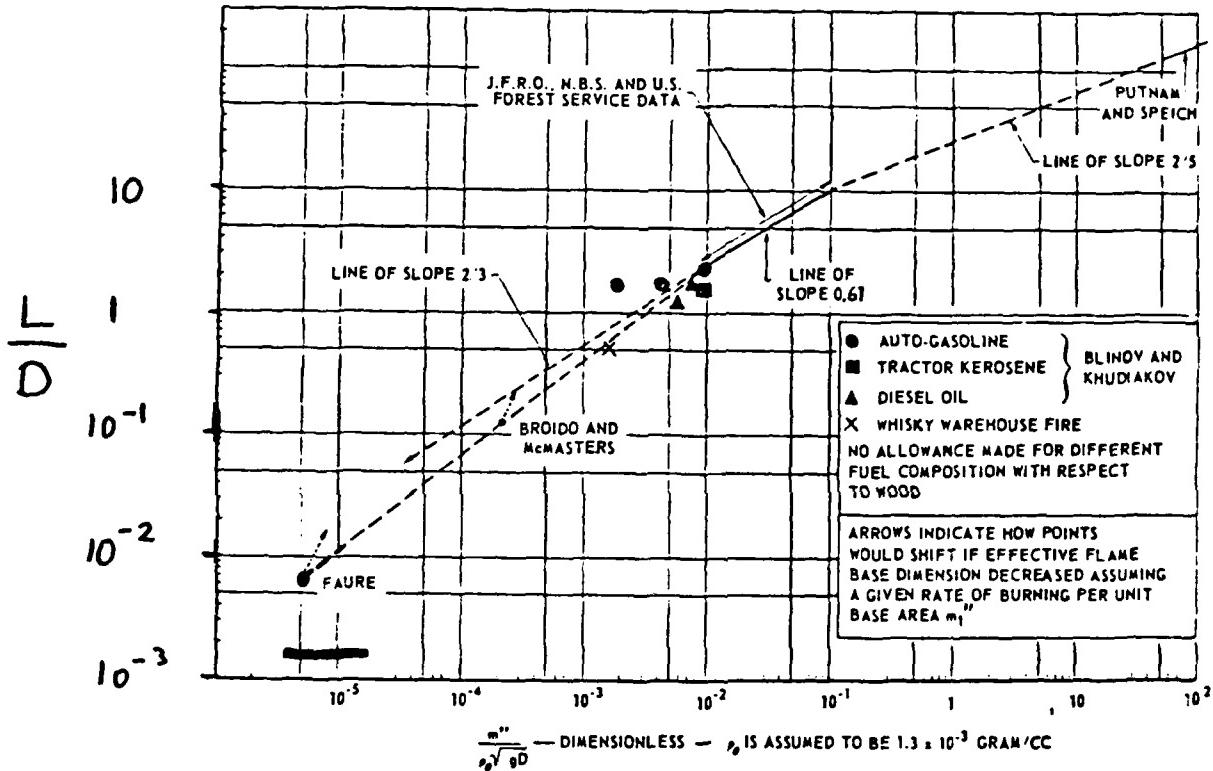


FIG. 4. Flame height correlation.

Figure 5. Flame height correlation according to P.H. Thomas (1963).

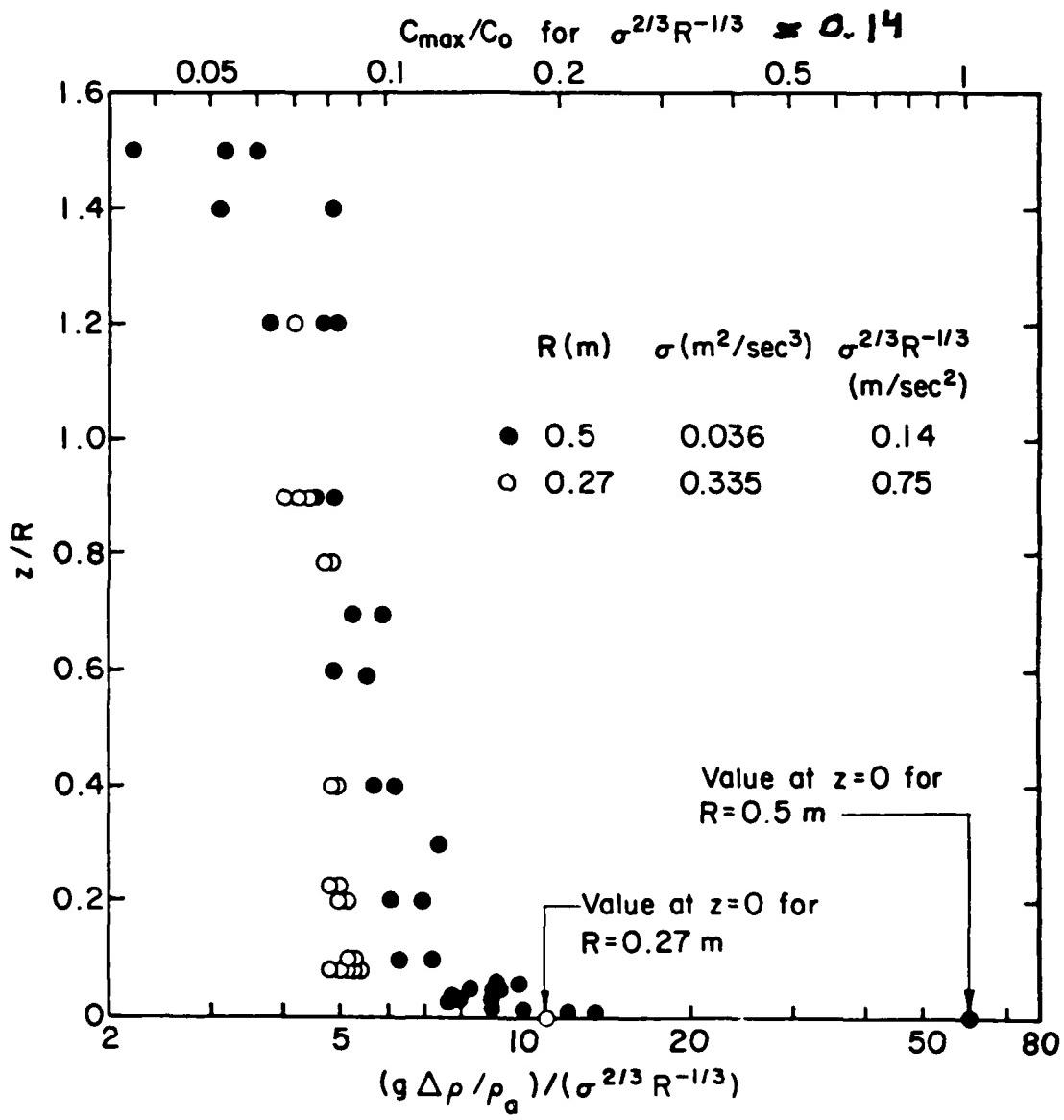


Figure 6. Measured dimensionless buoyancy parameter at the center of the plume in two physical models.

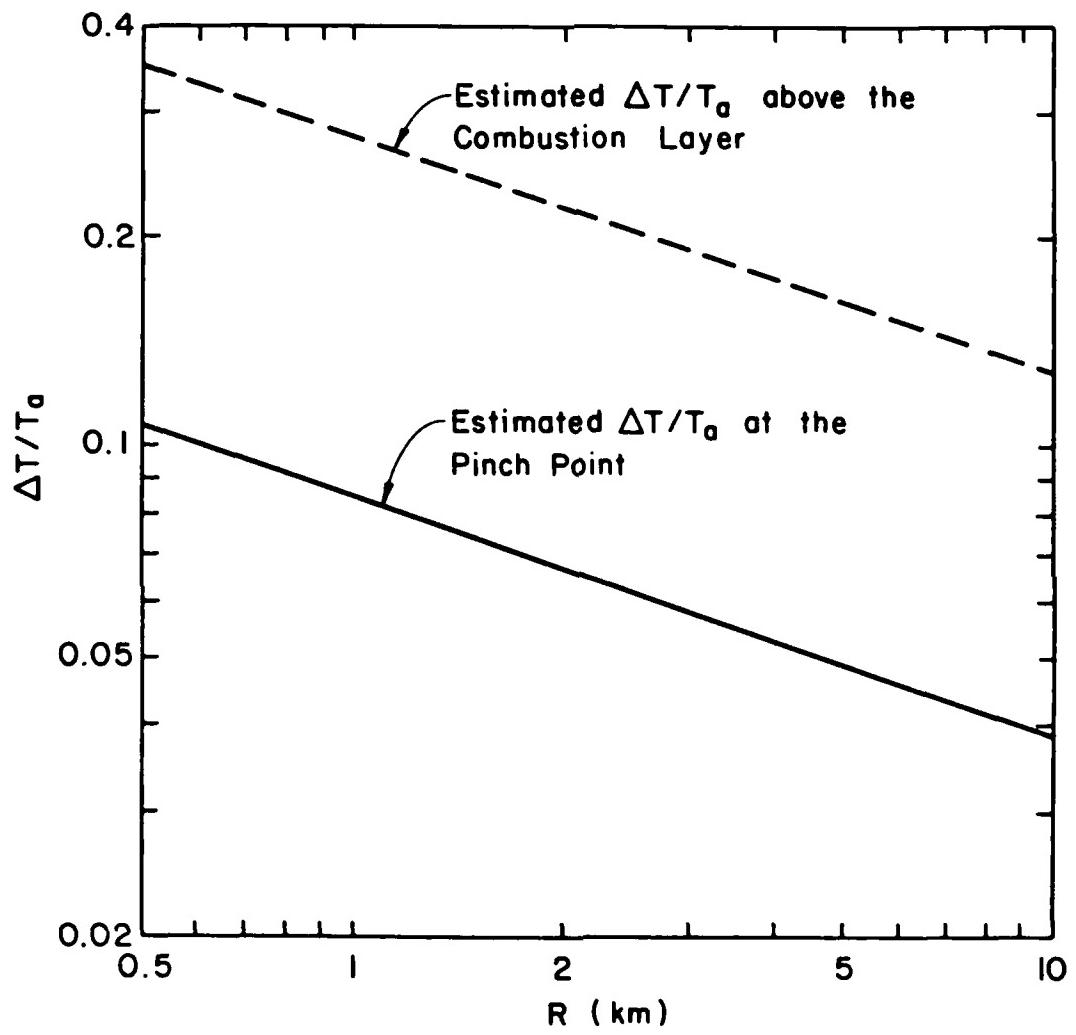
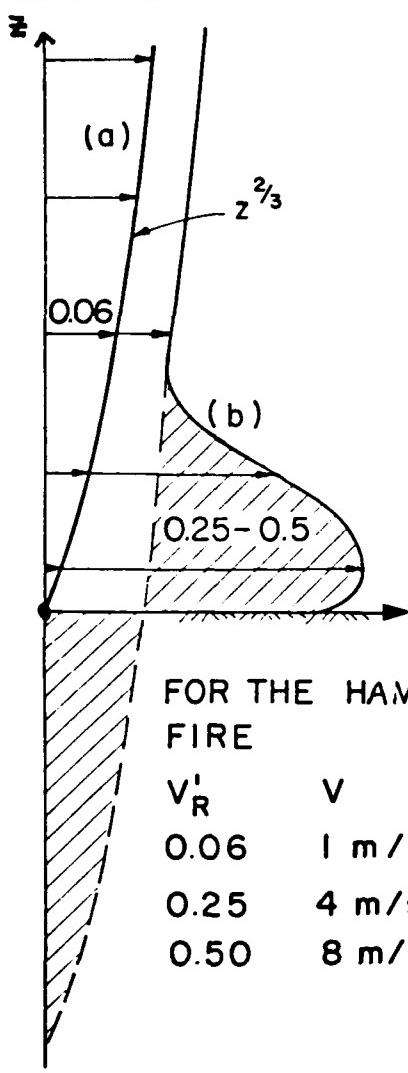


Figure 7. Estimated increase of plume temperatures for different size fires.

DIMENSIONLESS  
RADIAL VELOCITIES  
 $V'_R = V_R / (\sigma R)^{1/3}$   
 at  $r = R$



(a) POINT SOURCE  
at  $z = 0$

(b) AREA SOURCE

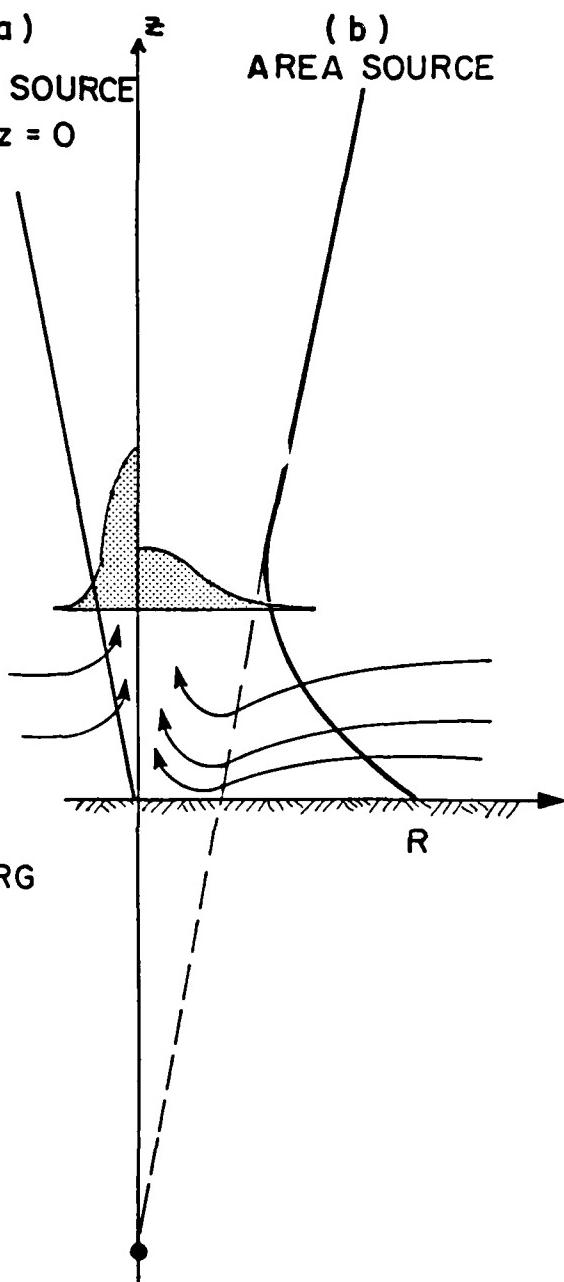


Figure 8. Schematic description of induced radial velocities for a point source (a) and an area source (b). (The correlation between the dimensionless velocities and expected velocities for the Hamburg Fire are shown in the figure).

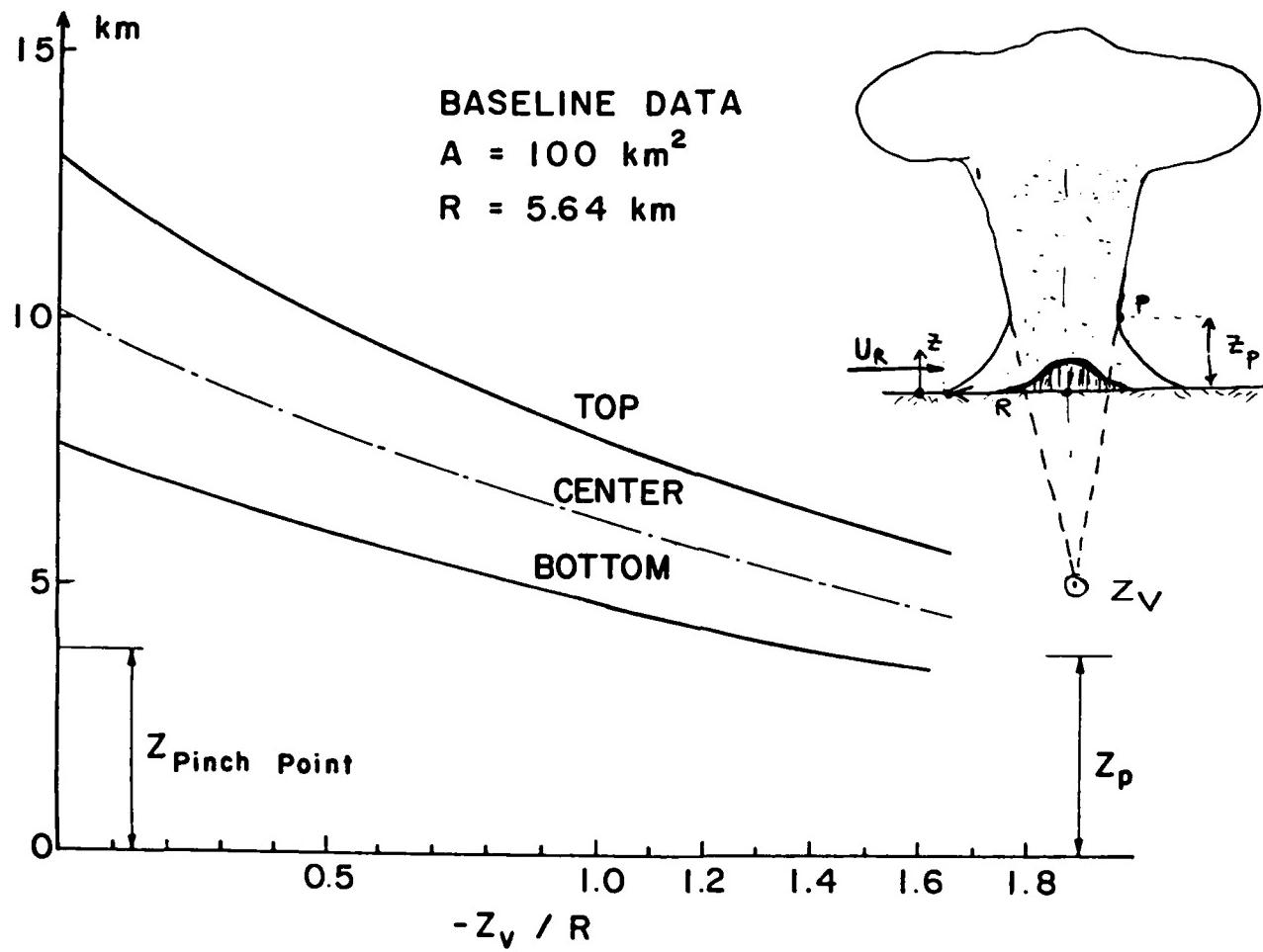
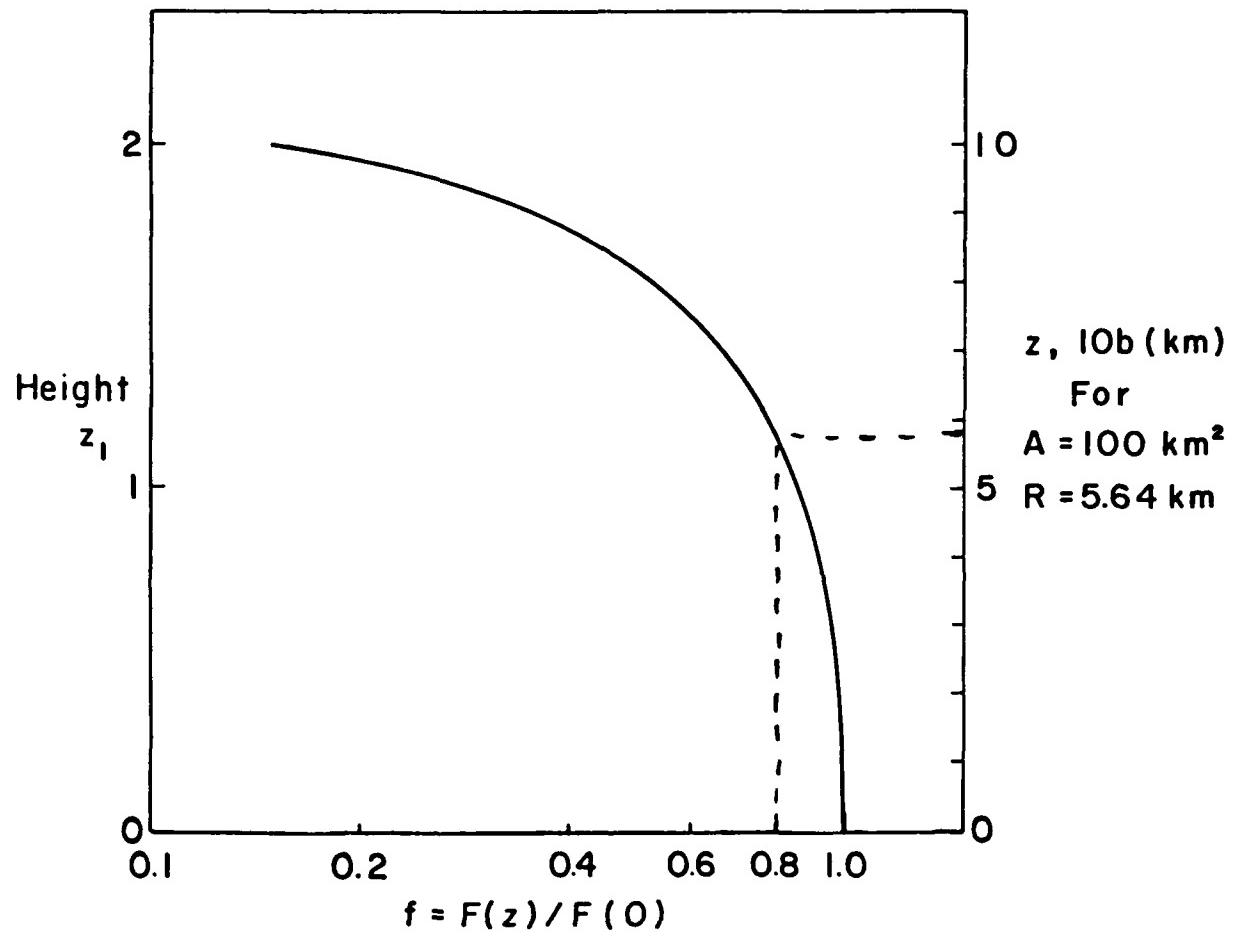


Figure 9. Estimate plume rise for a 100 km fire in a stably stratified atmosphere as a function of the position of the virtual source.



**Figure 10.** The relative reduction of buoyancy flux  $F$  with the height  $z$  in a stably stratified atmosphere according to Morton et al. (1956). (Note that most buoyancy flux disappears at large heights after the plume has increased its width.)

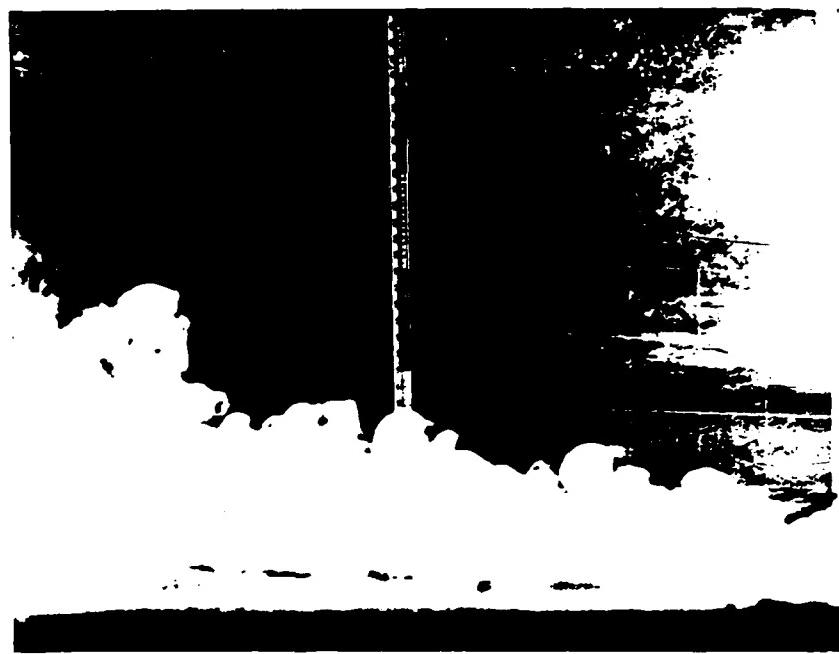
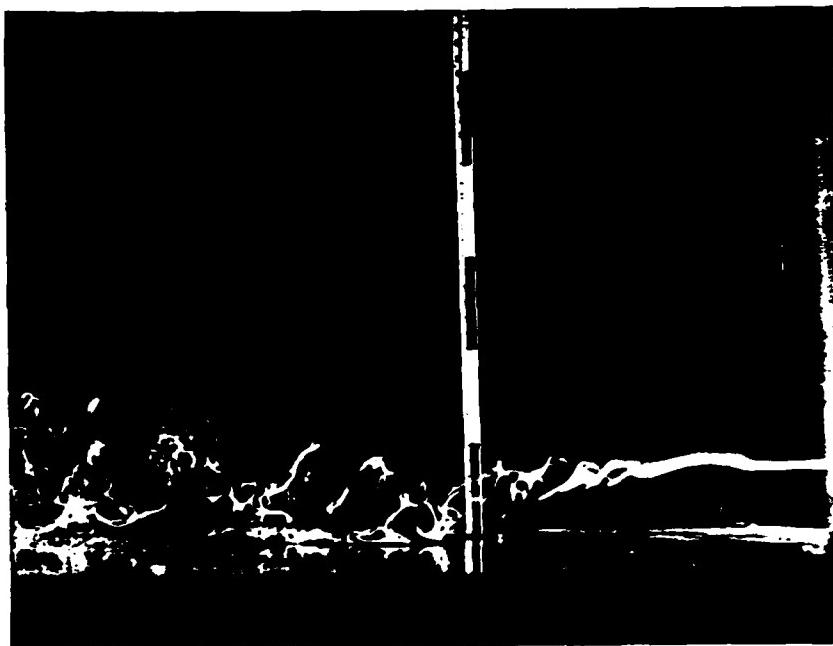
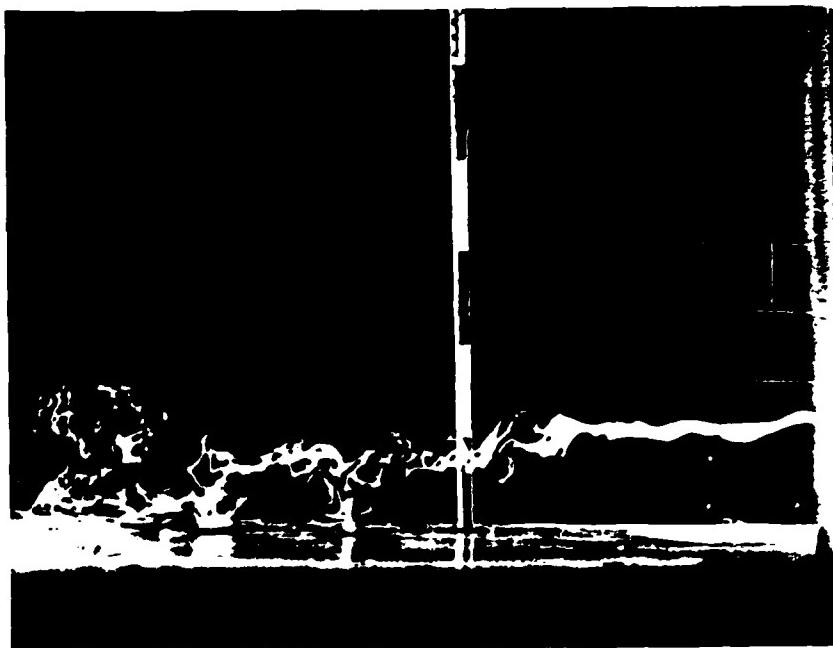


Figure 11. Photographs of the outer developing CBL in the two-dimensional fire model.



Figure 12. Photographs of the CBL in the two-dimensional fire model together with a smoke filament released above the layer.



**Figure 13.** Photographs of smoke filaments released above the CBL in the two-dimensional fire model. (Note the rapid decent of the filaments in the CBL.)

# **SMOKE PLUMES FROM LARGE AREA FIRES**

**A. KARAGOZIAN,**

**D. REMETCH**

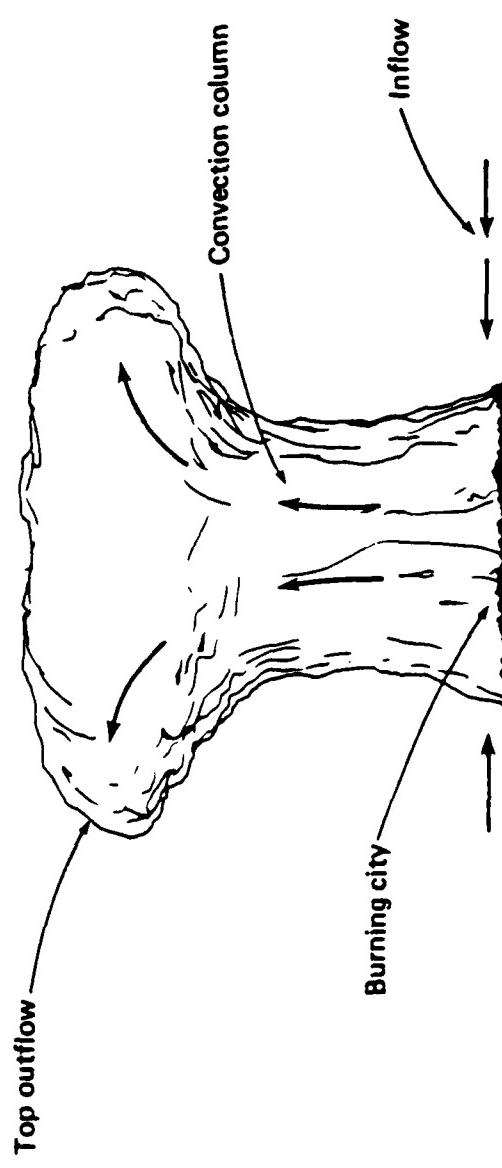
**R. D. SMALL**

**GLOBAL EFFECTS PROGRAM MEETING**

**February 26-27, 1986**

**PSR  
EATON**

## LARGE AREA FIRE IN THE ATMOSPHERE





212

## HYDROCODE APPROXIMATION OF LARGE SMOKE PLUMES

- VARIABLES
  - FIRE SIZE AND INTENSITY; BURNING TIME
  - ATMOSPHERE TEMPERATURE PROFILE
  - INVERSION HEIGHT
  - HUMIDITY PROFILE
- MODELING ASSUMPTIONS
  - HEAT RELEASE
  - TURBULENCE
  - TWO-DIMENSIONAL/AXISYMMETRIC
  - RESOLUTION
  - CITY STRUCTURE
- VALIDATION
  - FLAMBEAU
  - METEOTRON
  - CHAPEAU

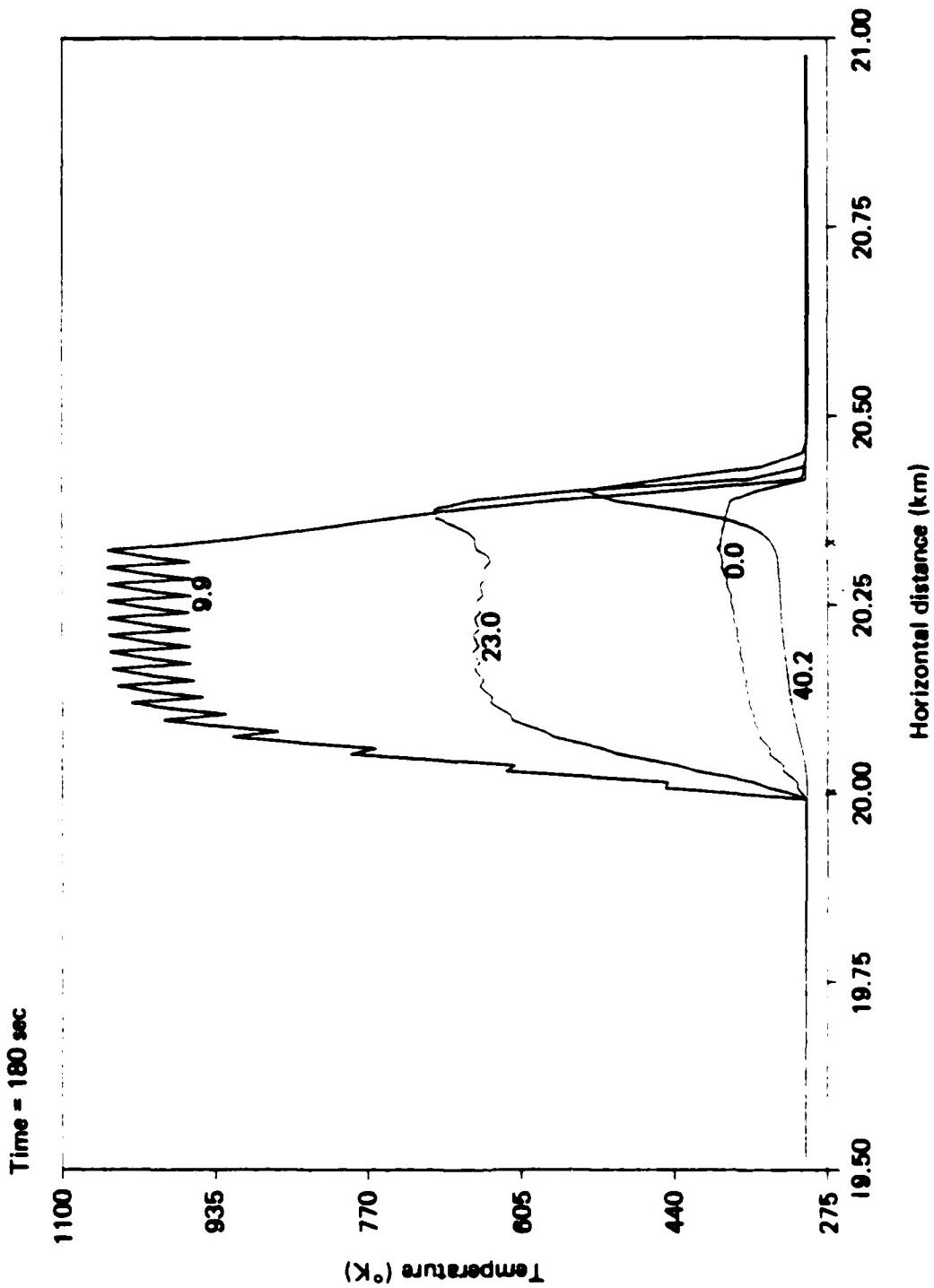
PSR  
E.T.N

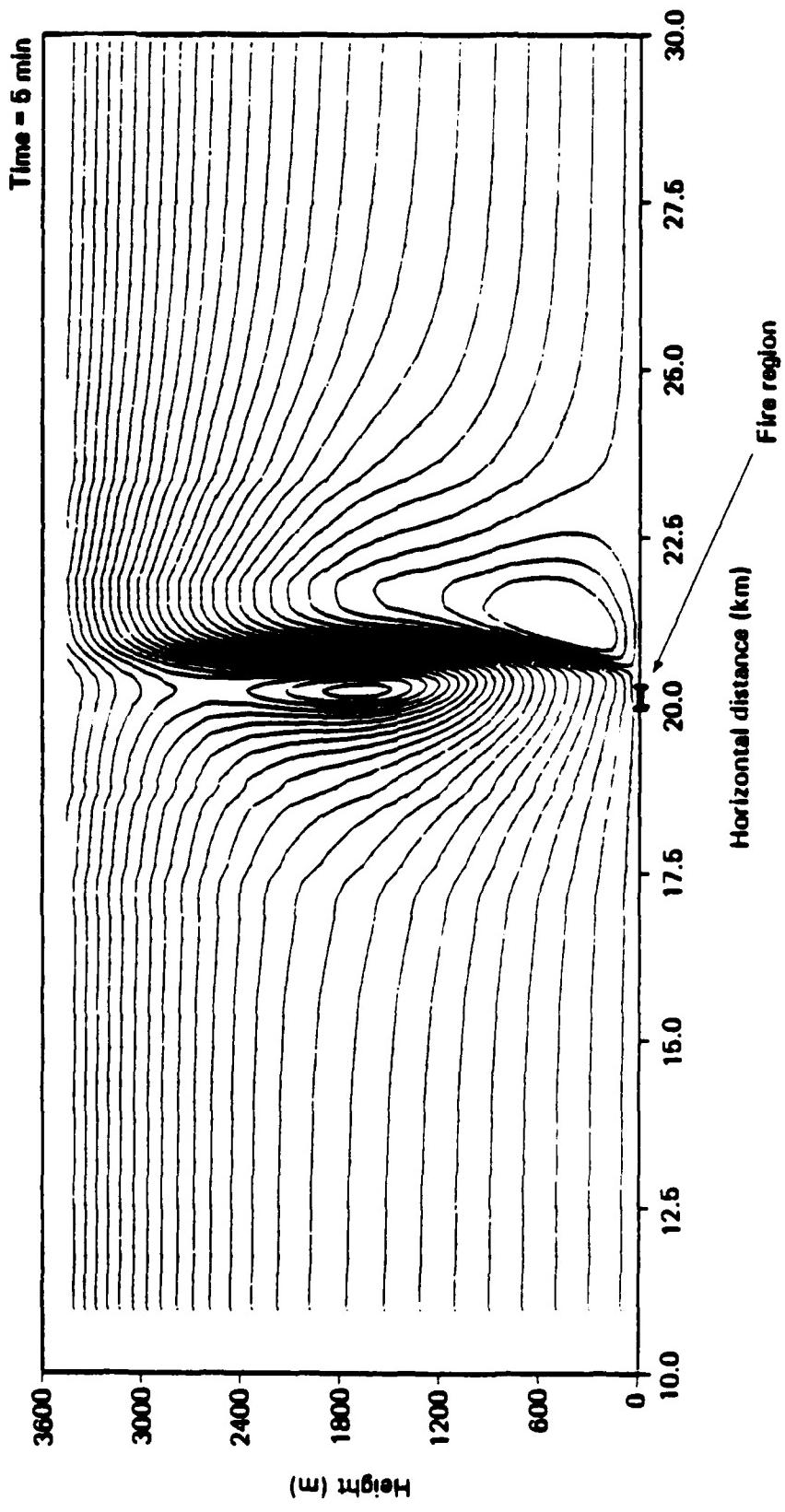
FLAMBEAU FIRE 460-7-66

- DATA
  - 30 ACRE TEST SITE
  - 240 15 m X 15 m FUEL PILES,
  - 7.5 m AISLES
  - 330 m CROSS SECTION
  - 40,000 LB/PILE - MIXED JUNIPER AND PINON
  - HEAT RELEASE MEASURED
  - VELOCITIES MEASURED
- SIMULATION
  - MEASURED HEAT RELEASE USED
  - TWO-DIMENSIONAL WITH AMBIENT WIND
  - FUEL BED - AISLES MODELED
- RESULTS
  - MERGING OF PLUMES CALCULATED
  - PLUME RISE IN CROSSWIND
  - FAIR AGREEMENT WITH MEASURED VELOCITIES

PSR  
F.T.N

## TEMPERATURE AT CONSTANT HEIGHTS





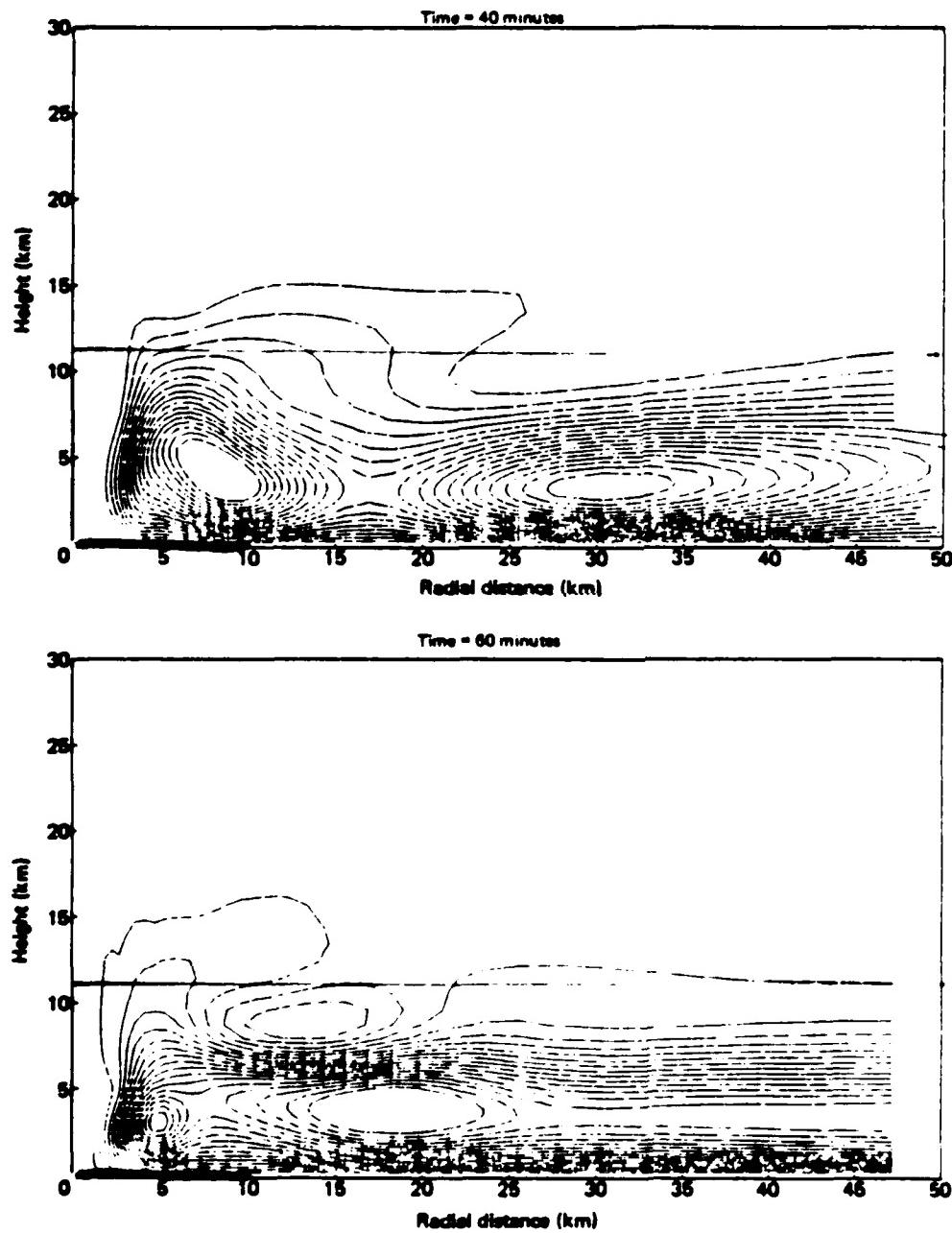
Streamlines: Flambeau fire 460~7~66.

EXAMPLE CASES

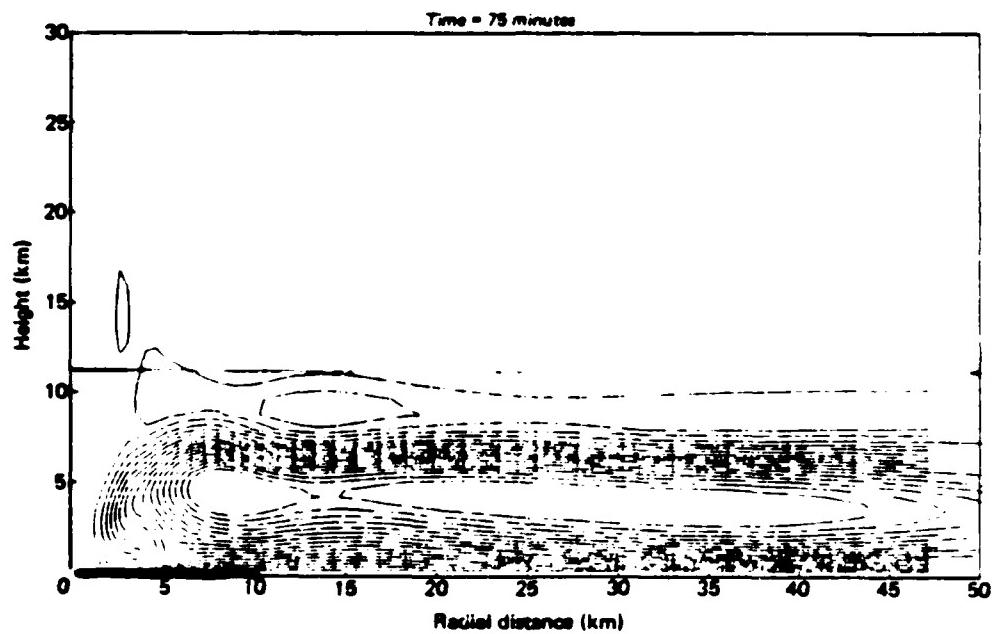
- FIRE RADIUS 10 km
- FIRE ZONE HEIGHT 100 m
- HEAT RELEASE RATES 1  $\text{kW}/\text{m}^3$  + 2.5  $\text{kW}/\text{m}^3$
- FUEL CONSUMPTION 2  $\text{g}/\text{cm}^2/\text{hr}$
- SMOKE EMISSION FACTOR 3 PERCENT
- RELATIVE HUMIDITY RATIOS 0, 0.5, 0.77
- U.S. STANDARD ATMOSPHERE TEMPERATURE PROFILE
- TROPOAUSE AT 11 km

PSR  
E.T.N

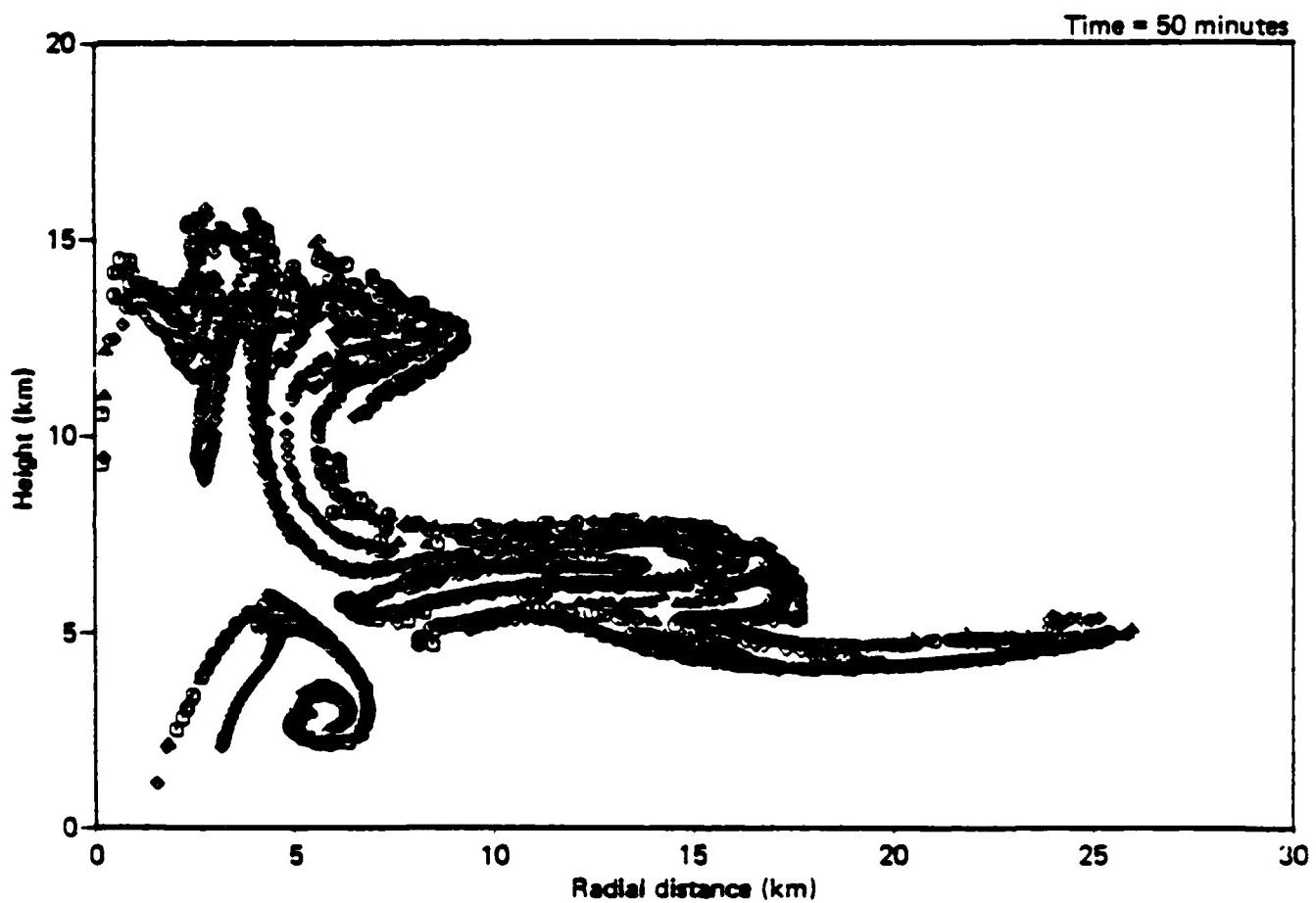
# **1 KW/M<sup>3</sup> LARGE AREA FIRE: U. S. STANDARD ATMOSPHERE**



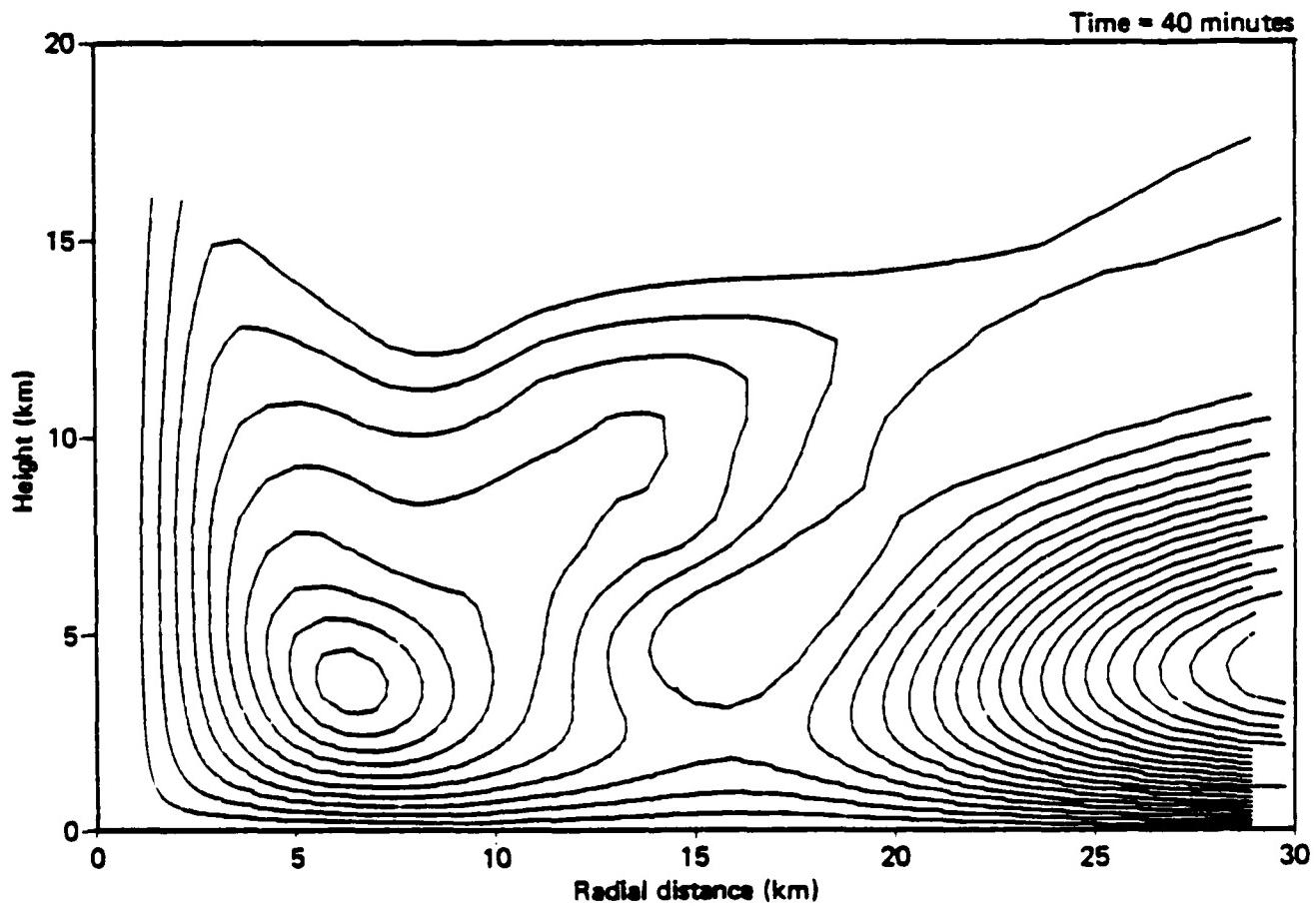
# **1 KW/M<sup>3</sup> LARGE AREA FIRE: U. S. STANDARD ATMOSPHERE**



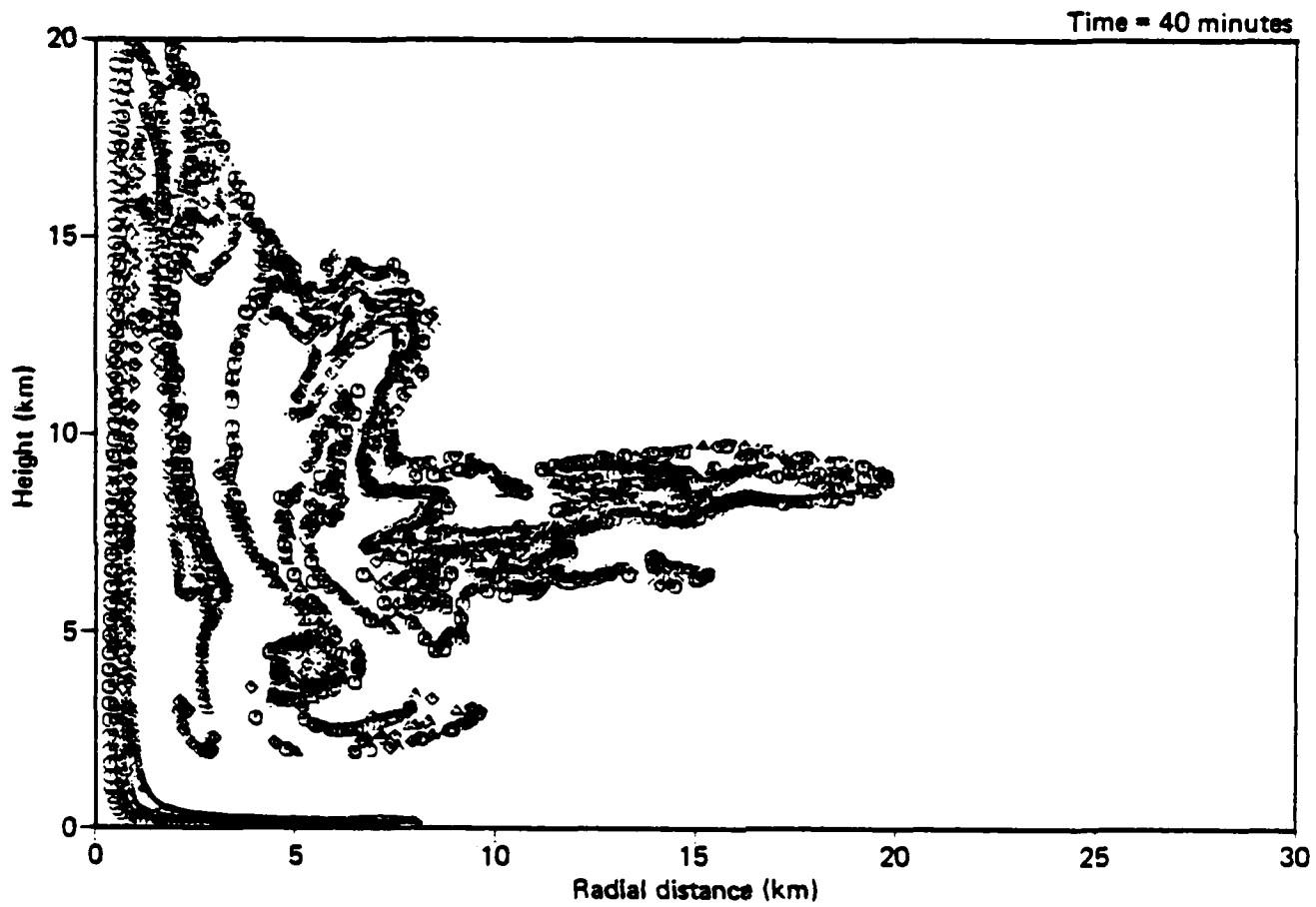
Smoke tracers: 10-km radius area fire;  $Q = 1\text{ kW/m}^3$ ;  $\phi_0 = 0.$



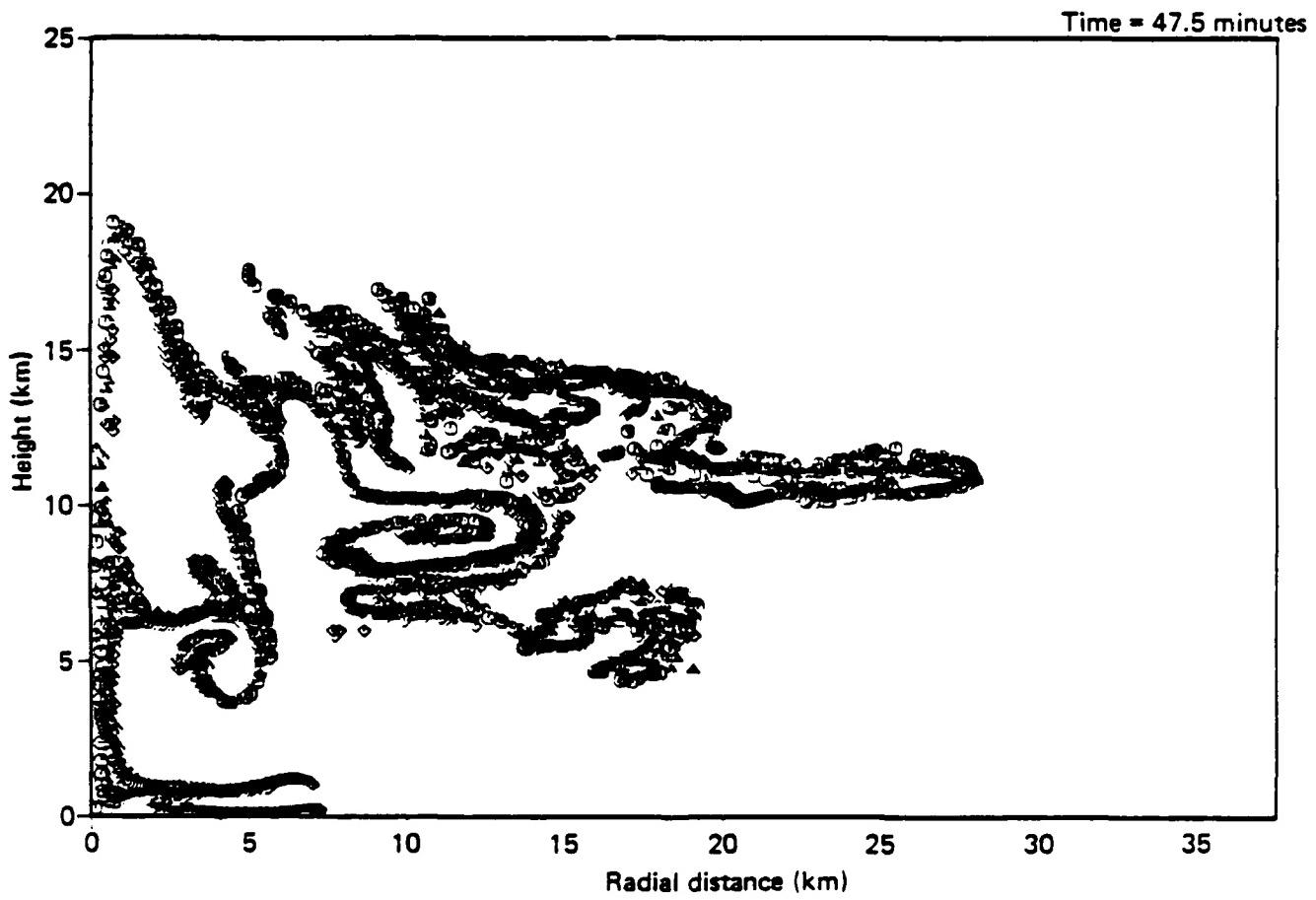
Streamlines: 10-km radius area fire;  $Q = 1 \text{ kW/m}^3$ ;  $\phi_0 = 0.77$ .



Smoke tracers: 10-km radius area fire;  $Q = 1 \text{ kW/m}^3$ ;  $\phi_0 = 0.77$ .



Smoke tracers: 10-km radius area fire;  $Q = 2.5 \text{ kW/m}^3$ ;  $\rho_0 = 0.77$ .



## CONCLUSIONS

- ATMOSPHERIC GRADIENTS CONTROL PLUME RISE
- MODERATE ENERGY PLUMES CONTAINED BY TROPOPAUSE
- MOISTURE INFLUENCES SMOKE CLOUD HEIGHT
- BETTER RESOLUTION OF CITY REQUIRED
- NEW HEAT RELEASE MODEL NEEDED
- LOWER HEAT RELEASE SIMULATIONS NEEDED

PSR  
E.T.N

TROPOPAUSE RESPONSE TO LARGE AREA FIRES

DAVID P BACON

SCIENCE APPLICATIONS  
INTERNATIONAL CORPORATION

FEBRUARY 26, 1986

PRESENTED AT THE DNA GLOBAL EFFECTS  
PROGRAM MEETING AT NASA/AMES

SAIC

Science Applications International Corporation

## The Terminal Area Simulation System (TASS)

Equations for:

Momentum

Pressure

Potential Temperature

Water Vapor

Cloud Water

Cloud Ice

Rain

Snow

Hail

Smoke (Massless Tracer)

Dust (Massive Tracer)

Turbulence:

Smagorinsky - shear  
- stratification

**Parameterizations :**

- Rain drop initiation
- Rain drop growth
- Evaporation of rain
- Freezing of supercooled cloud water and rain
- Initiation of cloud ice
- Ice growth
- Cloud ice becomes hail
- Deposition and sublimination of hail and cloud ice
- Hail growth
- Shedding of unfrozen water
- Initiation of snow
- Snow growth
- Melting of hydrometeors

**Additional Diagnostics :**

- Radar reflectivity
- Hydrometeor concentrations

## Boundary Conditions :

- Surface** - Non slip velocity
- Continuous thermodynamics
- Time-dependent heat source

Axial - Symmetric

Lateral - Open boundary

- Top** - Zero vertical velocity
- Other variables continuous

## Initial Conditions :

Ambient atmosphere initialized by temperature and dewpoint sounding.

Fireball modelled by a "Thermal Bubble" - a hemispherical region of over temperature (1 MT Nuclear = 675 KT Thermal)

**SAC**

Science Applications International Corporation

WOOD: 8500-9000 BTU/lb

8600 BTU/lb ~ 20 MJ/kg

2 % SMOKE EMISSION FACTOR = 20 g/kg

SMOKE EMISSION = 1 g/MJ

FIRE INTENSITY: UNIFORM OVER 10 km RADIUS

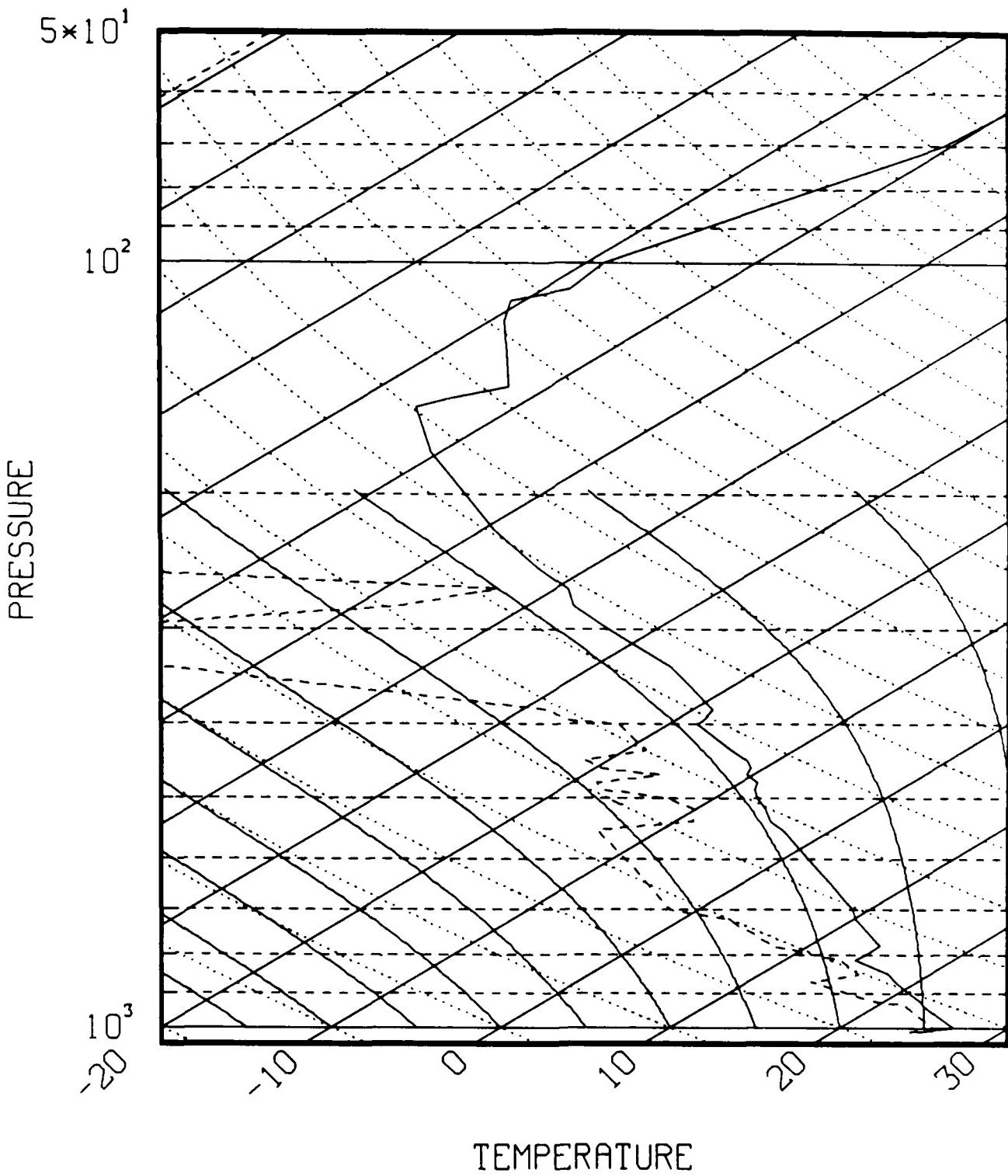
45 min LINEAR RISE  
60 min @ 50 kW/m<sup>2</sup>  
45 min LINEAR DECAY

TOTAL HEAT INPUT ~ 10<sup>17</sup> J  
TOTAL SMOKE INPUT ~ 10<sup>11</sup> g

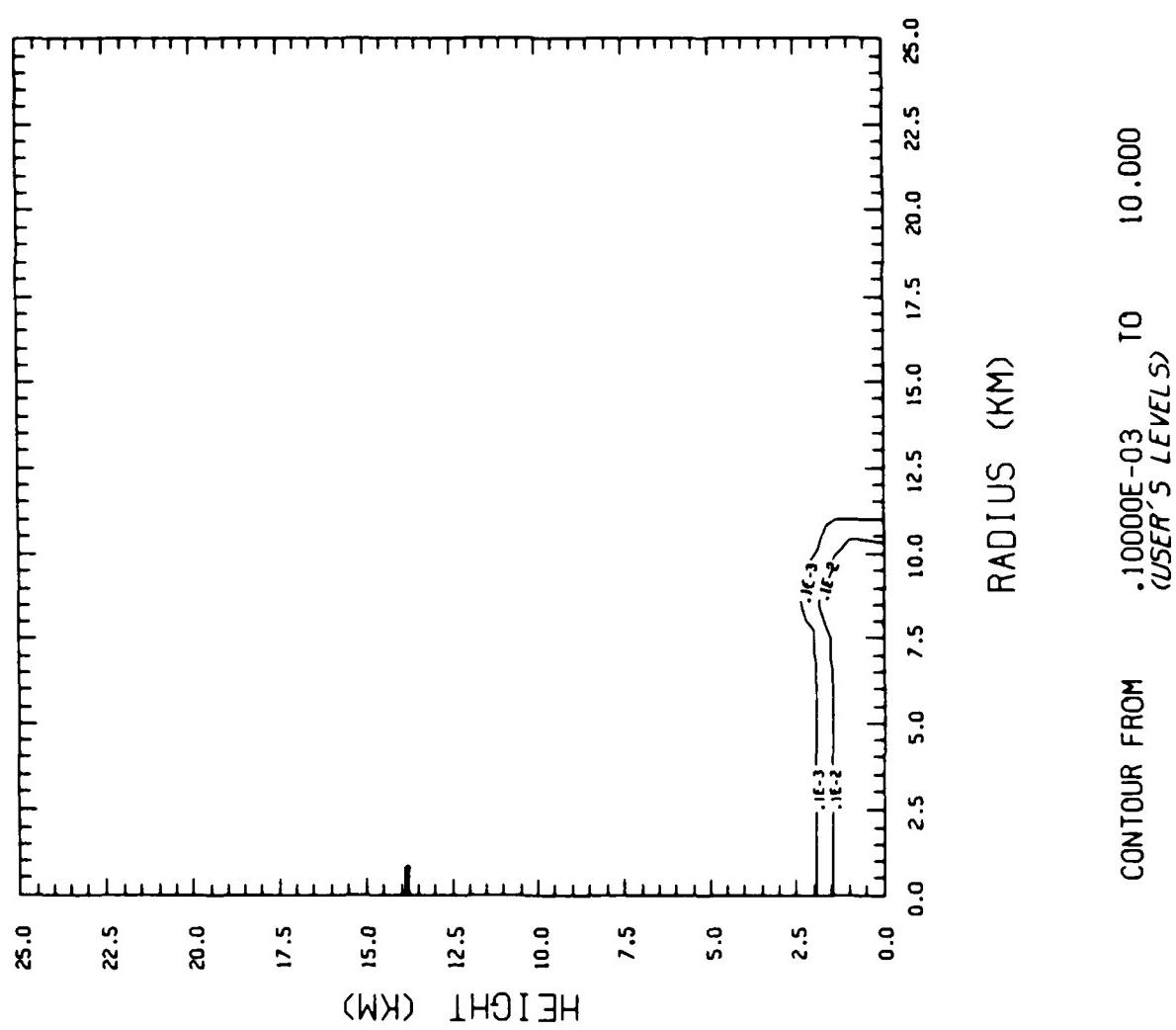
**SAC**

Science Applications International Corporation

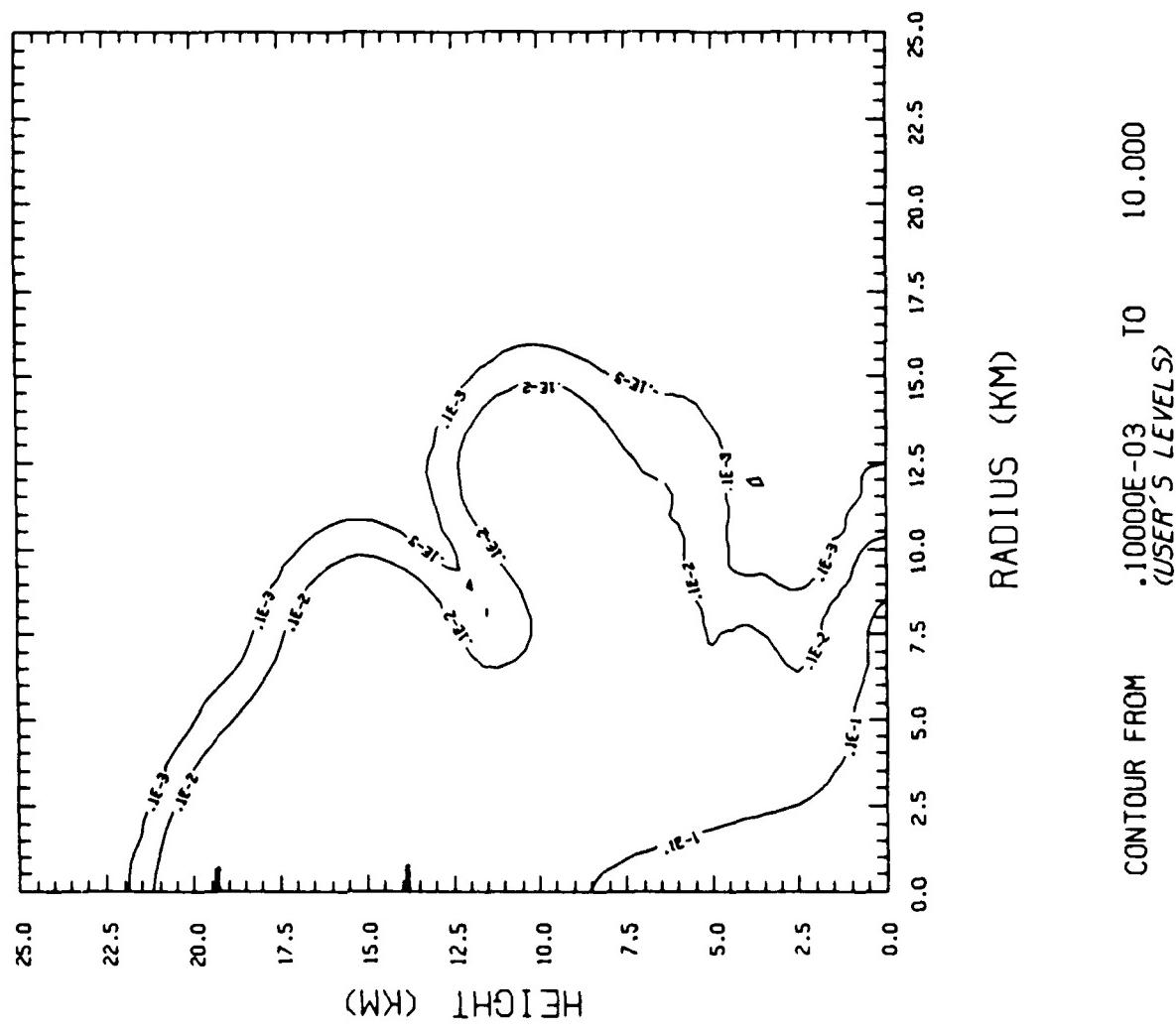
LCH DATA @ 2400Z 4-10-79



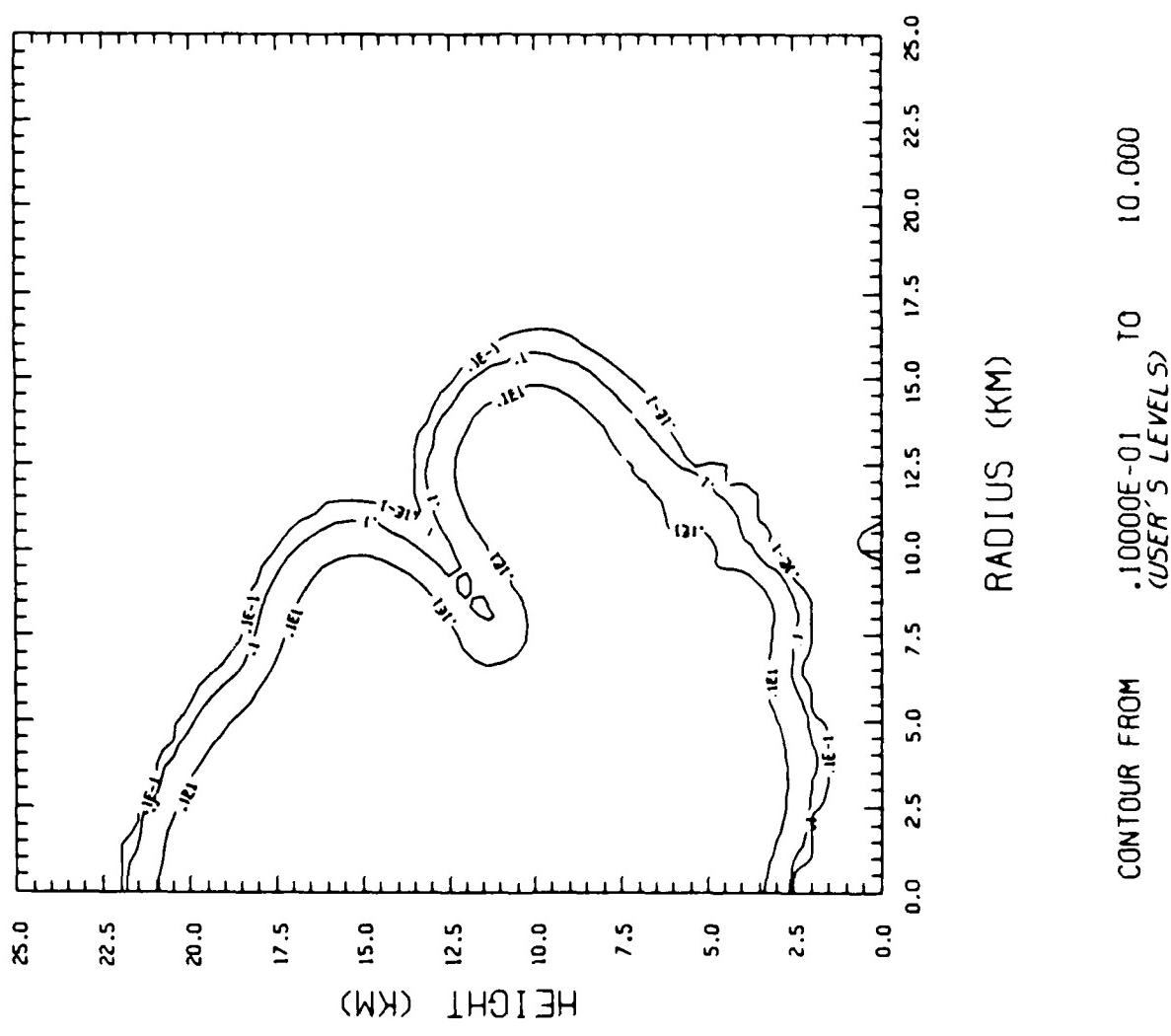
SMK 25 HR=-50 KW/M<sup>2</sup> @ 1CH TIME= 15.1



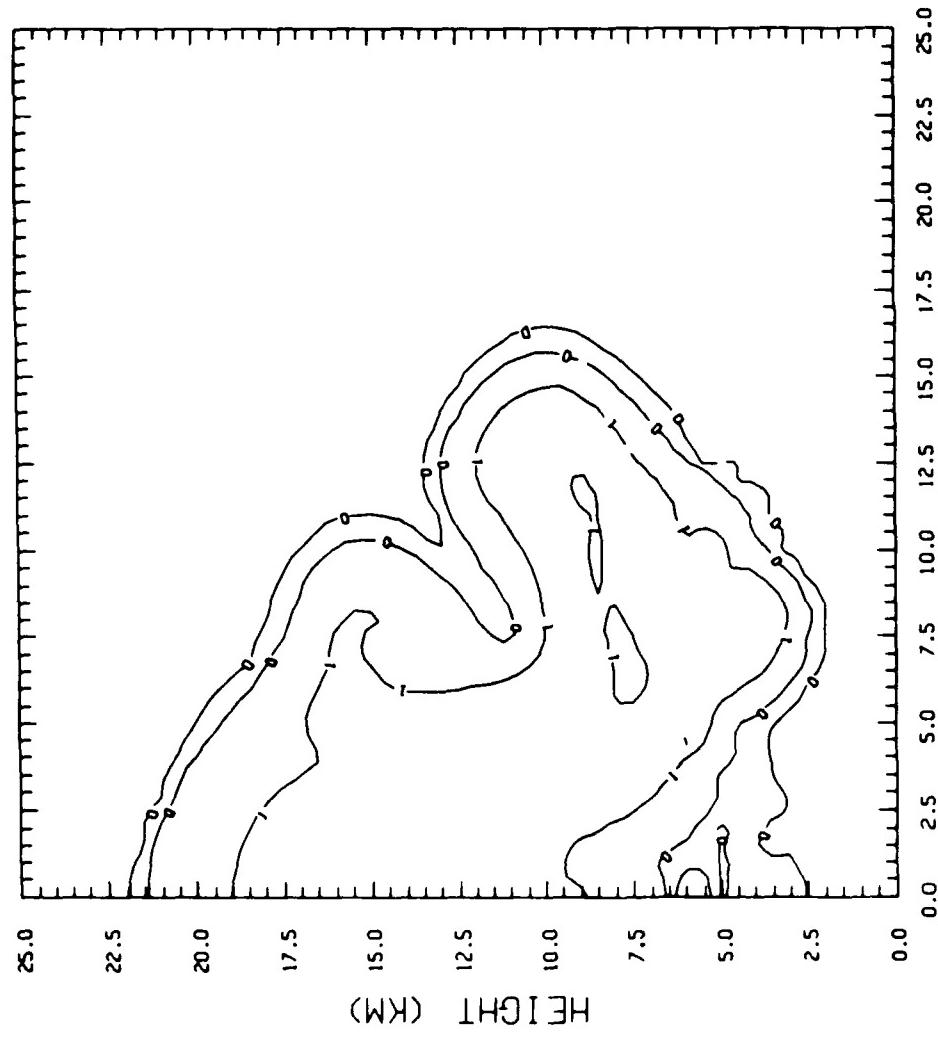
SMK 2.5 HR--50 KM/M\*\*2 • LCH TIME= 45.0



TCH 2.5 HR--50 KW/M\*\*2 @ LCH TIME= 45.0

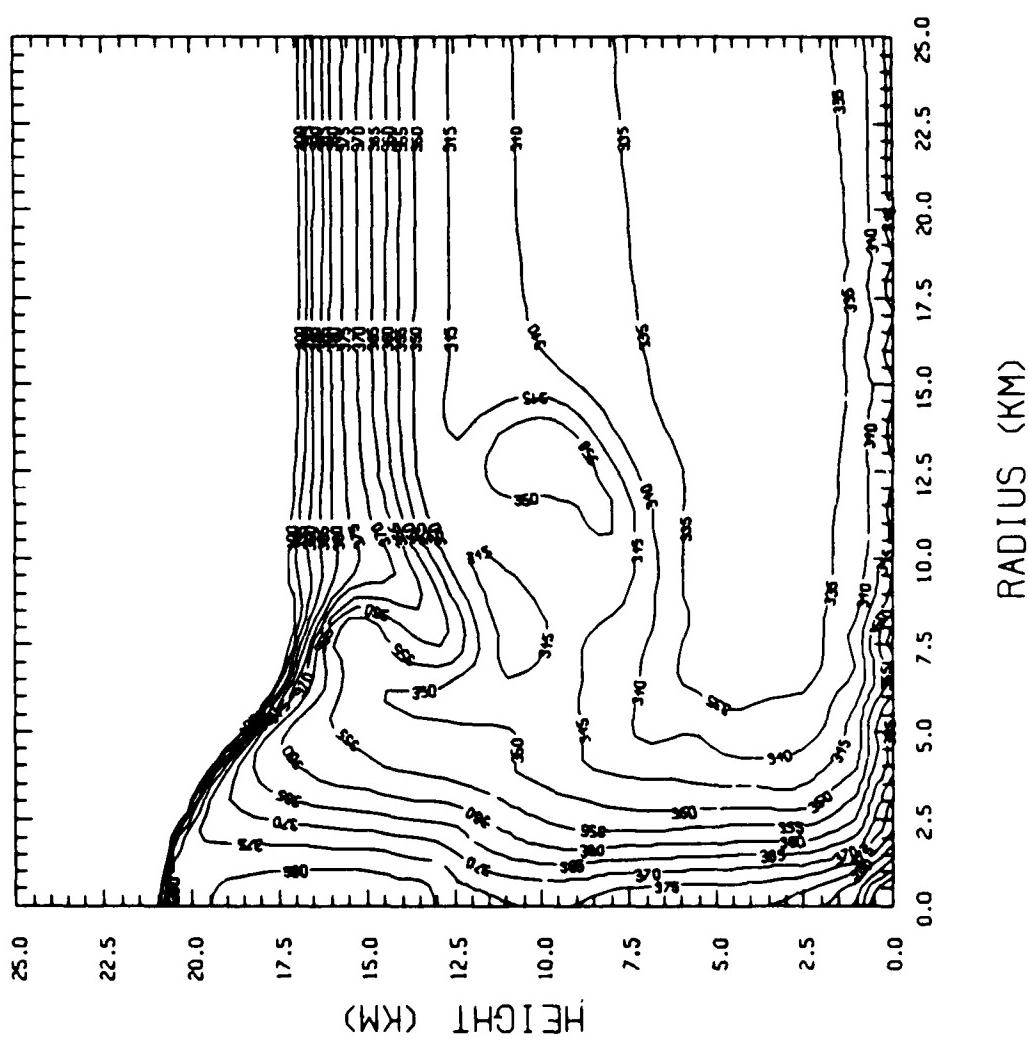


TPW 50 KW/M\*\*2 FIRE • LCH TIME= 45.0



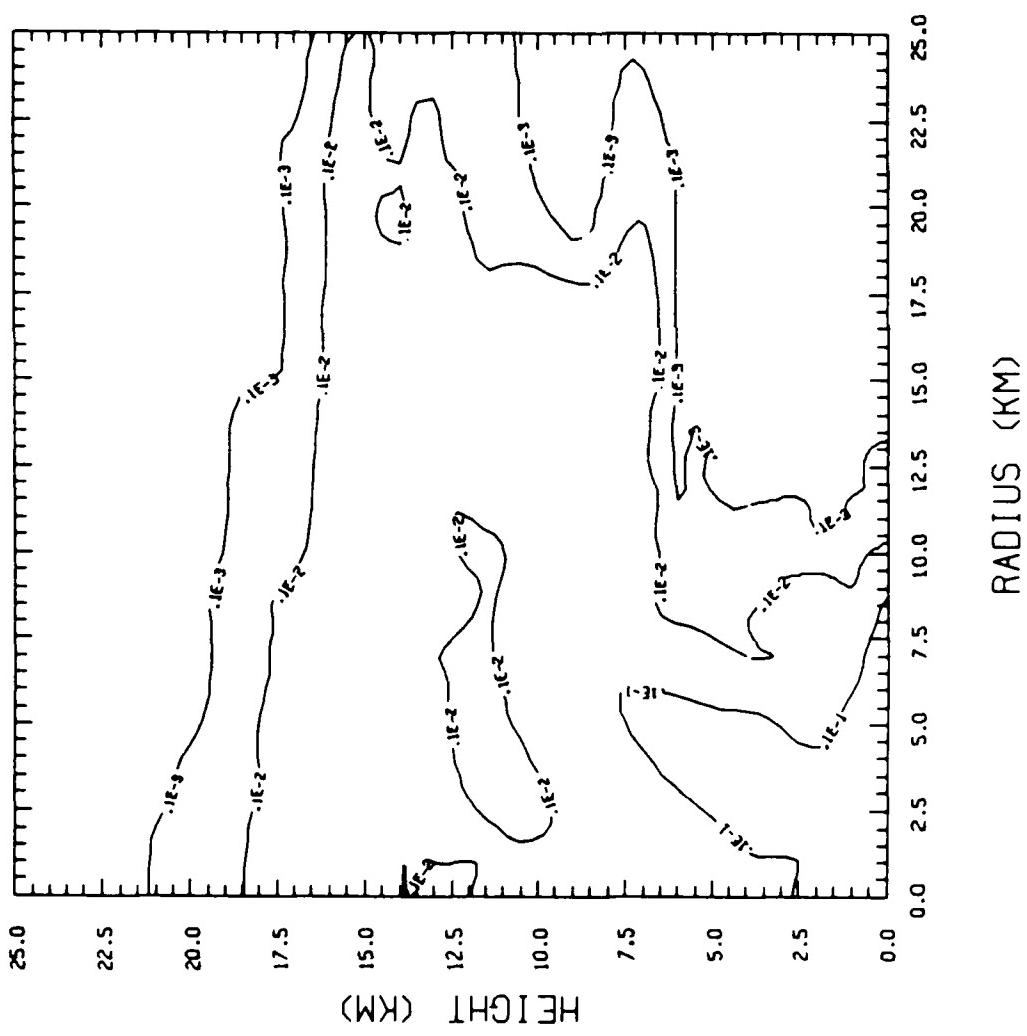
CONTOUR FROM .10000E-01 TO 1.0000  
(USER'S LEVELS)

EPOT 2.5 HR--50 KW/M\*\*2 @ LCH TIME= 45.0



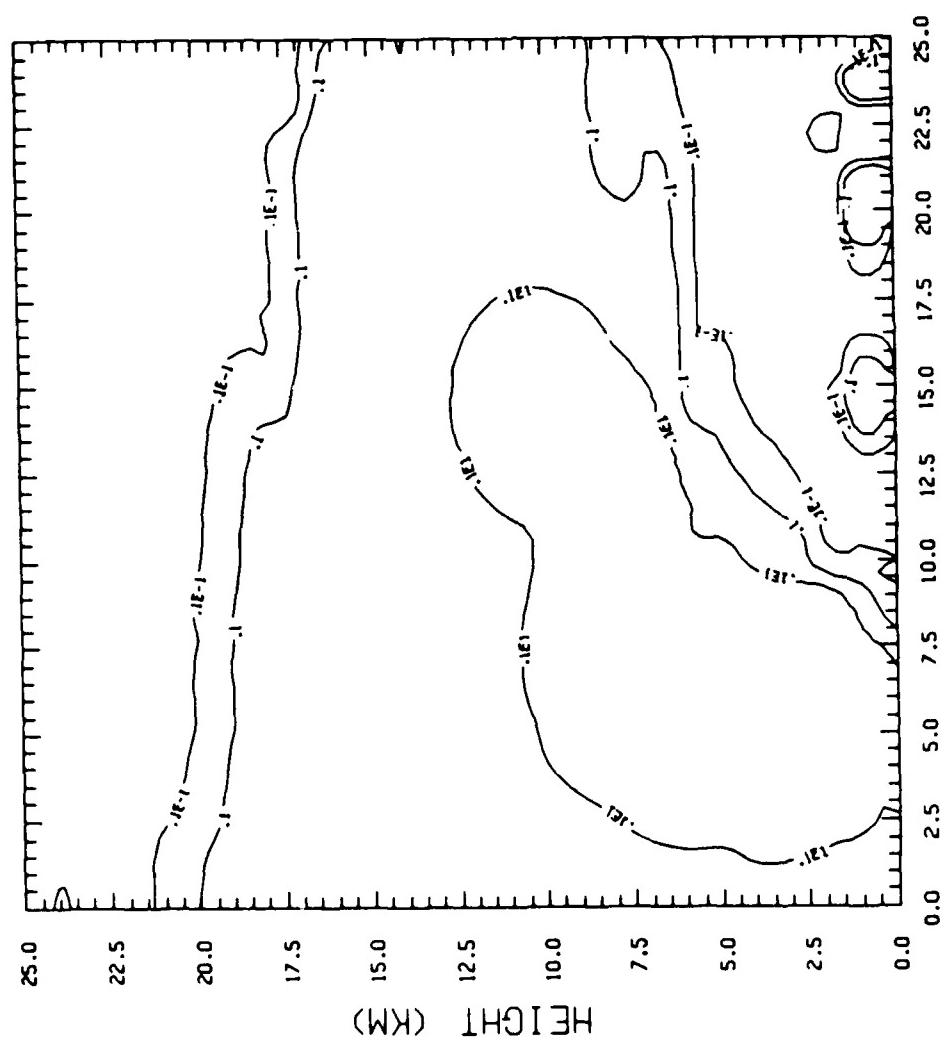
CONTOUR FROM 335.00 10 400.00  
(USER'S LEVELS)

SMK 2.5 HR--50 KW/M\*\*2 • LCH TIME= 90.0



CONTOUR FROM 10000E-03 •  
• (USER'S LEVELS) 10 10.000

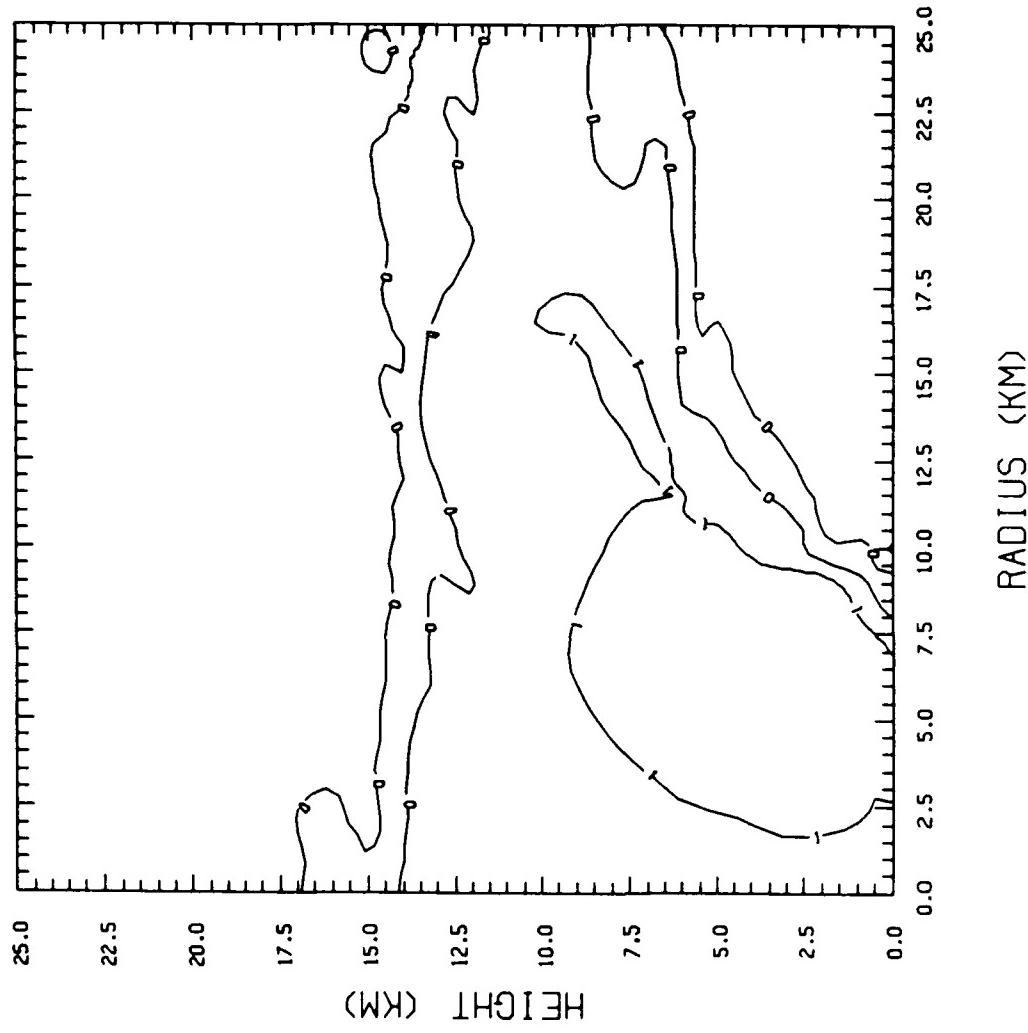
TCW 2.5 HR--50 KW/M\*\*2 @ LCH TIME = 90.0



RADIUS (KM)

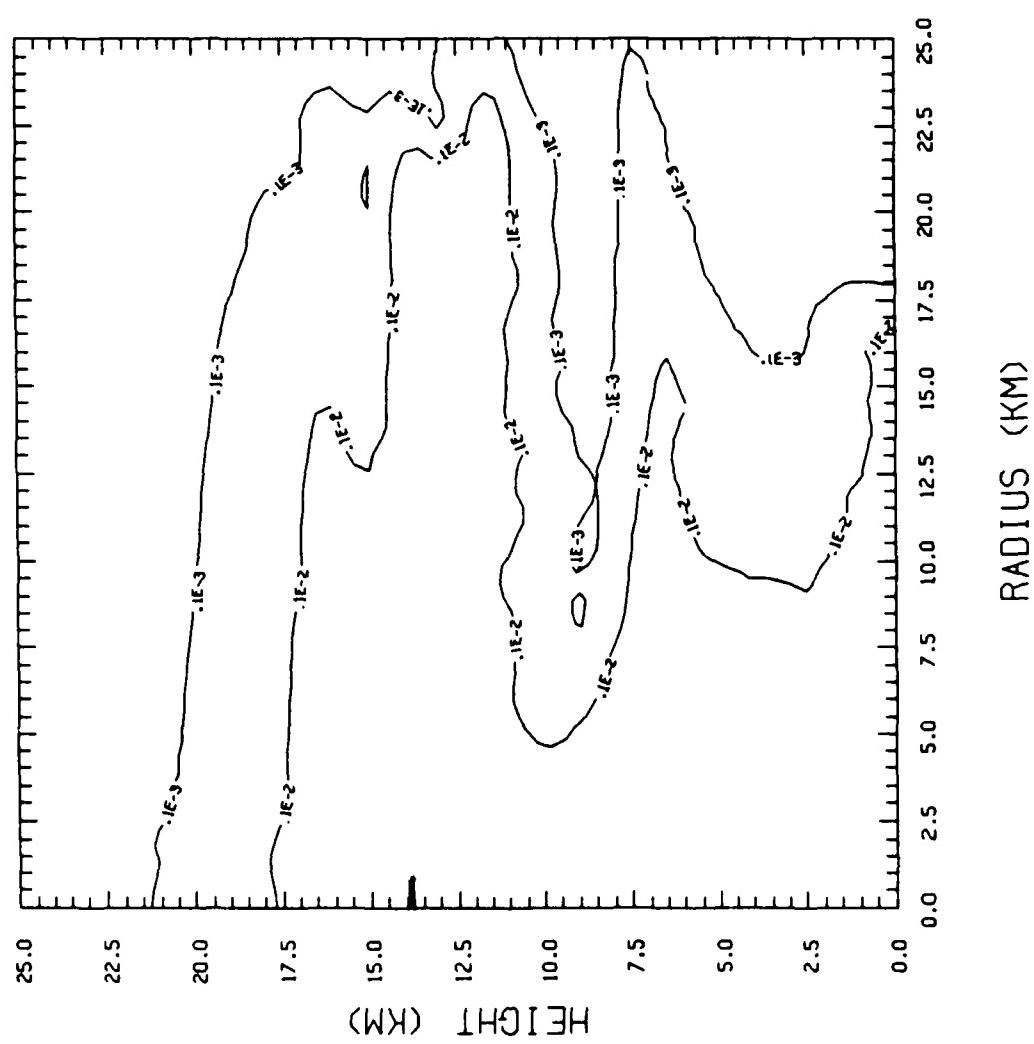
CONTOUR FROM 10000E-01 TO 10.000  
USER'S LEVELS

TPW 50 KW/M\*\*2 FIRE @ LCH TIME= 90.0



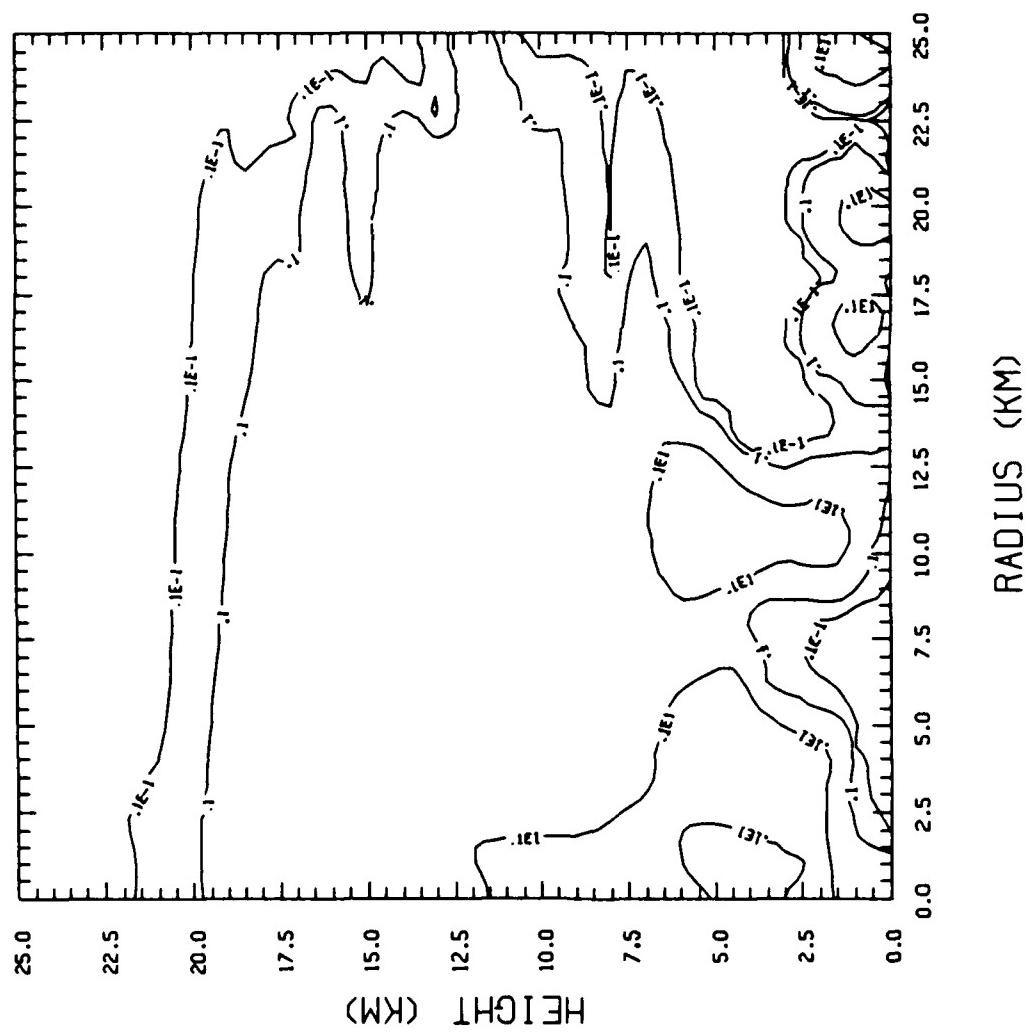
CONTOUR FROM .10000E-01 TO 1.0000  
(*USER'S LEVELS*)

SMK 2.5 HR--50 KW/M\*\*2 @ LCH TIME=180.0



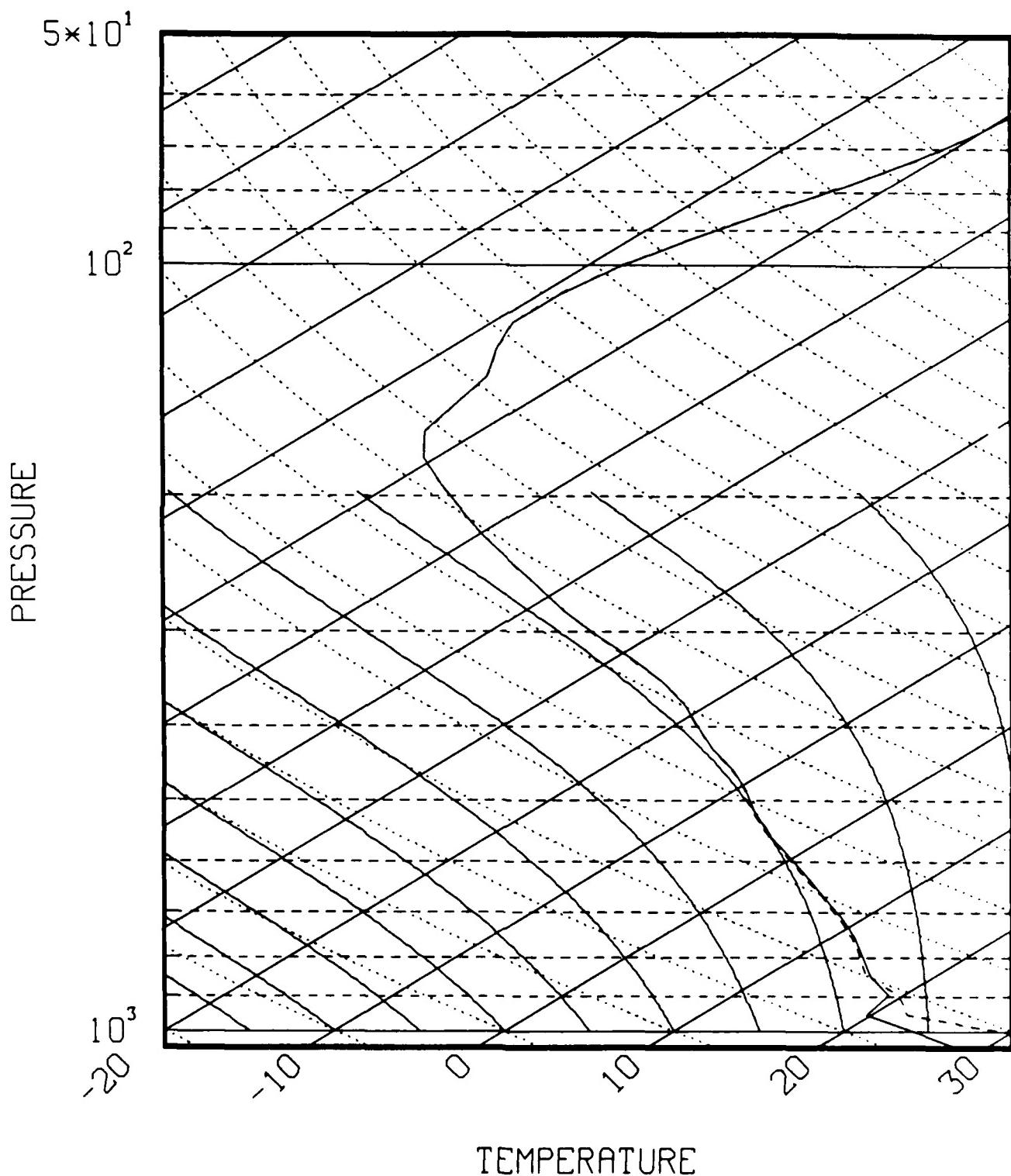
CONTOUR FROM 10000E-03 TO 10.000  
(USER'S LEVELS)

TCH 2.5 HR--50 KW/M\*\*2 @ LCH TIME=180.0

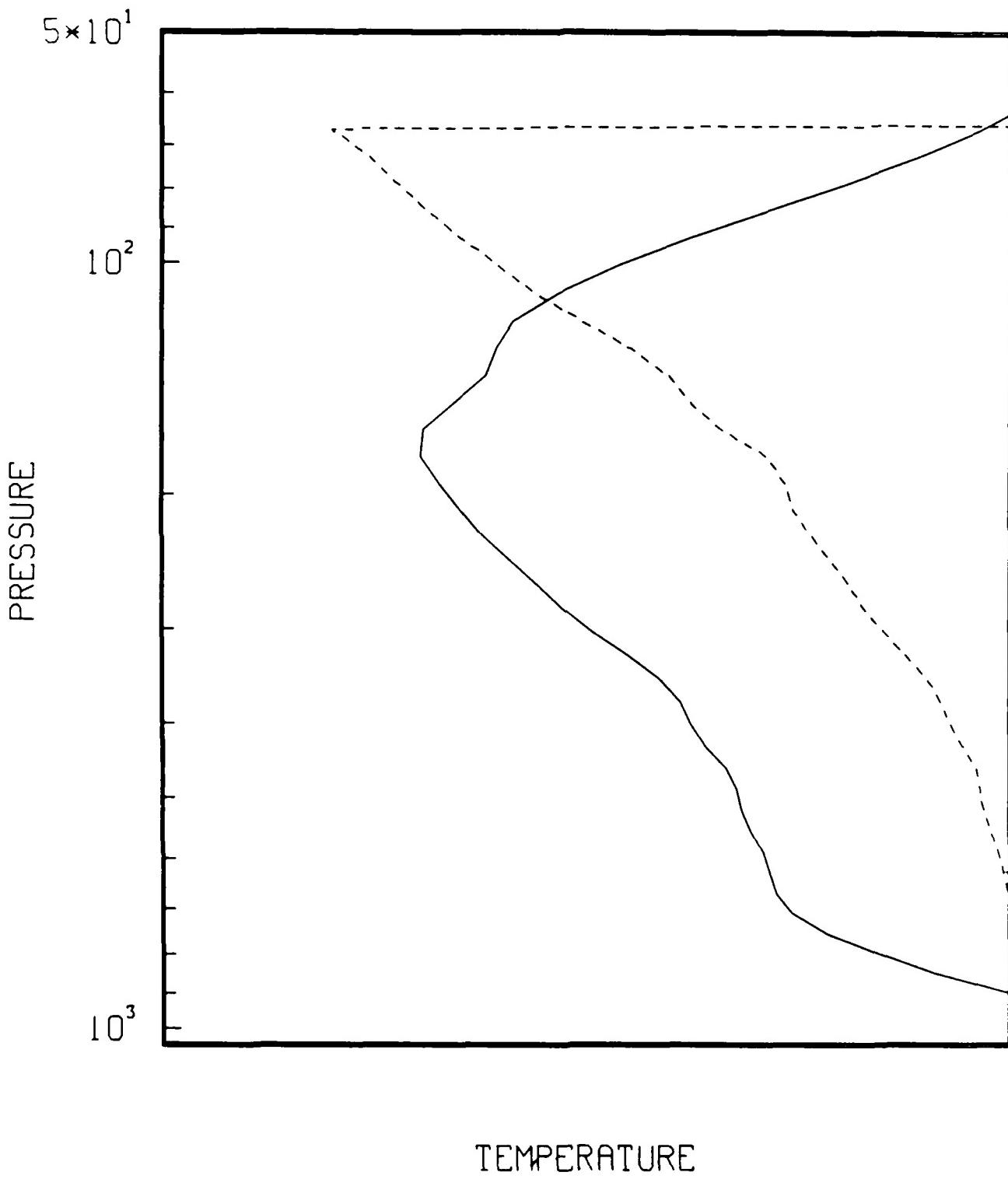


CONTOUR FROM .10000E-01 (USER'S LEVELS) TO 10.000

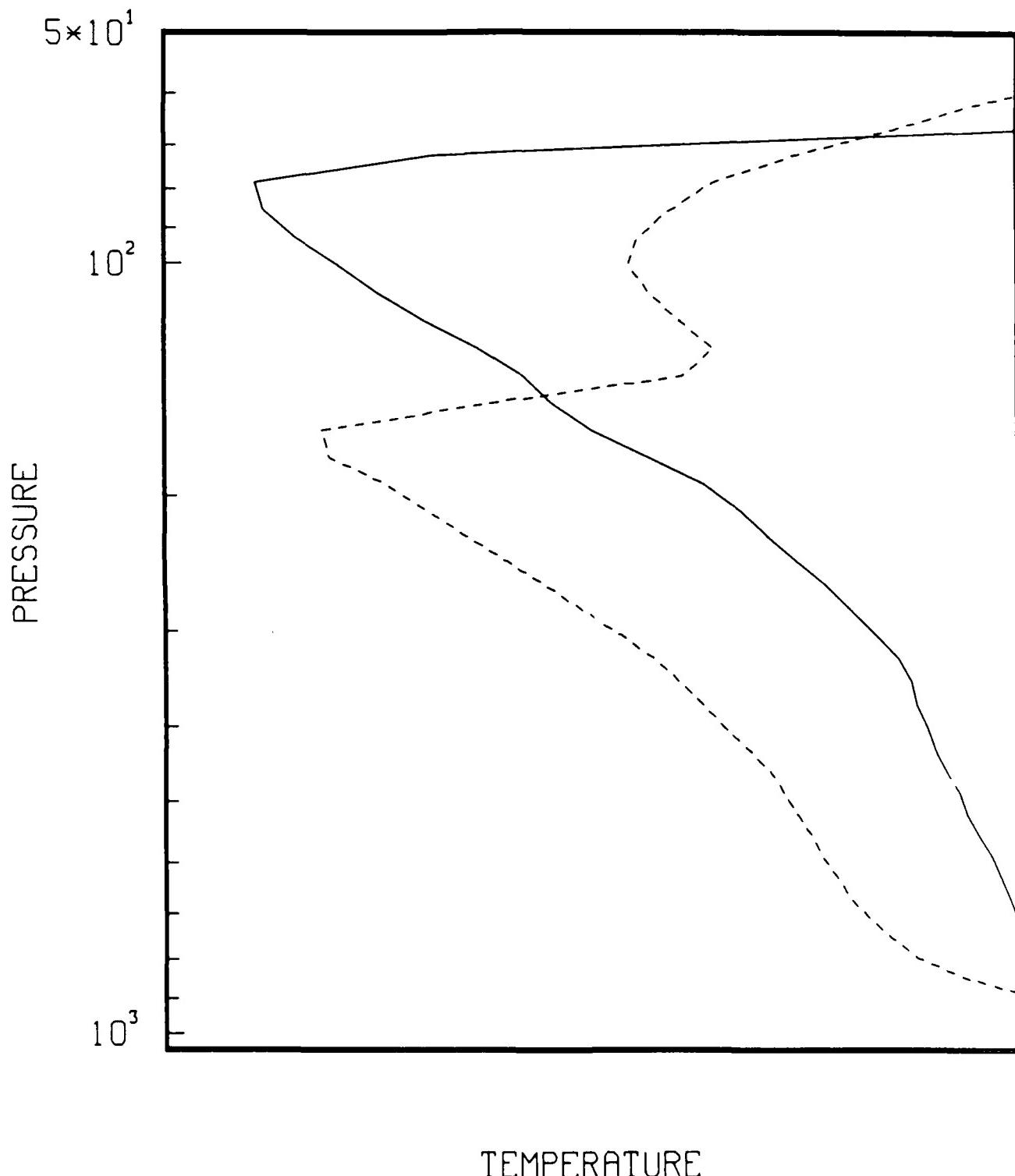
ANALYZED LCH DATA @ 0 & 15 MIN



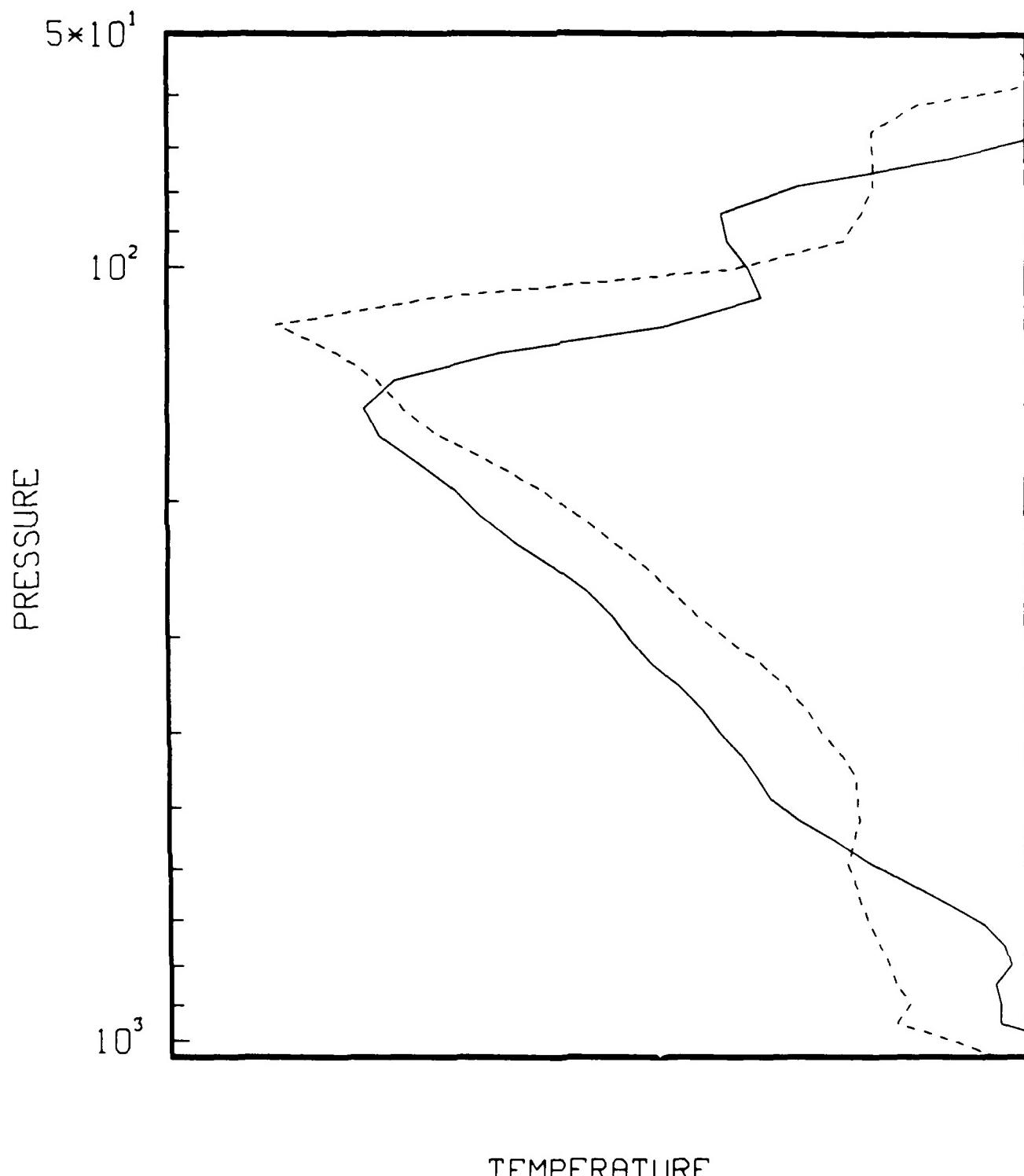
ANALYZED LCH DATA @ 30 & 45 MIN



ANALYZED LCH DATA @ 60 & 90 MIN

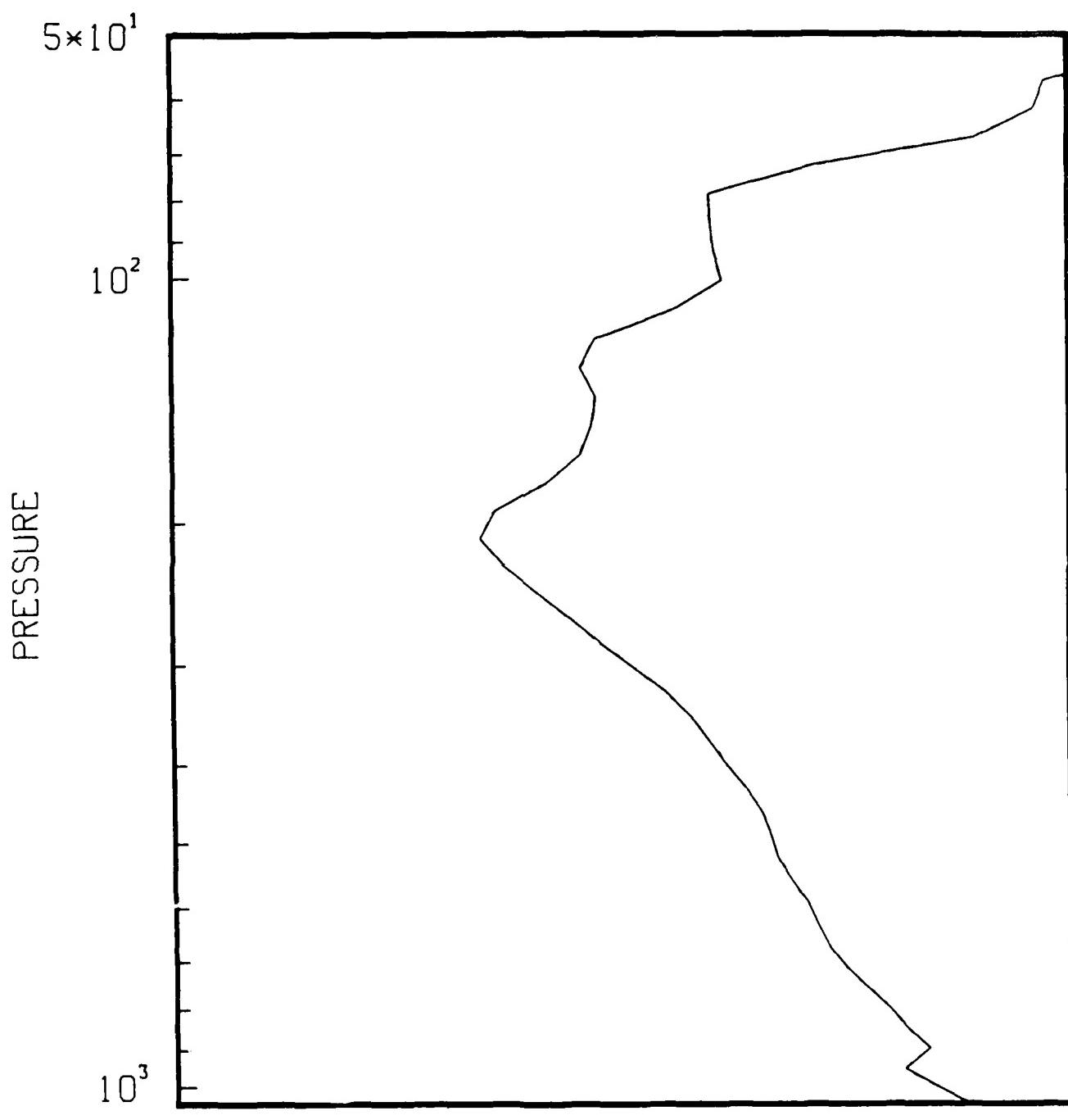


ANALYZED LCH DATA @ 120 & 150 MIN

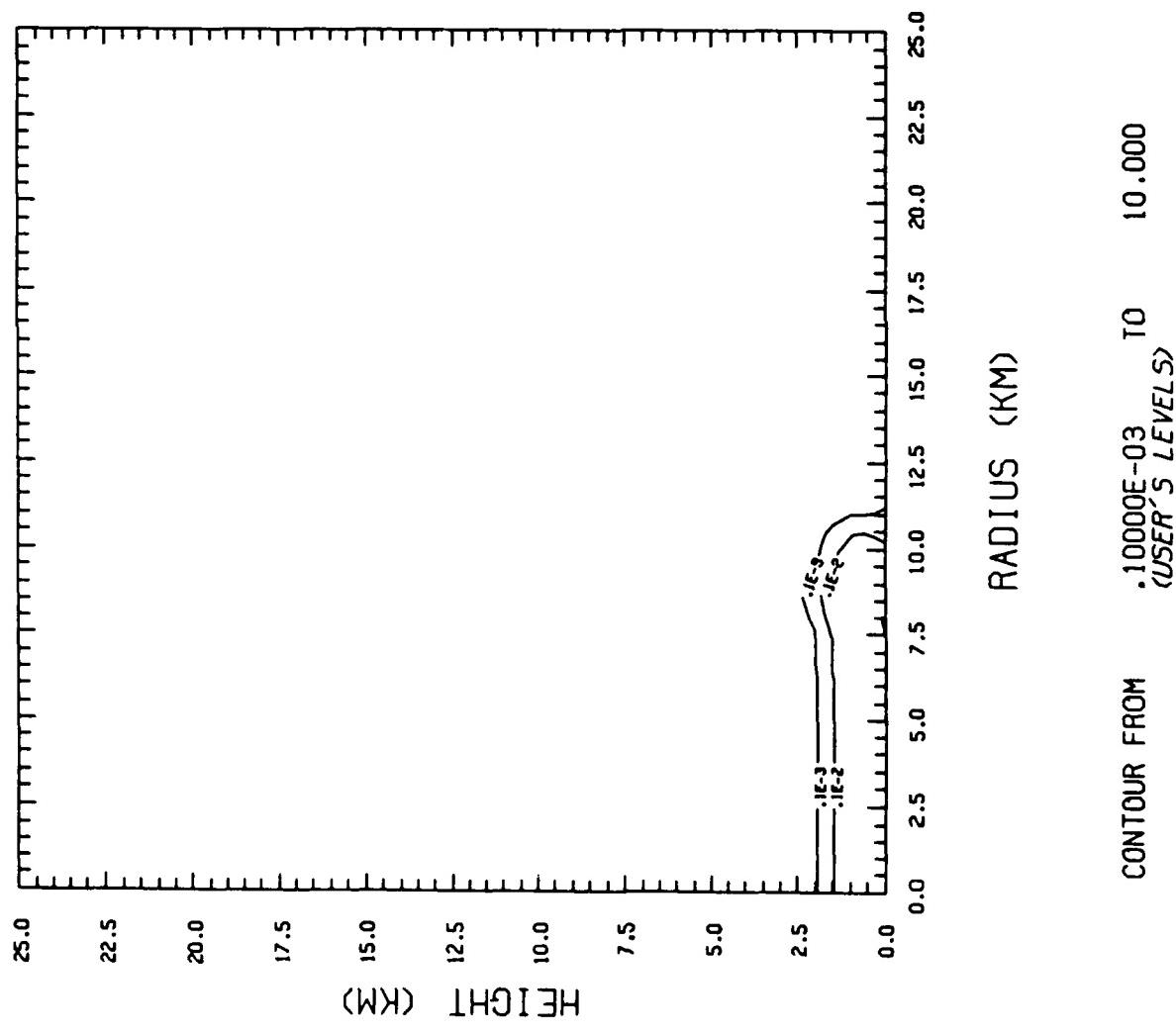


ANALYZED LCH DATA @

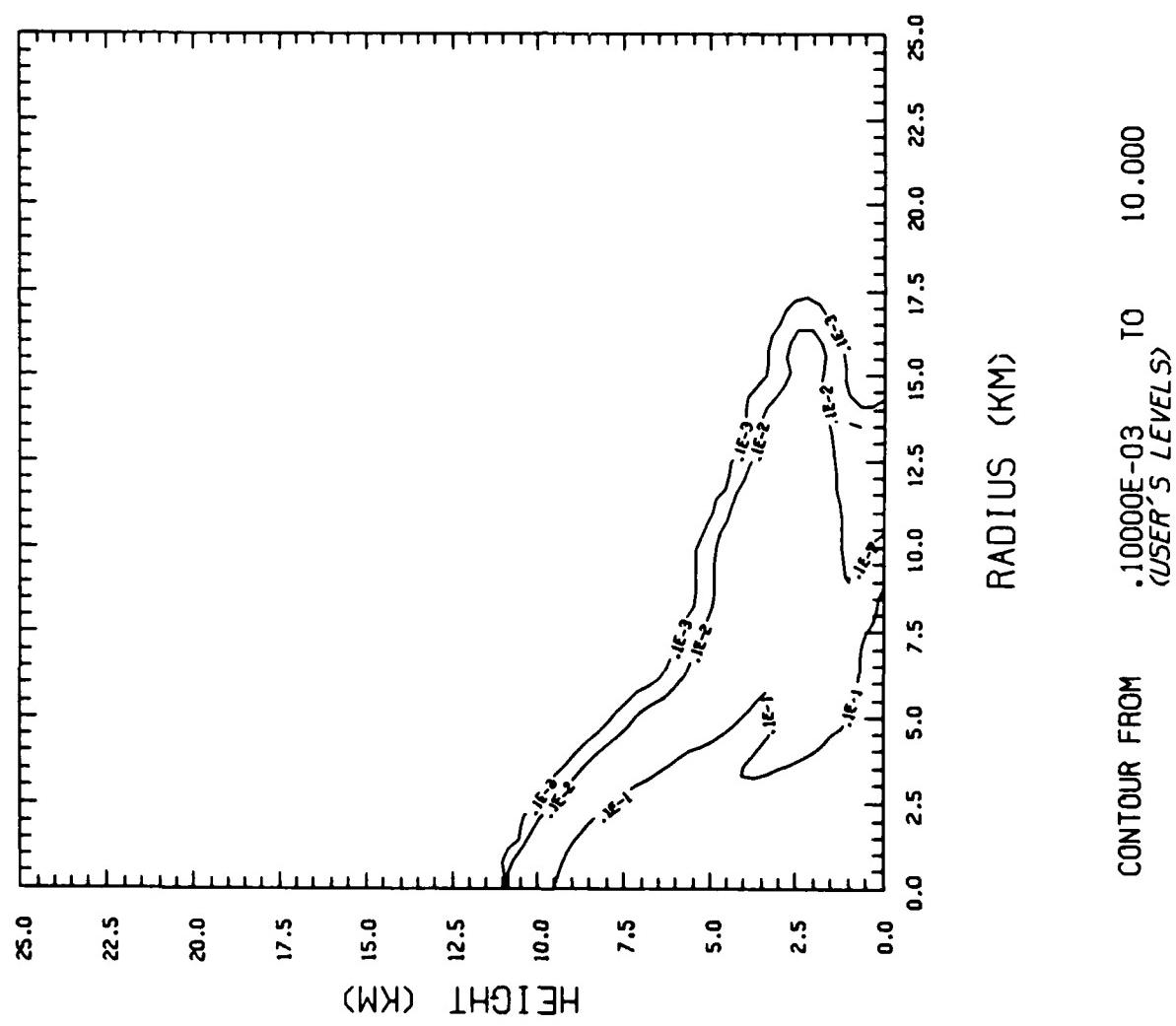
180 MIN



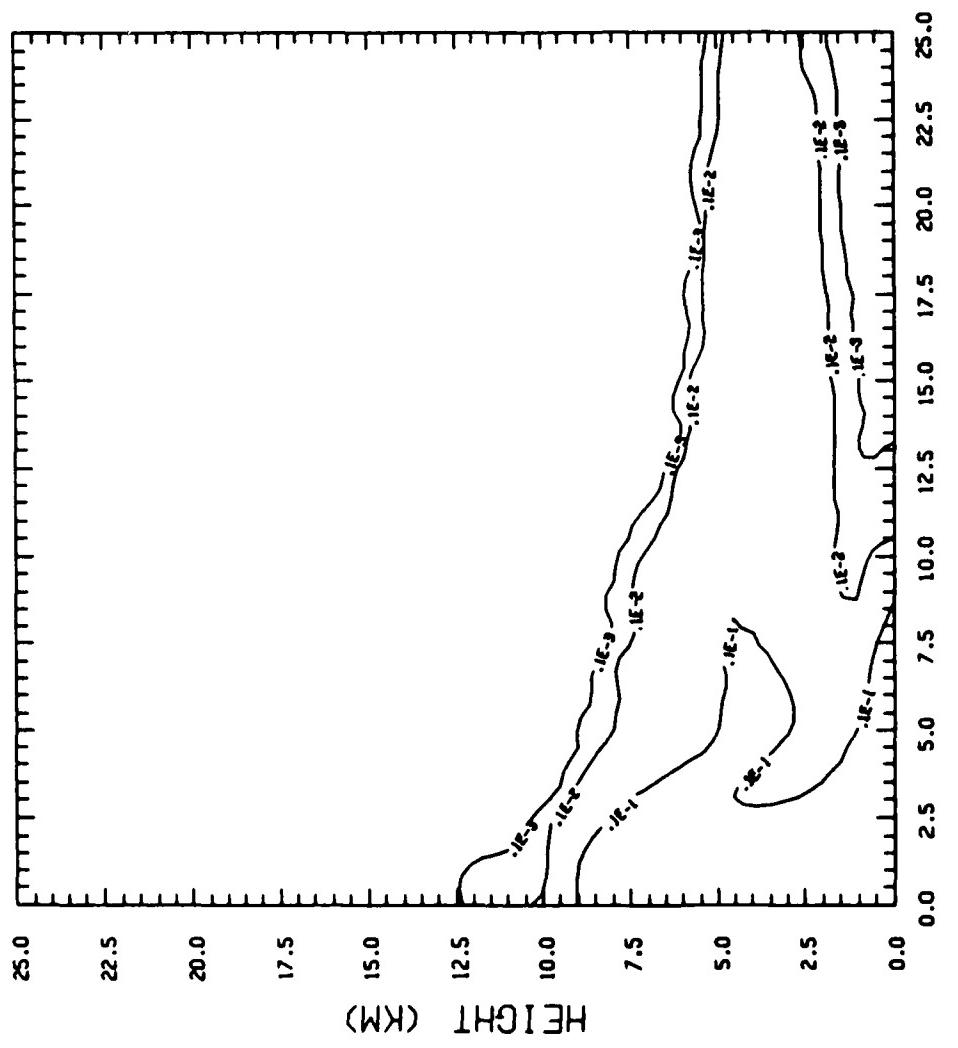
SMK 50 KW/M\*\*2 FIRE @ LCH-DRY TIME= 15.1



SMK 50 KW/M\*\*2 FIRE @ LCH-DRY TIME= 45.0



SMK 50 KW/M\*\*2 FIRE • LCH-DRY TIME= 90.0



### CONCLUSIONS :

- WITHOUT THE PRESENCE OF WATER, THE SMOKE CANNOT STABILIZE ABOVE THE TROPOAUSE.
- WITHIN MOISTURE PRESENT, THERE IS  $O(10^{-3})$  TIMES MORE CONDENSATE THAN SMOKE AND  $O(10^{-2})$  TIMES MORE PRECIPITATION. EARLY TIME SCAVENGING MUST BE CONSIDERED.
- SMOKE CAN BE INJECTED INTO THE "STRATOSPHERE", BUT ONLY ALONG WITH WATER WHICH WILL AFFECT ITS RESIDENCE TIME.

**NEXT EFFORTS :**

- ADD SOUTH DAKOTA SCHOOL OF MINES  
SCAVENGING MODEL TO TASS .
- ADD DIAGNOSTICS FOR OPTICAL DEPTH  
AND TOTAL SMOKE ABOVE A GIVEN  
ALTITUDE v . ALTITUDE .

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# The Numerical Simulation of Smoke Plume Dynamics in Intense Fires

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Michael M. Bradley

Lawrence Livermore National Laboratory

## The numerical model

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- Two-dimensional
- Time-dependent
- Nonhydrostatic
- Compressible
- Eulerian
- Terrain-following coordinate system
- Cloud and rain parameterization
- Dynamic equations are fully nonlinear
- Flux form of substance transport equations  
(using fourth-order spatial differencing)

## **Benchmark Case**

---

- U.S. standard spring-fall atmosphere
- No background wind
- Ambient moisture only - none from fire
- Smoke is passive tracer

Vertical Smoke Distributions After 1 hr

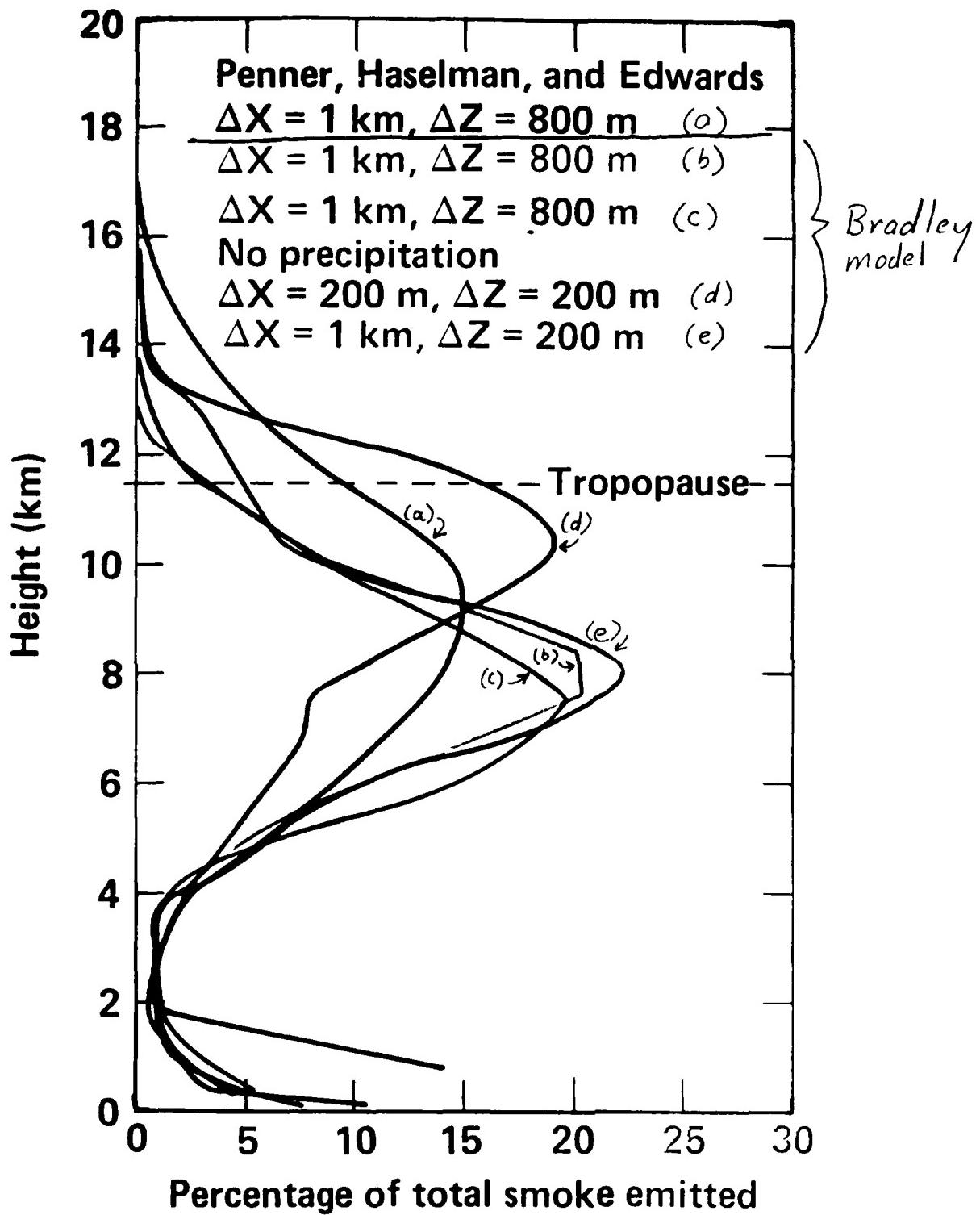


Figure 1 Normalized smoke profiles for intense fire ( $89 \text{ w m}^{-2}$ ) with no wind and ambient moisture only.

$\Delta x = \Delta z = 200 \text{ m}$   
time = 1 hr  
Bradley model

U.S. Standard spring-fall atmosphere  
No ambient wind  
Moisture-ambient only  
89  $\text{kW/m}^2$  heat source

## integrated smoke distribution \*

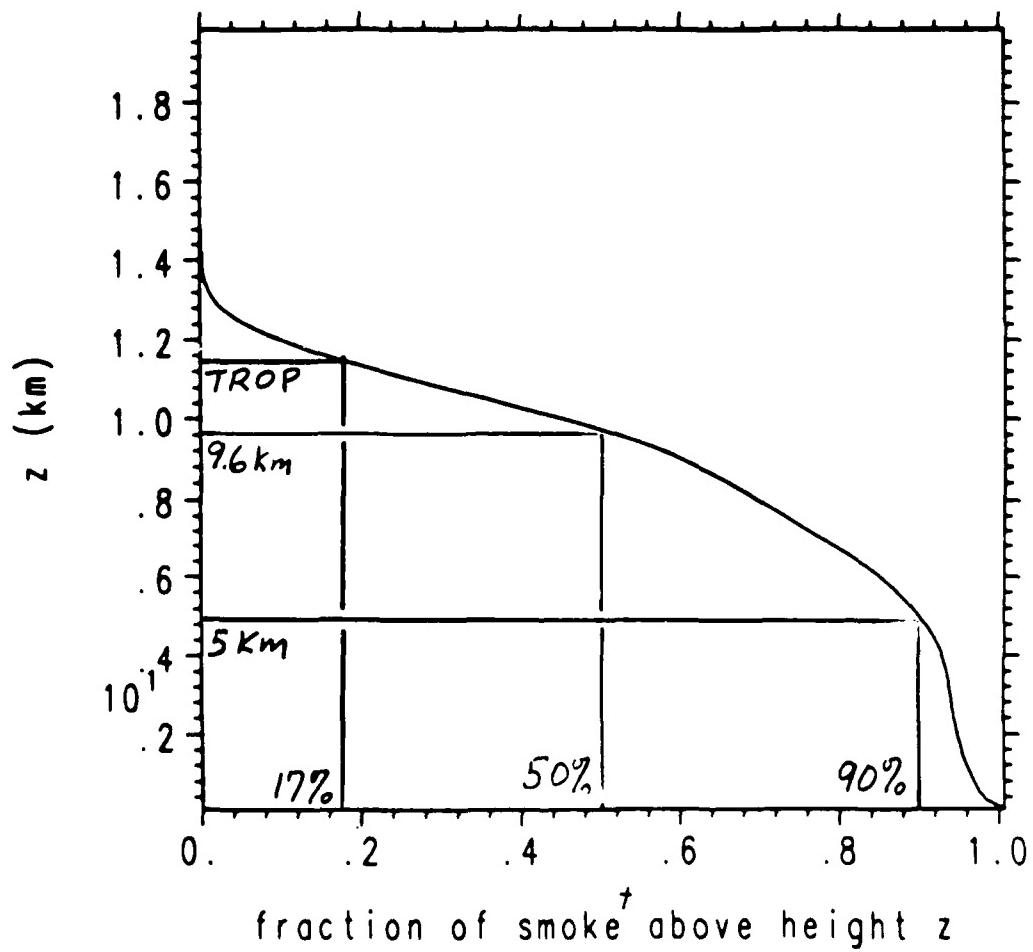


Figure 2

\* Top-downward integration of curve (d) on Figure 1.

<sup>t</sup> Fraction of smoke emitted from source. Conservation is good to within less than  $\frac{1}{2}\%$ .

### A Note on Linearization

Model uses this equations:

$$\frac{\partial u}{\partial t} = -c_p \Theta_m \frac{\partial \pi}{\partial x} + \dots$$

where:

$$\Theta_m = \Theta(1 + .61 q_v)(1 - q_c - q_r - q_s)$$
$$\Theta = \bar{\Theta}(z) + \theta' \quad q_v = \bar{q}_v(z) + q_v'$$

Some models linearize this to:

$$\frac{\partial u}{\partial t} = -c_p \bar{\Theta}_v \frac{\partial \pi}{\partial x} + \dots$$

where:

$$\bar{\Theta}_v = \bar{\Theta}(z) (1 + .61 \bar{q}_v(z))$$

At least for the benchmark case, there was no significant difference in the vertical smoke profile using the linearized form.

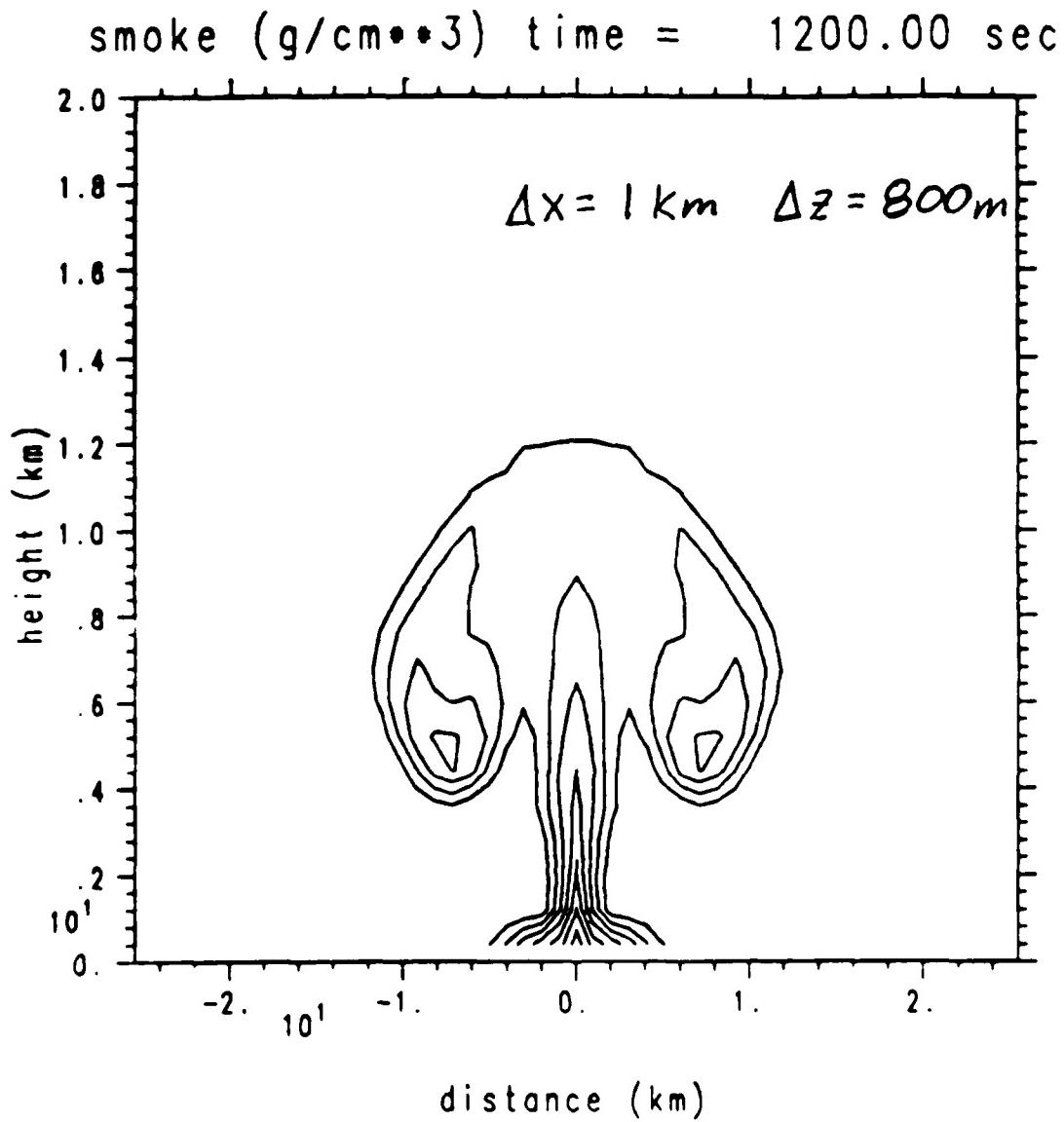


Figure 3. Smoke density contours for low resolution ( $\Delta x = 1000 \text{ m}$ ,  $\Delta z = 800 \text{ m}$ ) simulation. Contour interval is  $5 \times 10^{-10} \text{ g cm}^{-3}$ . Compare with Figure 4 and curve (b) in Figure 1.

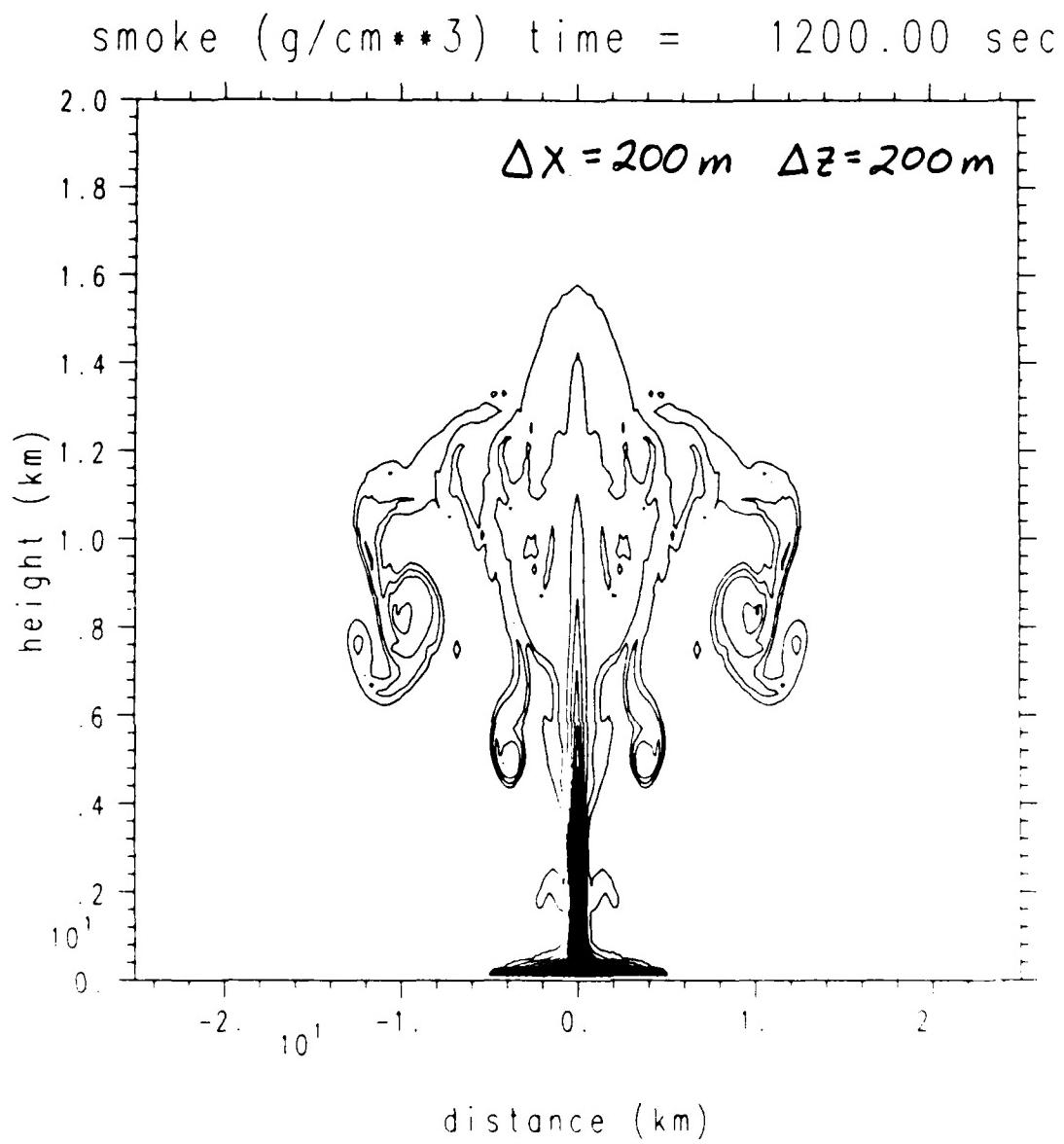


Figure 4. Smoke density contours for high resolution ( $200 \text{ m} \times 200 \text{ m}$ ) simulation. Contour interval is  $5 \times 10^{-10} \text{ g cm}^{-3}$ . In addition to exhibiting fine structure, the smoke plume rises higher than in the low resolution simulation. Smoke distribution is shown in Figure 1, curve (d) (at 1 hr).

## Sources of Differences

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- Equations
- Numerical methods
- Resolution, especially horizontal
- Turbulence parameterizations
- Atmospheric soundings

## **Future Research**

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- Condensation scavenging parameterization
- Interface with sophisticated condensation nucleation model
- Three-dimensional model

SECTION 3  
MESOCALE MODELING

# **REGIONAL-SCALE INTERACTIONS BETWEEN ATMOSPHERIC DYNAMICS AND AEROSOLS**

**Douglas L. Westphal**

*The Pennsylvania State Univ.,  
Department of Meteorology,  
Univ. Park, PA 16802*

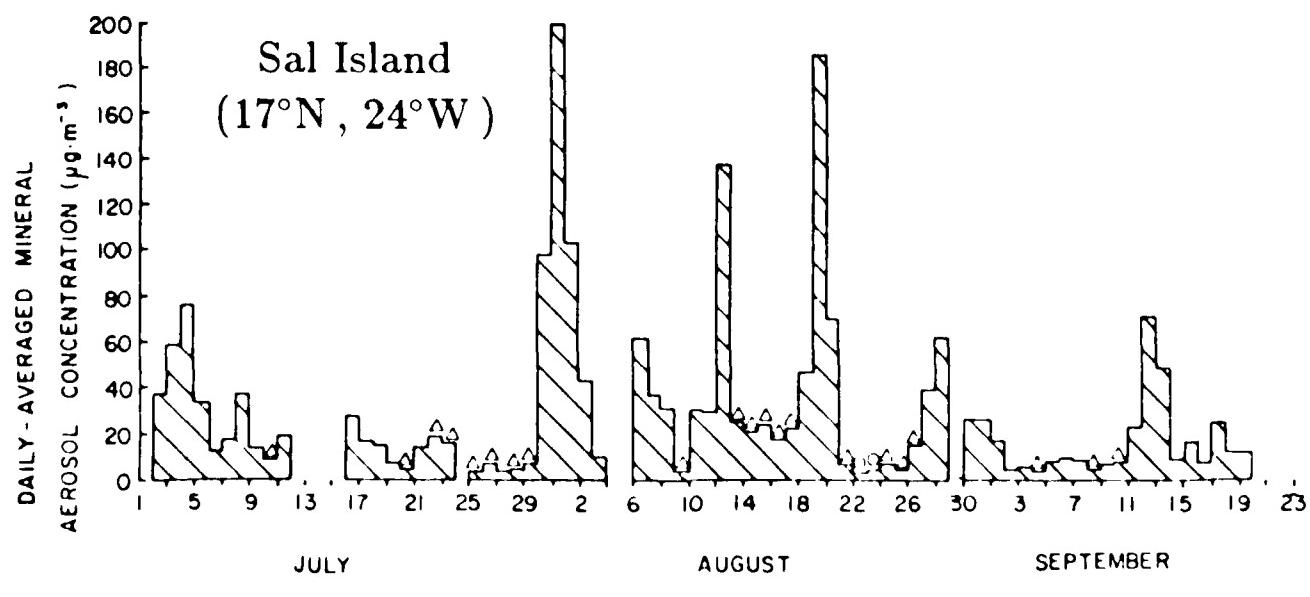
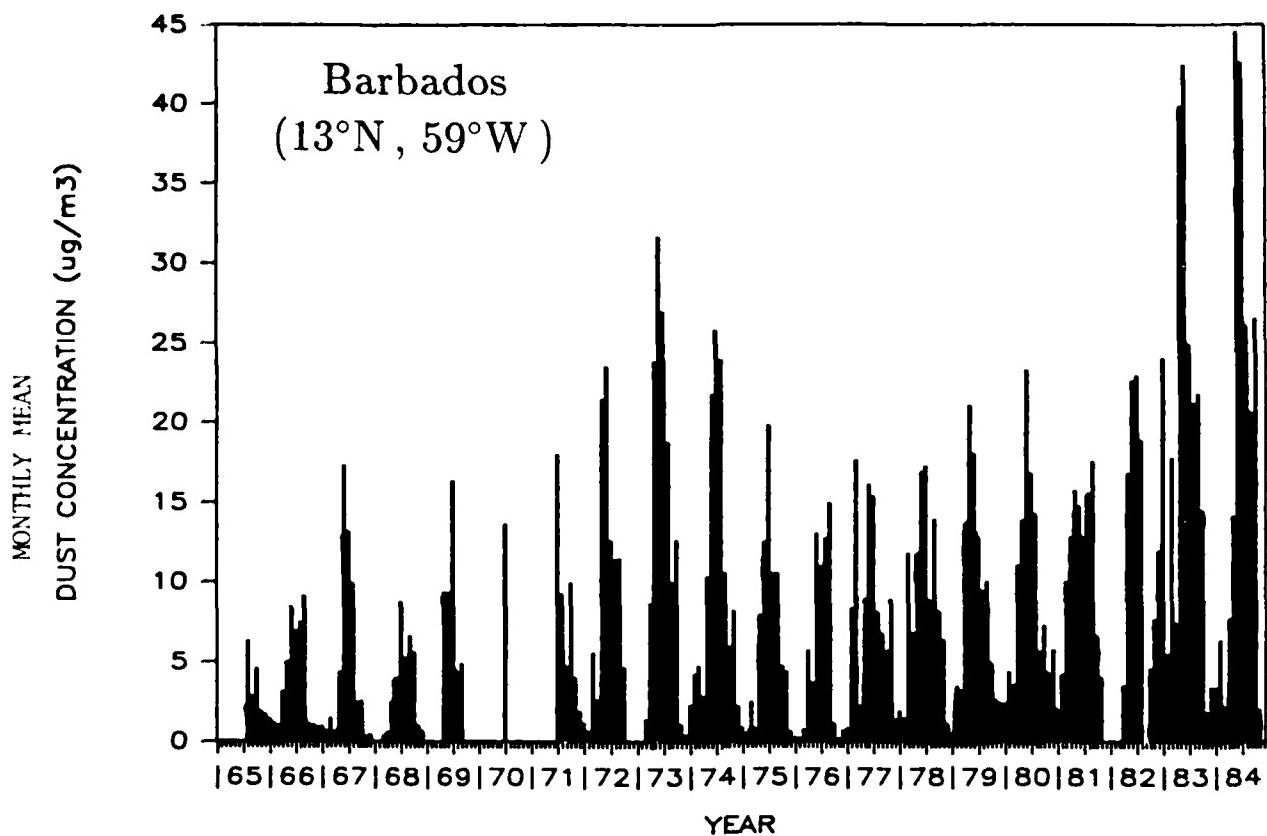
**and**

**Owen B. Toon**

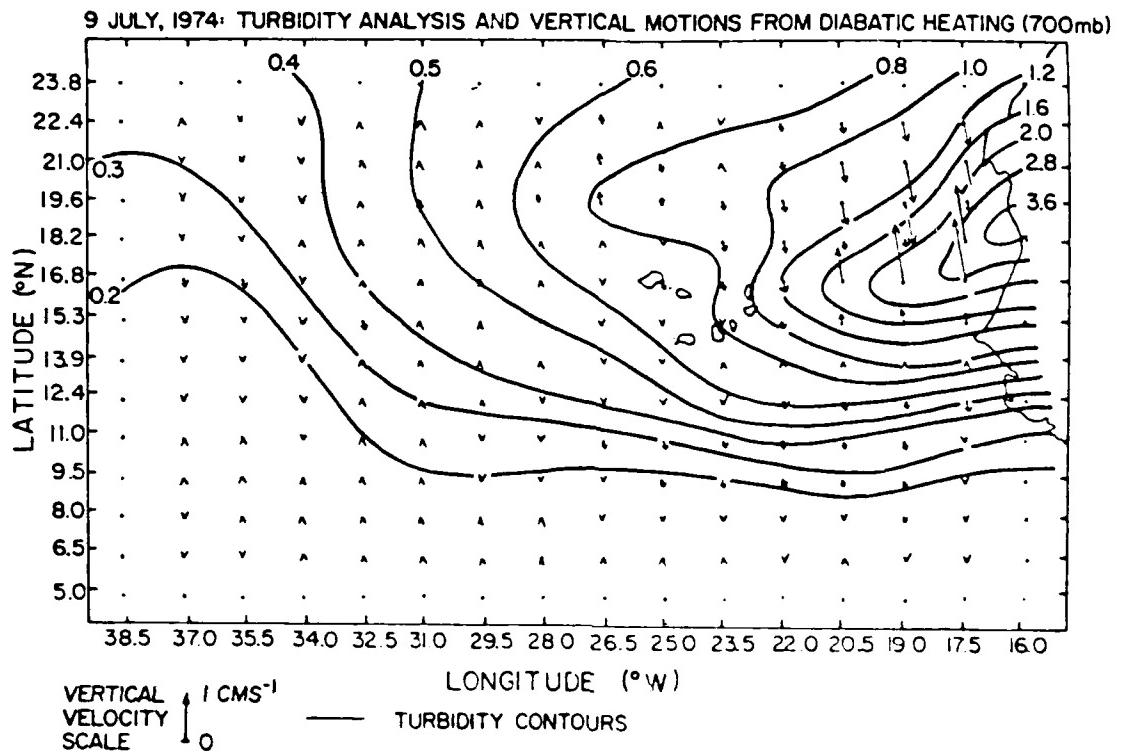
*NASA Ames Research Center,  
Space Sciences Division,  
Moffett Field, CA 94035*

## **NUCLEAR WINTER: MESOSCALE TOPICS**

- Vertical lofting of aerosol by the fire and cloud
- Coagulation and wet removal in the convective plume
- Dynamical perturbation by the fire and cloud
- Self-generated buoyancy of soot clouds; cloud formation in buoyant soot clouds; cycling of soot through cloud droplets
- Spread of individual plumes; formation of larger, “GCM-scale” clouds
- Influence of diurnal radiative forcing on the response of the atmosphere
- Influence of geographical location and season on the above items



1974



Turbidity isopleths determined from satellite irradiances and diagnosed 700 mb vertical motions associated with radiative heating of dust for 1200 GMT, July 9, 1974 (Karyampudi, 1979).

## **NUMERICAL MODELS**

**Atmospheric** — Penn State/NCAR Limited-Area Model

- Primitive equations, hydrostatic
- Predicted ground temperature
- Subgrid-scale vertical fluxes
- Precipitation: Large-scale and subgrid-scale (cumulus) precipitation

**Aerosol** — Ames Research Center's Aerosol Model

- Continuity equation for each size aerosol and each type of aerosol
- Dry and wet removal
- Coagulation
- Condensation
- Surface or gridpoint source, or initial cloud
- Radiation

## SCAVENGING PARAMETERIZATION

Scavenging rate  $\Lambda$  in ( $s^{-1}$ ) is given by

$$\begin{aligned}\Lambda &= 4.2 \times 10^{-4} E R^{0.79} & R^{0.63} \leq 7000/EH \\ \Lambda &= 3R^{0.16}/H & R^{0.63} > 7000/EH\end{aligned}$$

where:

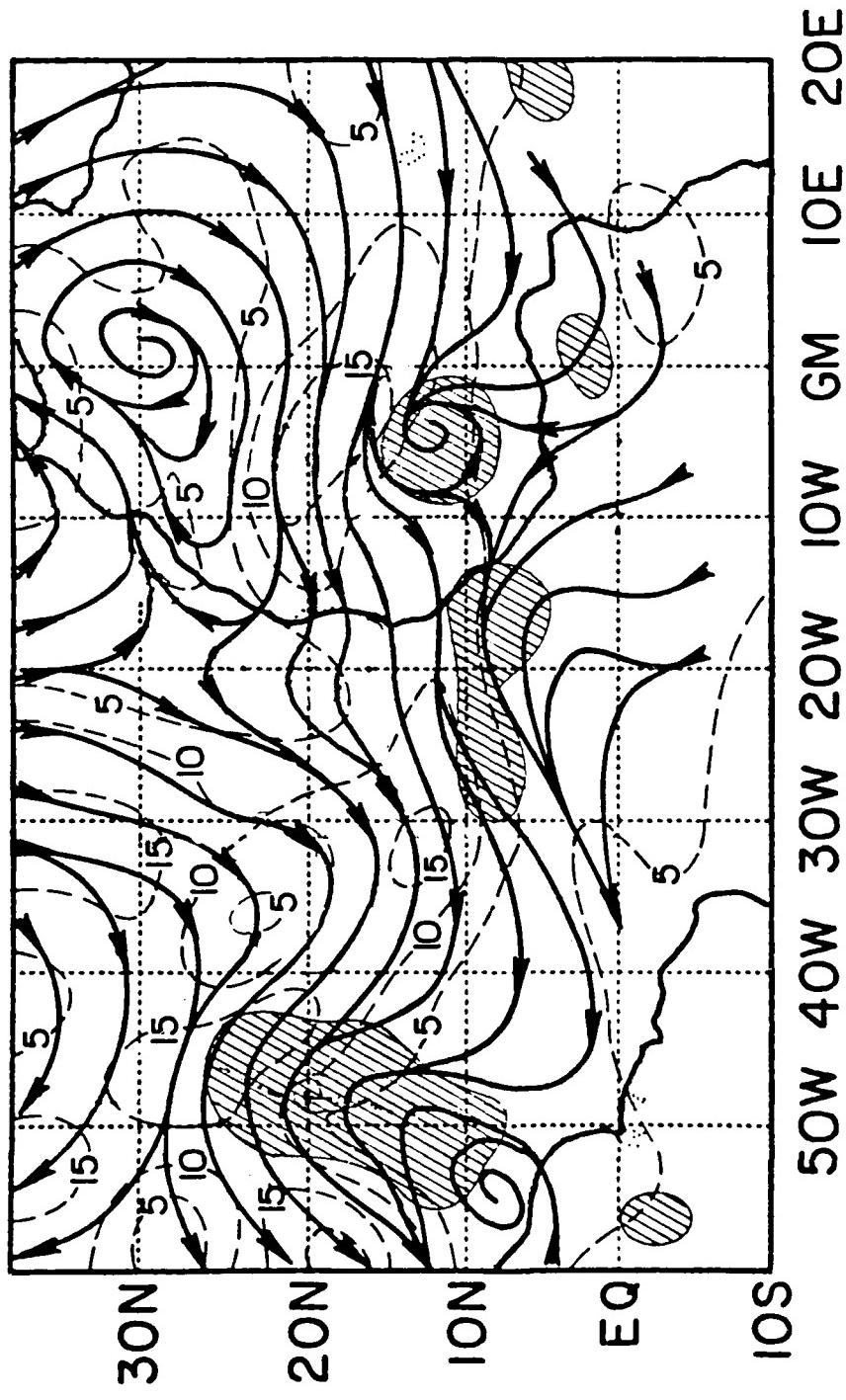
- $E$  is the collection efficiency, 0.83
- $H$  the depth of the precipitating cloud (m)
- $R$  the rainfall rate ( $mm\ hr^{-1}$ )

- Marshall-Palmer droplet size distribution assumed
- No dependence on aerosol size
- Non-convective rainfall  $R_n$  is due to supersaturation in the grid box
- Convective rainfall  $R_c$  and the fraction  $a_p$  covered by convection are determined by a cumulus cloud parameterization

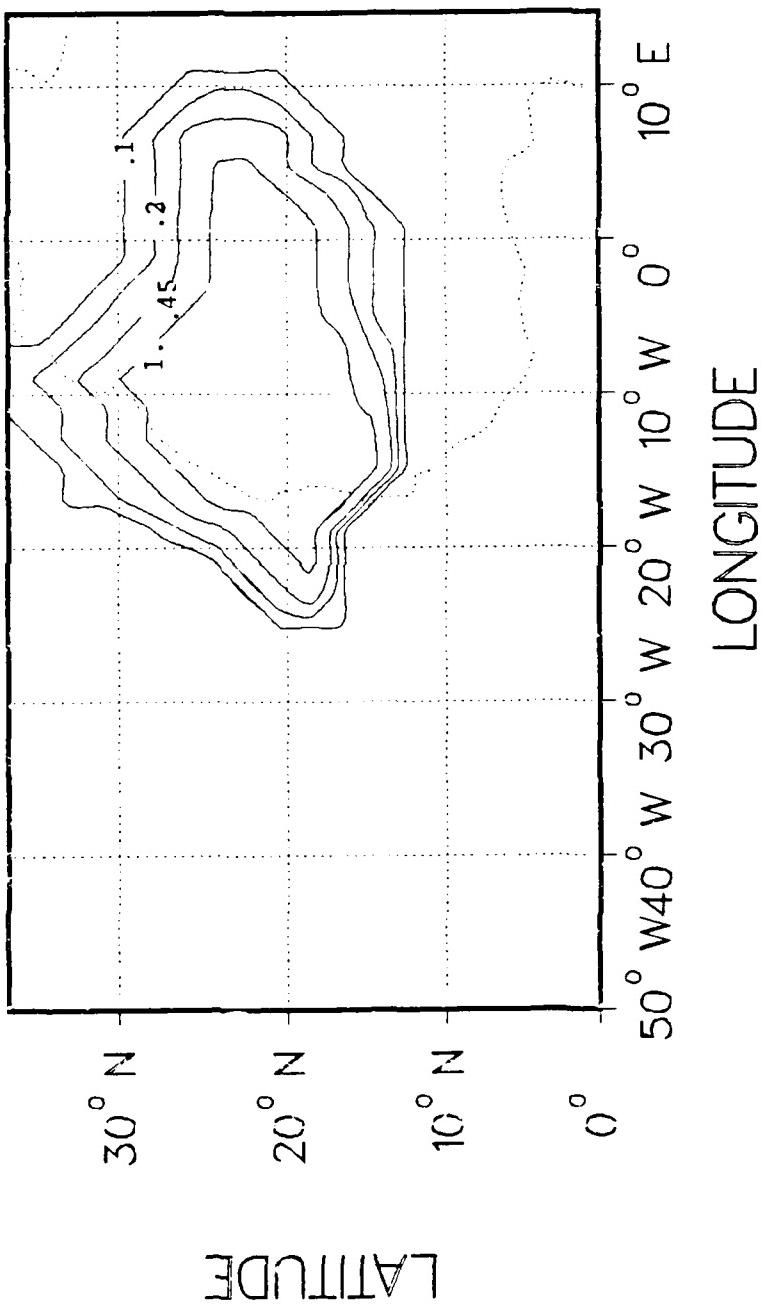
The total scavenging rate for a grid box is

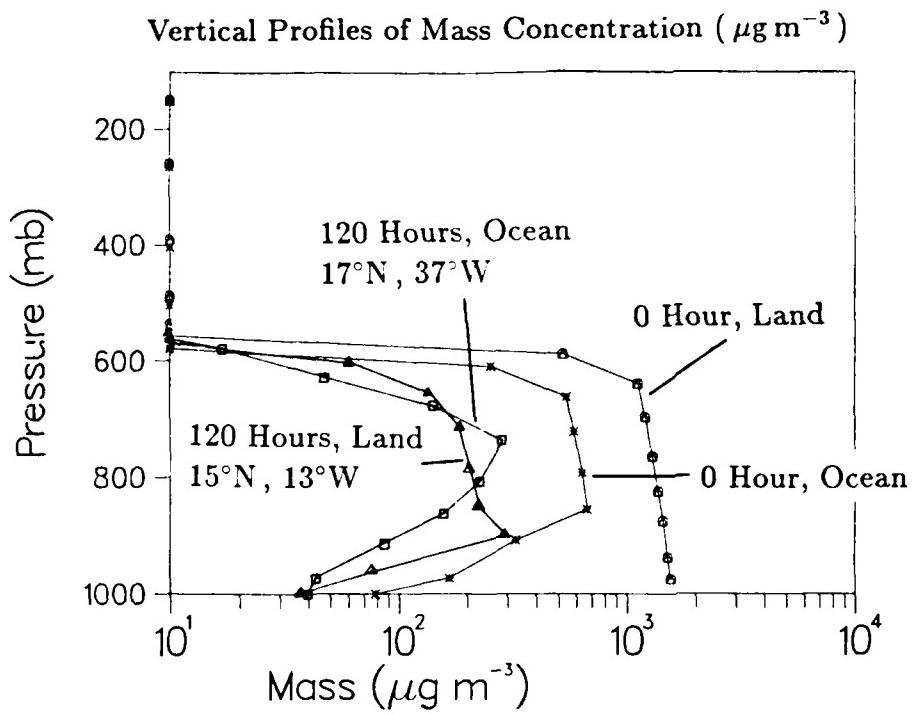
$$\Lambda = (1 - a_p)\Lambda(R_n) + a_p\Lambda(R_n + R_c)$$

120 hr. simulated 700 mb streamlines (solid), isotachs (dashed, in  $\text{ms}^{-1}$ ), and areas of precipitation (cross-hatched).

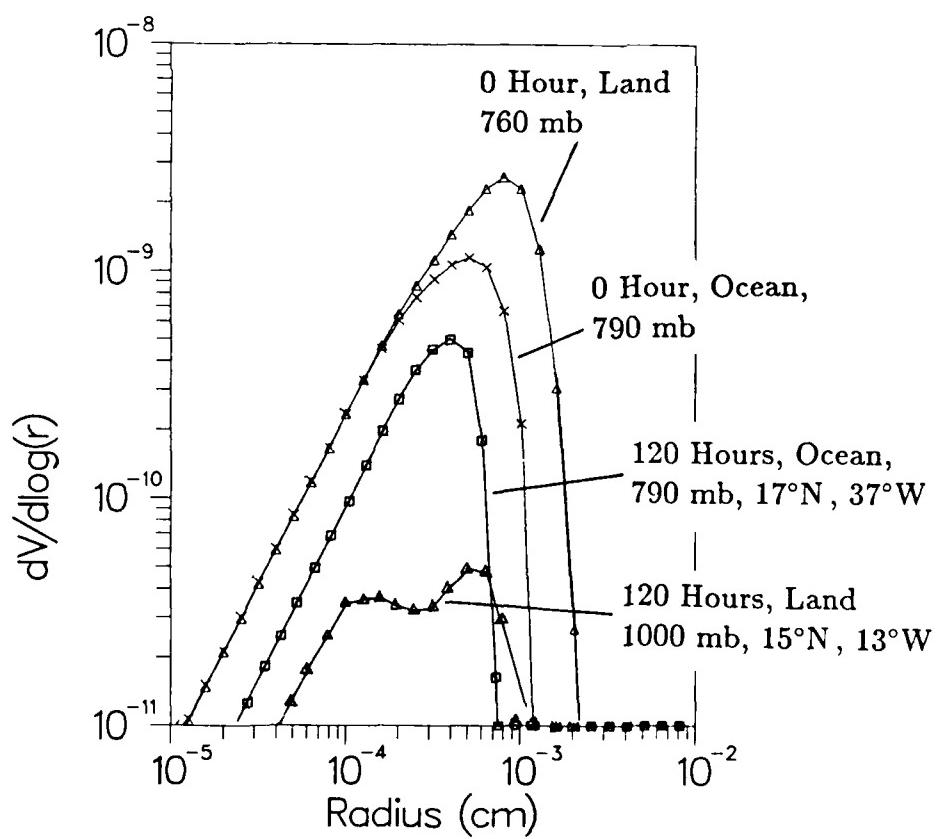


0 hr. Optical Depth, 23 Mt

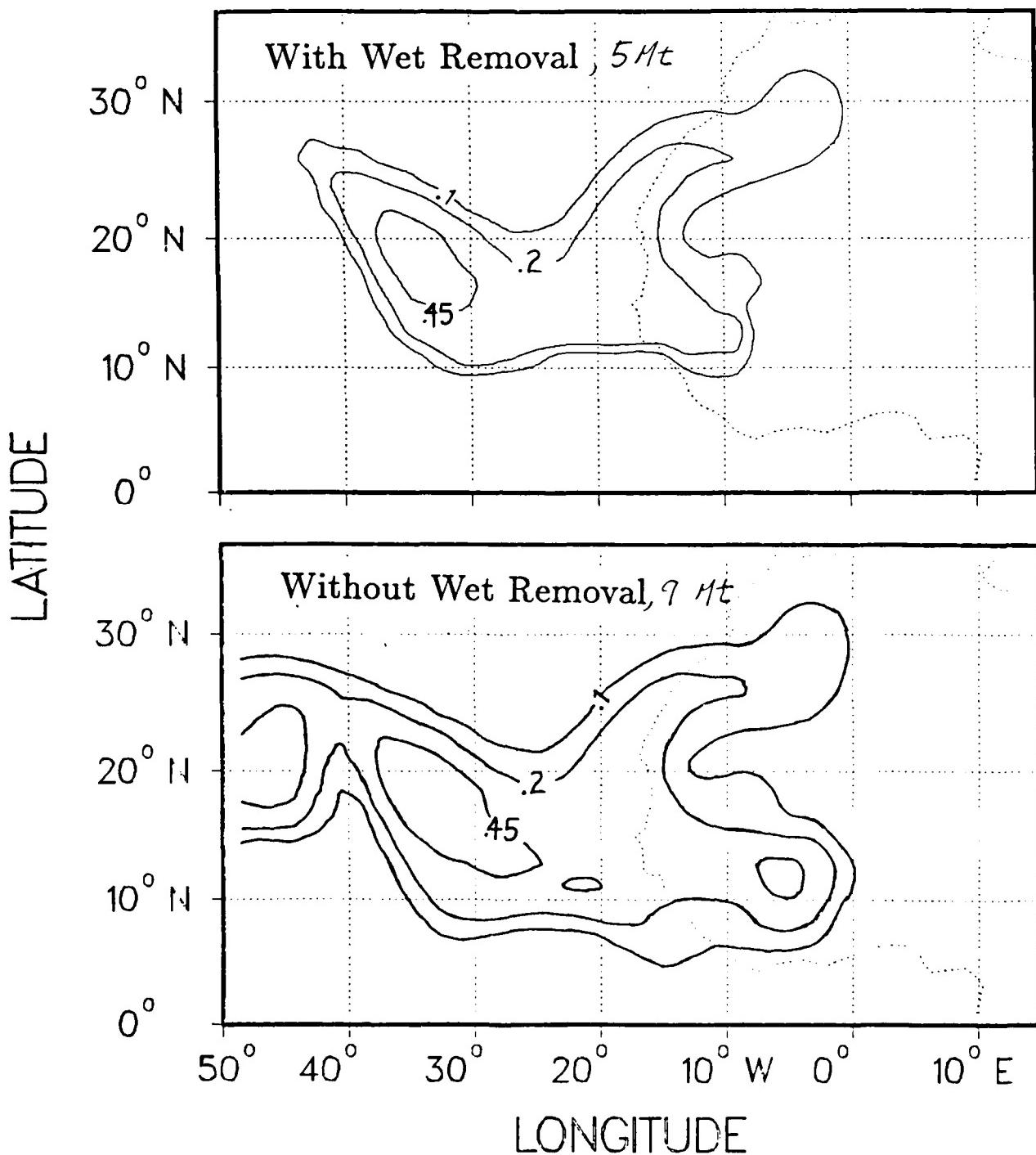




Size Distributions by Volume



120.1 Hours  
Optical Depth



## FUTURE WORK

- Include radiative heating in the Saharan dust simulations
- Conduct tests to determine the resolution requirements of soot cloud modeling
- Use a non-hydrostatic model to develop an input dataset of the initial spatial and size distribution of a soot cloud
- Carry out experiments that address the outstanding problems of nuclear winter

**Algorithms for Computations of Aerosol Physics  
and Radiative Transfer**

Owen B. Toon  
NASA Ames Research Center  
Space Sciences Division  
Moffett Field, CA 94035

## **Overview of Ames Program**

### **I Develop Natural Analogs to Nuclear Winter**

**A. Saharan Dust Storms**

**B. Volcanic Clouds**

### **II Determine Fuel Loading**

### **III Develop Algorithms for Aerosol Physics**

### **IV Perform Regional Scale Modeling**

## **Algorithms for Aerosol Physics**

**I Aerosol Microphysics and Transport**

**II Radiative Transfer in Aerosols**

**III Interactions Between Aerosols and Clouds**

## **Aerosol Microphysics and Transport**

### **I General Properties**

**A. In Arbitrary Coordinates So Can Be Used  
In A Variety of Dynamical Models**

**B. Description in Preparation**

### **II Transport**

**A Timesplit So Can Be Used in 1,2,or 3-D Models**

**B. Finite Elements in Horizontal, Curve Fitting  
Scheme in Vertical. Exactly Conserve Mass, Have  
Little Numerical Diffusion, No Negative Numbers,  
Numerically Efficient**

### **III Microphysics**

**A. Resolves Size Distribution**

**B. Includes: Sedimentation, Coagulation, Multiple  
Aerosol Constituents, Condensational Growth,  
Gases**

$$\frac{\partial c}{\partial t} + \nabla_z \cdot c_z v_z + \frac{\partial}{\partial z} (w_z + v_{fall}) c_z - \nabla_z \cdot g^k H \cdot \nabla c_z / S$$

$$- \frac{\partial}{\partial z} g^k c_z \frac{\partial}{\partial z} (c / \beta) = (P - L) \quad (2)$$

$$\frac{\partial c}{\partial t} + \frac{\partial u c}{\partial x_1} + \frac{\partial v c}{\partial x_2} + \frac{\partial w c}{\partial x_3} - \frac{\partial}{\partial x_1} g^* k_1 \frac{\partial c / g^*}{\partial x_1}$$

$$\frac{\partial}{\partial x_2} g^* k_2 \frac{\partial c / g^*}{\partial x_2} - \frac{\partial}{\partial x_3} g^* k_3 \frac{\partial c / g^*}{\partial x_3} = (P - L) v_m' H_{m1} H_{m2} \quad (1)$$

$$\frac{\partial C_z}{\partial t} + \nabla_z \cdot C_z V_z + \frac{\partial}{\partial z} (w_z + v_{fall}) C_z - \nabla_z \cdot g K_H \cdot \nabla C_z / S \\ - \frac{\partial}{\partial z} g K_z \frac{\partial}{\partial z} (C / g) = (P - L) \quad (2)$$

$$\frac{\partial C}{\partial t} + \frac{\partial u C}{\partial x_1} + \frac{\partial v C}{\partial x_2} + \frac{\partial w C}{\partial x_3} - \frac{\partial}{\partial x_1} g^* K_1 \frac{\partial C}{\partial x_1} \\ - \frac{\partial}{\partial x_2} g^* K_2 \frac{\partial C}{\partial x_2} - \frac{\partial}{\partial x_3} g^* K_3 \frac{\partial C}{\partial x_3} = (P - L) v_m h_{m1} h_{m2} \quad (1)$$

TABLE 1  
Conversion scaling factors for various coordinate systems

Coordinate Systems	$H_{m_1}$	$H_{m_2}$	$H_{m_3}$	$ds_1$	$ds_2$	$V_m$	$ds_3$
Spherical Coordinates	$\sin \phi$	1	a	$d\lambda$	$d\phi$		
Mercator	$\frac{\cos \phi}{\cos \phi_0}$	$\frac{\cos \phi}{\cos \phi_0}$	1	$dx$	$dy$		
Polar Stereographic	$\left(\frac{1+\sin \phi}{2}\right)^{-1}$	$\left(\frac{1+\sin \phi}{2}\right)^{-1}$	1	$dx$	$dy$		
Lambert Conformal	$\frac{(1+\sin \phi) \cos \phi_0}{(1+\sin \phi_0) \cos \phi}$	$\frac{(1+\sin \phi) \cos \phi_0}{(1+\sin \phi_0) \cos \phi}$	1	$dx$	$dy$		
Rectangular	1	1	1	$dx$	$dy$	1	$dz$
Pressure						$1/g g$	$dp$
Sigma = $P - P_{top} = P - P_{top}$ $P_{bottom} - P_{top} P^*$						$P^*/g g$	$d\sigma$
Log Pressure $Z = H \ln (P/P_{sfc})$						$T/T_0$	$dz$

$$U = U_c/H_{m_1} V = V_c/H_{m_2} W = ds_3/dt + V_{fall} V_m^{-1} C = C_z V_m \alpha \alpha_2 H_{m_1} H_{m_2}$$

$$Kx_1 = 1/H_{m_1}^2 K, \quad Kx_2 = 1/H_{m_2}^2 K, \quad Kx_3 = V_m^{-2} K_z$$

$$S^* = S V_m H_{m_1} H_{m_2} \quad dx_1 = H_{m_3} ds_1 \quad dx_2 = H_{m_3} ds_2 \quad dx_3 = ds_3$$

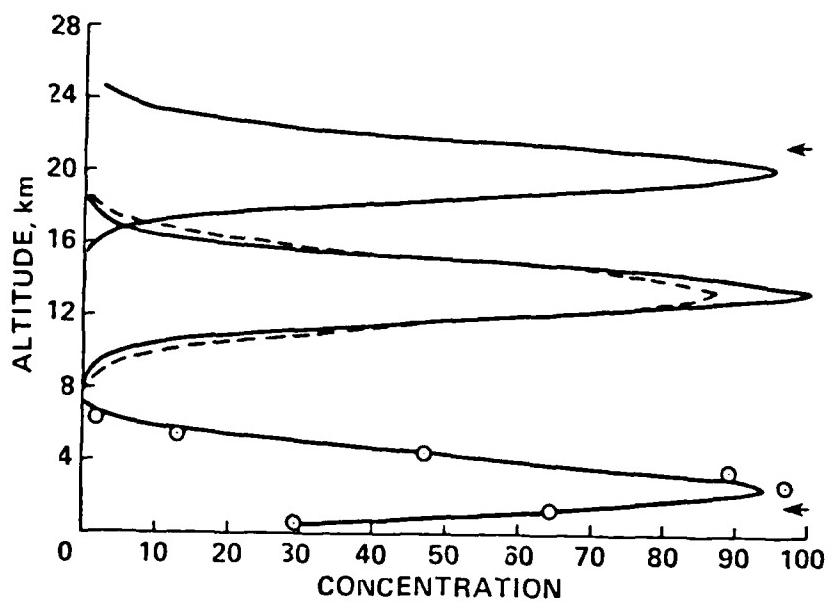
$\bar{\phi}$  = colatitude

$\phi$  = latitude at which projection is true

$\lambda$  = longitude  $n$  = constant which can be adjusted to make the projection true at a second latitude  $\phi_1$

$$g = \text{air density} \quad n = (k(k-1))^{-1} \quad k = \ln \left( \frac{\cos \phi_0}{\cos \phi} \right) \div \ln \left( \frac{\tan(\pi/4 - \phi_0/2)}{\tan(\pi/4 - \phi_1/2)} \right)$$

$$g = \text{gravitational acceleration} \quad H = \text{scale height} = R T_0 / g$$



AD-A185 150

TECHNICAL PAPERS PRESENTED AT THE DEFENSE NUCLEAR  
AGENCY GLOBAL EFFECTS R. (U) DOD NUCLEAR INFORMATION  
AND ANALYSIS CENTER SANTA BARBARA CA. 15 MAY 86

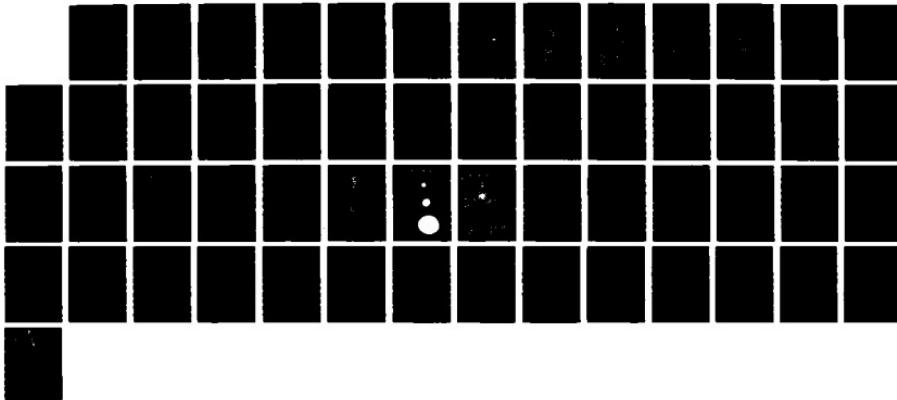
4/4

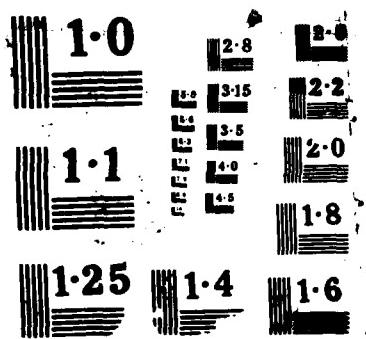
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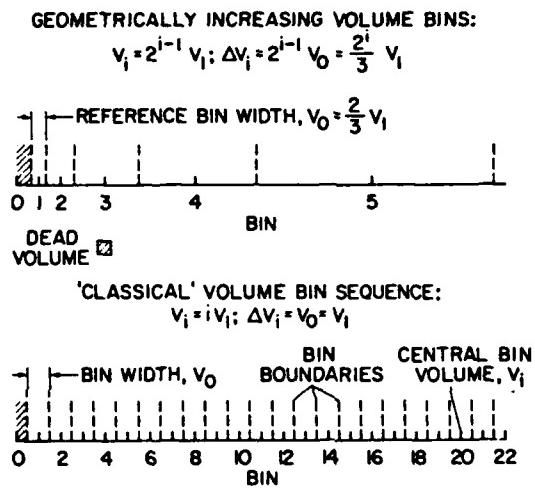
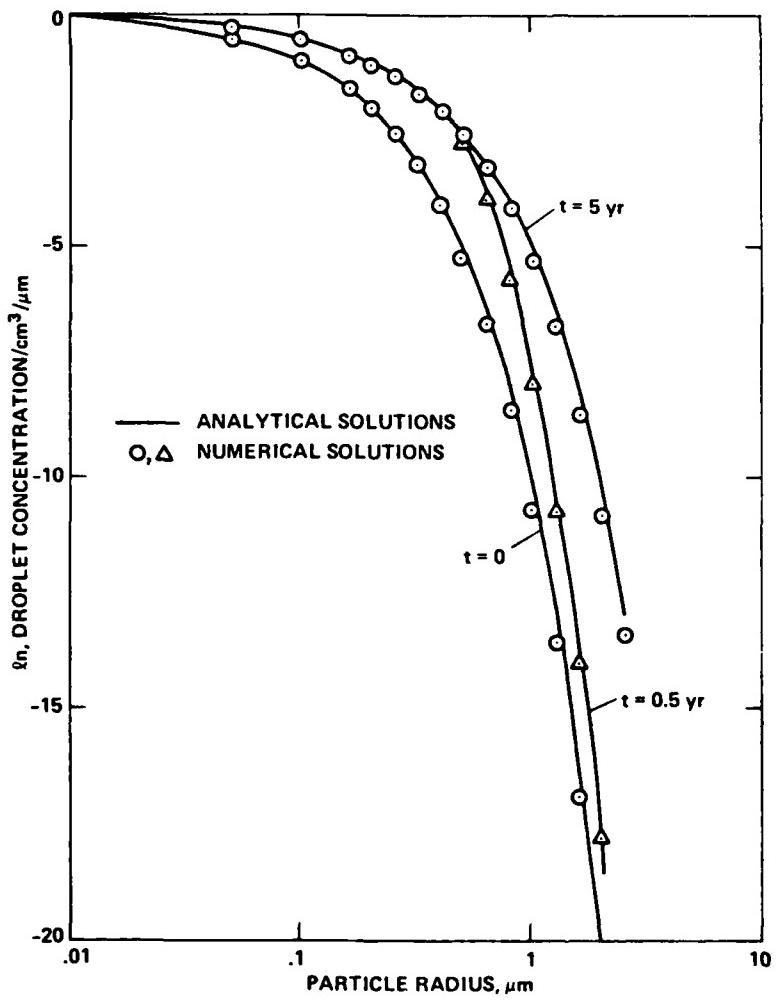


Figure 7.- Aerosol particle volume bin structures for numerical models.



**Figure 8.-** Aerosol droplet growth. The evolution of a droplet size distribution following a change in the droplet growth rate is shown. Both analytical and numerical solutions are given. The conditions for the calculation are described in the text.

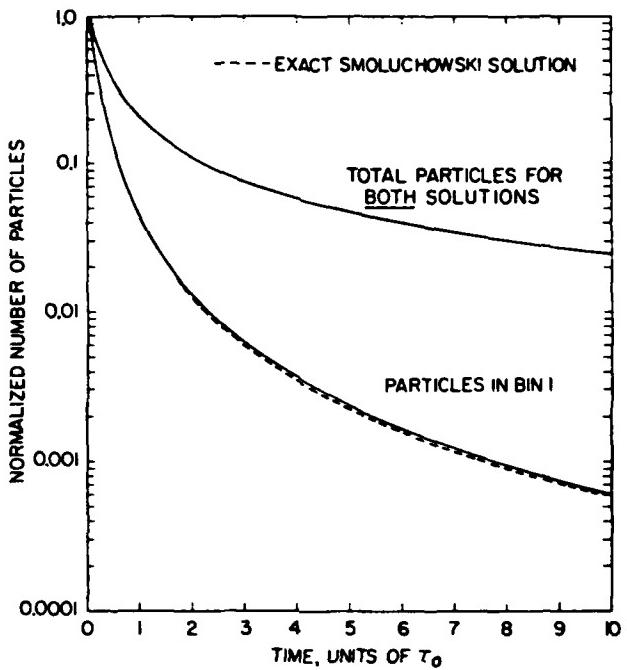


Figure 9.- Aerosol coagulation. Calculated transient decay curves are shown for the number of particles in the smallest model size bin, and for the total number of particles of all sizes, when one particle is initially placed in the smallest model bin. Time is measured relative to the characteristic coagulation time,  $\tau_0$ . The exact solutions of Smoluchowski are shown for comparison.

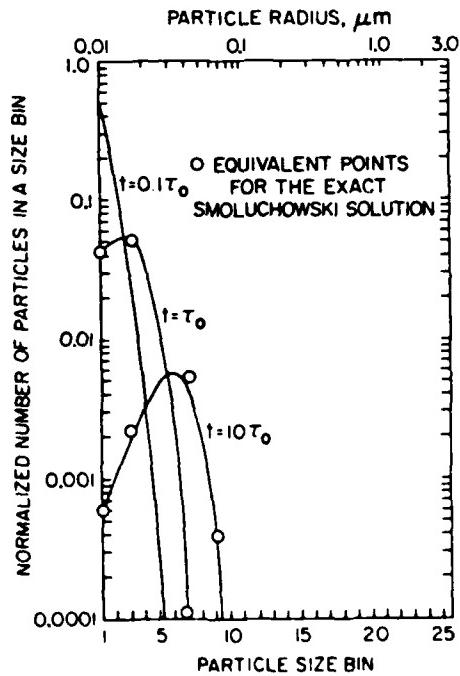


Figure 10.- The development of an aerosol size spectrum by coagulation. The conditions for the calculation are the same as for figure 9.

## **Radiative Transfer**

### **I Solar Wavelengths**

**Use Two-Stream Approximation for Scattering with Tridiagonal Solver. Numerically Efficient, Accurate to 10% or Better for Heating Rates.**

### **II Infrared Wavelengths**

**Use Two Stream Source Function Technique for Scattering. Numerically Efficient, Exact in Limit of No Scattering, 10% Accuracy with Scattering.**

### **III Status**

**A. Algorithms Developed and Checked for Speed and Accuracy**

**B. Absorption Coefficients Under Development  
Based on McClatchey tape.**

$$\mu \frac{\partial I_\nu}{\partial \tau}(\tau_\nu, \mu, \phi) = I_\nu(\tau_\nu, \mu, \phi) - S_\nu(\tau_\nu, \mu, \phi) - \frac{\omega_0 \nu}{4\pi} \int_0^\pi \int_0^{2\pi} P_\nu(\mu, \mu l, \phi, \phi l) I_\nu(\nu, \mu l, \phi l) d\mu l d\phi l. \quad (1)$$

$$\frac{\partial F_n^+}{\partial \tau_n} = \gamma_{1n} F_n^+ - \gamma_{2n} F_n^- - S_{+n} \quad (10)$$

$$\frac{\partial F_n^-}{\partial \tau_n} = \gamma_{2n} F_n^+ - \gamma_{1n} F_n^- - S_{-n} \quad (11)$$

$$F_n^+ = k_{1n} \exp(\lambda_n \tau) + \Gamma_n k_{2n} \exp(-\lambda_n \tau) + C_n^+(\tau) \quad (18)$$

$$F_n^- = \Gamma_n k_{1n} \exp(\lambda_n \tau) + k_{2n} \exp(-\lambda_n \tau) + C_n^-(\tau). \quad (19)$$

Mesoscale Modeling of Coastal Flows

During Periods of  
Extended Solar Obscuration

by

*Charles R. Molenkamp*



**Basic Question:**

**If a layer of smoke obscures the ground, will a temperature difference develop between land and sea that could modify the flow field so that the likelihood of precipitation would be increased?**

**Preliminary Answer:**

**Not in the first few days after solar obscuration.**

## **Approach:**

**Numerical Simulation Using the Colorado State University  
Mesoscale Model (Pielke, et. al)**

- CSU Mesoscale Model

- Hydrostatic incompressible flow
- Radiational and surface heat budget
- Surface flux, planetary boundary, and synoptic layers
- Vertical diffusion depends on atmospheric stability

**Procedure:**

**Run CSU Mesoscale Model**

- 4 hour initialization
- 24 hour normal diurnal cycle starting at dawn (05:18)
- 48 hours with no solar incident radiation

**Initial Conditions**

- West Coast with 5 m/s synoptic flow from West
- Sounding for the Oregon Coast on 23 Aug 72
- Sea Surface Temperature is 289 K

## **CSU Mesoscale Model -- Radiation Calculation**

- Solar flux absorbed in the atmosphere by water vapor
- Longwave flux divergence includes absorption, emission by water vapor and CO<sub>2</sub>
- Surface temperature calculated by a heat balance equation
  - incoming solar radiation
  - incoming longwave radiation
  - latent heat flux
  - sensible heat flux
  - soil heat flux
  - outgoing surface longwave radiation

## **MODIFICATIONS TO THE CSU MESOSCALE MODEL**

- CLOUDS

- When the atmosphere is saturated any excess water vapor is assumed to exist as cloud water (fog)
- Evaporative heating/cooling changes temperature

- RADIATION (Long Wave)

## MODIFICATIONS TO THE CSU MESOSCALE MODEL

- CLOUDS

- RADIATION (Long Wave)

- New model for clear air long wave radiation  
The new version calculates fluxes at each level  
The old version assumed for each level that the entire atmosphere had the temperature of that level
- Effect of clouds on long wave radiation  
The upward and downward flux out of each cloudy layer is given by:

$$F_{out} = (1 - \epsilon) F_{in} + \epsilon \sigma T^4$$

where  $\epsilon$  is the emissivity of the layer  
and  $T$  is the temperature at the outgoing edge of the layer

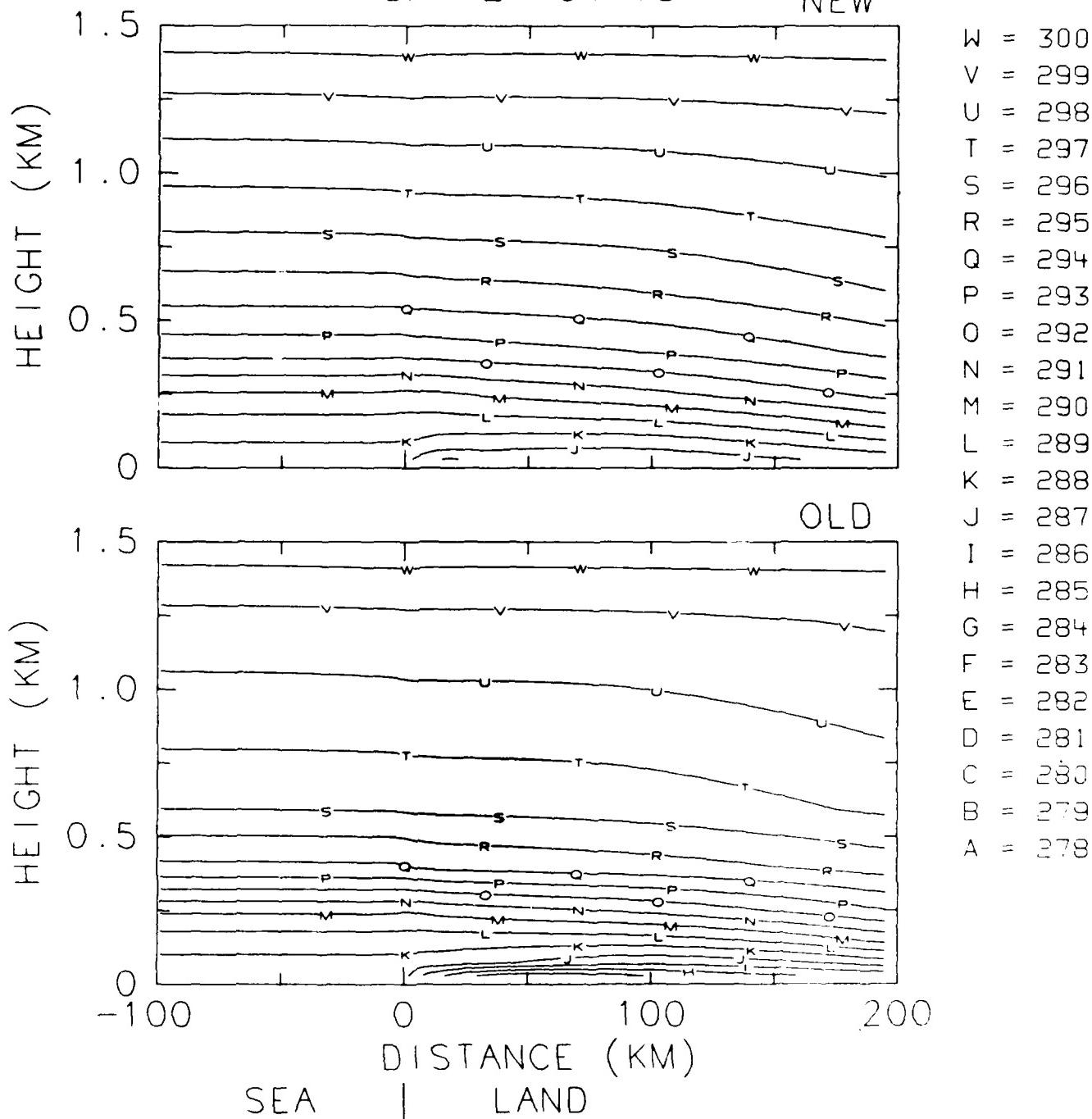
## **2 Runs of the Model**

- OLD
  - no clouds
  - old radiation parameterization
- NEW
  - clouds where atmosphere is saturated
  - new radiation parameterization, including the effect of clouds

# POTENTIAL TEMPERATURE

DAY 2 04:18

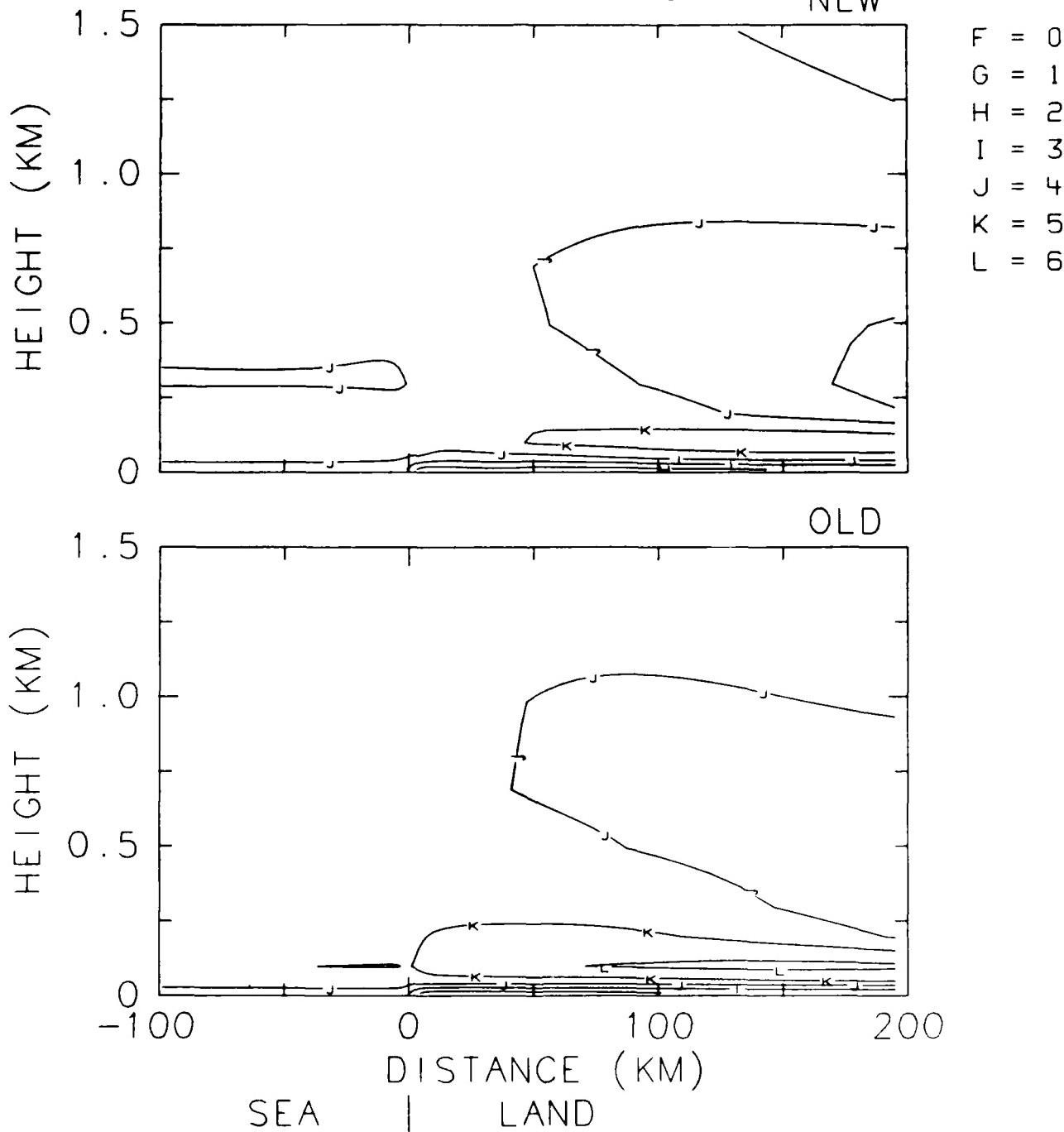
NEW



EAST-WEST VELOCITY (M/S)

DAY 2 04:18

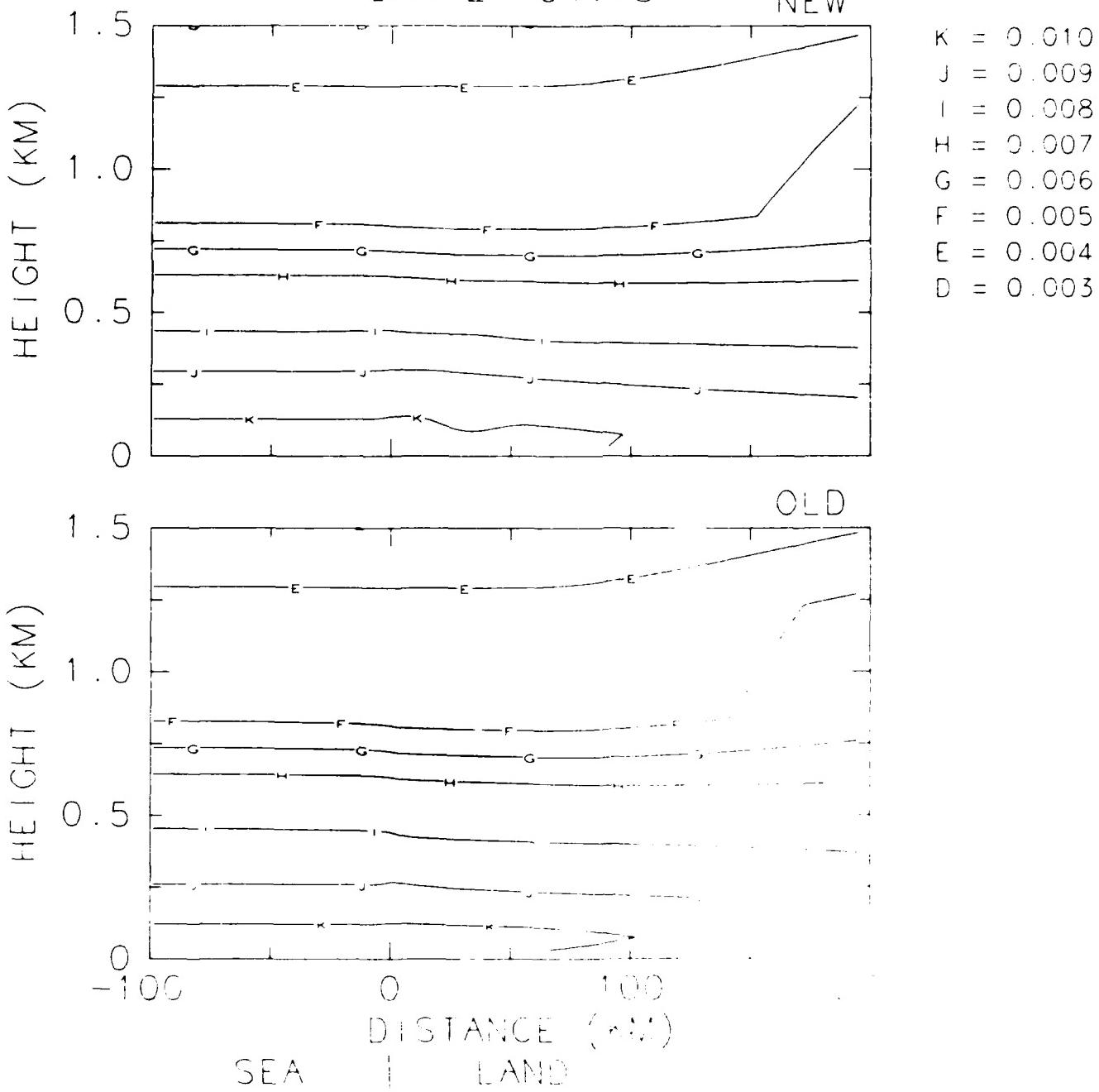
NEW

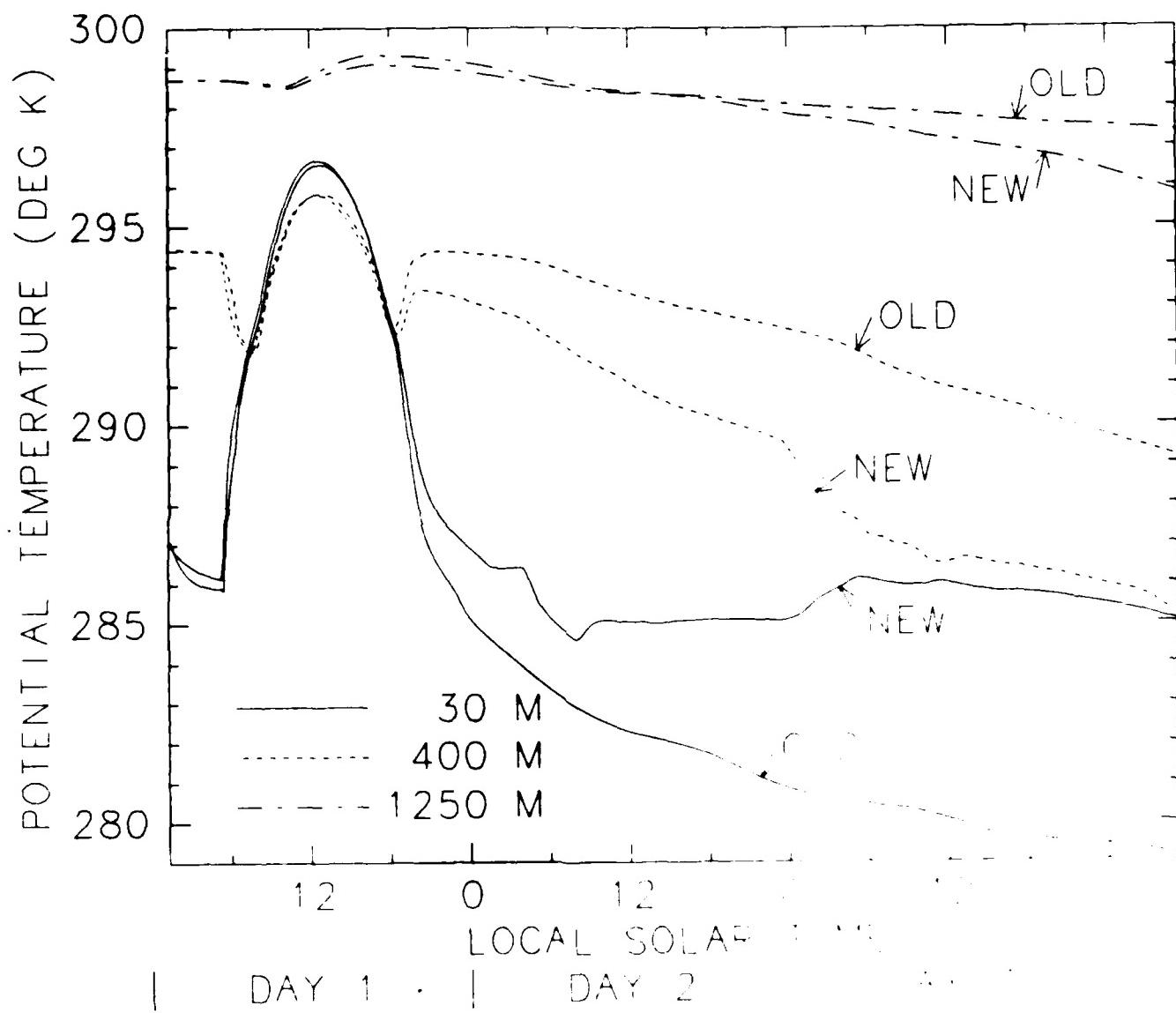


SPECIFIC HUMIDITY (G/G)

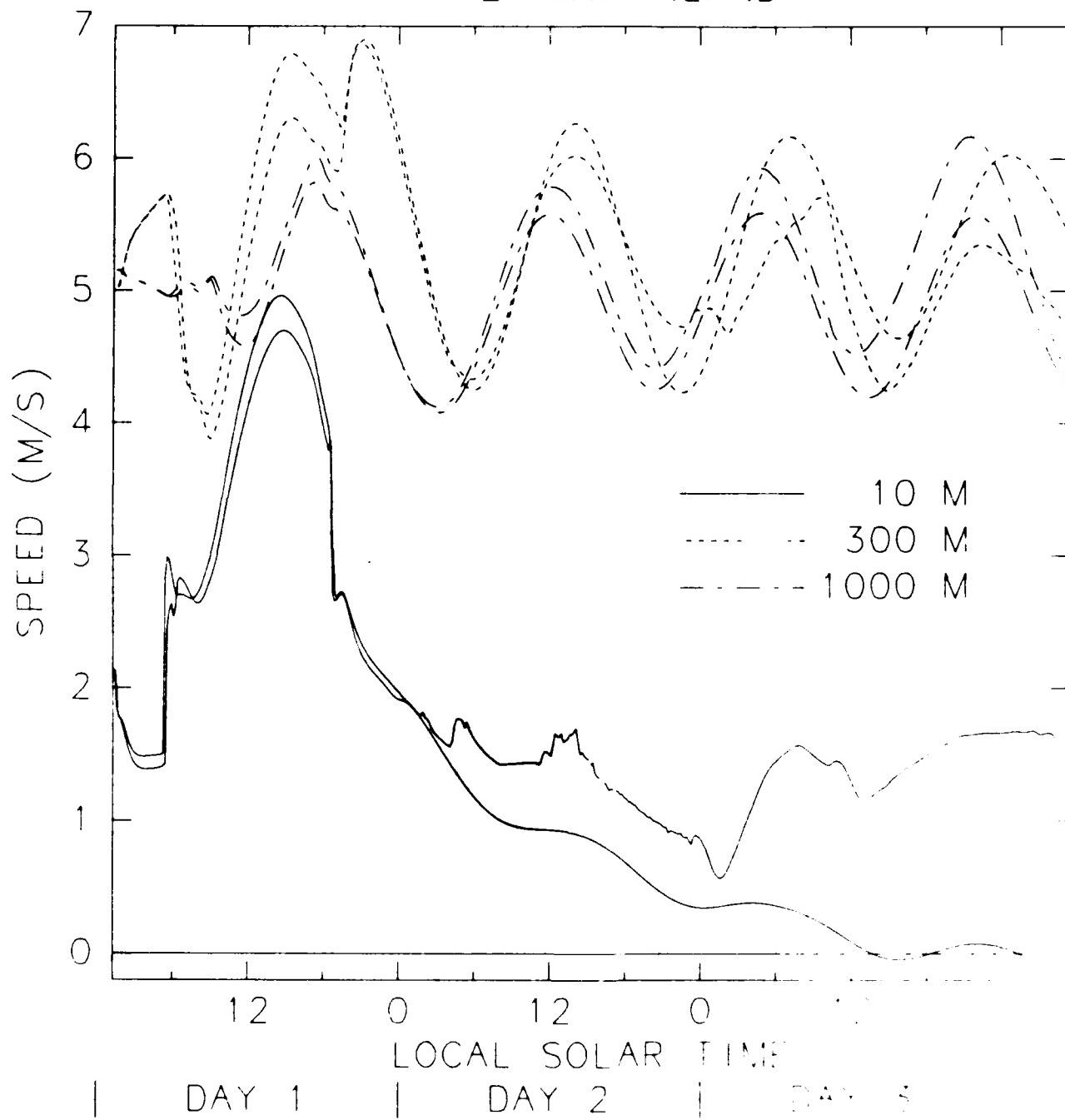
DAY 2 04:18

NEW

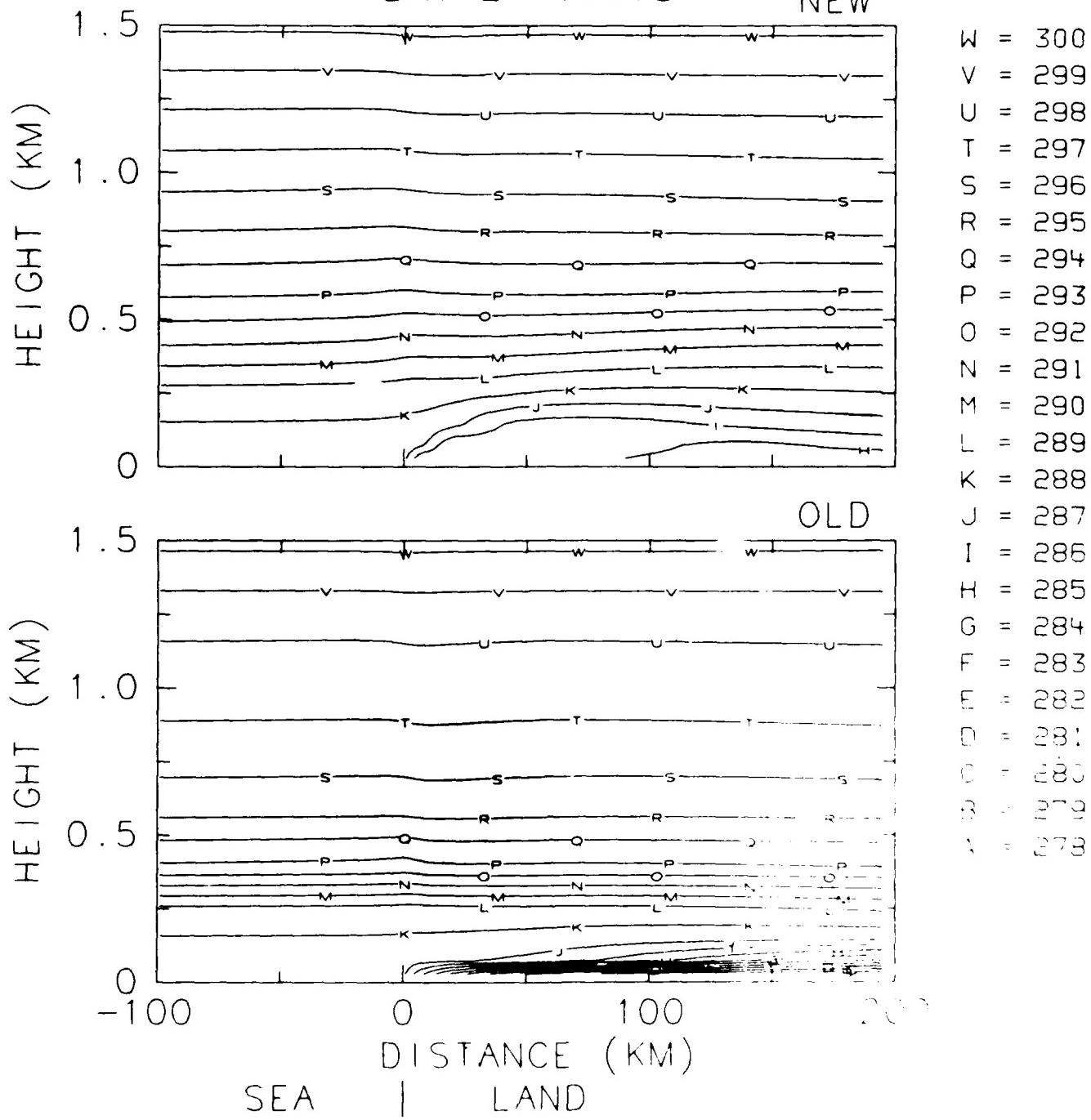


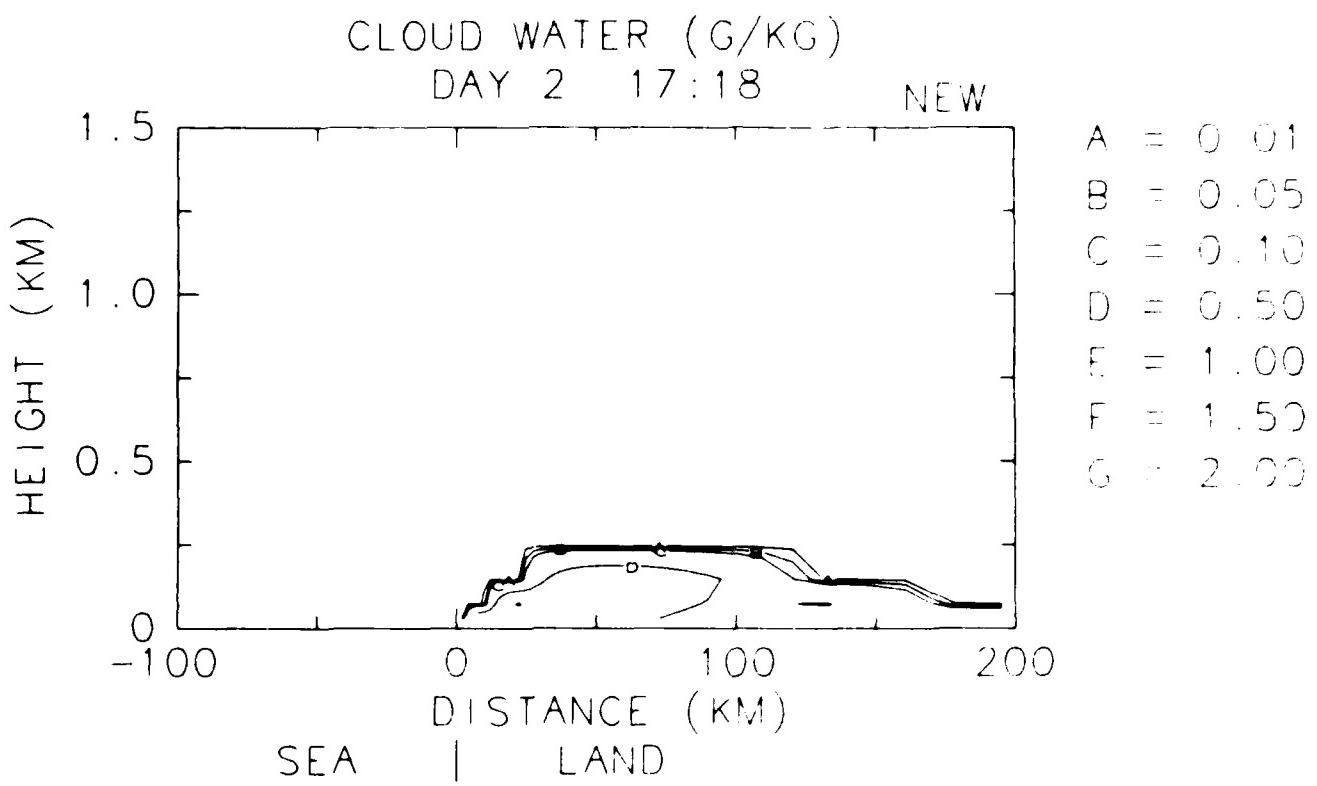


WIND NORMAL TO COASTLINE  
AT 27 KM INLAND



POTENTIAL TEMPERATURE  
DAY 2 17:18 NEW

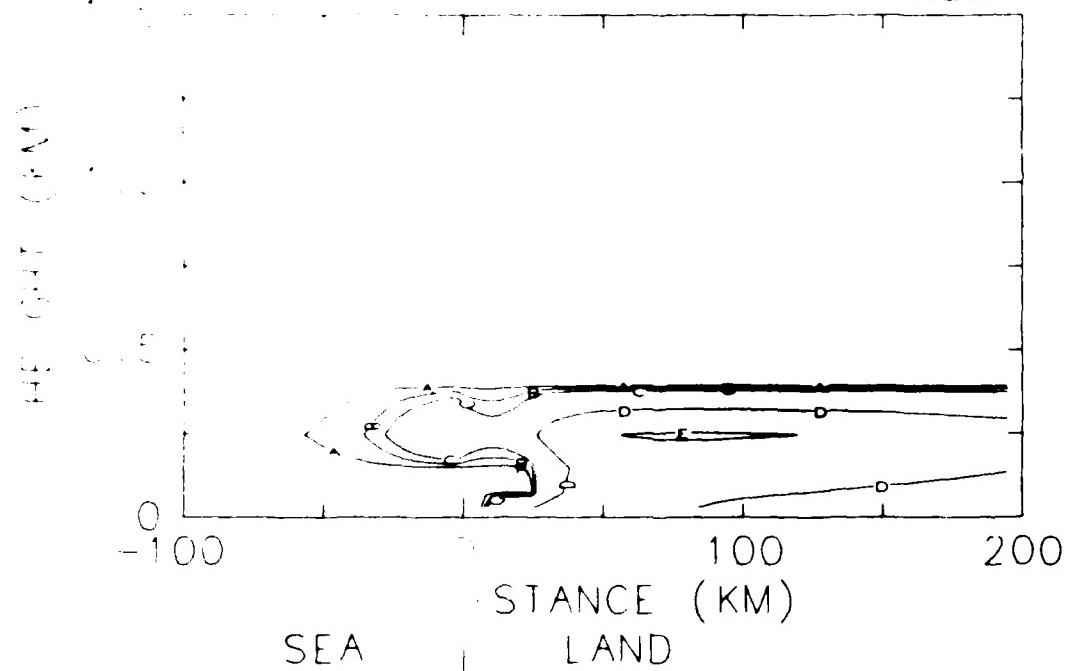




CLOUD WATER (G/KG)

DAY 3 01:18

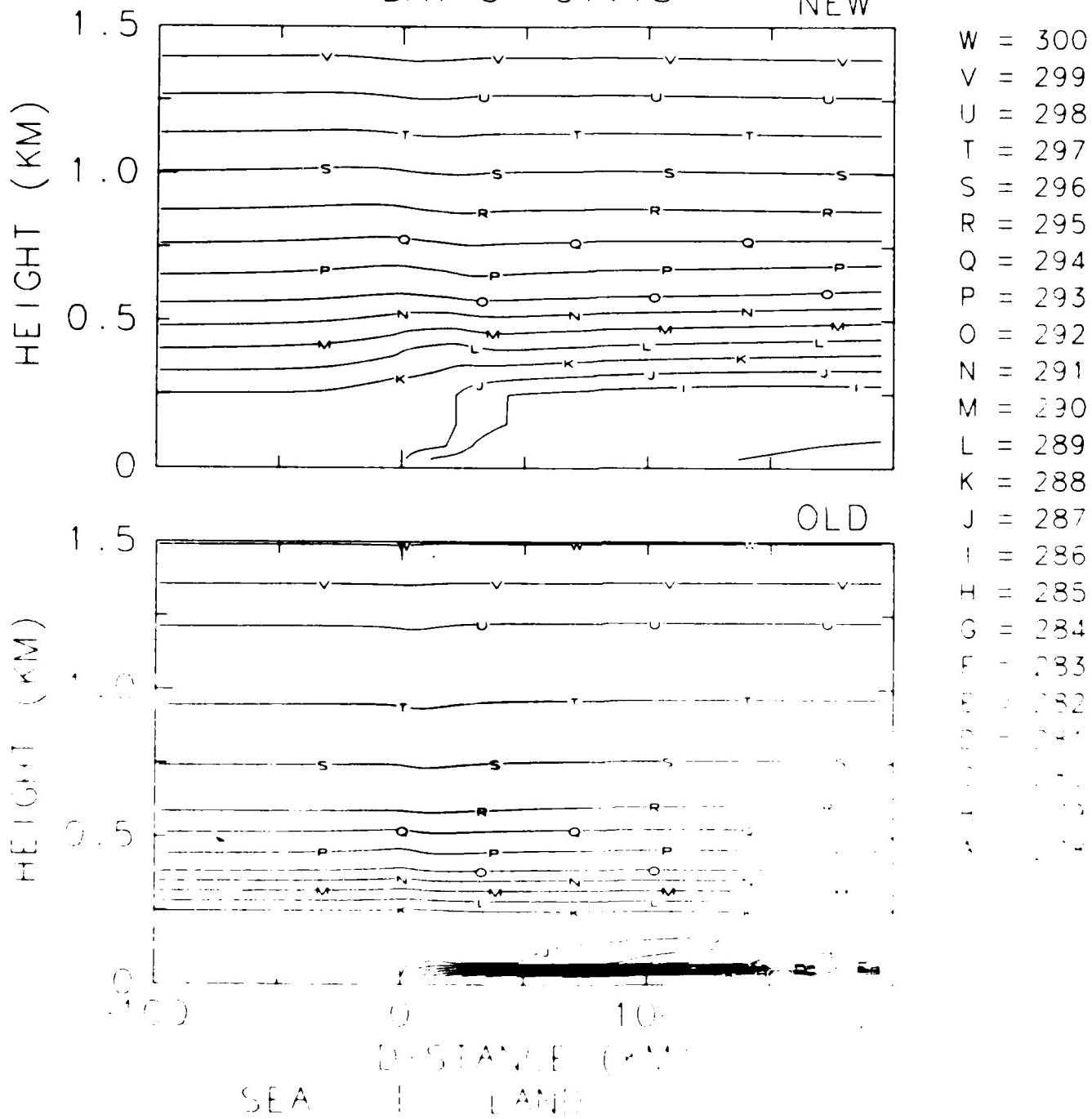
NEW



POTENTIAL TEMPERATURE

DAY 3 01:18

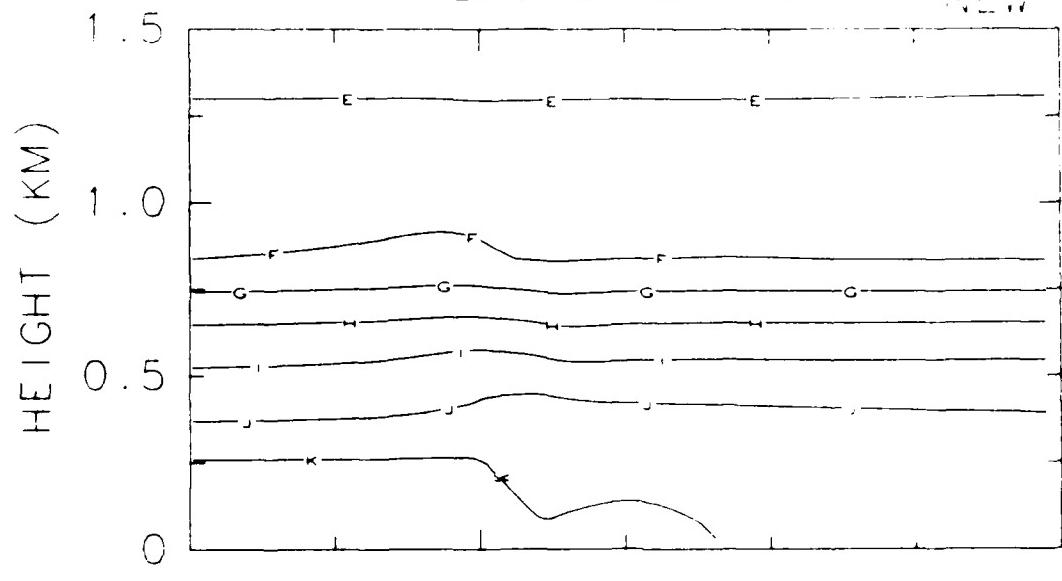
NEW



SPECIFIC HUMIDITY (G/G)

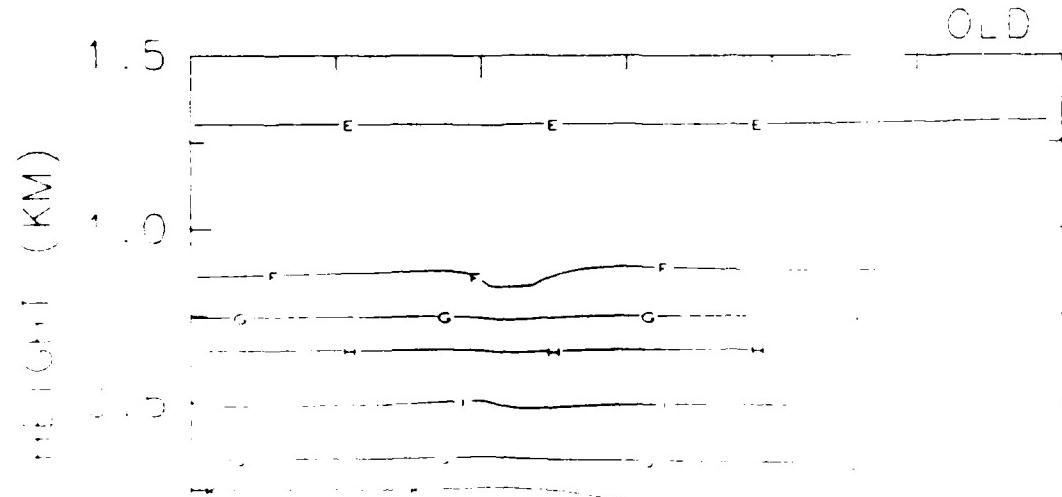
DAY 3 01:18

NEW



CD = 0.006  
CD = 0.005  
CD = 0.004  
CD = 0.003

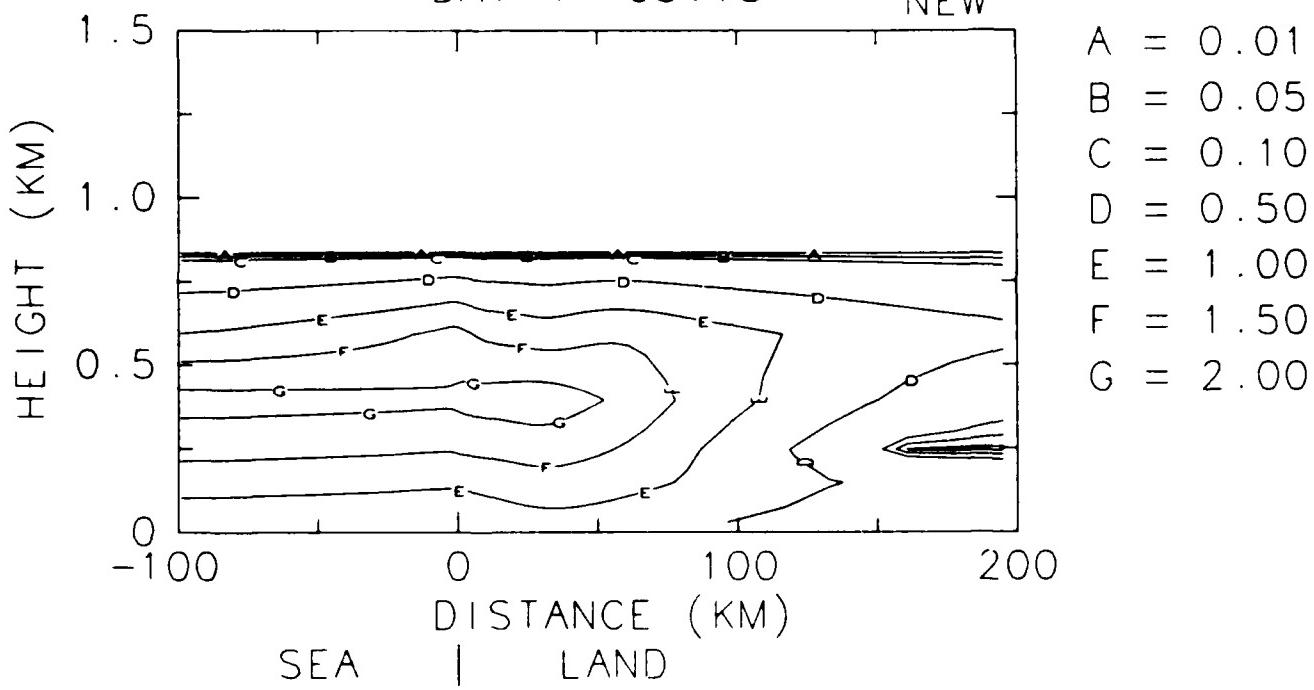
OLD



CLOUD WATER (G/KG)

DAY 4 05:18

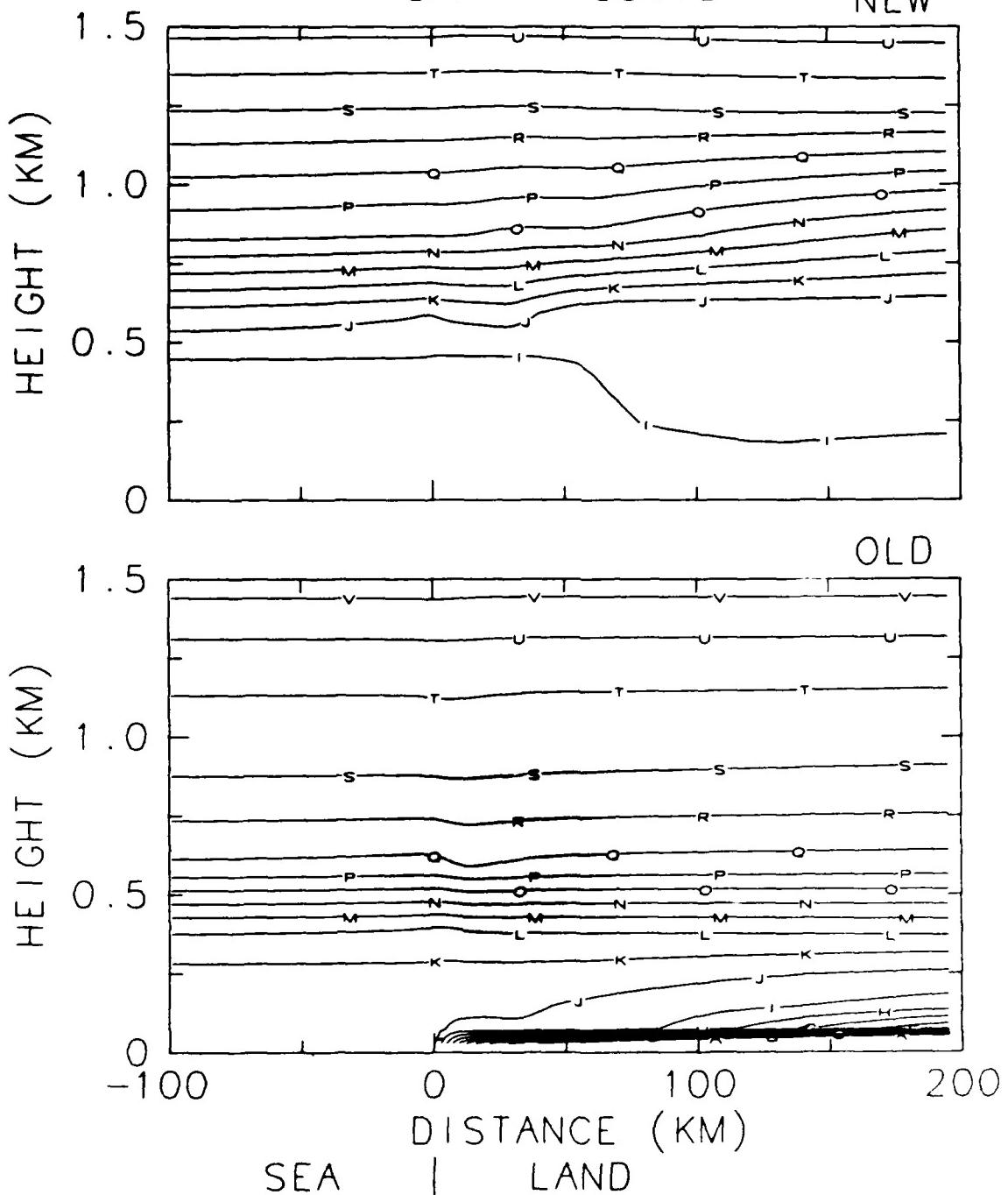
NEW



# POTENTIAL TEMPERATURE

DAY 4 05:18

NEW

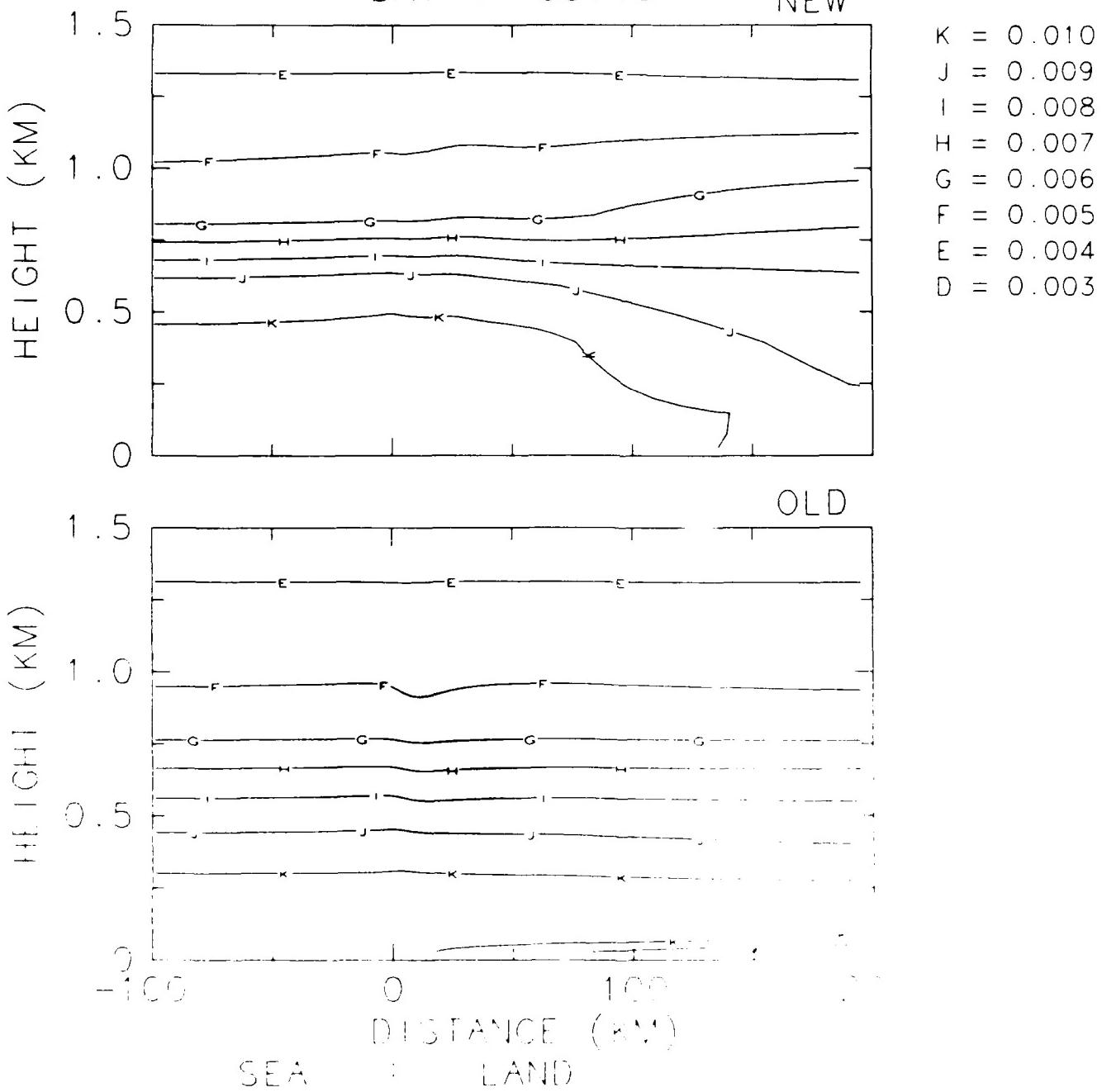


W	= 300
V	= 299
U	= 298
T	= 297
S	= 296
R	= 295
Q	= 294
P	= 293
O	= 292
N	= 291
M	= 290
L	= 289
K	= 288
J	= 287
I	= 286
H	= 285
G	= 284
F	= 283
E	= 282
D	= 281
C	= 280
B	= 279
A	= 278

SPECIFIC HUMIDITY (G/G)

DAY 4 05:18

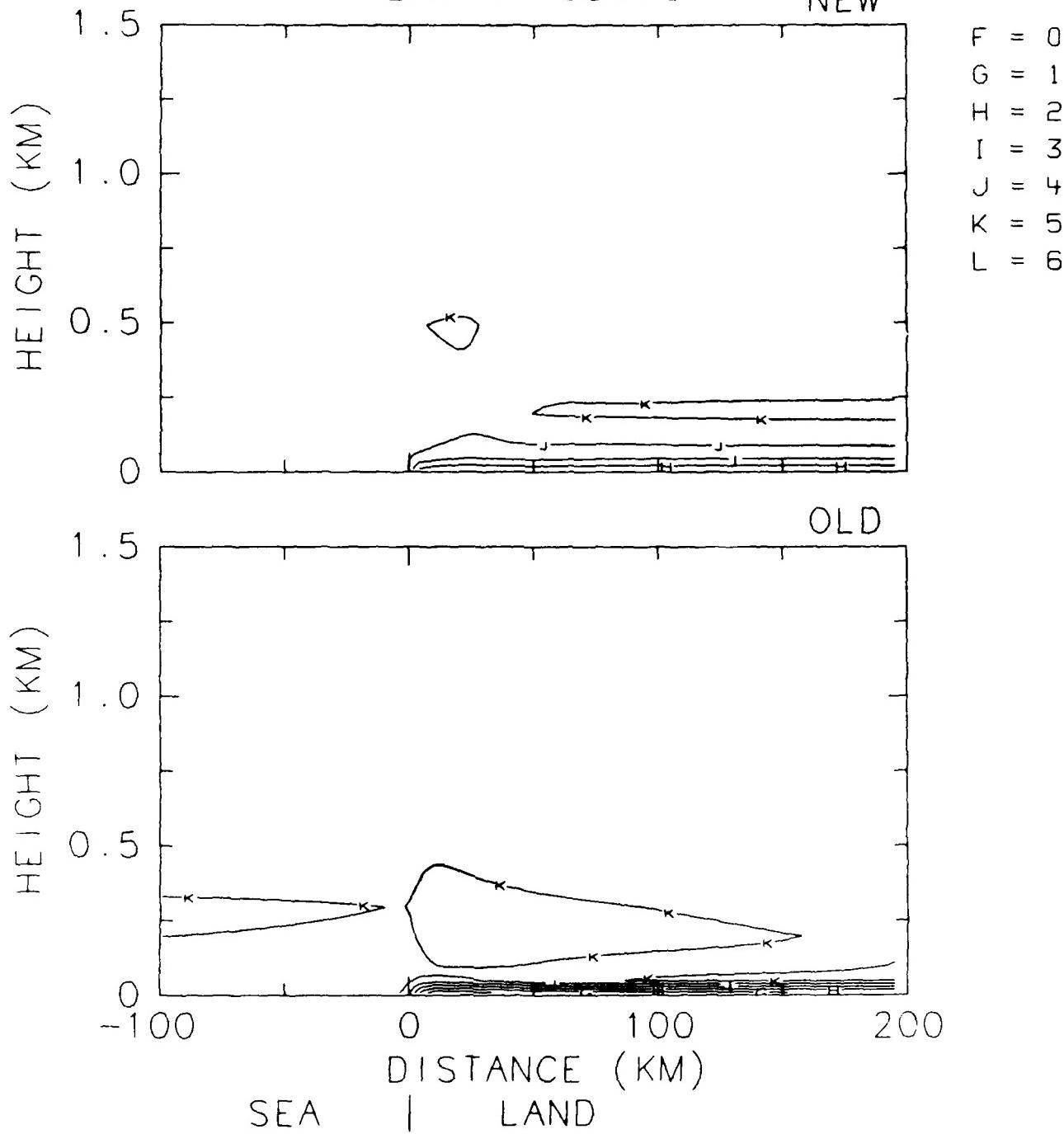
NEW



EAST-WEST VELOCITY (M/S)

DAY 4 05:18

NEW



- OLD
  - + Very Cold Layer Forms over Land
  - + Cold Layer is Very Thin
  - + No Significant Perturbation above 200 M
  - + No Significant Vertical Motion
  
- NEW
  - + Fog Forms over Land at Dawn and Thickens with Time
  - + Fog Forms over Ocean about 20 Hours Later
    - Vertical Mixing of Moisture from Sea
    - Radiative Cooling Aloft
  - + Fog over Land and Sea Up to 800 M in 2 1/2 Days
  - + Fog Eliminates Land/Sea Surface Temperature Contrast
    - + No Significant Vertical Motion

A Numerical Simulation  
of  
The Possible Effects of  
Nucleation Scavenging  
on  
Plume Injection

by  
Greg Tripoli,  
Chang Chen  
William Cotton

R. A. N. S.

# Regional Atmospheric Modeling System

Time dependent P.D.E. for

$\nabla$  ( $u, v, w$ ) (velocity)

$P$  (pressure)

$\Theta$  (potential temperature.)

$r_v$   
 $r_c$   
 $r_r$   
 $r_i$   
 $r_g$   
 $r_a$

}

Mixing ratio ( $\frac{\text{mass water}}{\text{mass air}}$ )

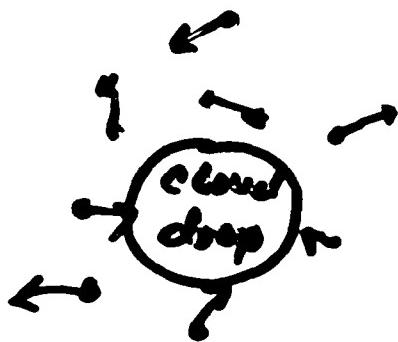
for

vapor, cloud, rain, crystals,  
graupel, aggregates

# Phoretic Scavenging



# Brownian Scavenging



# Hydrostatic Capture



# Electric Effects



# *Nucleation Scavenging*

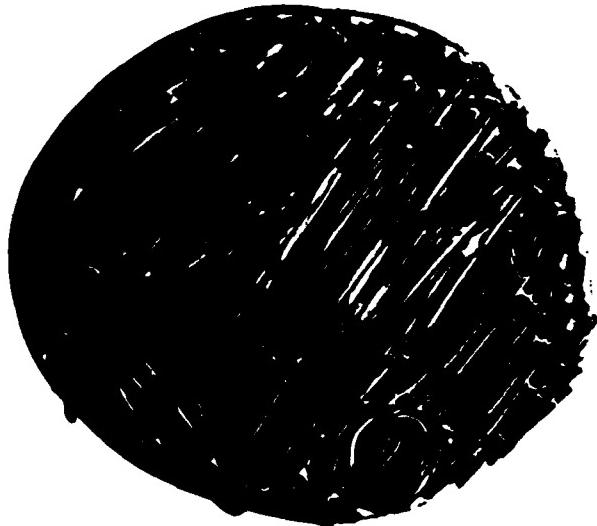
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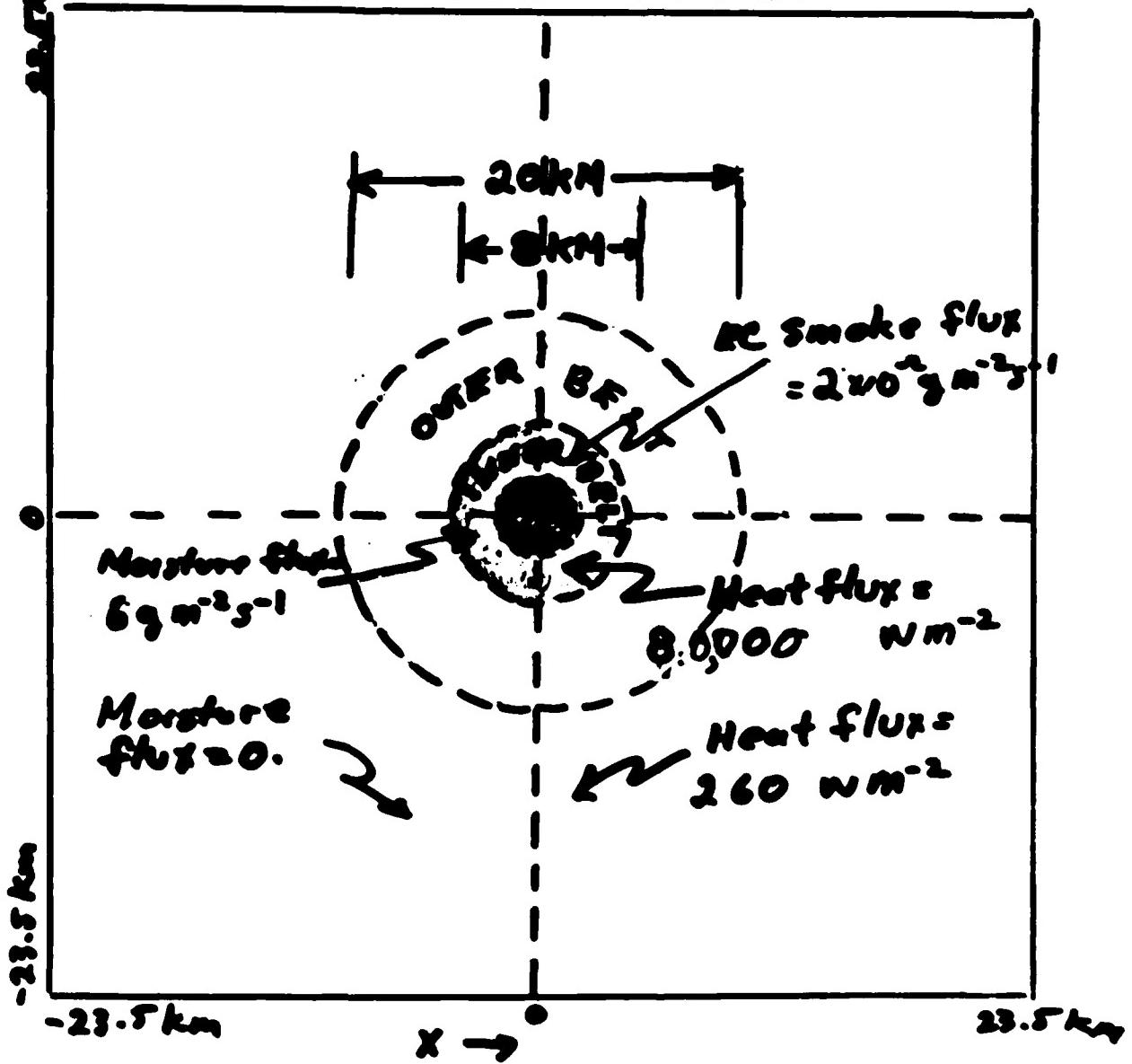
B



C



## EXPERIMENTAL DESIGN



$$\begin{aligned}\Delta X &\approx 1 \text{ km} \\ \Delta Y &\approx 1 \text{ km} \\ \Delta Z &\approx .75 \text{ km}\end{aligned}$$

EC Produced is  
20% of smoke which is  
1% of fuel  
(Coutam, et.al, 1997, for wood)

### Assumptions:

1. Homogeneous fuel density of  $110 \text{ kg/m}^3$  (10% water content) within inner belt.
2. Neglect fire in outer belt.
3. City center same as inner belt

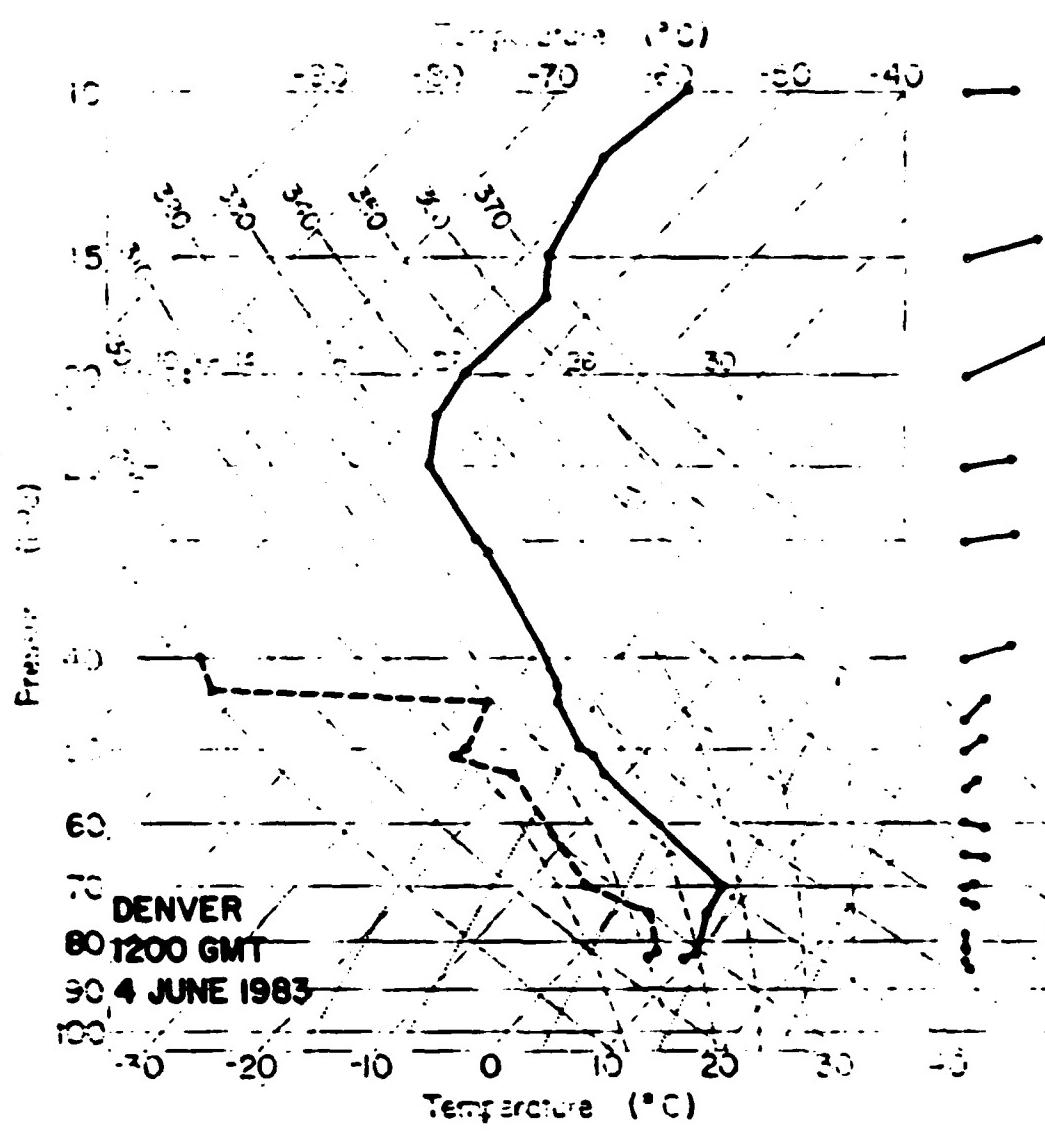


Figure 3

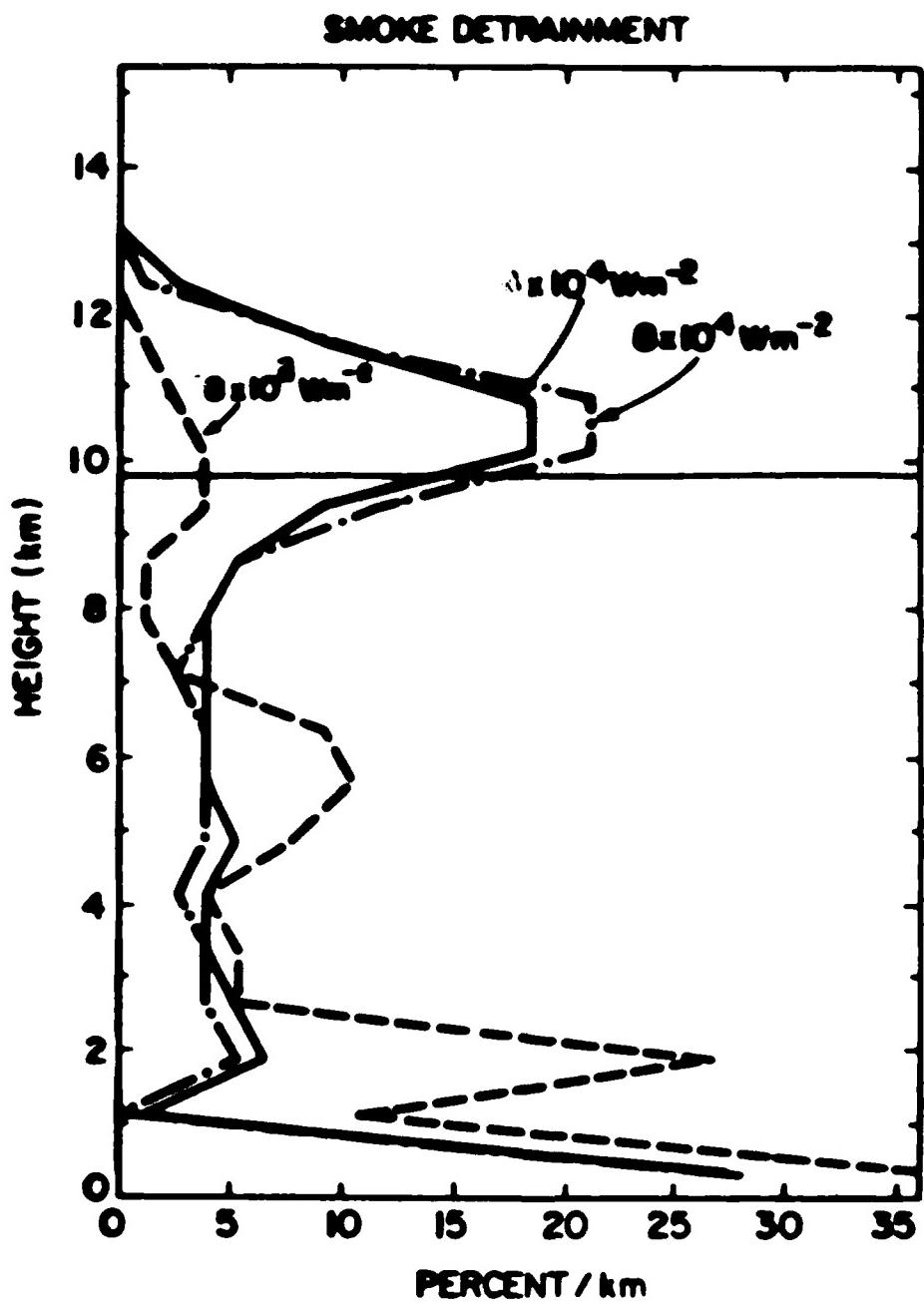


Figure 4

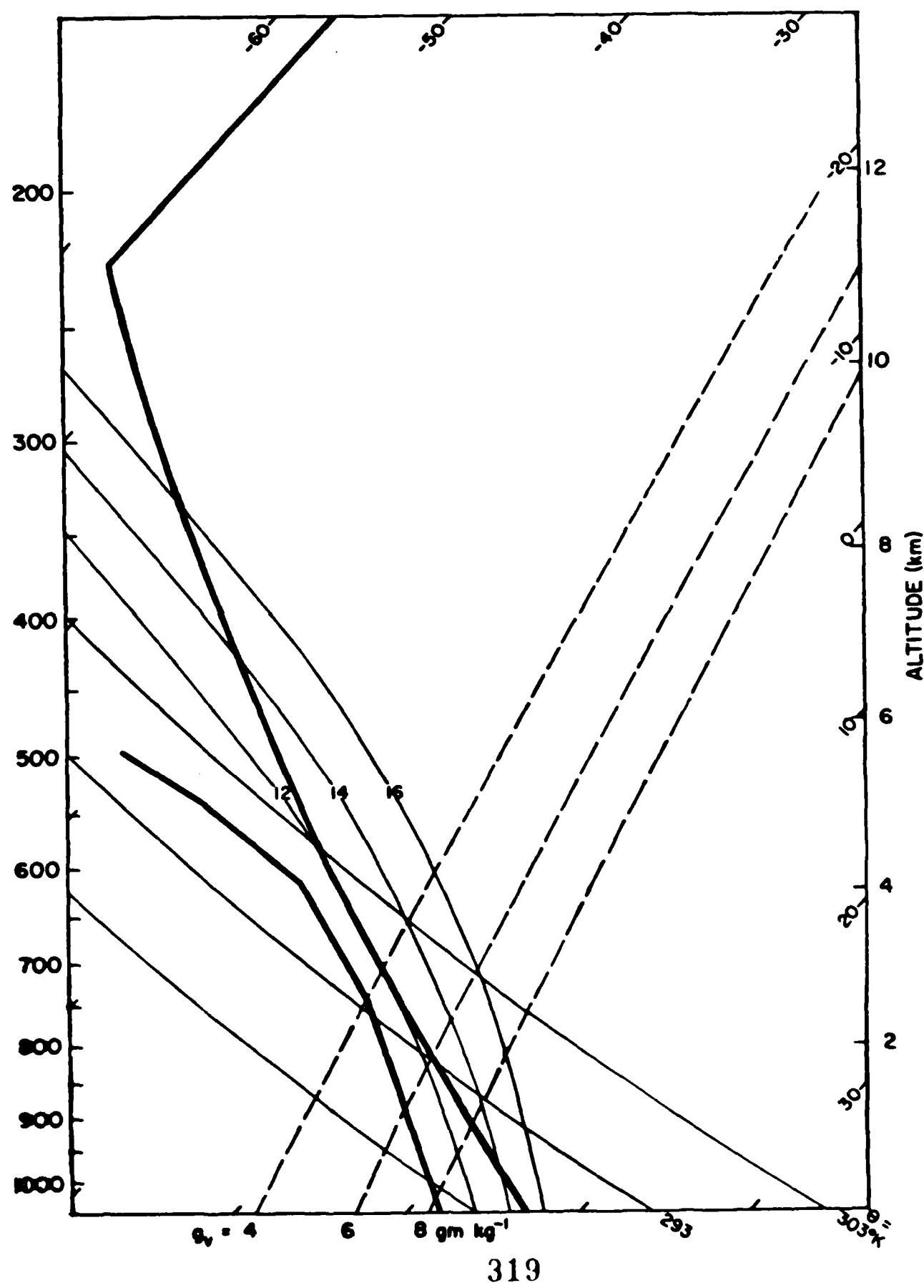
DESIGN OF EXPERIMENT  
TO TEST EFFECTIVENESS OF  
NUCLEATION SCAVENGING

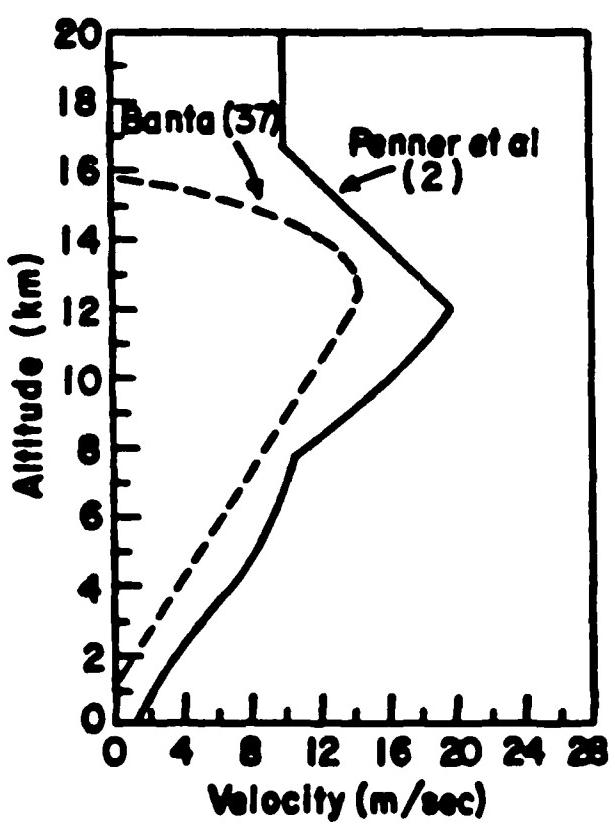
- 1) Use conservative standard spring/fall atmosphere, wind as Penner et.al. (1985)
- 2) Use Penner et.al. (1985) high intensity case (low intensity case results trivial since deep cloud never forms)
- 3) Use Cotton (1985) phoretic/Brownian scavenging model.
- 4) Nucleation scavenging:

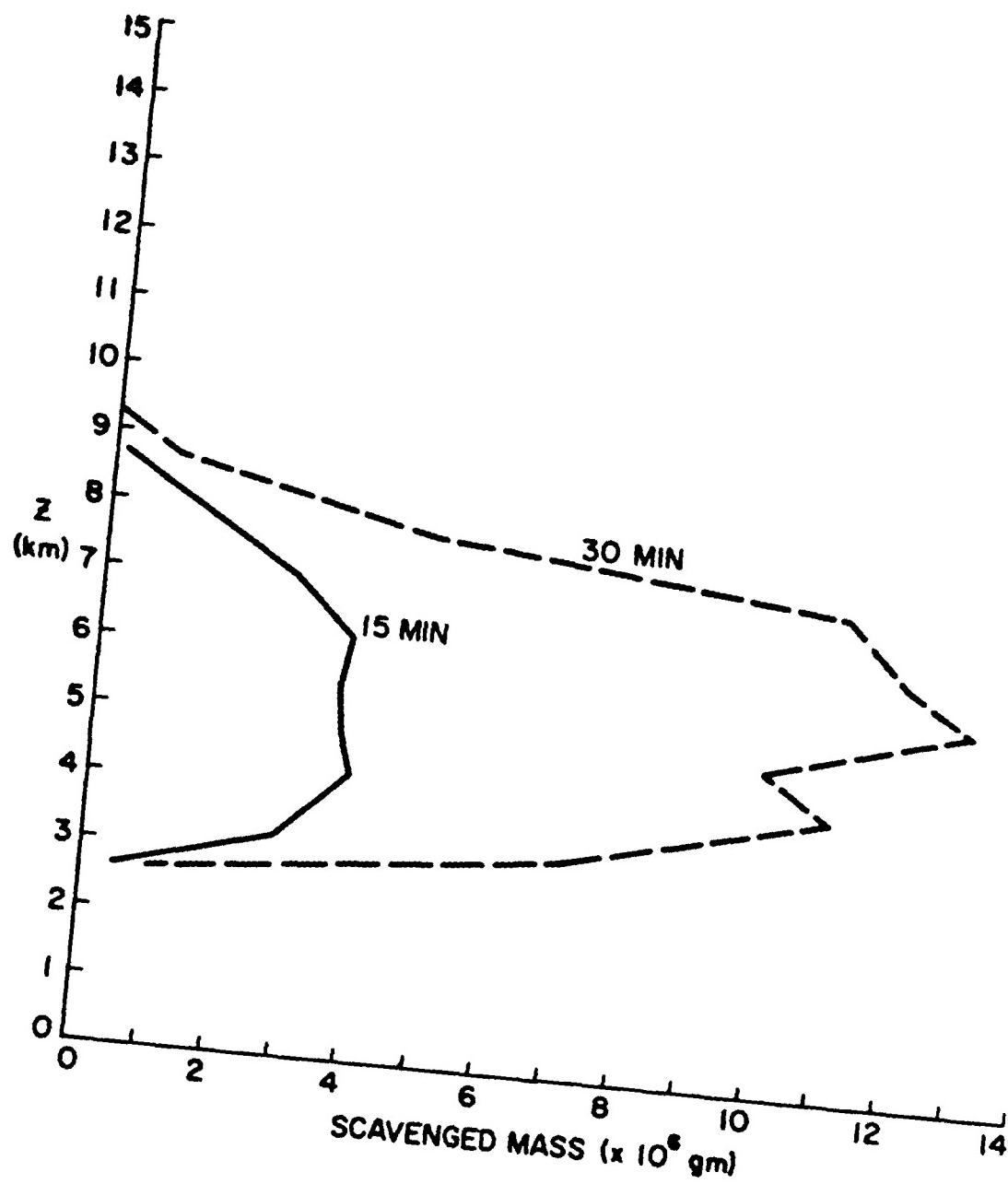
$$S_N = \frac{CL_{CP} + C_{CI} + C_{CB} + C_{CA}}{r_C} r_{EC}$$

INHERENT ASSUMPTIONS FOR NUCLEATION SCAVENGING

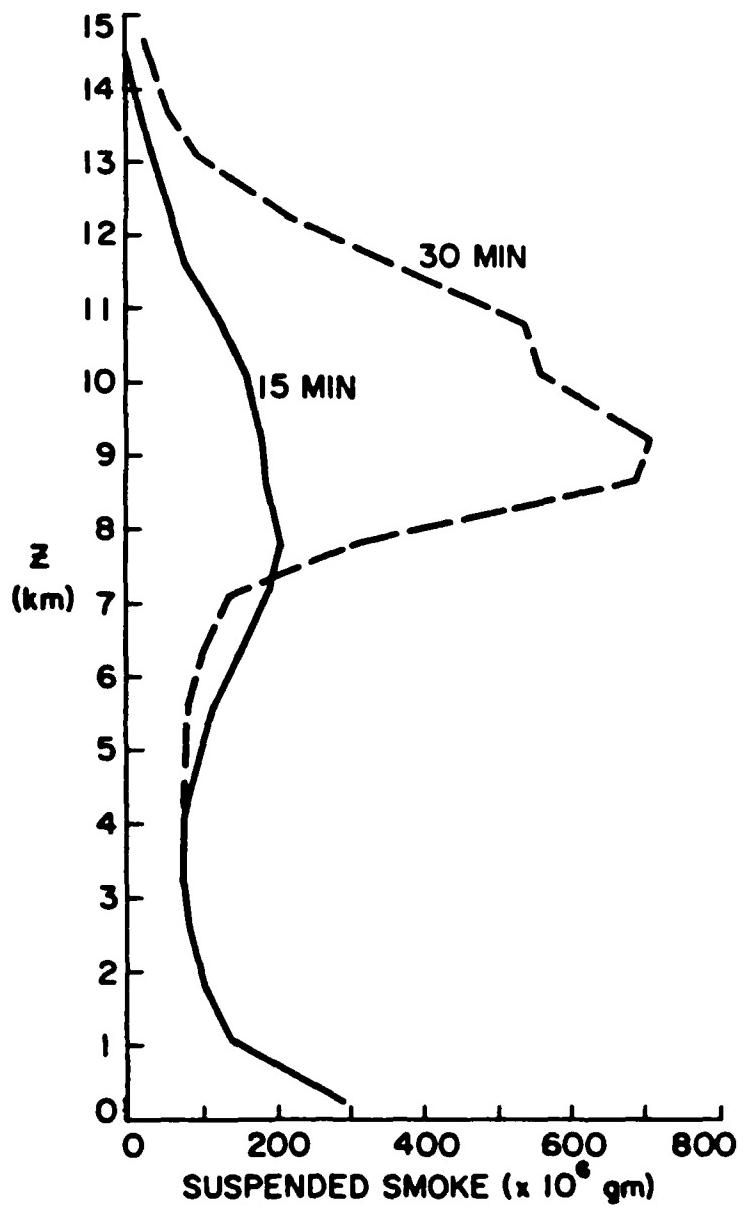
- 1) All smoke particles will nucleate.
- 2) Smoke collected onto precipitation particle will be lost from system, even if precipitation re-evaporates. (We assume many will be combined to make large precipitating smoke particle).
- 3) Collection efficiency of cloud droplets unity despite smaller size.







(No Nucleation Scavenging)



(No Nucleation Scavenging)

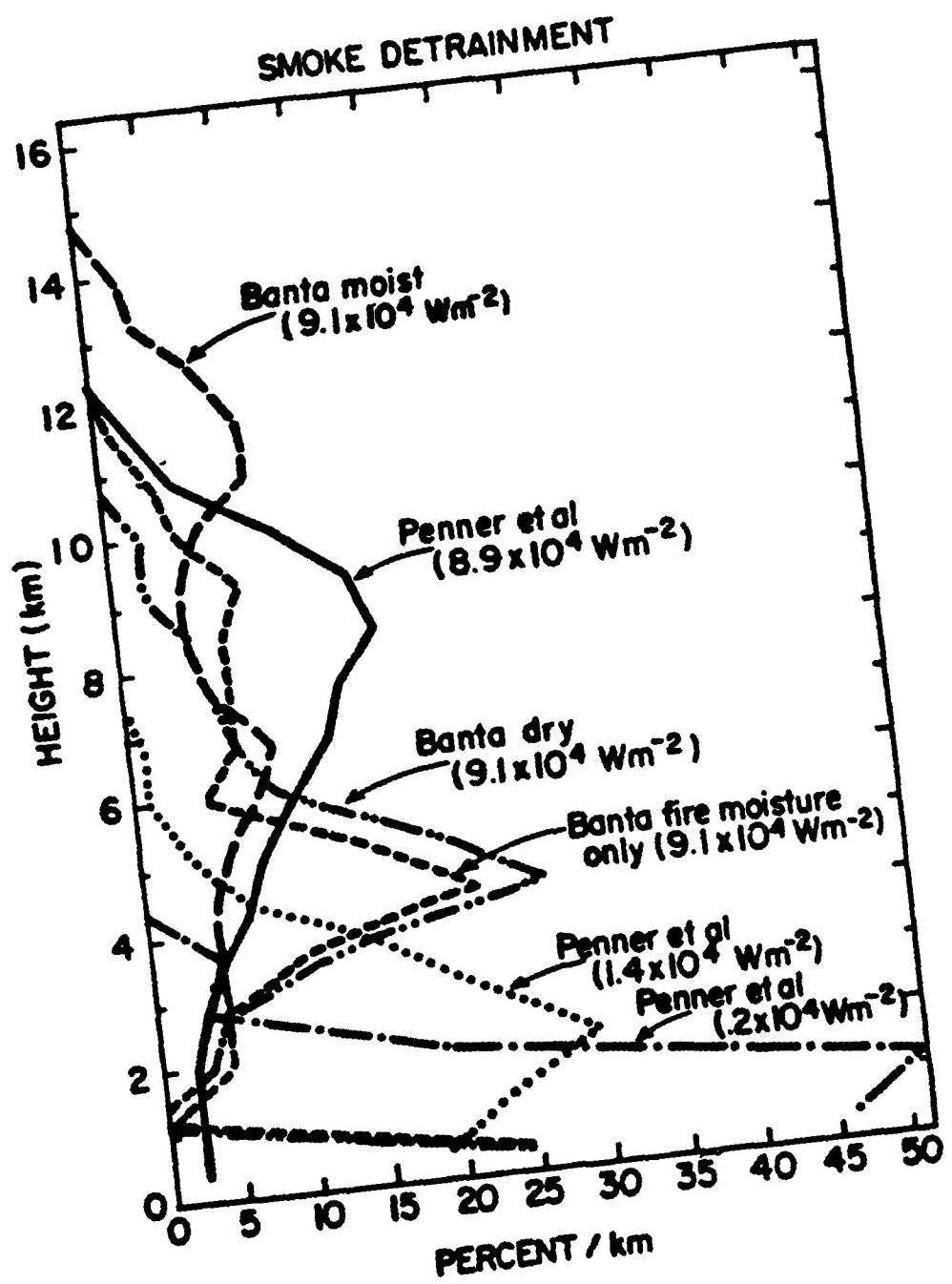
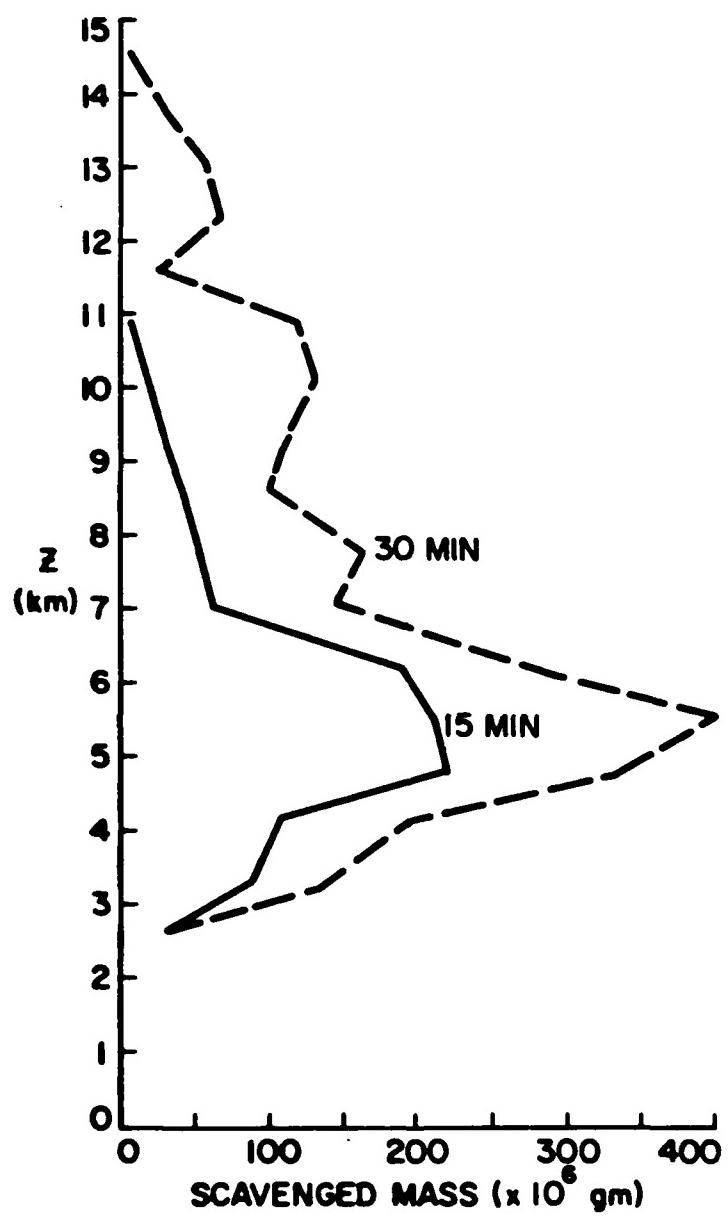
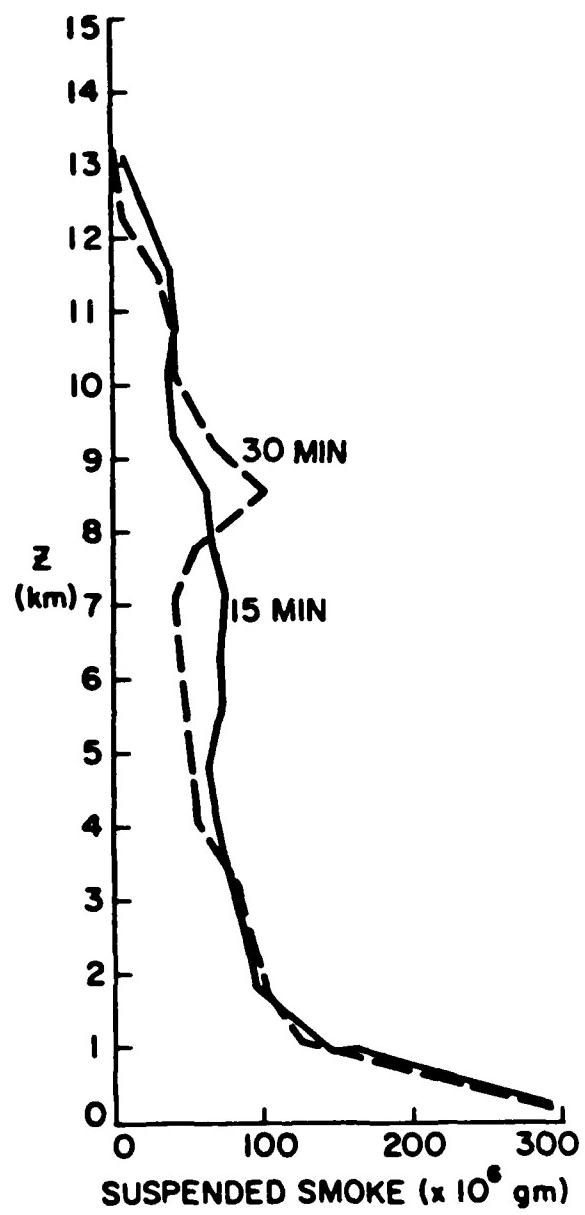


Figure 10

(Results of other investigators)



Nucleation Scavenging



Nucleation Scavenging

## CONCLUSIONS ABOUT SCAVENGING

- A. Nucleation scavenging might remove up to 80% of smoke if effective.
- B. Phoretic scavenging is of minor importance.
- C. There are some major uncertainties:
  - 1) How effective is smoke as a cloud condensation nuclei?
  - 2) Will the very large numbers of CCN render the accretion of cloud droplets by precipitation particles ineffective (1 g/kg cloud ~ 1 micron cloud drops).
  - 3) How much larger will the smoke particles captured by precipitation be?

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